Investigating the Earth's Core with Neutrinos

Rebekah Pestes

Center for Neutrino Physics Virginia Tech



CNP Research Day - May 6, 2022

Table of Contents

1 Introduction

- Inside the Earth
- Neutrinos in Matter

2 Paper with P. Denton (arXiv:2110.01148)

3 Summary

1 Introduction

- Inside the Earth
- Neutrinos in Matter

2 Paper with P. Denton (arXiv:2110.01148)

3 Summary

Looking inside the Earth

"View" with Seismography



Measurement depends on composition, pressure, and temperature of the material, as well as exact location and depth of the earthquake

Neutrino Hamiltonian in Matter:

$$H = \frac{1}{2E} \left(U^{\dagger} \begin{bmatrix} 0 & 0 & 0 \\ 0 & \Delta m_{21}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 \end{bmatrix} U + 2\sqrt{2}G_F N_e E \begin{bmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{bmatrix} \right)$$

Oscillation probability:

$$P_{\alpha \to \beta}(t) = |\langle \nu_{\beta} | \nu_{\alpha}(t) \rangle|^{2} = \left| \sum_{i,j} U_{\beta i} U_{\alpha j}^{*} \left\langle \nu_{i} \right| e^{-iHt} \nu_{j} \right\rangle \right|^{2}$$

Atmospheric Neutrinos



1 Introduction

- •

2 Paper with P. Denton (arXiv:2110.01148)

3 Summary

Neutrino oscillations through the Earth's core

Peter B. Denton^{1,*} and Rebekah Pestes^{1,2,†}

¹High Energy Theory Group, Physics Department, Brookhaven National Laboratory, Upton, New York 11973, USA

²Center for Neutrino Physics, Department of Physics, Virginia Tech, Blacksburg, Virginia 24061, USA

(Received 12 October 2021; accepted 8 December 2021; published 29 December 2021)

Neutrinos have two properties that make them fairly unique from other known particles: extremely low cross sections and flavor changing oscillations. With a good knowledge of the oscillation parameters soon in hand, it will become possible to detect low-energy atmospheric neutrinos sensitive to the forward elastic scattering off electrons in the Earth's core providing a measurement of the core properties and the matter effect itself. As the dynamics of the Earth's core are complicated and in a difficult to probe environment, additional information from upcoming neutrino experiments will provide feedback into our knowledge of geophysics as well as useful information about exoplanet formation and various new physics scenarios including dark matter. In addition, we can probe the existence of the matter effect in the Earth ad constrain the nonstandard neutrino interaction parameter e_{ee}^{\oplus} . We show how DUNE's sensitivity to low-energy atmospheric neutrino oscillations can provide a novel constraint on the density and radius of the Earth's core at the 9% level and the Earth's matter effect at the 5% level. Finally, we illuminate the physics behind low-energy atmospheric neutrino resonances in the Earth.

DOI: 10.1103/PhysRevD.104.113007

DUNE: Deep Underground Neutrino Experiment



LArTPC (Liquid Argon Time Projection Chamber) for far detector \Rightarrow Can determine direction neutrino came from Energy range (0.1 GeV-8.0 GeV) includes atmospheric neutrino energies

Simulating DUNE Set Up

- Detector with DUNE's far detector specs
 - 40 kton
 - 10 years
- Honda flux model averaged over angles for source

Simulating DUNE Set Up

- Detector with DUNE's far detector specs
 - 40 kton
 - 10 years
- Honda flux model averaged over angles for source
- Used nuSQuIDS to calculate oscillated flux
- Assumed NO for simulated data
- Fixed oscillation parameters

Simulating DUNE Set Up

- Detector with DUNE's far detector specs
 - 40 kton
 - 10 years
- Honda flux model averaged over angles for source
- Used nuSQuIDS to calculate oscillated flux
- Assumed NO for simulated data
- Fixed oscillation parameters
- χ^2 calculated for fits

$$\chi^{2} = \sum_{i} \frac{aty(\phi_{true_{i}} - \phi_{fit_{i}})^{2}}{\phi_{true_{i}}} + \sum_{j} \left(\frac{s_{j}}{\sigma_{j}}\right)^{2}$$

- Minimized over systematic parameters
- 1st Fit: Varied ε_{ee}
- 2nd Fit: Varied radius of Earth's core

Simulating DUNE Earth's Density Profile

Accepted model: Preliminary Reference Earth Model (PREM)

Assumes spherical Earth

When changing core radius, scaled core density to keep Earth's mass constant



• Flux uncertainties

$$\Phi_{\alpha} = \Phi_{\alpha,0} f_{\alpha}(E) (E_{\nu}/E_0)^{\gamma}$$

- $\Phi_{\alpha,0}$ = flux normalization, 1 ± 40% for $\alpha = \nu_e, \nu_\mu, \bar{\nu}_e, \bar{\nu}_\mu$
- γ = spectral index, 0 \pm 0.2

• Flux uncertainties

$$\Phi_{\alpha} = \Phi_{\alpha,0} f_{\alpha}(E) (E_{\nu}/E_0)^{\gamma}$$

- $\Phi_{\alpha,0}$ = flux normalization, 1 ± 40% for $\alpha = \nu_{e}, \nu_{\mu}, \bar{\nu}_{e}, \bar{\nu}_{\mu}$
- γ = spectral index, 0 \pm 0.2
- Exclude partially contained ν_{μ} events (reduce ν_{μ} events by 25%)
- Assume good flavor discrimination, but no $\nu/\bar{\nu}$ discrimination

• Flux uncertainties

$$\Phi_{\alpha} = \Phi_{\alpha,0} f_{\alpha}(E) (E_{\nu}/E_0)^{\gamma}$$

- $\Phi_{\alpha,0}$ = flux normalization, 1 ± 40% for $\alpha = \nu_e, \nu_\mu, \bar{\nu}_e, \bar{\nu}_\mu$
- γ = spectral index, 0 \pm 0.2
- Exclude partially contained ν_{μ} events (reduce ν_{μ} events by 25%)
- Assume good flavor discrimination, but no $\nu/\bar{\nu}$ discrimination
- For resolutions:
 - 10 cos(*θ*_Z) bins
 - 43 log(E_{ν}) bins (10% resolution for E_{ν})











1 Introduction

2 Paper with P. Denton (arXiv:2110.01148)

3 Summary

- Seismology measurements of Earth's layers depend on many variables
- Can measure size of Earth's core using atmospheric neutrinos
 - Depends on fewer variables that are not measured using other means
 - Affected by new physics scenarios
- DUNE can measure Earth's matter effect to 5% and the size of the Earth's core to 9%

Questions?

This research was funded by the DOE.

DUNE Simulation Earth's Core Sensitivity - Full Range



