

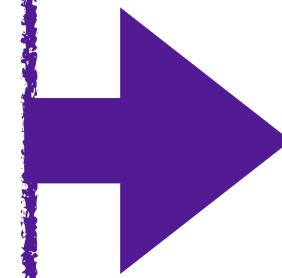
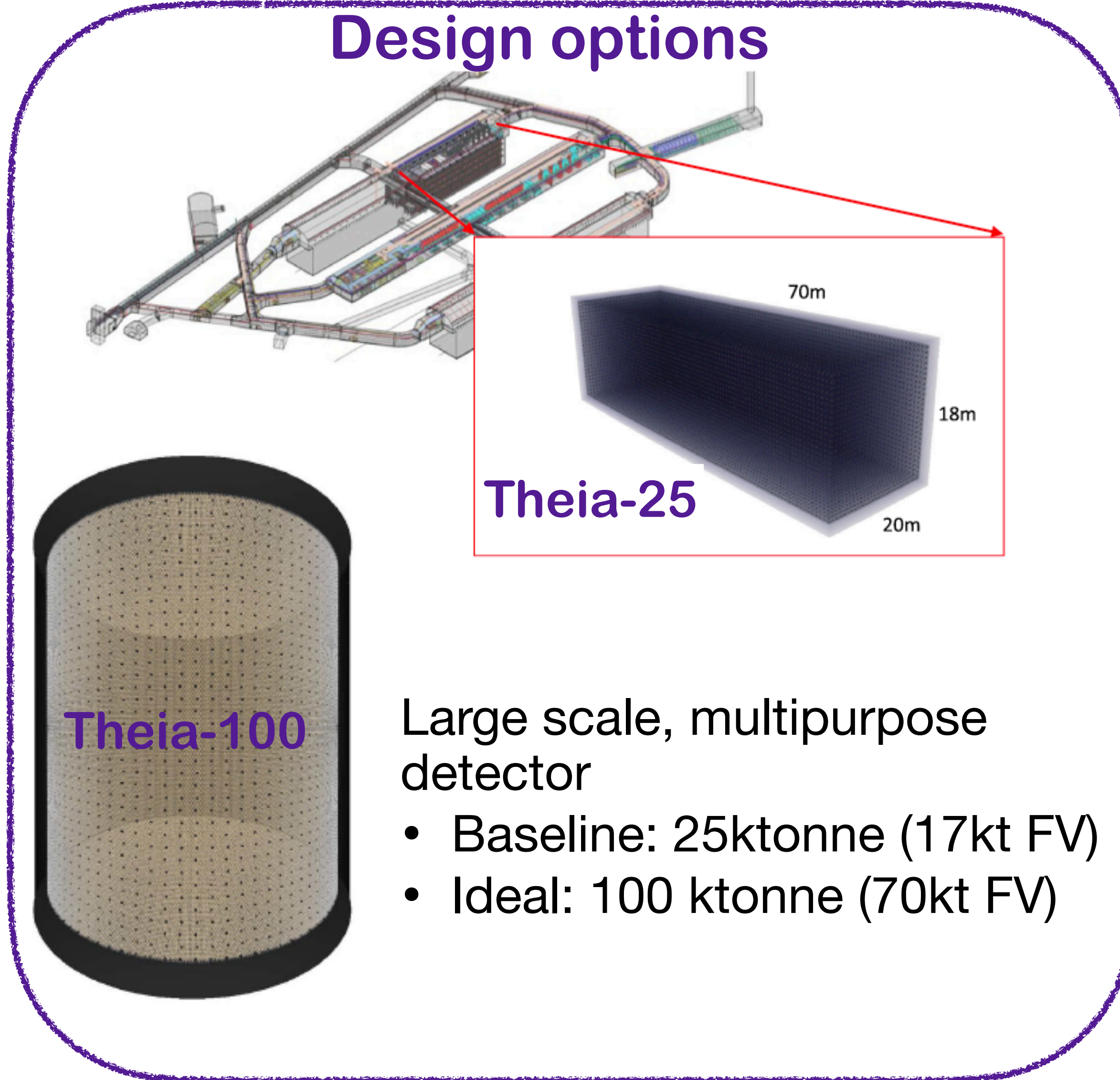
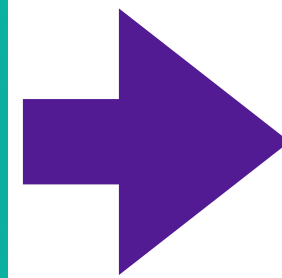
THEIA: An advanced optical neutrino detector

Zara Bagdasarian for the THEIA collaboration
University of California, Berkeley

NuSTEC workshop
March 18th 2021

Theia: advanced optical multipurpose neutrino detector

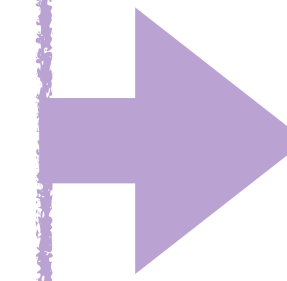
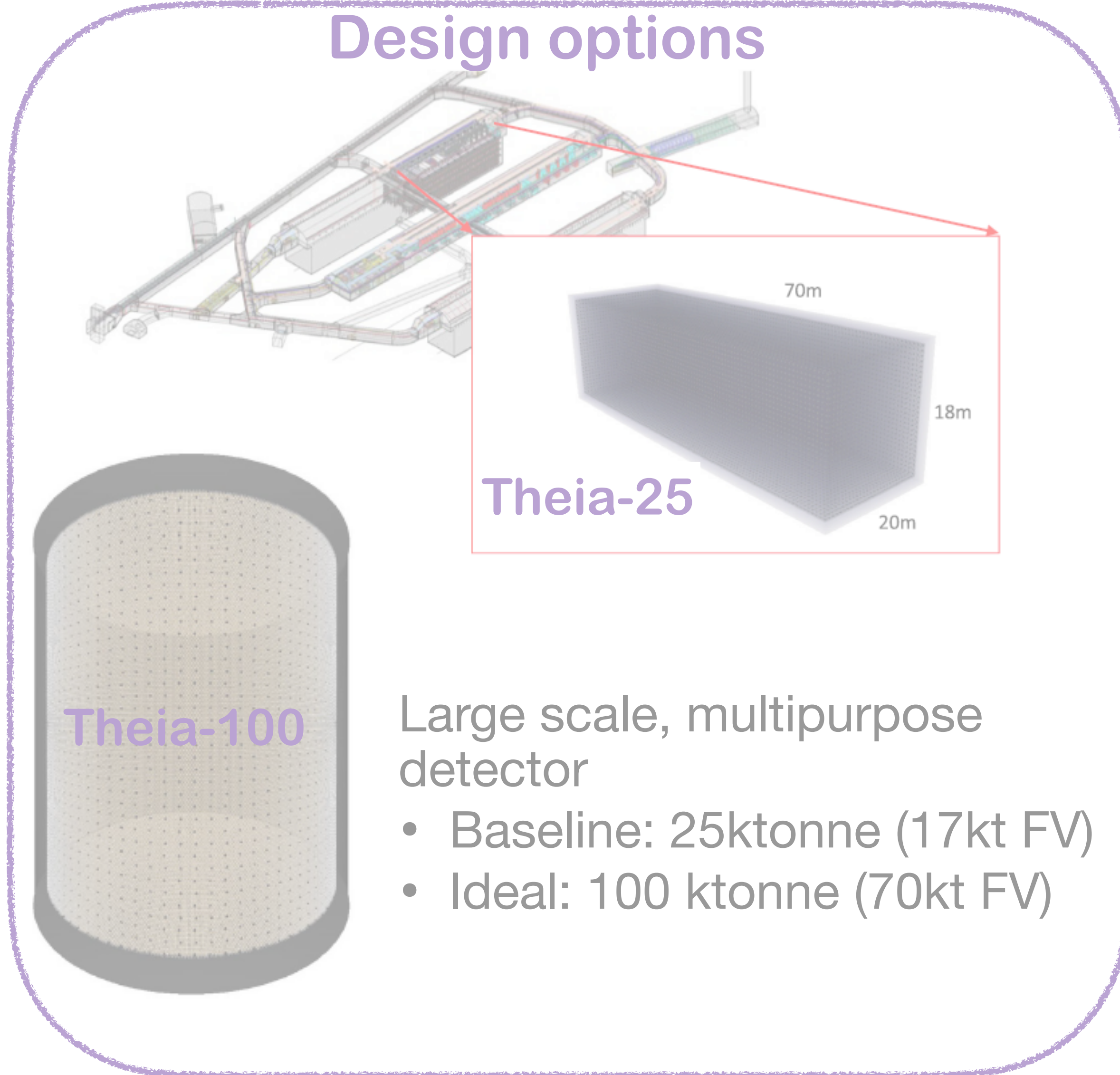
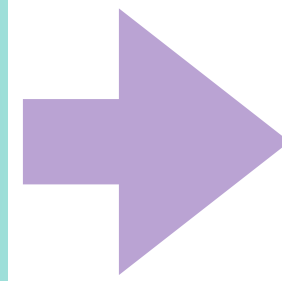
Cutting edge developments in the target material and photodetection



Broad physics program:
Studying neutrino
fundamental properties
and astrophysical
objects

Theia: advanced optical multipurpose neutrino detector

Cutting edge developments in the target material and photodetection



Broad physics program:
Studying neutrino
fundamental properties
and astrophysical
objects

How to broaden the current physics reach



Scintillation Detectors:

- ✓ High light yield
- ✓ Low energy threshold
- ✓ Good energy and position resolutions
- ✗ Limited in size by absorption and cost
- ✗ No directionality

Cherenkov Detectors:

- ✓ Directional information
- ✓ Can be very large (low absorption)
- ✓ Particle ID at high energies
- ✗ No access to physics below the Cherenkov threshold
- ✗ Low light yield

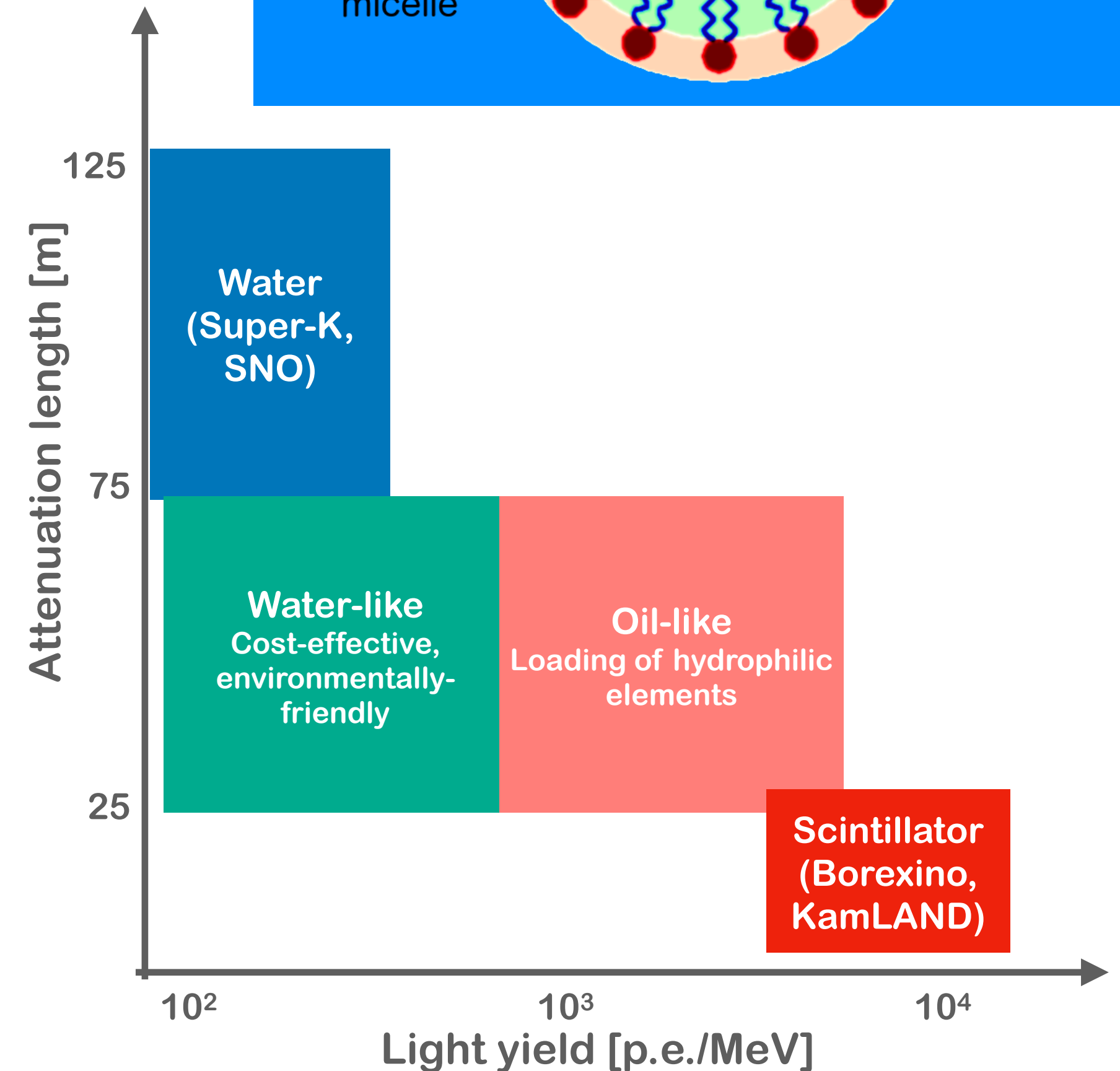
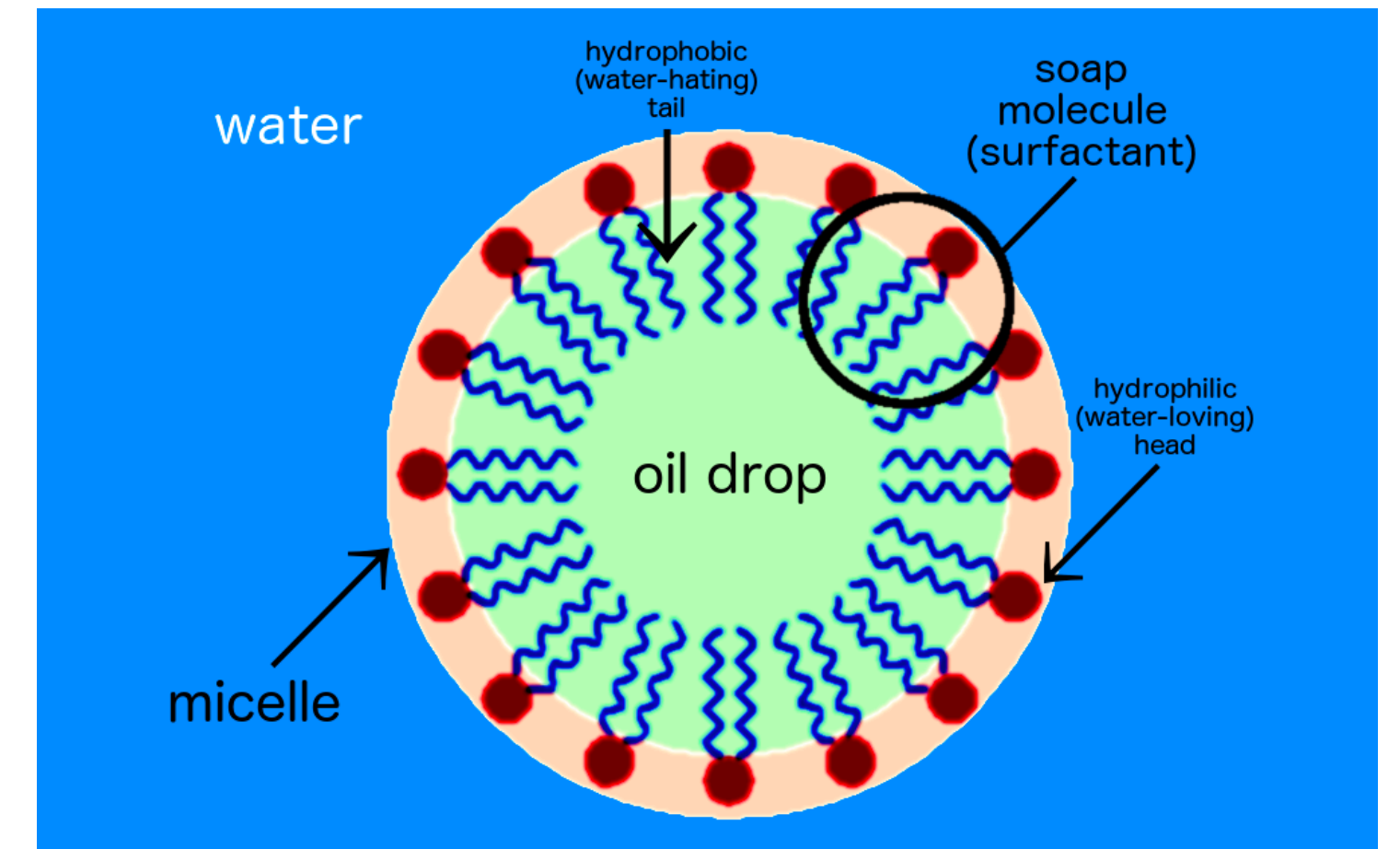
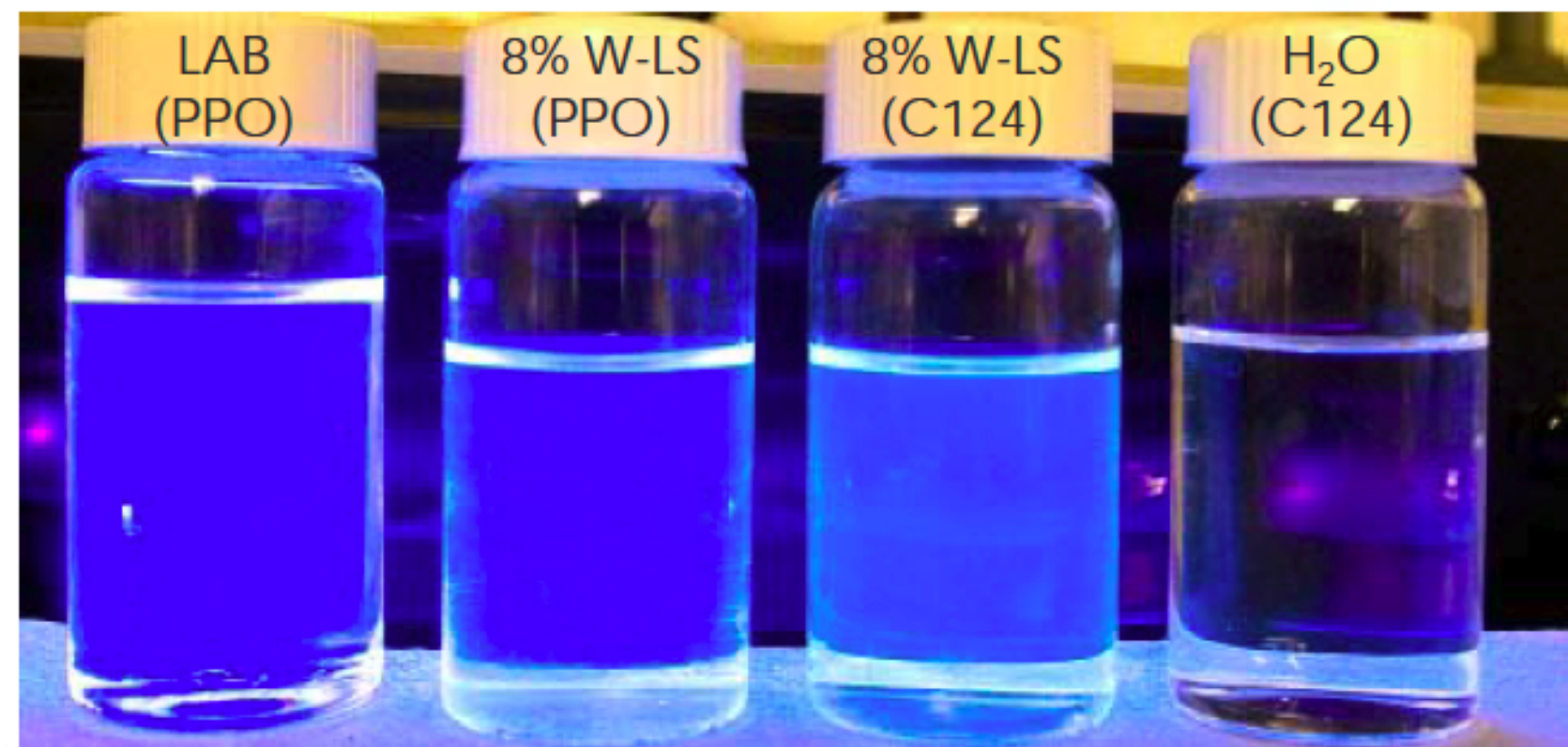


Water-based Liquid Scintillation (WbLS) Detectors: Get best of two worlds



Water-based Liquid Scintillator - Basics

- Water-based Liquid Scintillator (WbLS) is a mixture of pure water and oil-based liquid scintillator
- WbLS is made using a surfactant (soap-like) such as PRS* (hydrophilic head and hydrophobic tail) to hold the scintillator molecules in water in a “micelle” structure
- Combines the advantages of water (transparency, low cost) and liquid scintillator (high light yield)



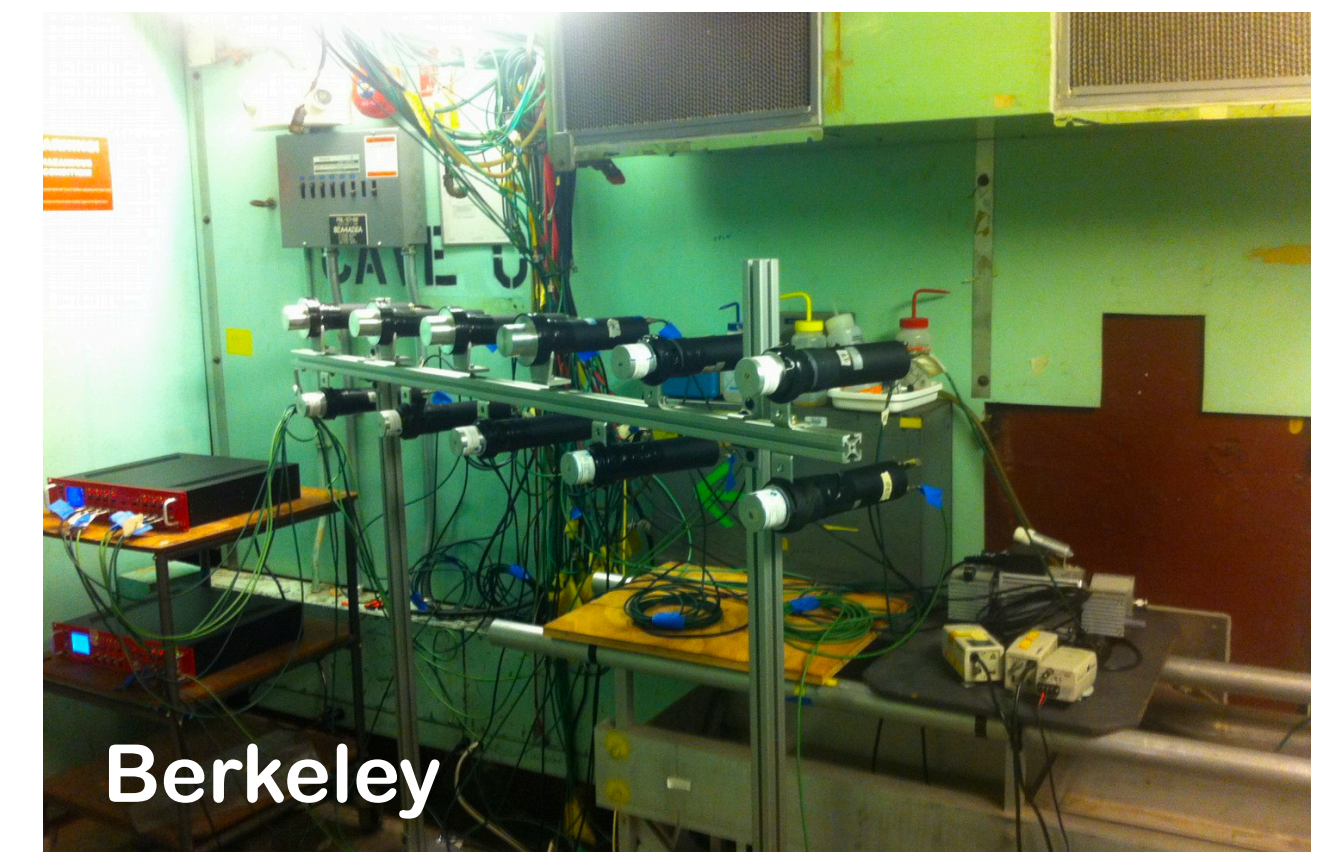
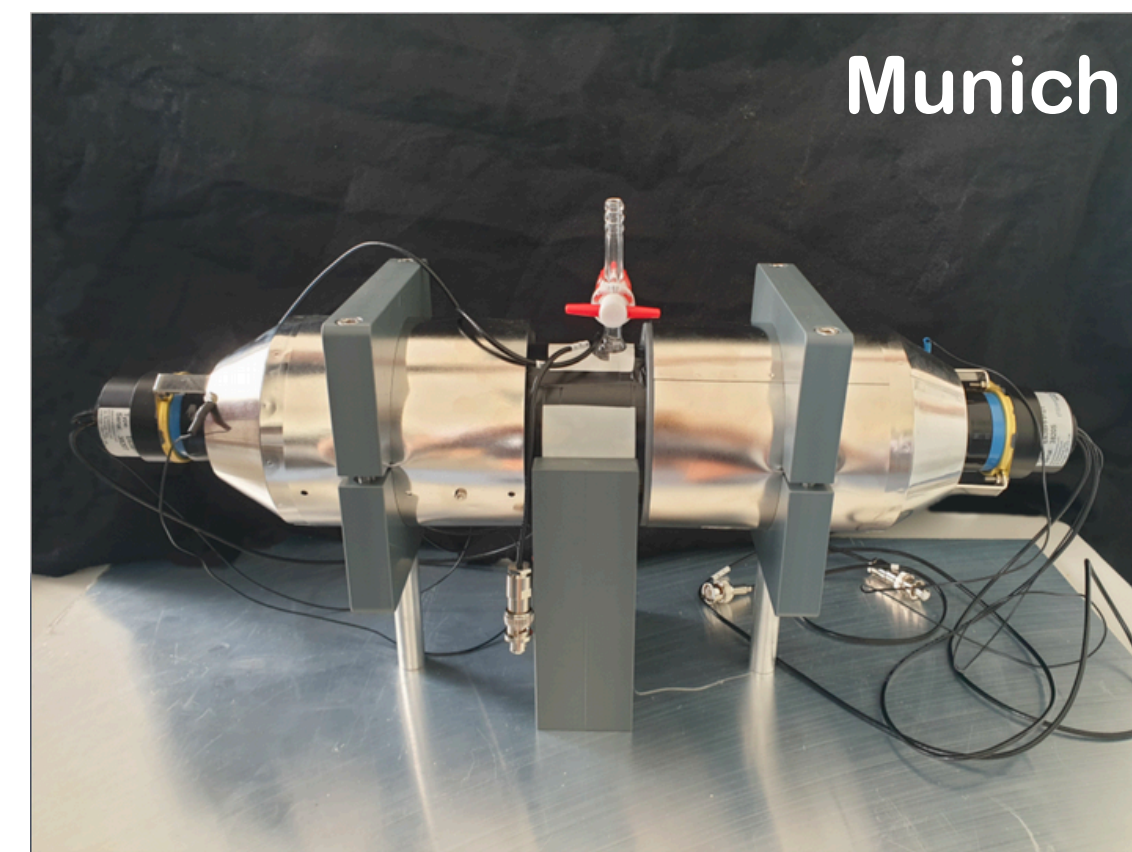
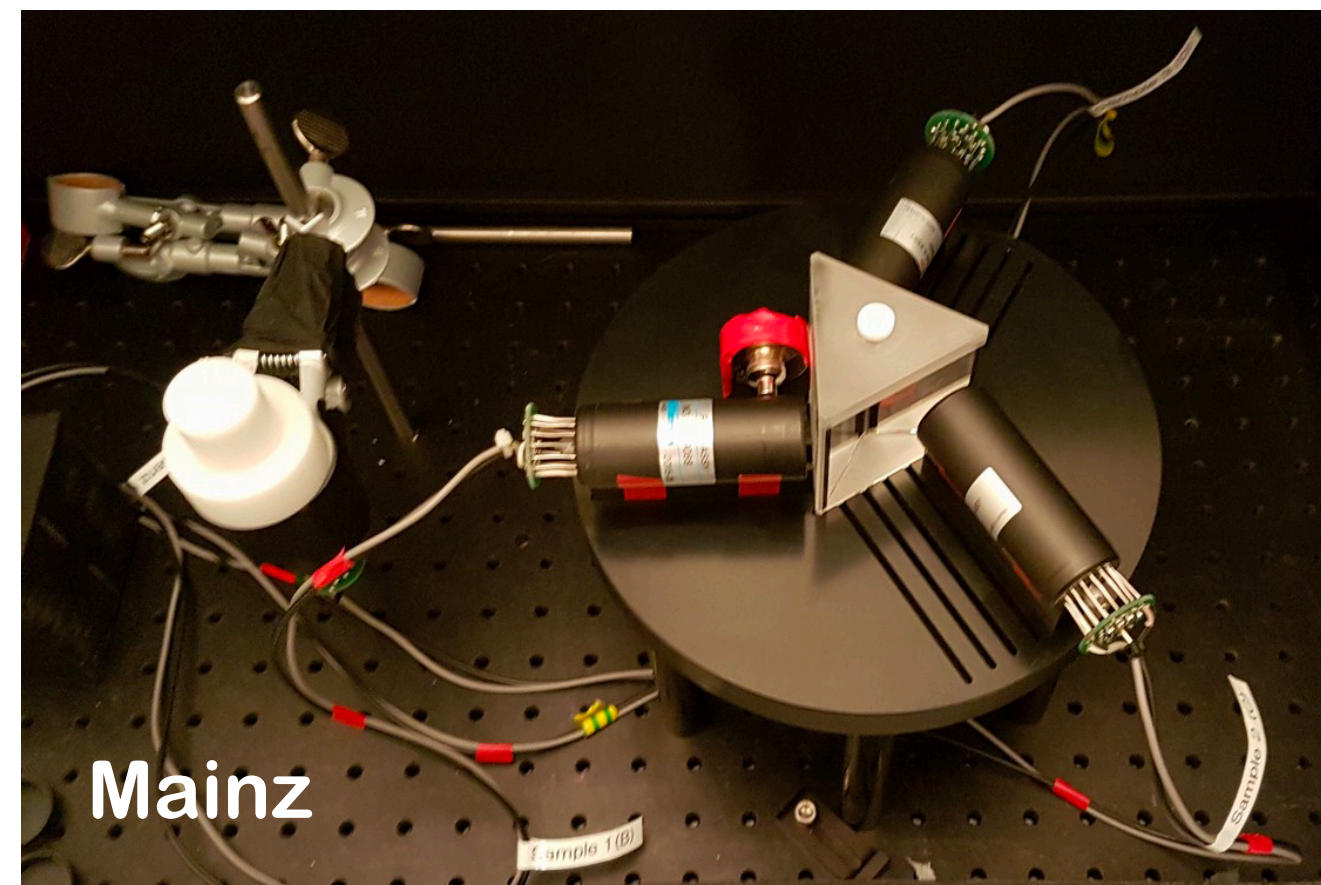
Water-based Liquid Scintillator - Advanced

Developed Water-based Liquid Scintillator (WbLS) cocktails require extensive characterization:

- Light Yield
- Emission spectrum
- Scintillation time profile
- Scattering and attenuation lengths

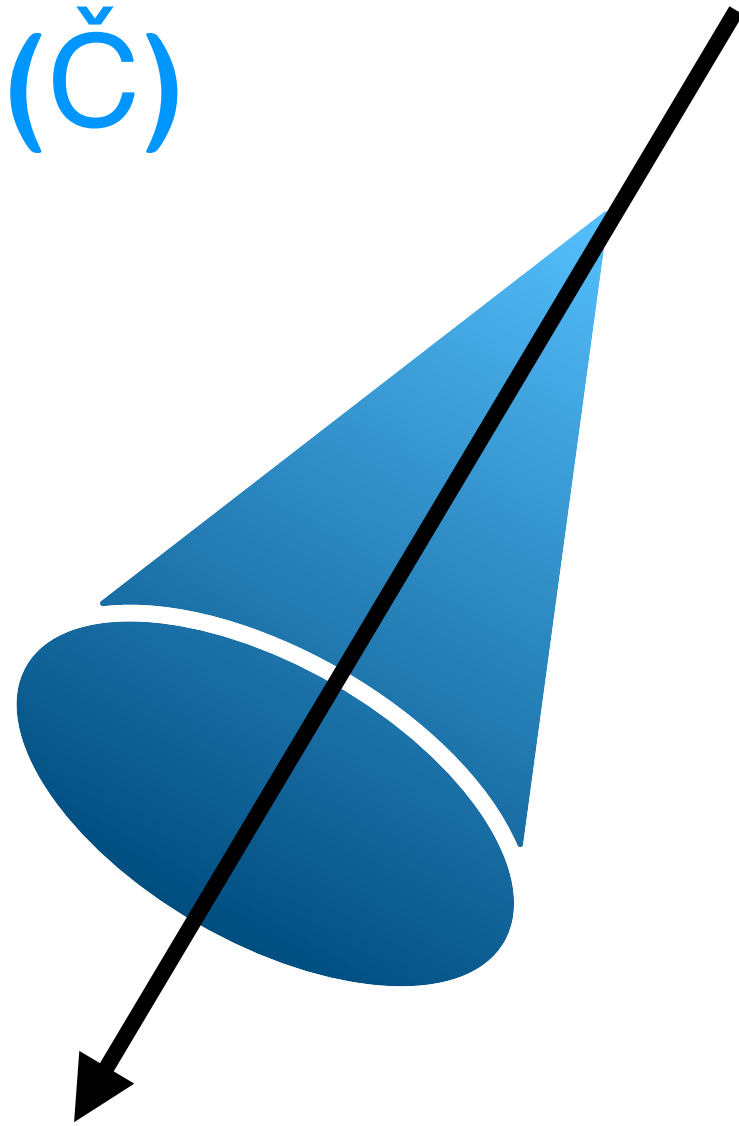
Other relevant developments:

- Nanofiltration
- Advanced reconstruction techniques, including machine learning
- Cherenkov/Scintillation separation demonstration

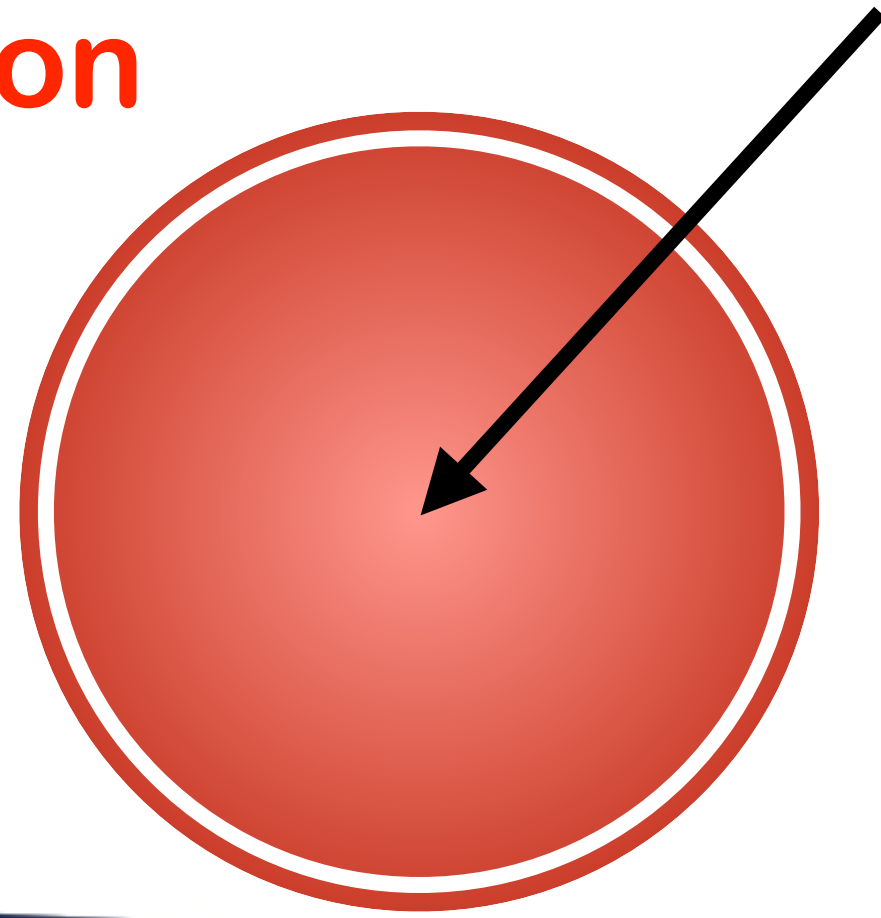


Cherenkov/scintillation photons separation

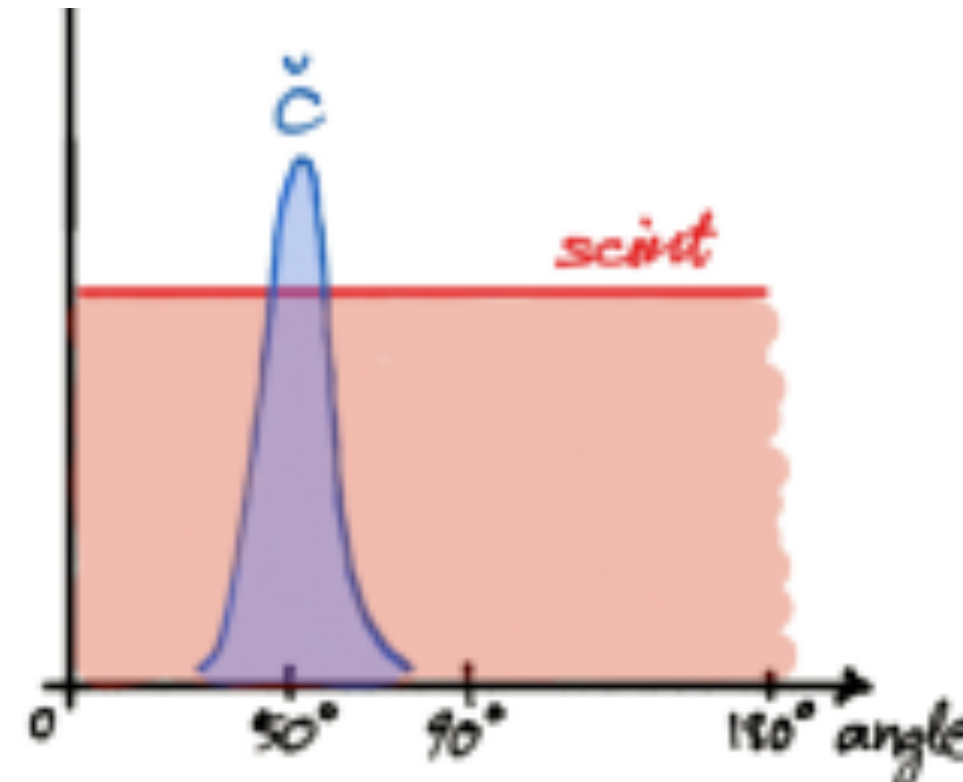
Cherenkov (\check{C})



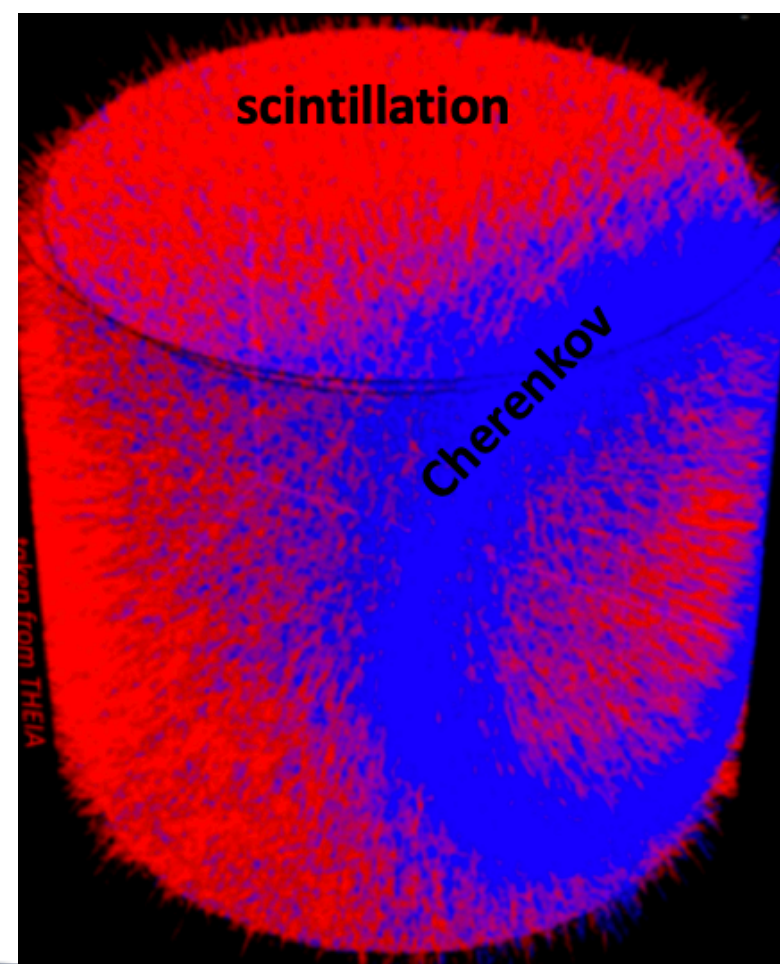
Scintillation



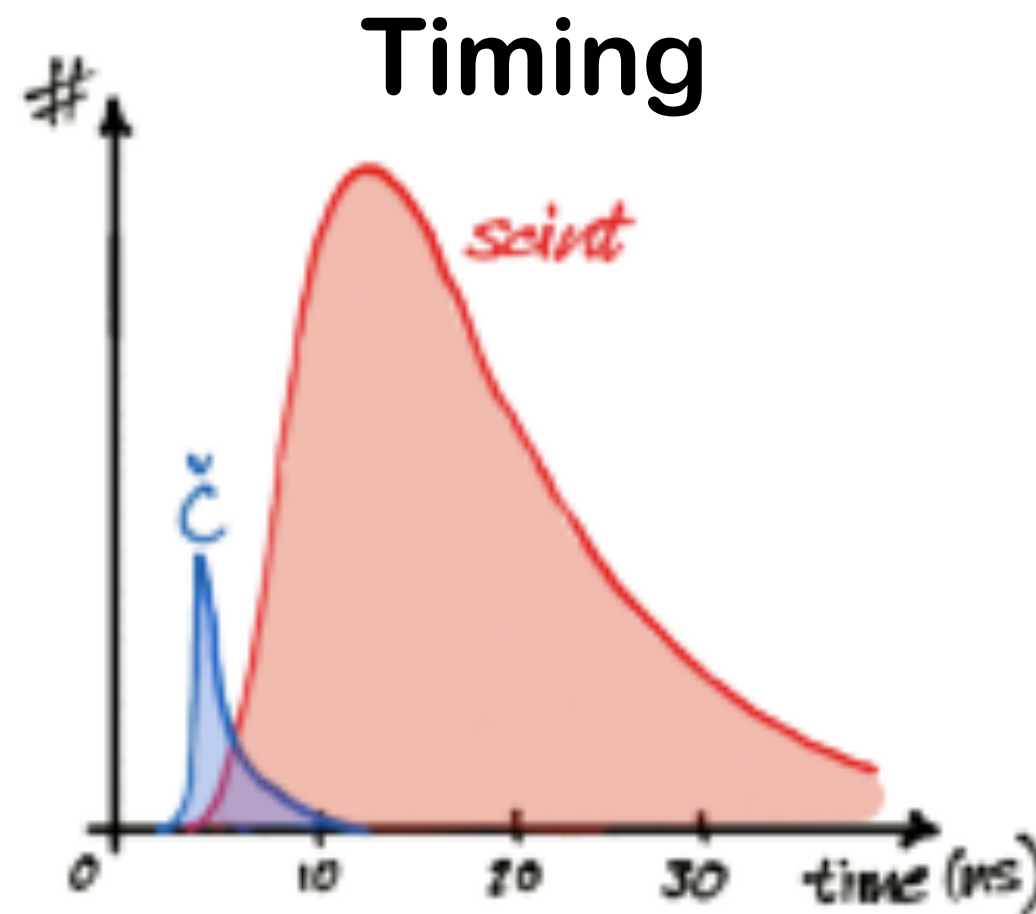
Angular distribution



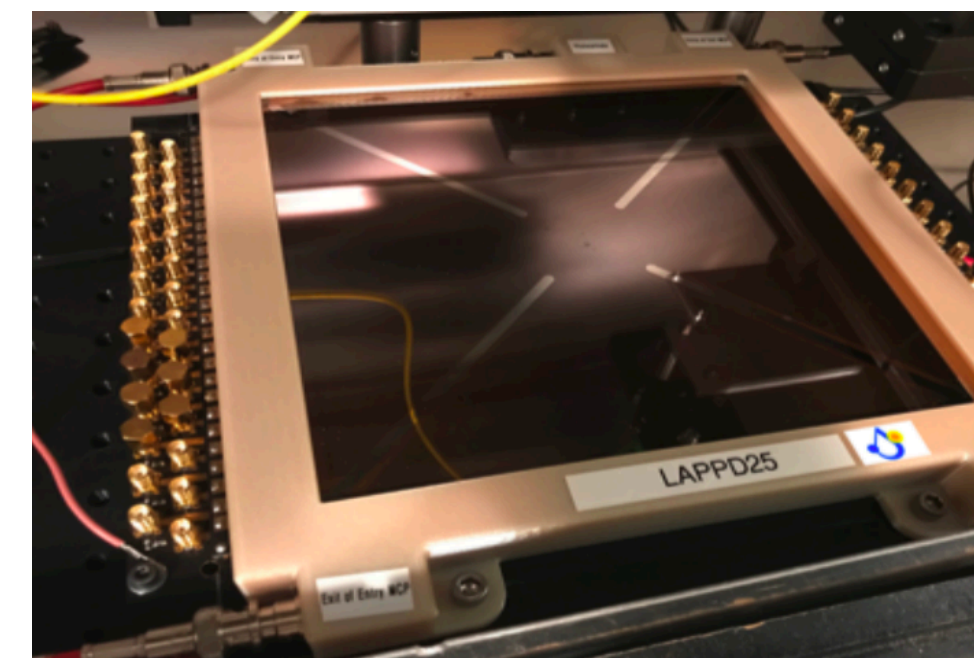
Angular resolution



Timing

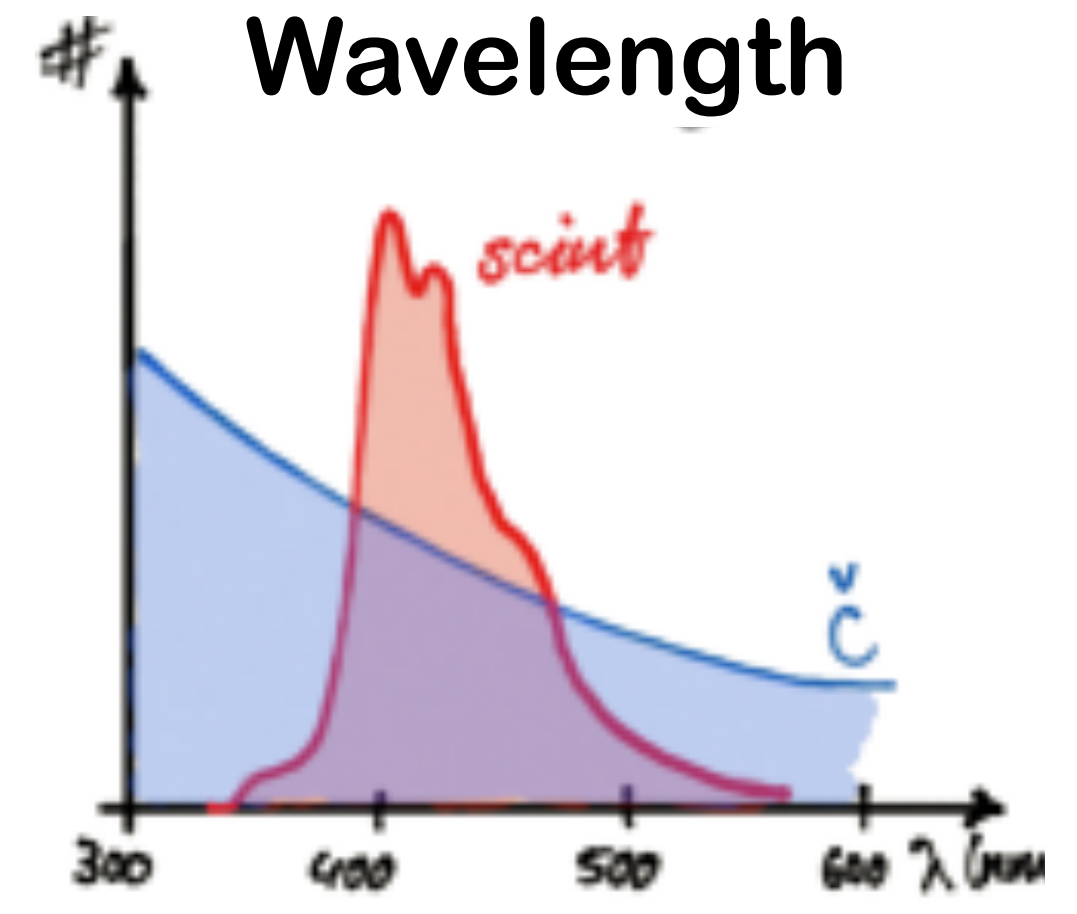


Large area picosecond photodetectors LAPPDs (~70 ps TTS) or other fast photodetectors



B.W.Adams et al. NIM A Volume 795, 1 (2015)

Wavelength



- Dichroic filters
- Red-sensitive PMTs
- Filtering



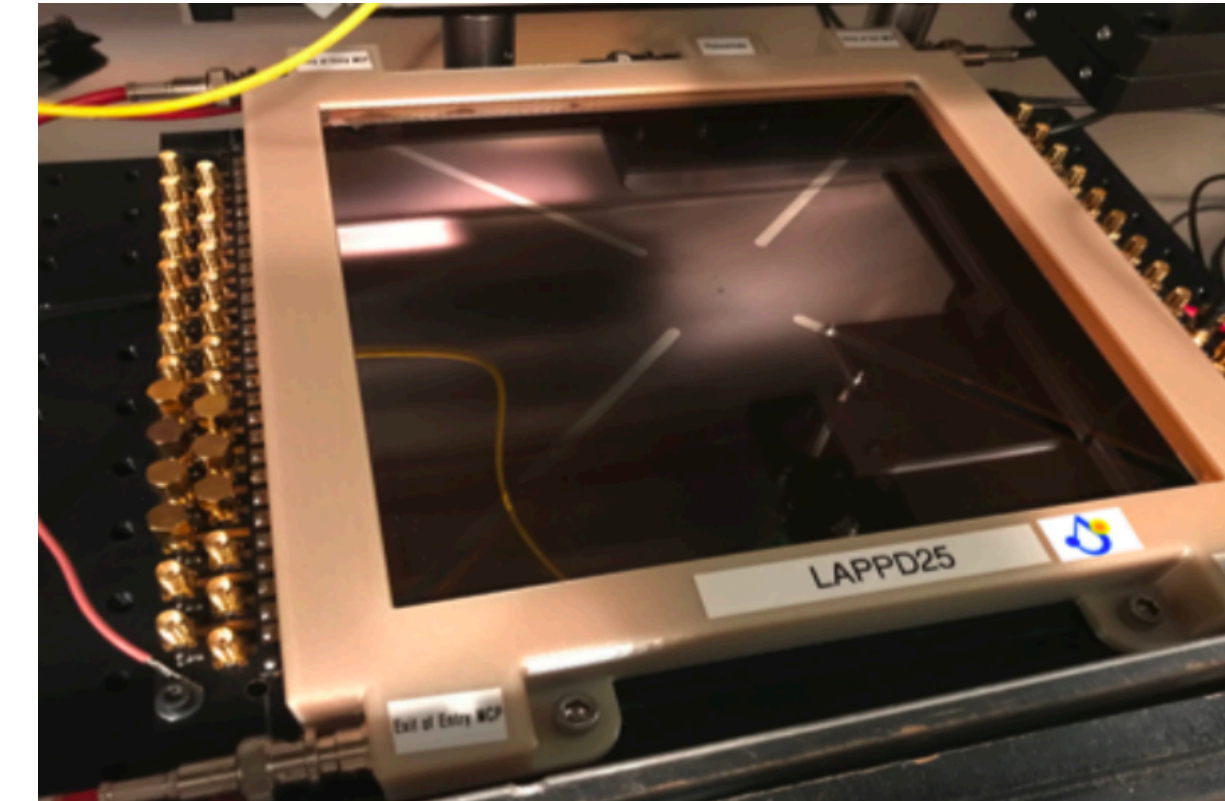
T. Kaptanoglu et al. Phys. Rev. D 101, 072002 (2020)

New Generation Photodetectors

Large area picosecond photodetector (LAPPD):

Micro-channel plate, fast-timing photodetectors

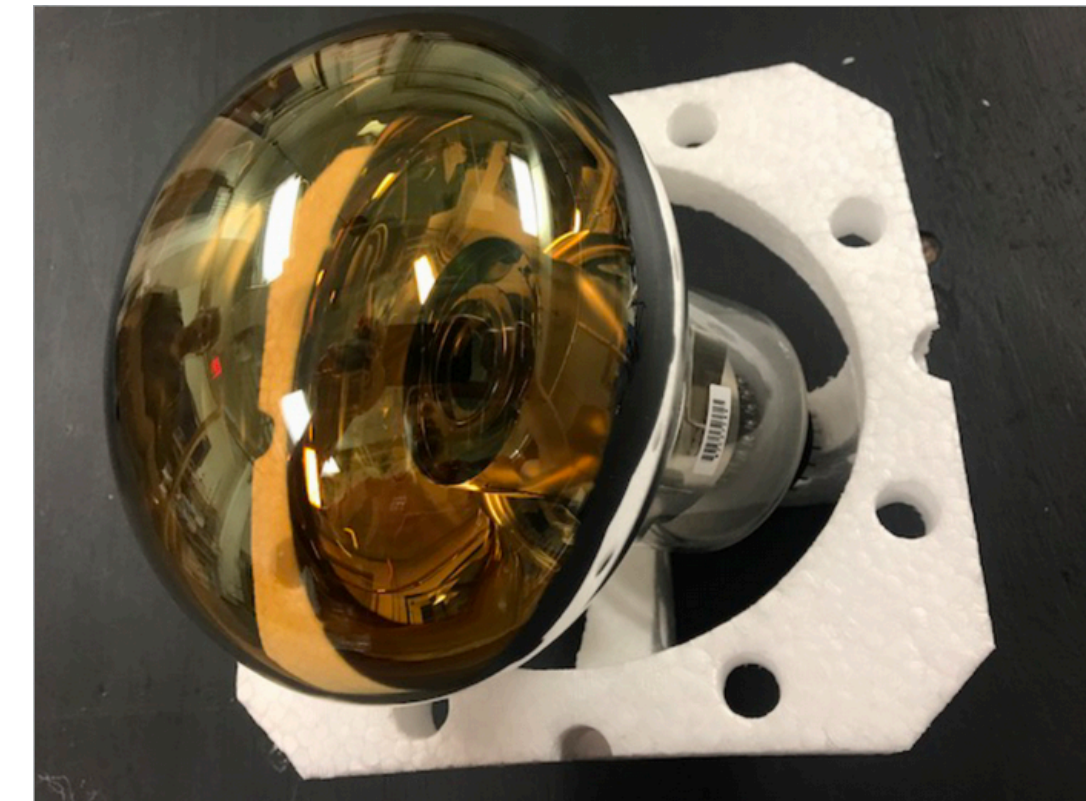
- Large-area: 20 × 20 cm
- Fast timing: ~70 ps time resolution
- High quantum efficiency (QE): >20-30 %
- Position resolution: mm scale



Very fast large-area & HQE PMT

TTS~500ps

Q.E. > 30%



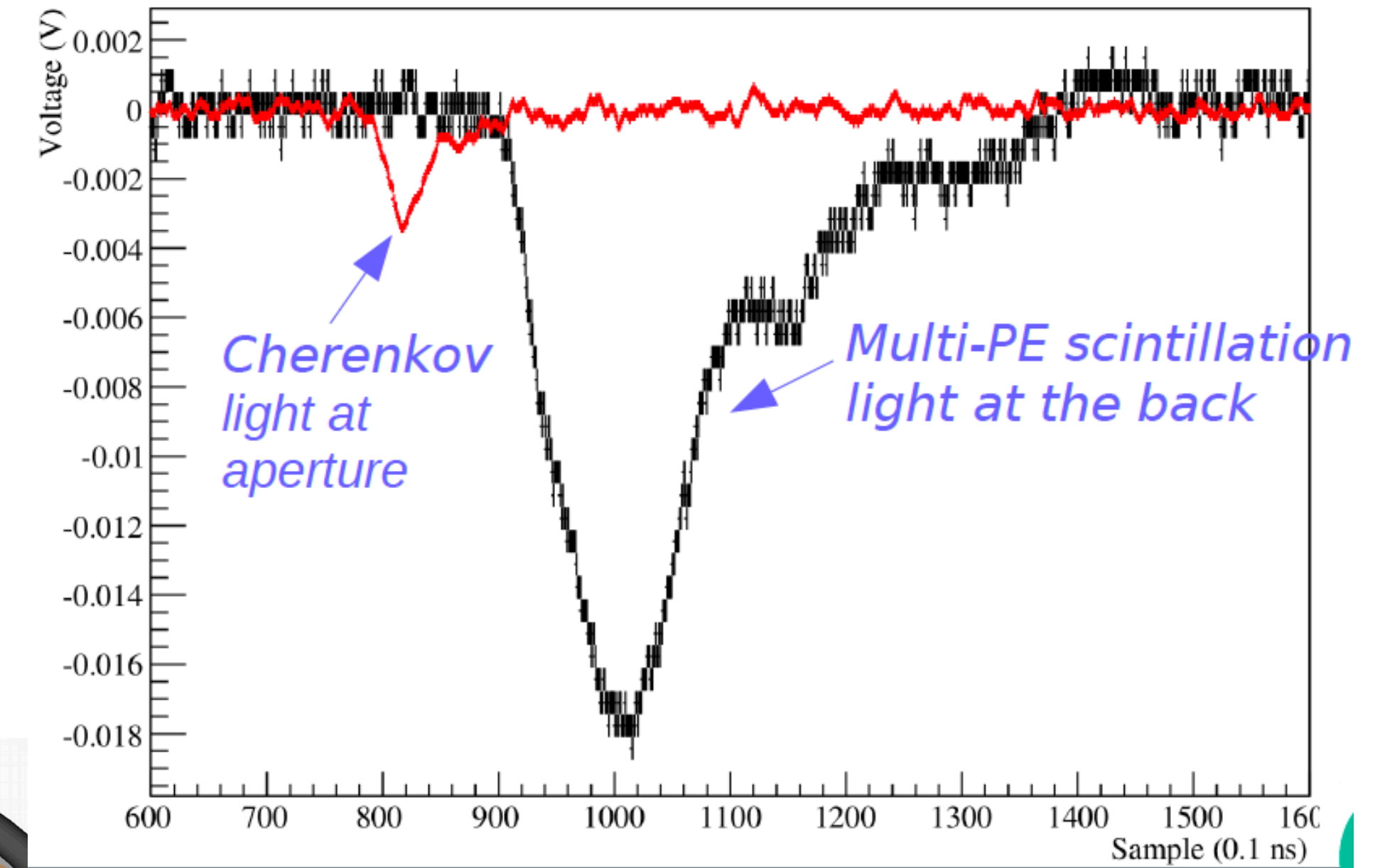
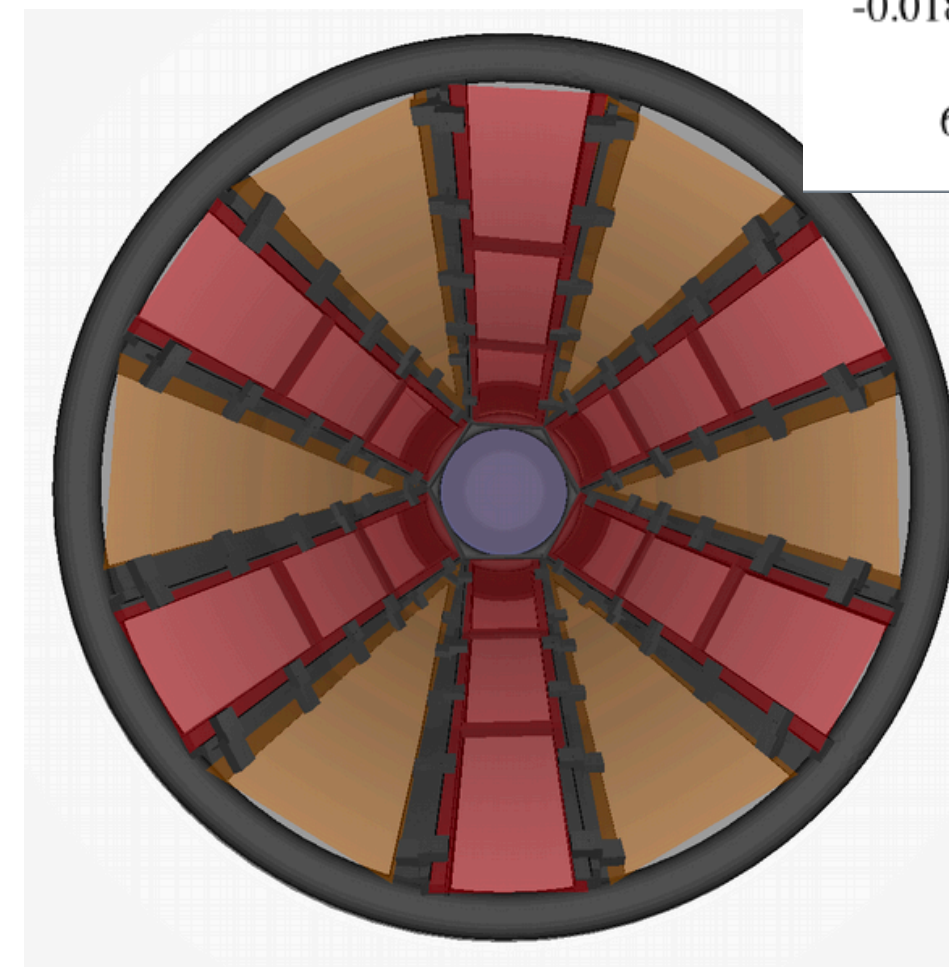
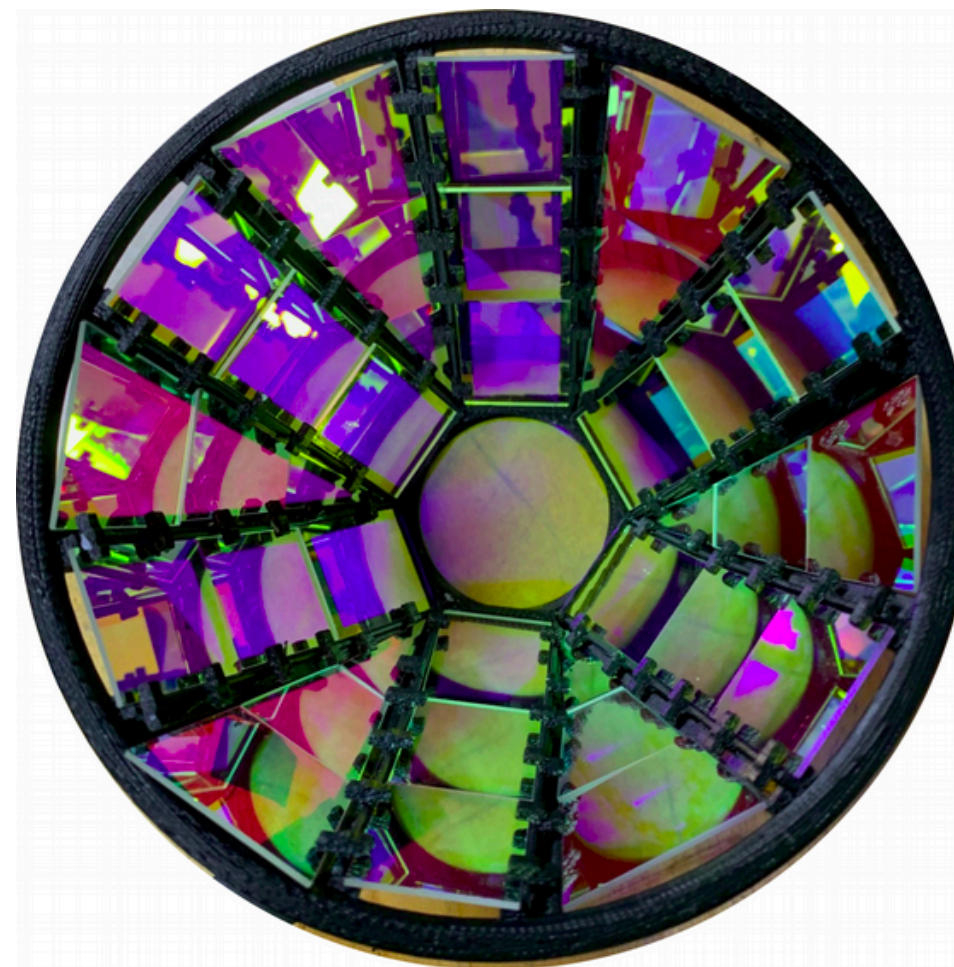
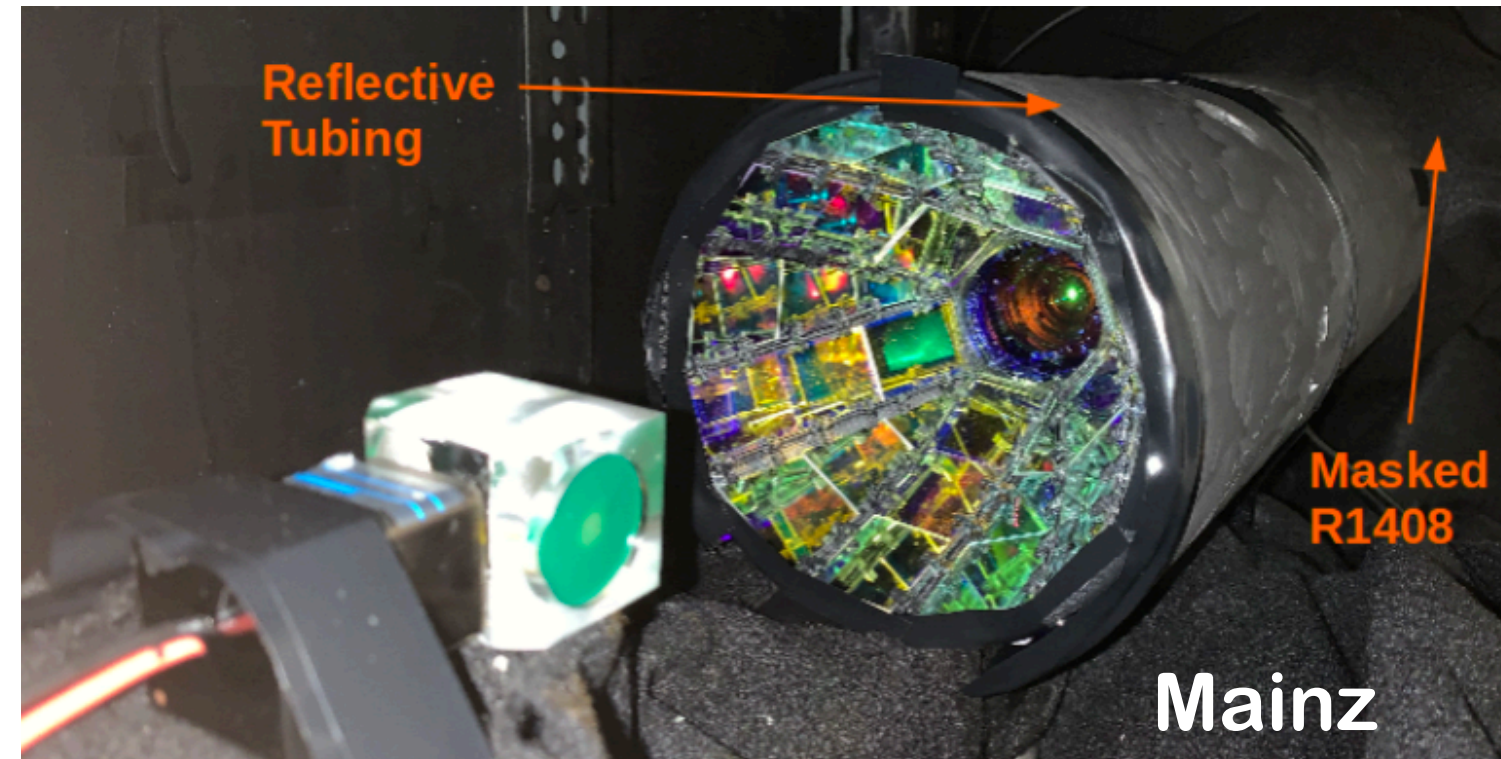
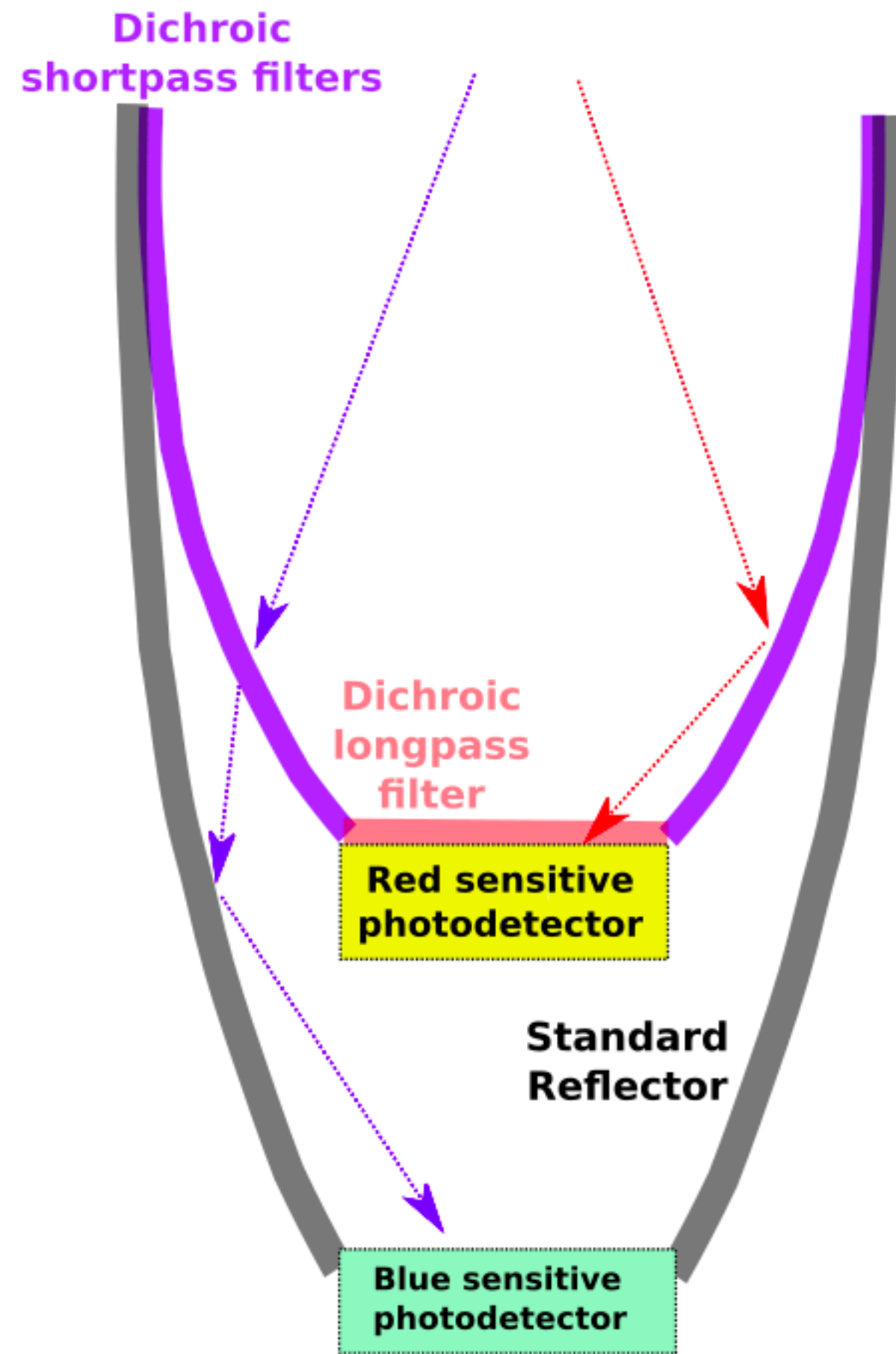
Large-area red-sensitive PMTs

TTS~500ps

Q.E. > 30%



Dichroic filters (Wavelength discrimination)



T. Kaptanoglu et al JINST 14 T05001 (2019)
T. Kaptanoglu et al. Phys. Rev. D 101, 072002 (2020)

Performance measurements and MC studies

WbLS characterization:

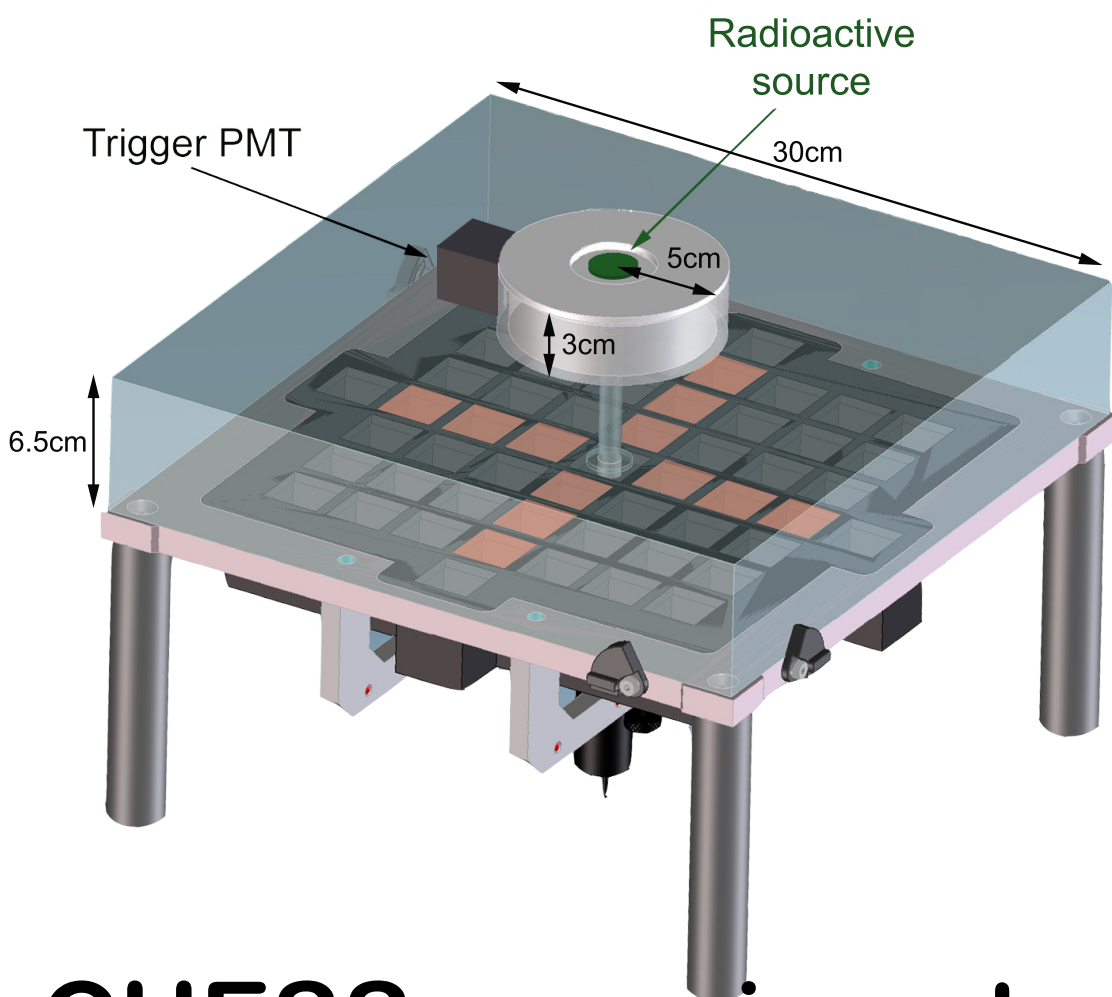
- Scintillation light yield
- Emission time profile with betas and X-rays
- Emission spectra

Monte Carlo model construction

Cherenkov/scintillation separation at small scale using muons

Monte Carlo model validation

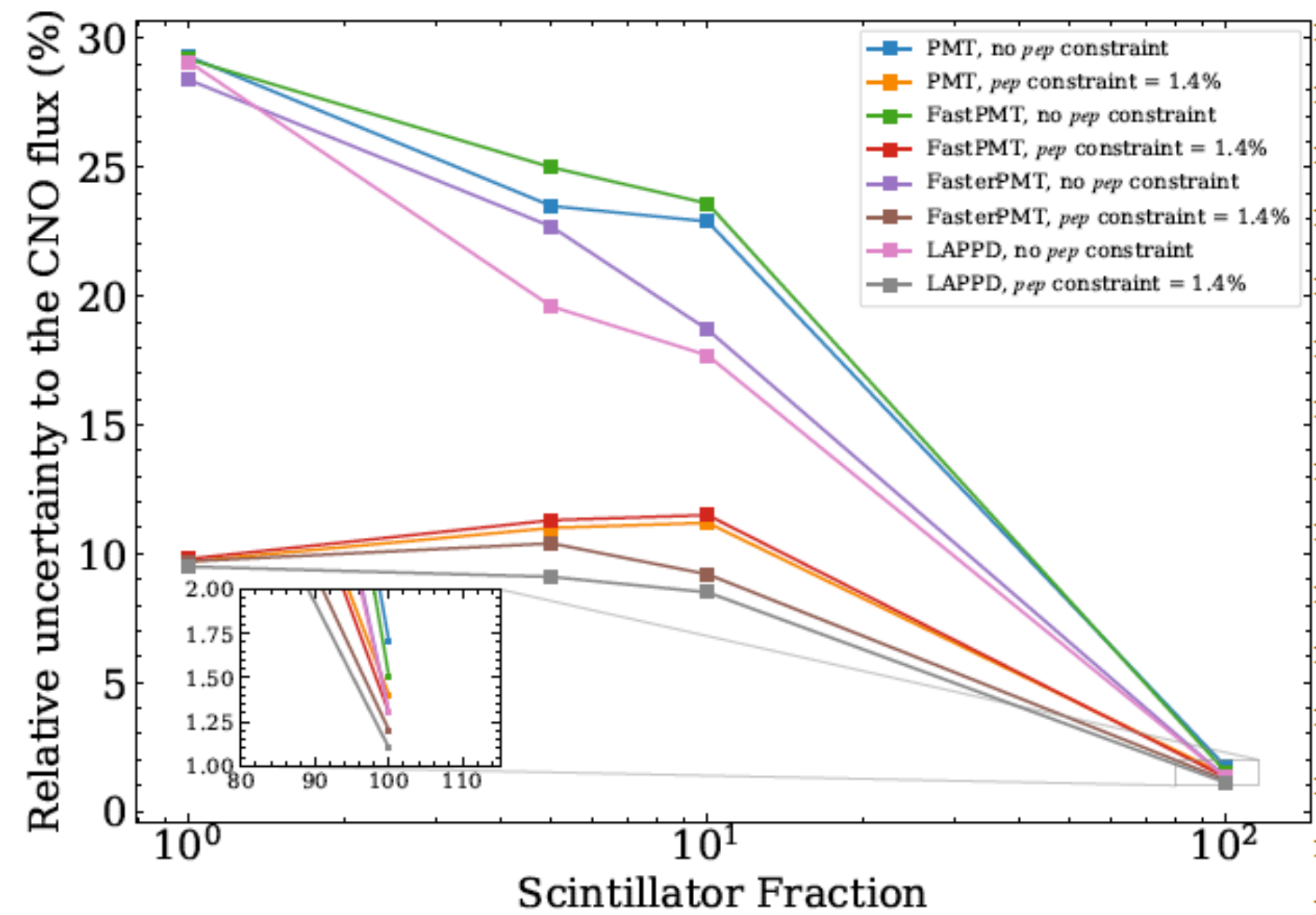
Prediction of impact in large-scale detectors



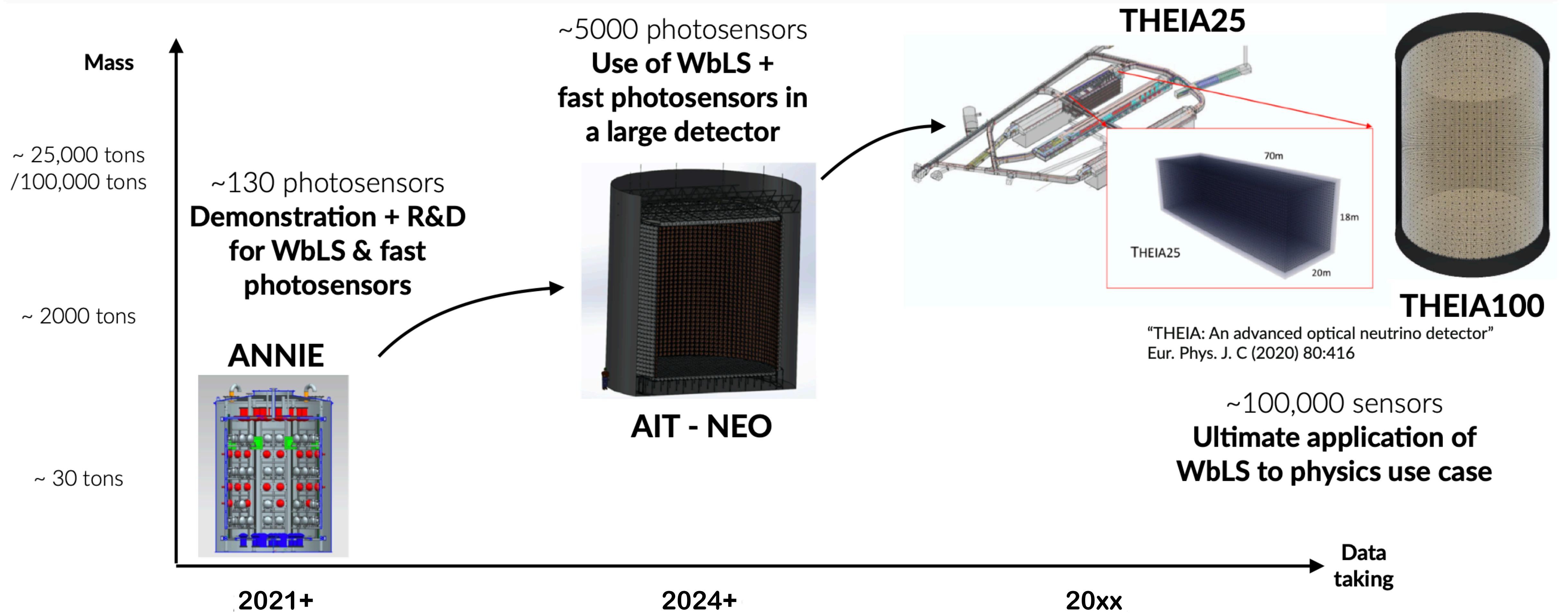
CHES experiment at UC Berkeley

J. Caravaca et al. Eur. Phys. J. C 80, 867 (2020)
 D. Onken et al., Mater. Adv., 1, 71-76 (2020)
 J. Caravaca et al. Eur. Phys. J. C 77, 811 (2017)
 J. Caravaca et al. Phys. Rev. C 95, 055801 (2017)

Example of impact on the CNO measurement precision:



Scaling up



ANNIE @FermiLab



Main Goal:

understanding neutrino-nucleus interactions, focusing on production and multiplicity of final-state neutrons

Show feasibility of WbLS in a neutrino-beam environment

CV

Start of data taking: ~2016
(WbLS ~2021-2022)

Location: Booster Neutrino Beam @ FermiLab (USA)

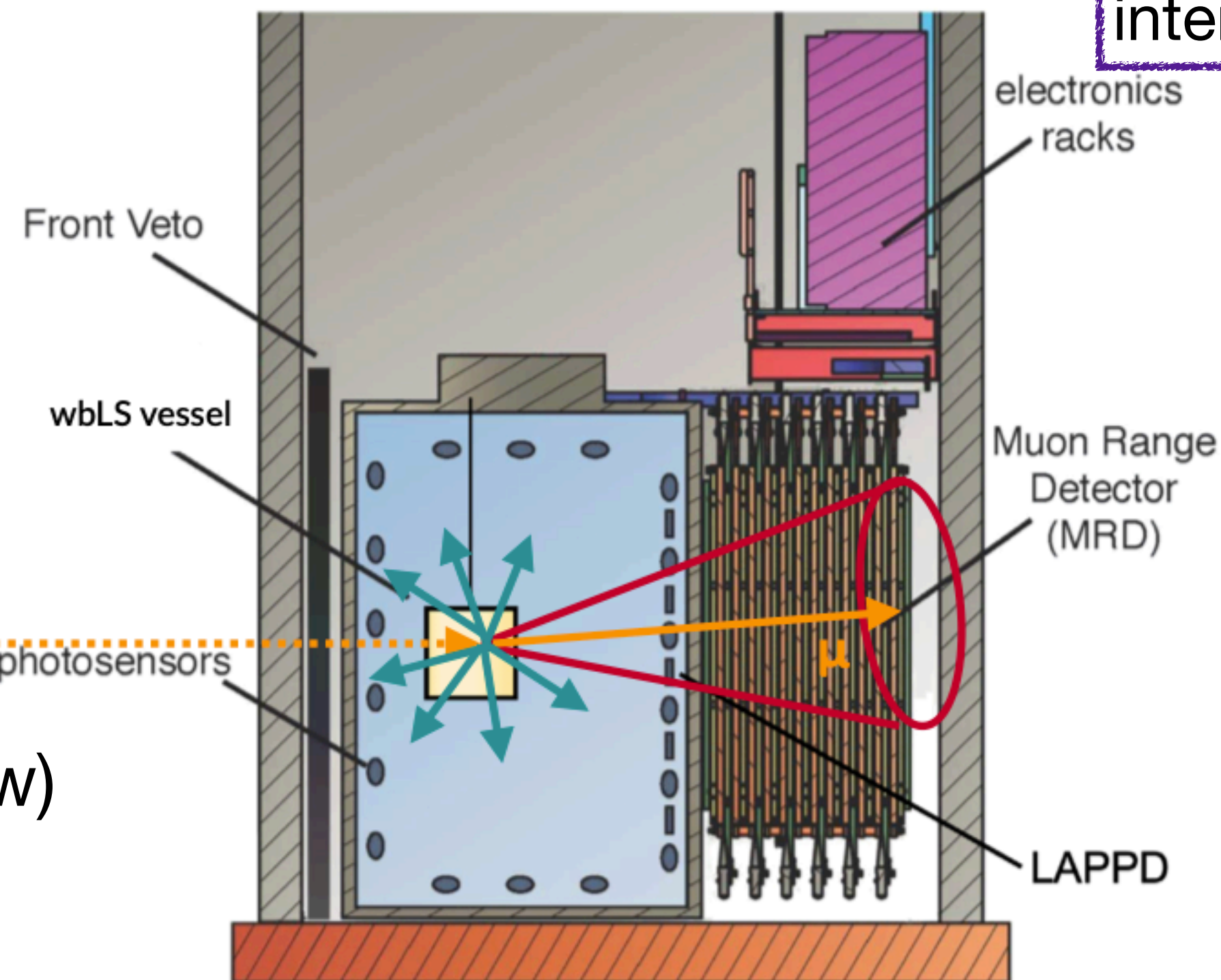
Water-volume: 26t

WbLS volume: 0.5t

Interests: neutrino-nucleus interactions



Beam neutrino



WbLS vessel



Rough visualization of vessel
(NOT final design)

Interest to Theia:

- First deployment of LAPPDs (happening now)
- Deployment of 0.5 t WbLS
- C/S separation in a large-scale experiment
- High and low-energy events reconstruction
- Neutron detection

Advanced Instrumentation Testbed Neutrino Experiment One (AIT/NEO)



Neutrino Experiment One (NEO) the first demonstration of reactor monitoring in the far-field

Main Goal:

non-intrusively detect the ON/OFF power cycle of a single reactor

Interest to Theia:

- Deployment of kt scale WbLS
- Low energy antineutrinos detection in WbLS

CV

Start of data taking:

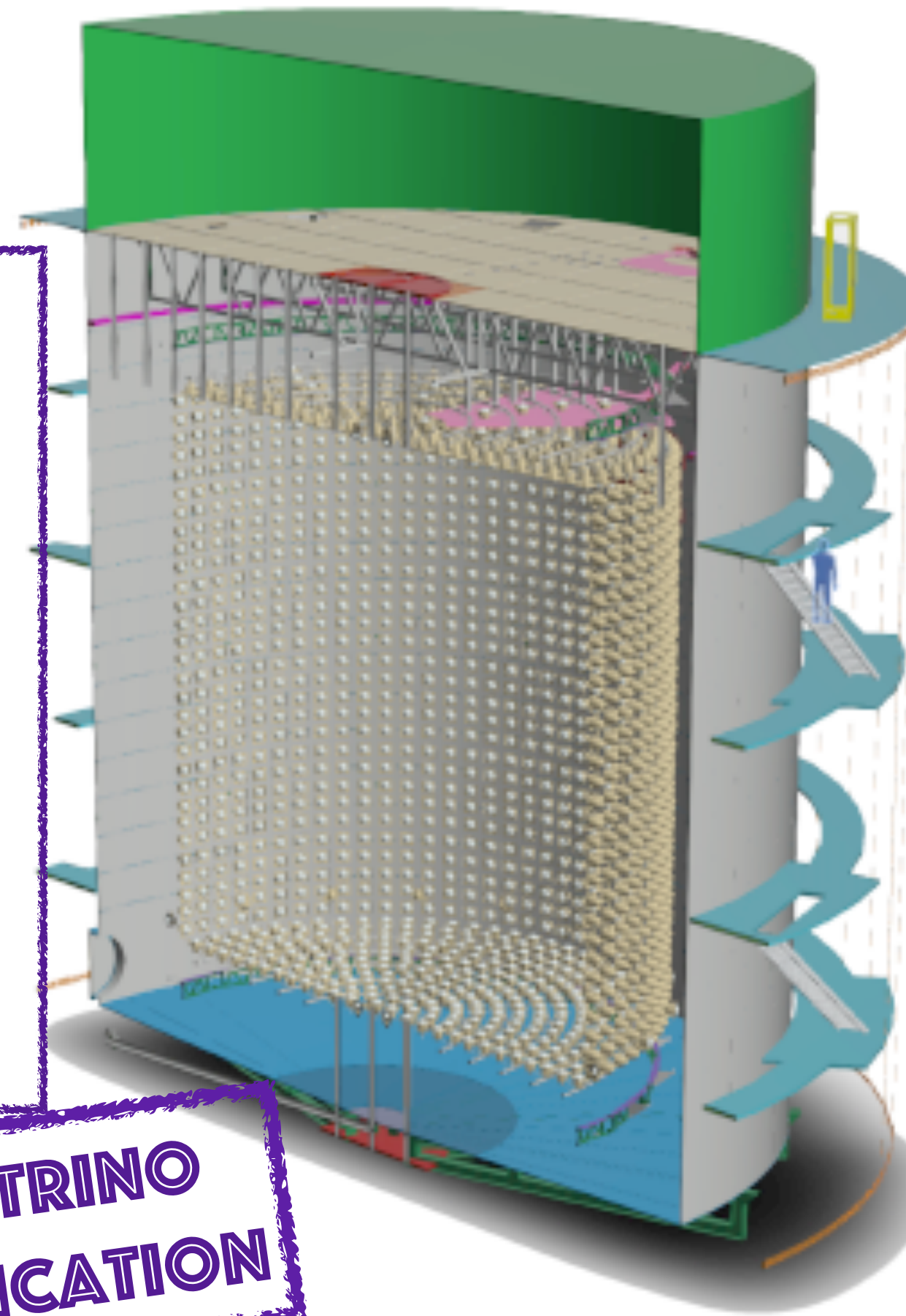
~2024

Location: Boulby Underground Lab (UK)

WbLS volume: 1kt

Interests:

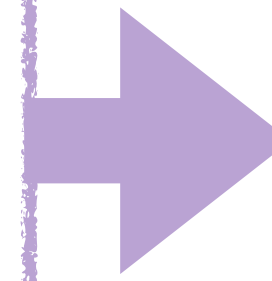
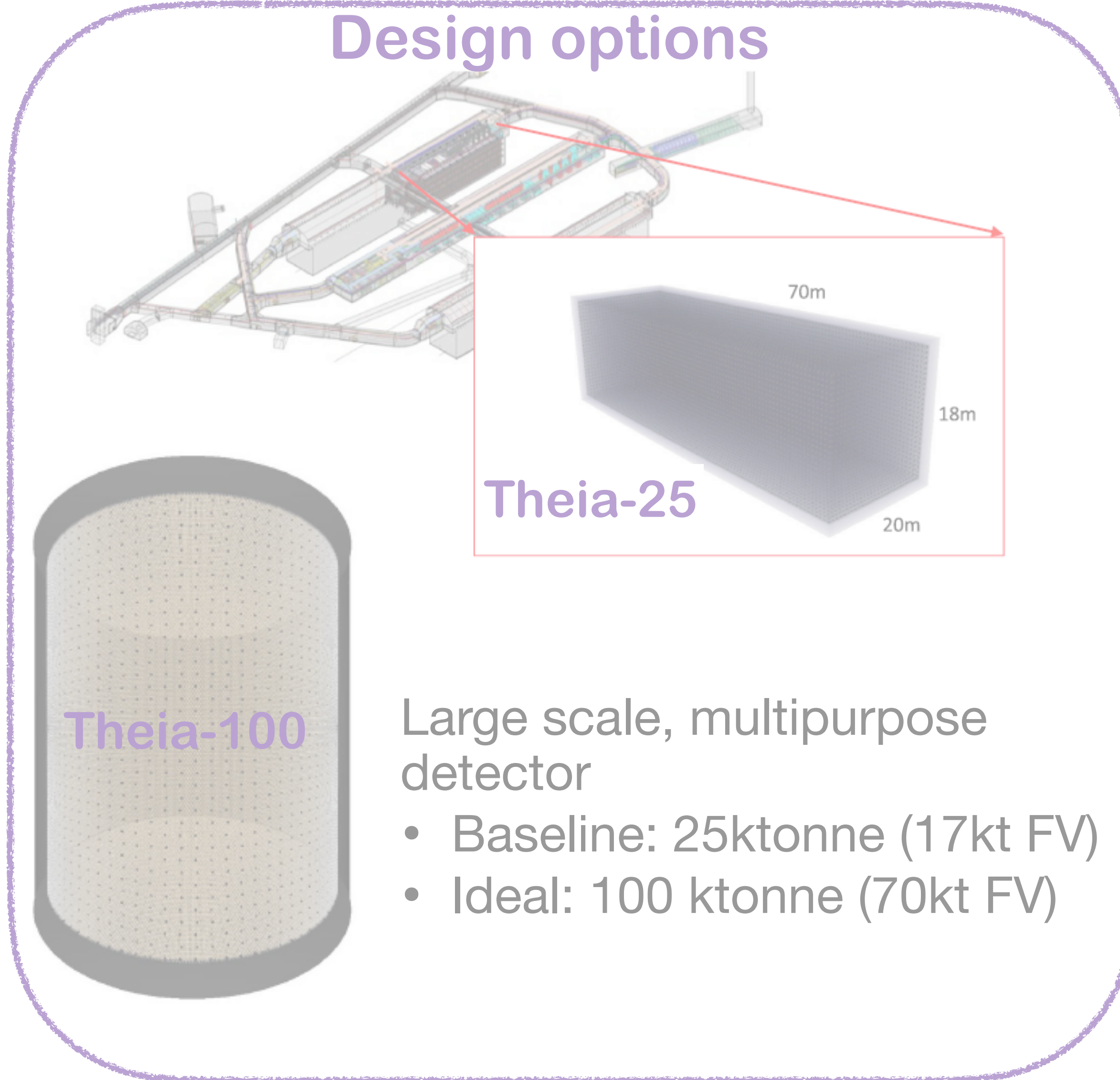
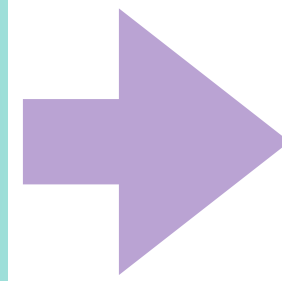
Non-proliferation, Solar neutrinos, NLDBD



NEUTRINO APPLICATION

Theia: advanced optical multipurpose neutrino detector

Cutting edge developments in the target material and photodetection



Broad physics program:
Studying neutrino
fundamental properties
and astrophysical
objects

Theia: multipurpose neutrino detector

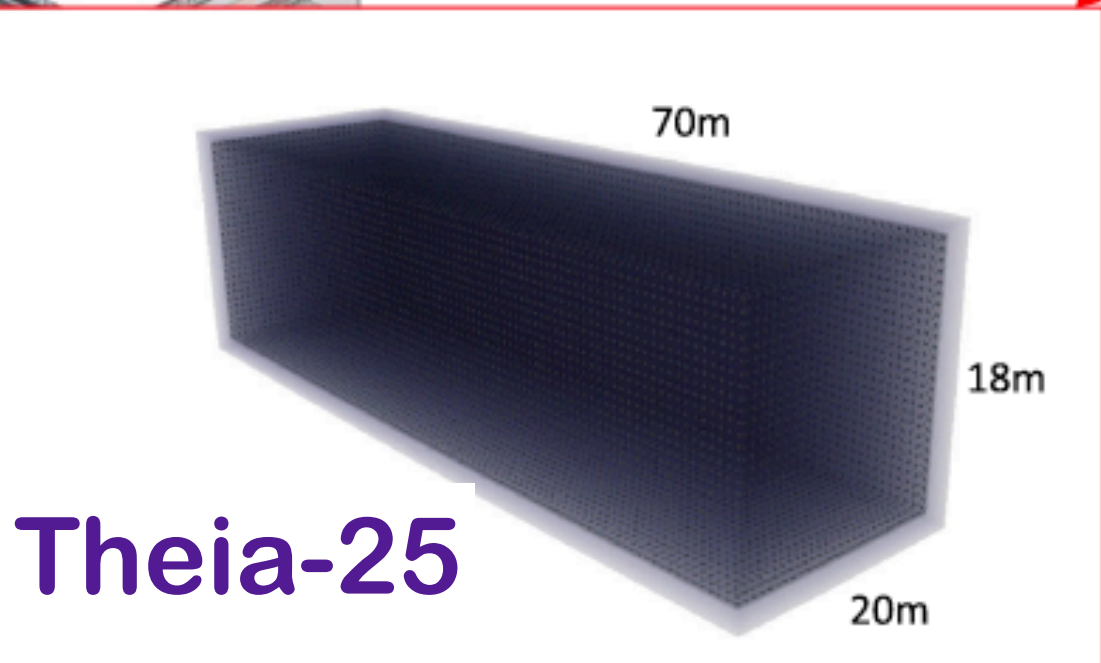
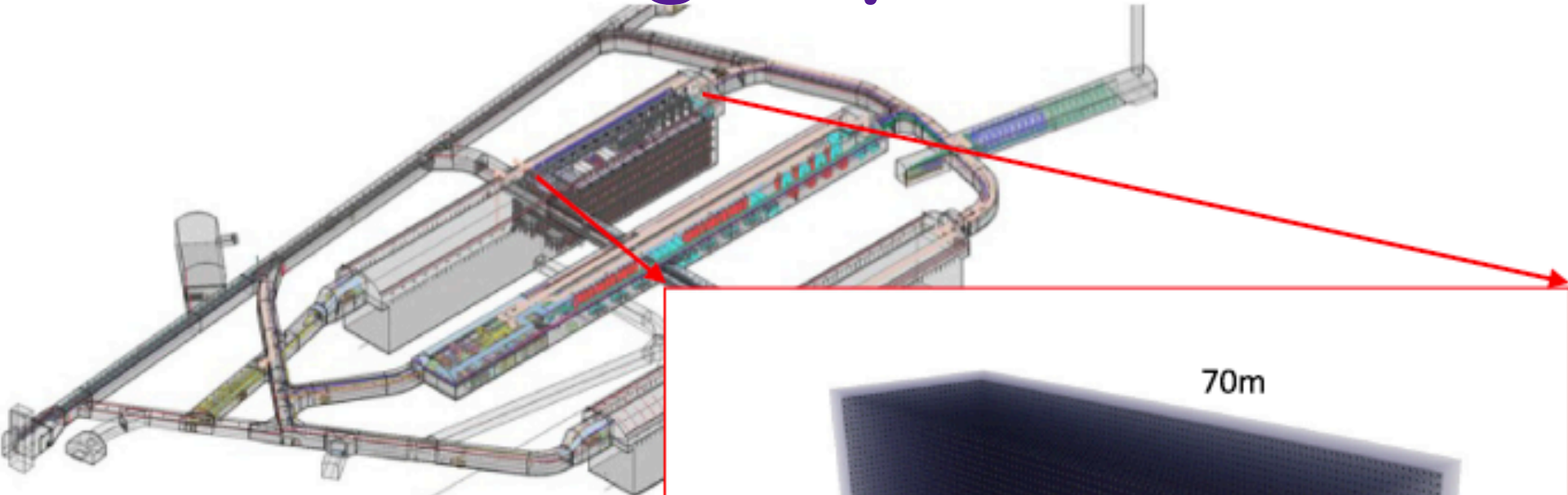
solar neutrinos
(CNO, ^8B)

geoneutrinos


diffuse supernova
neutrinos (DSNB)

supernova burst
neutrinos

Design options



Theia-25



Theia-100

Large scale, multipurpose detector

- Baseline: 25ktonne (17kt FV)
- Ideal: 100 ktonne (70kt FV)

Scintillator fraction tunable depending on the physics goal
->staged approach

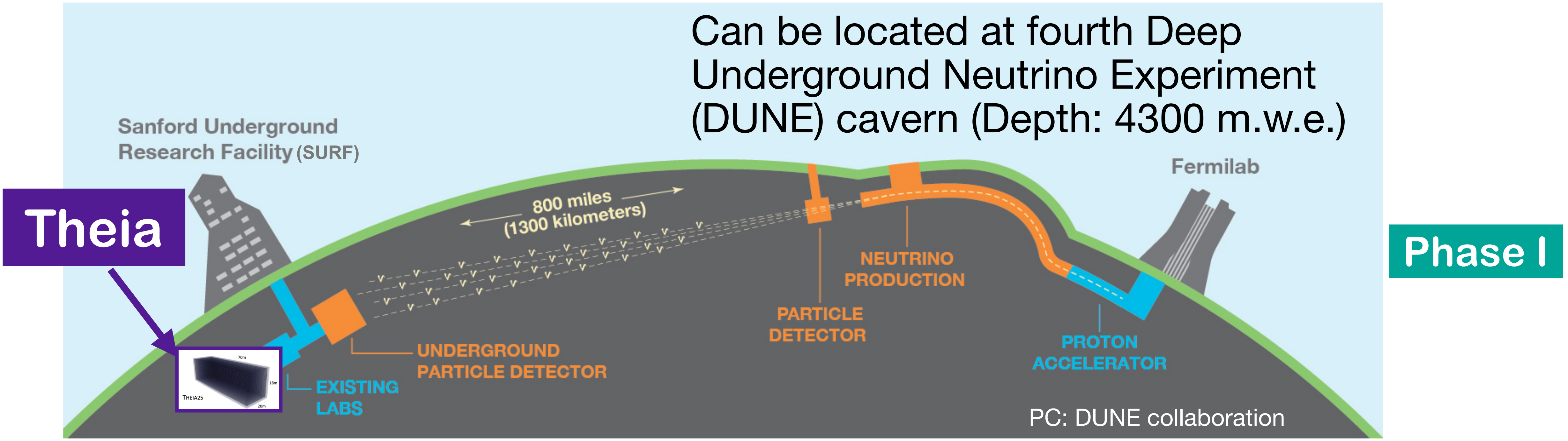
neutrino mass
ordering

neutrino CP-
violating phase δ

neutrinoless
double beta decay

nucleon decay

Theia-25 at long baseline neutrino facility (LBNF)



- Using Fermilab's LBNF neutrino beam for **long-baseline neutrino oscillation measurements**
- Precision measurement of neutrino CP-violating phase δ

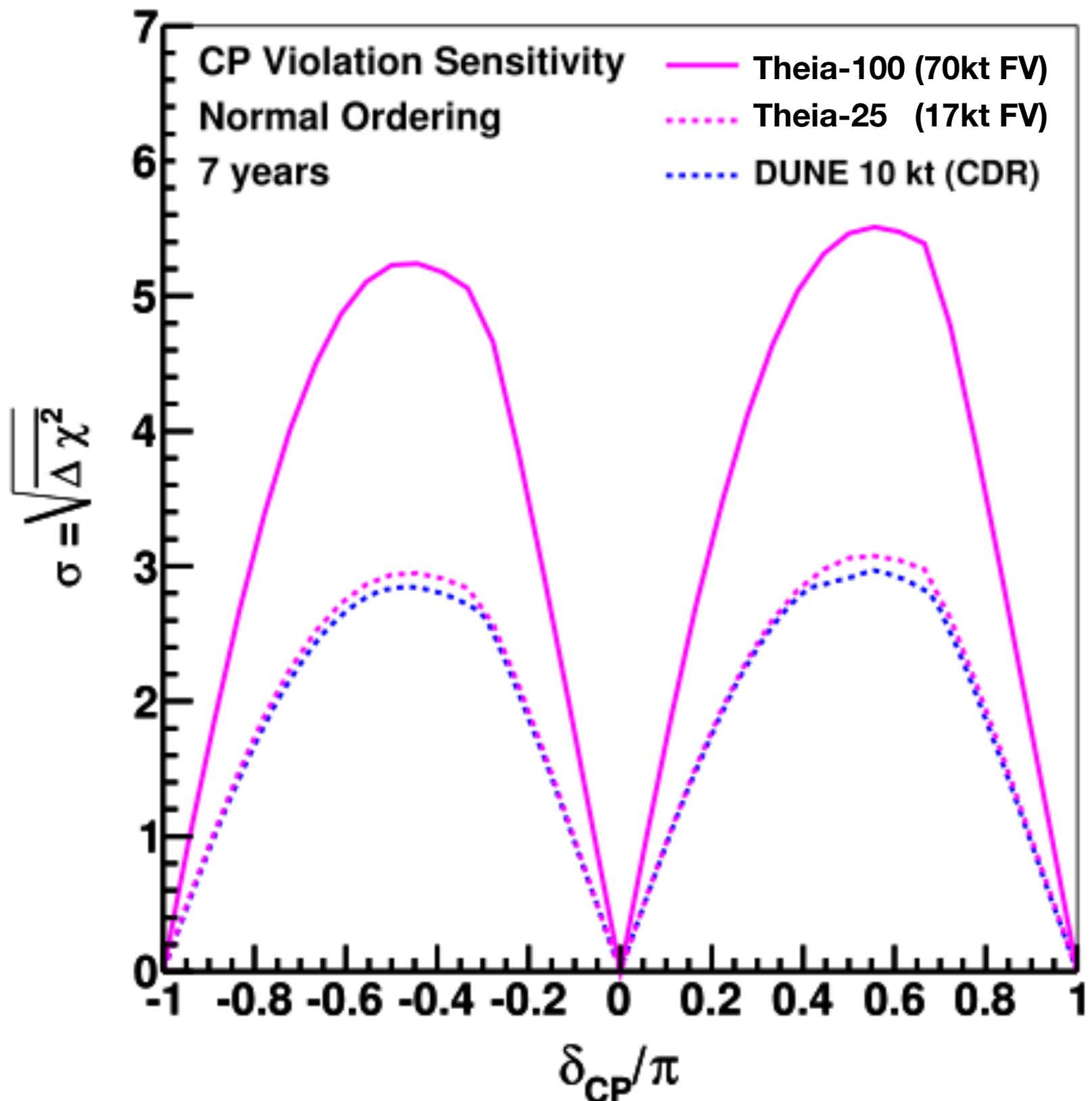
Theia: oscillation parameters

Theia can complement DUNE measurements (same location, different target, systematics) - important cross-check

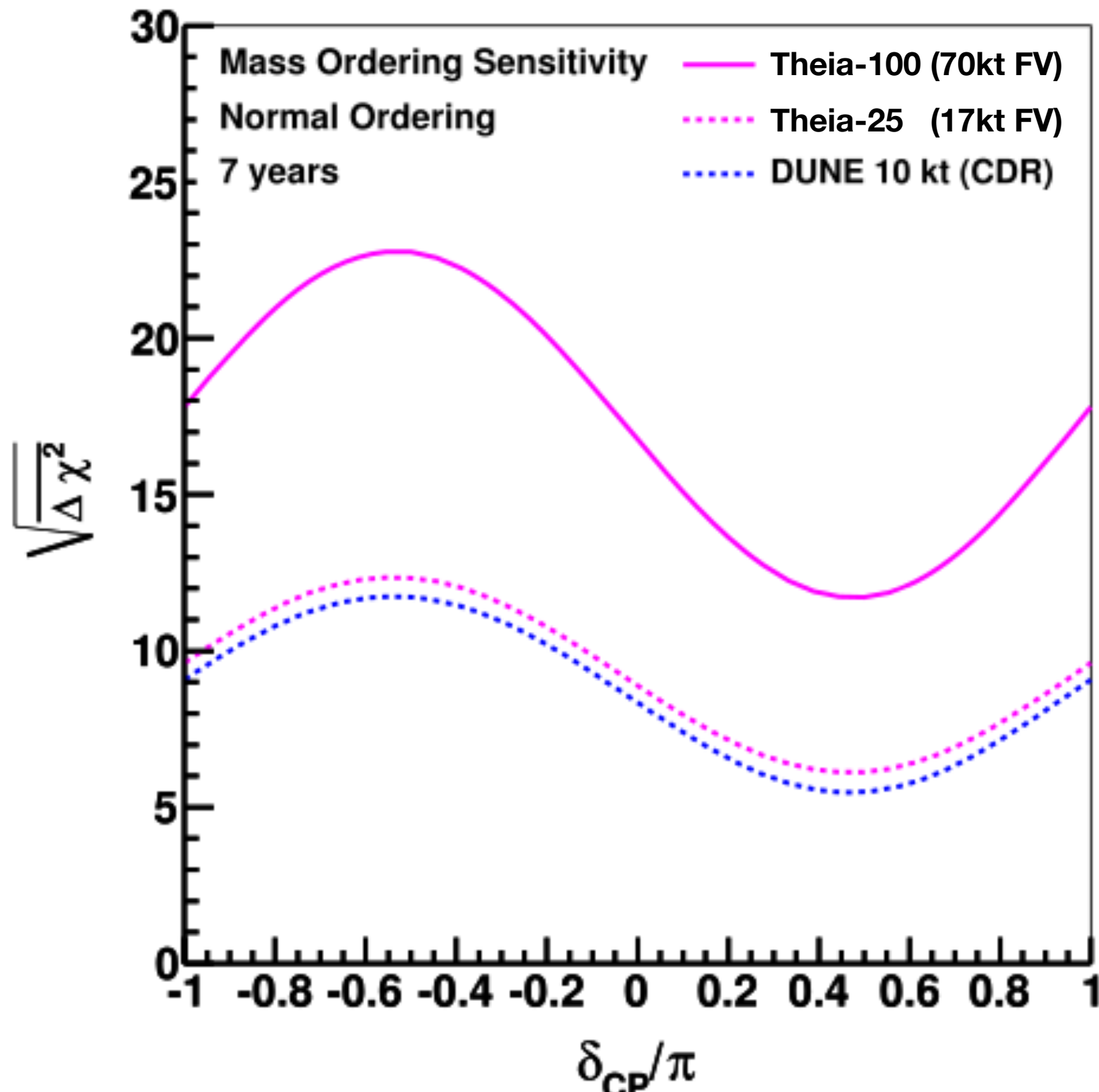
$$P(\nu_\mu \rightarrow \nu_\mu) \simeq 1 - 2\sin(2\theta_{23})\sin^2\frac{1.27\Delta m_{32}^2 L}{E} + f(\delta_{CP}) + \mathcal{O}(\theta_{13}^2)$$

- Comparison of the unoscillated flux (measured close to the beam source) and oscillated flux at far distance.
- Combination of 3D scintillator tracker with Theia more similar to T2K
- Long baseline (1300km): more matter -> more sensitivity to mass hierarchy
- > 5σ for 30% of δ_{CP} values (524 kt-MW-year)

CP Violation Sensitivity



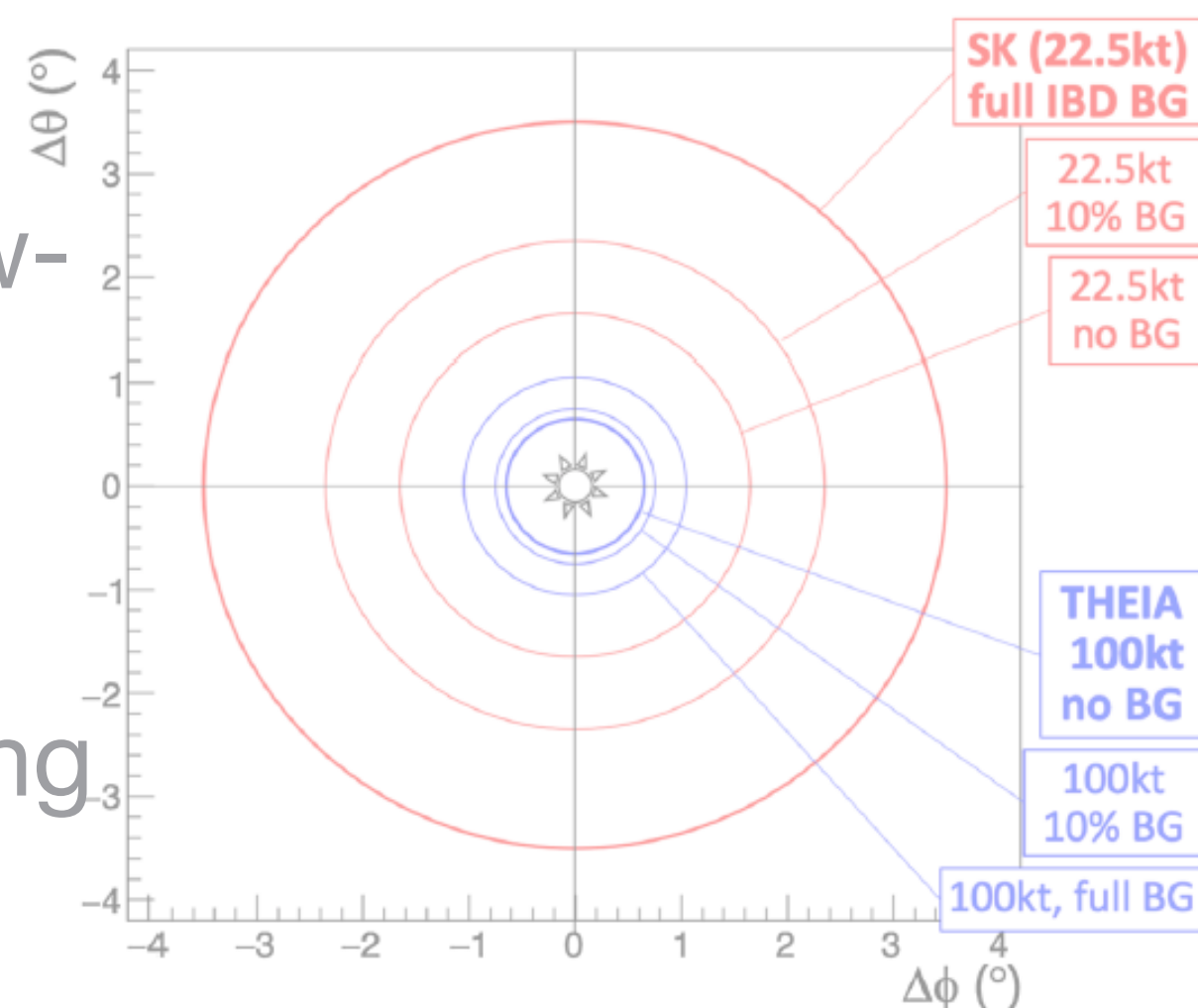
Mass Ordering Sensitivity



Theia: supernova neutrinos

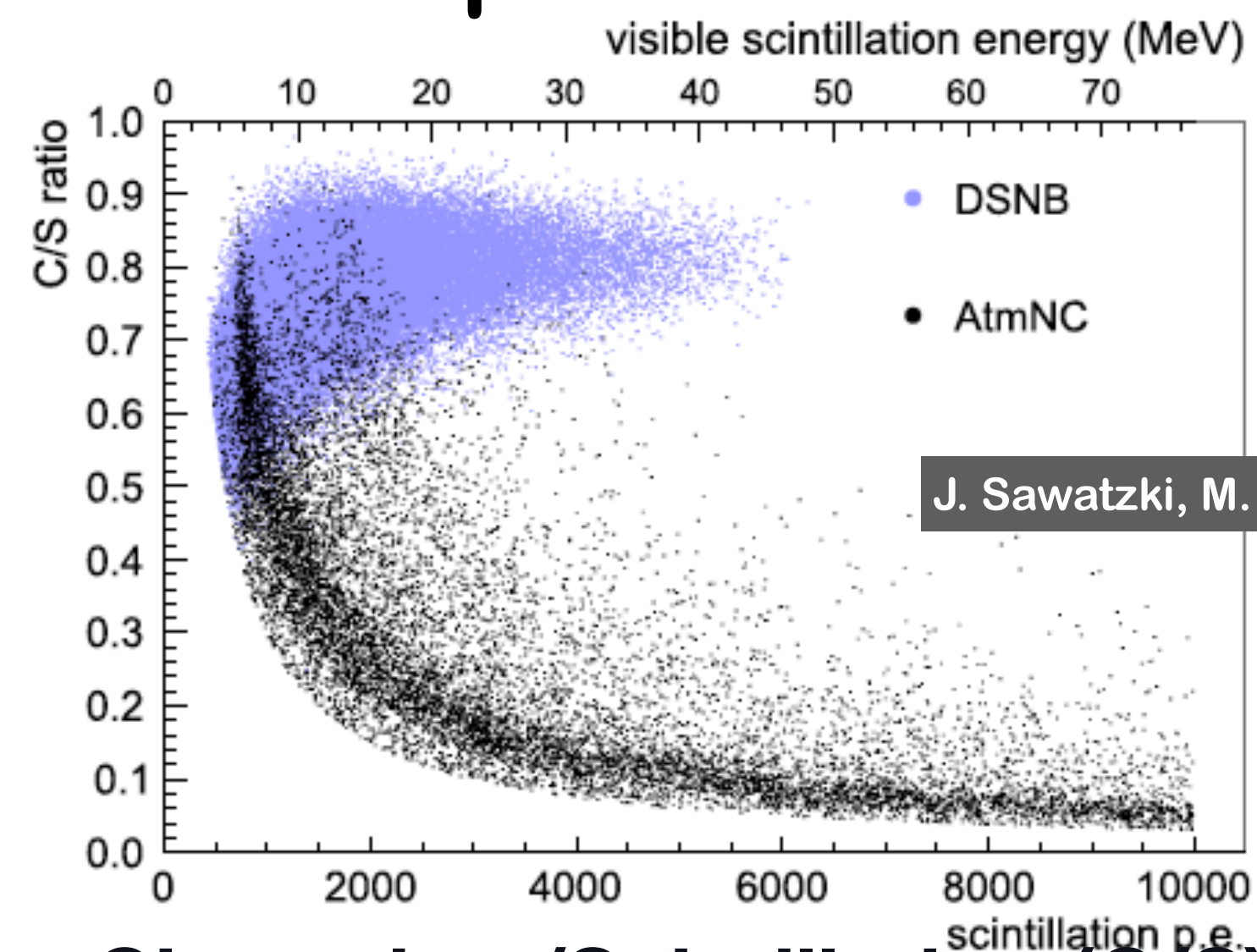
Supernova (SN) burst neutrinos

- High statistics low-threshold
- Flavor-resolved neutrino spectra
- Supernova pointing



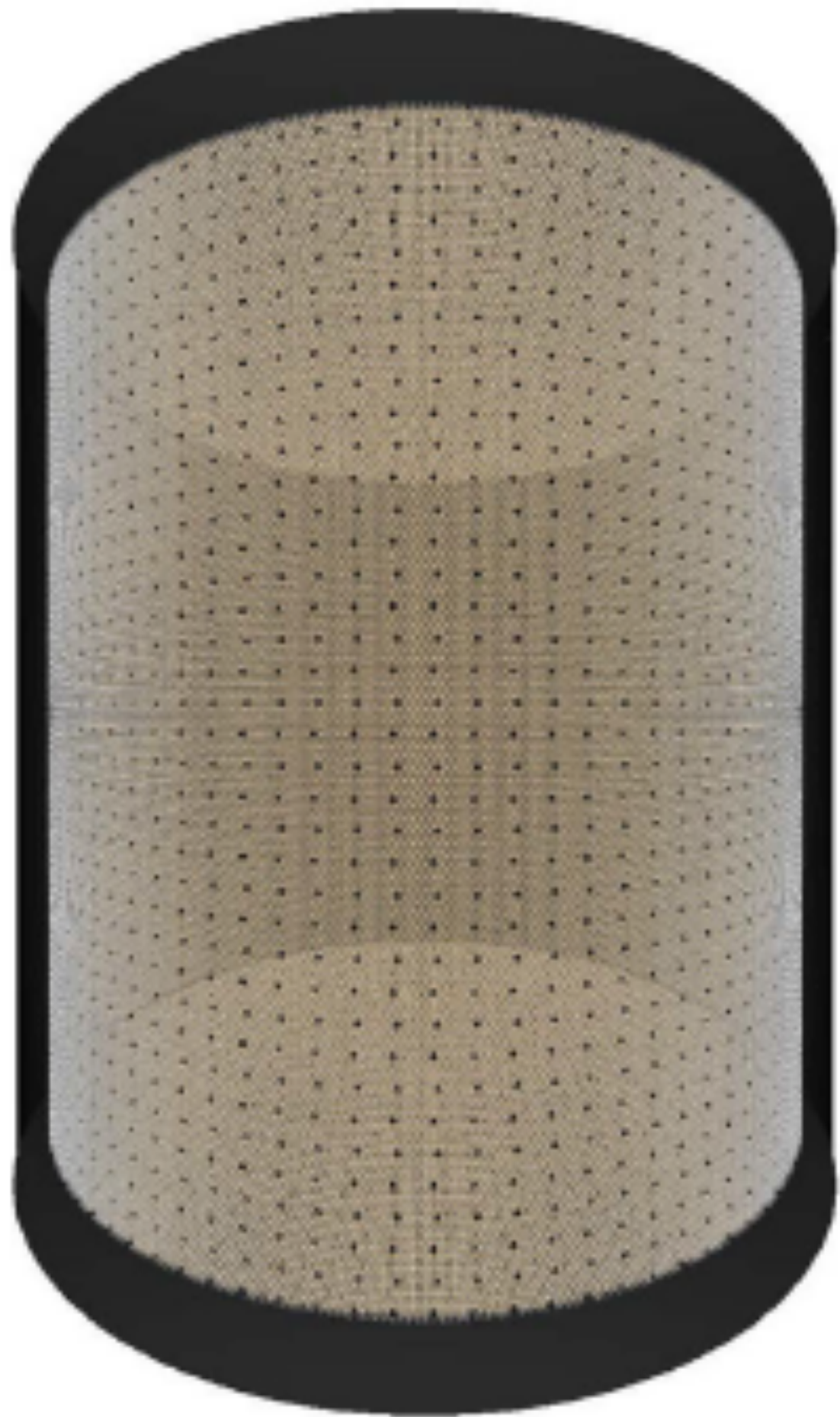
- **At LBNF:** the combination of WbLS (THEIA) and liquid argon (DUNE) detectors at the same site -> high-statistics co-detection of neutrinos and antineutrinos.
- Complementarity to JUNO and Hyper-K: opposite side of the Earth -> Earth matter effects
- Pre-supernova neutrinos

Diffuse supernova neutrino background (DSNB) Diffuse, isotropic flux of ν from all SN explosion in the Universe.



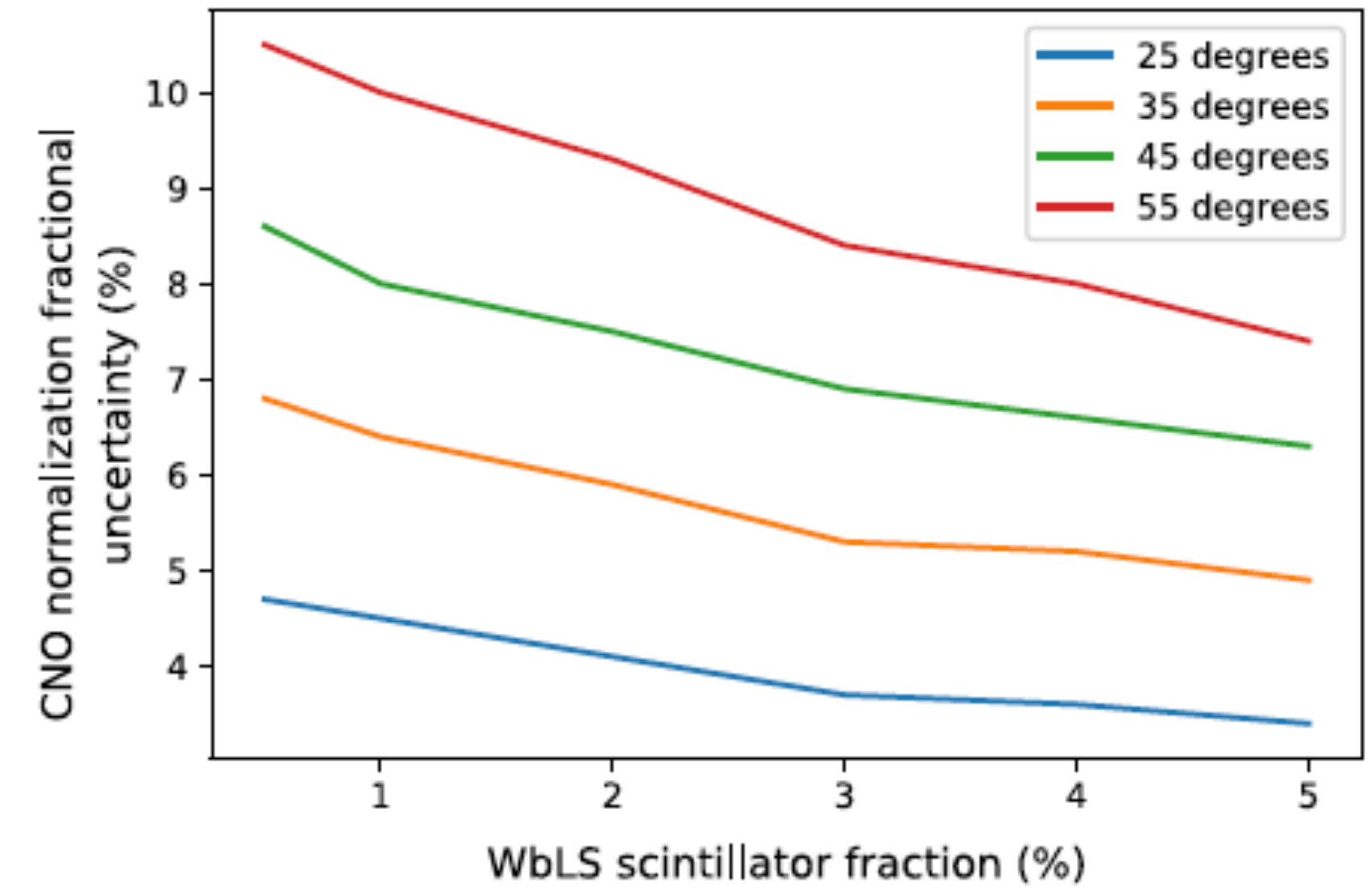
- **Cherenkov/Scintillation (C/S) ratio** gives a powerful handle to discriminate atmospheric neutral current background signals;
- substantial increase in event statistics when added to Super-K and JUNO;
- 5σ discovery (125 kton-year): ~ 8 years (Theia-25) or ~ 2 years (Theia-100)

Theia: solar neutrinos

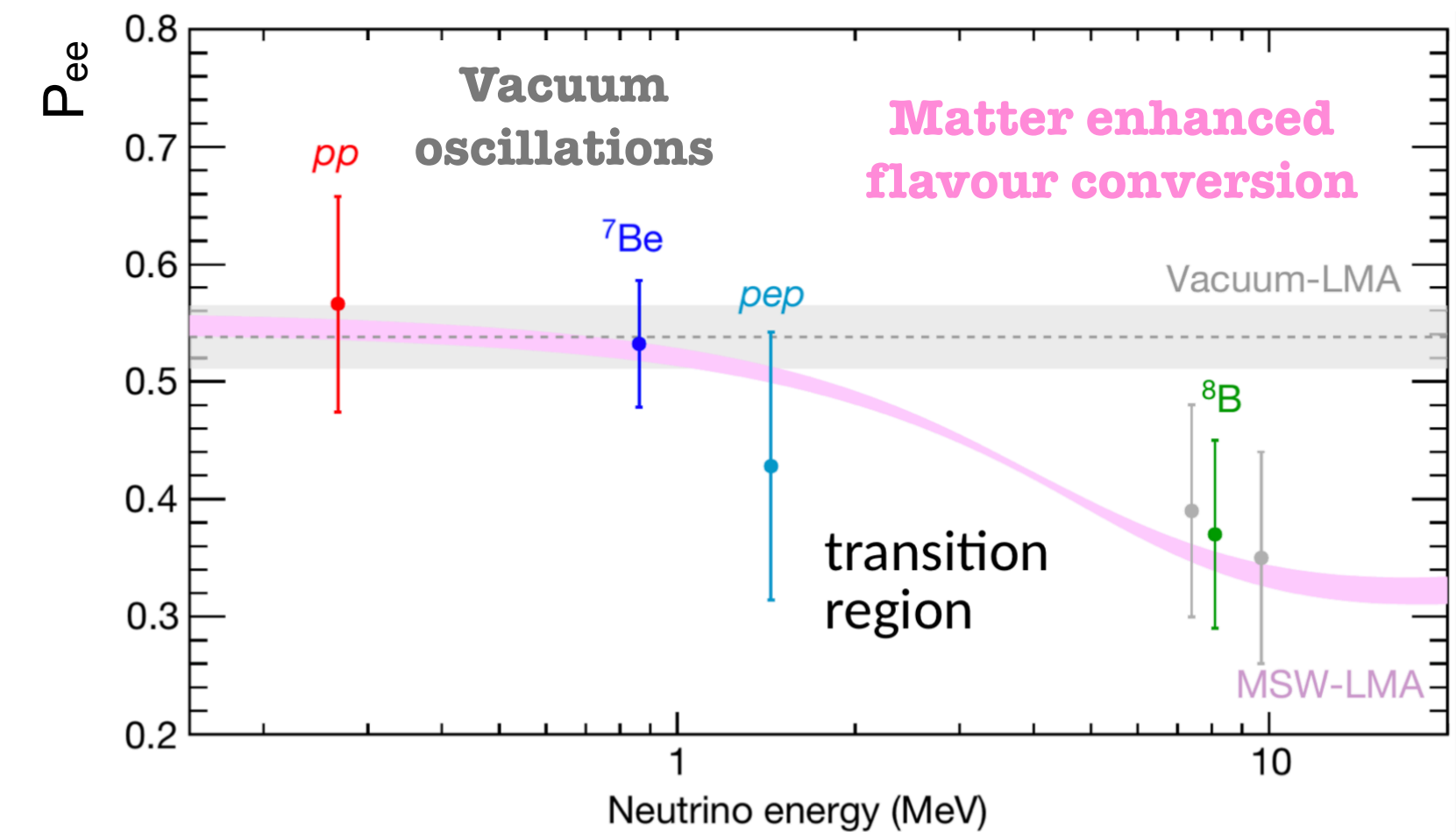


Theia can continue Borexino's solar legacy:

- CNO neutrinos (directionality based background rejection, solar metallicity puzzle)
- ^8B solar neutrinos high-statistics, low-threshold \rightarrow new physics in the MSW-vacuum transition region

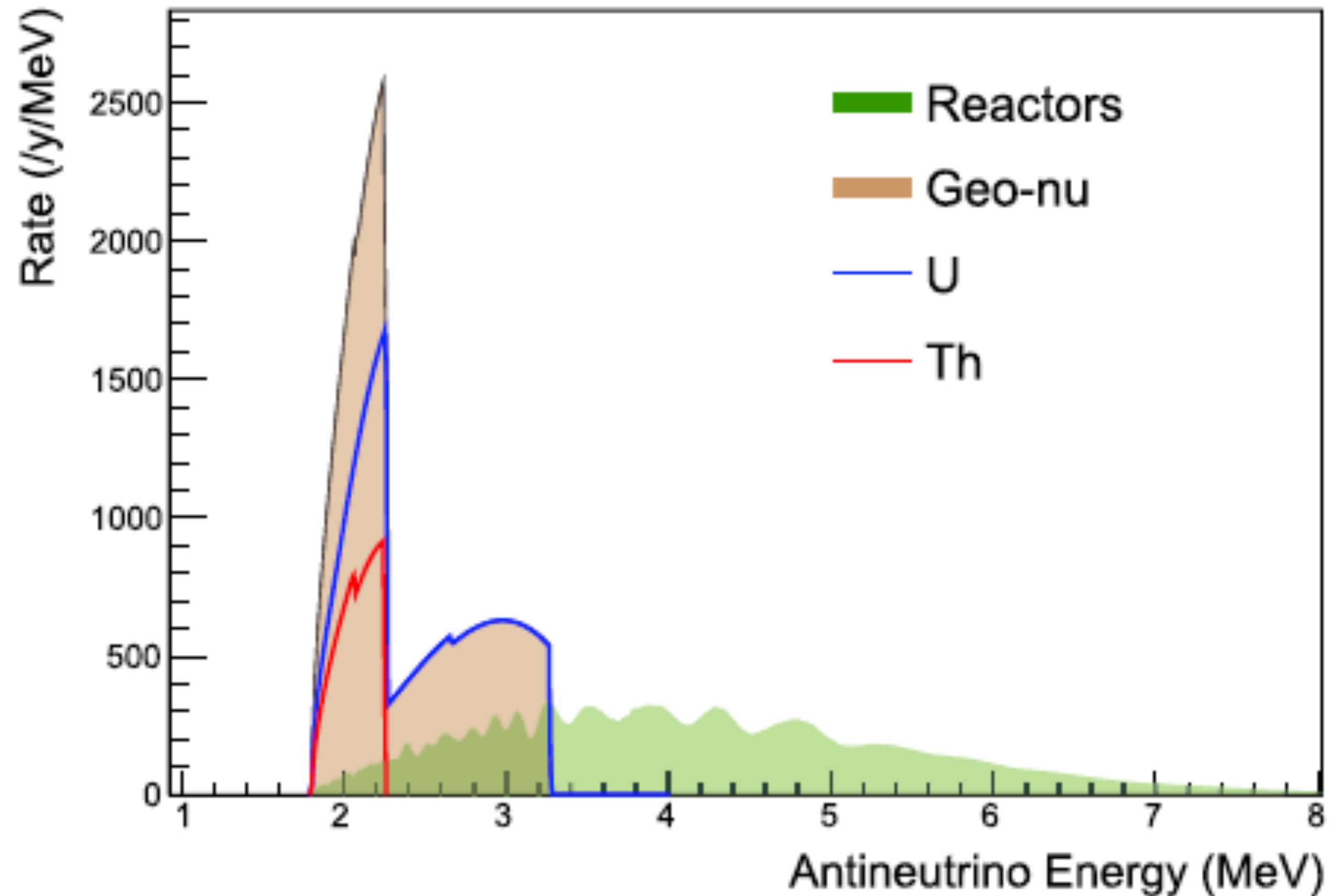


M. Askins, Z. Bagdasarian et al Eur. Phys. J. C 80, 416



Borexino measurements Nature 562, p 505

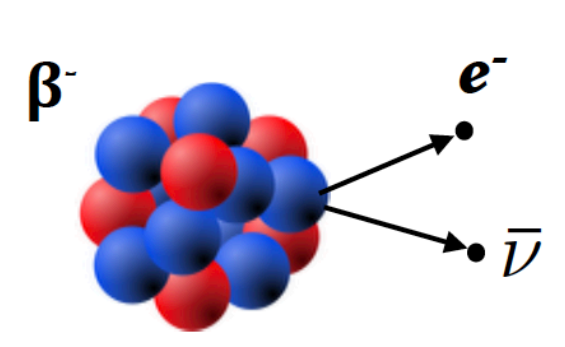
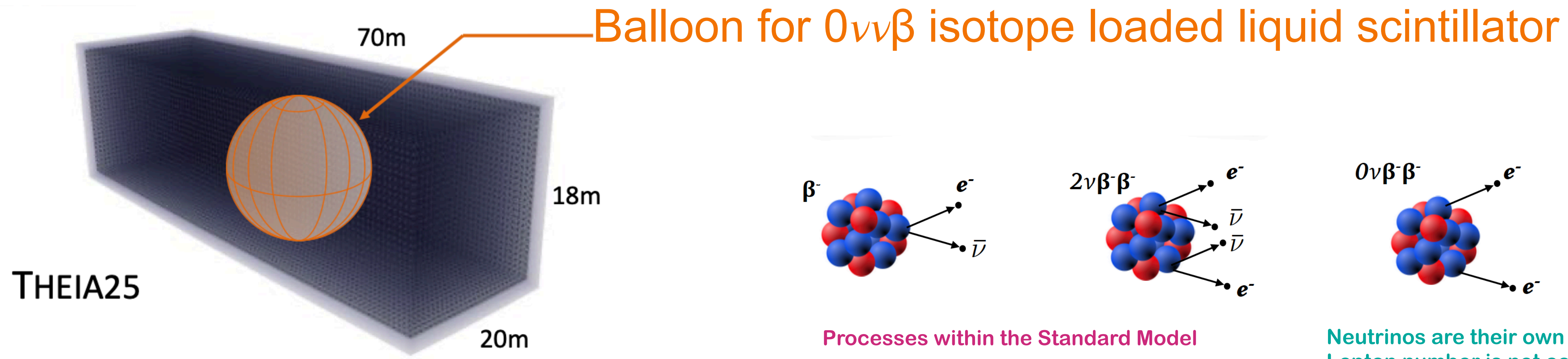
Theia: Geoneutrinos



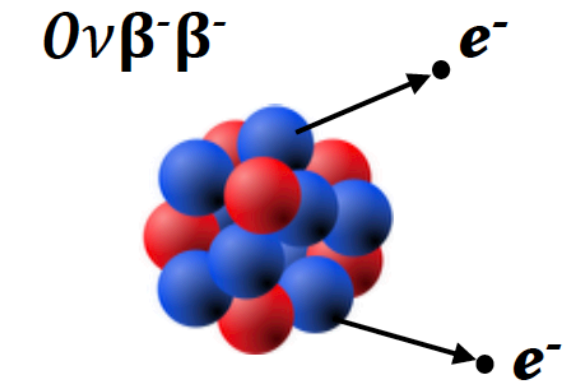
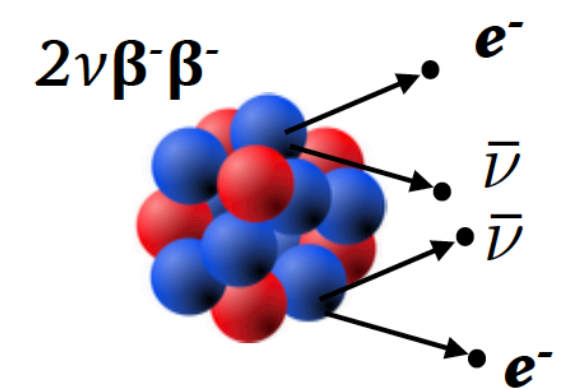
- Rate at Sanford Underground Research Facility (SURF): 26.5 interactions per kT-year
- High statistics (in comparison with existing two measurements)
- Explore geographical variations of the geoneutrino flux

Analysis of antineutrino capabilities of Theia is in preparation

Theia: Neutrinoless Double Beta Decay

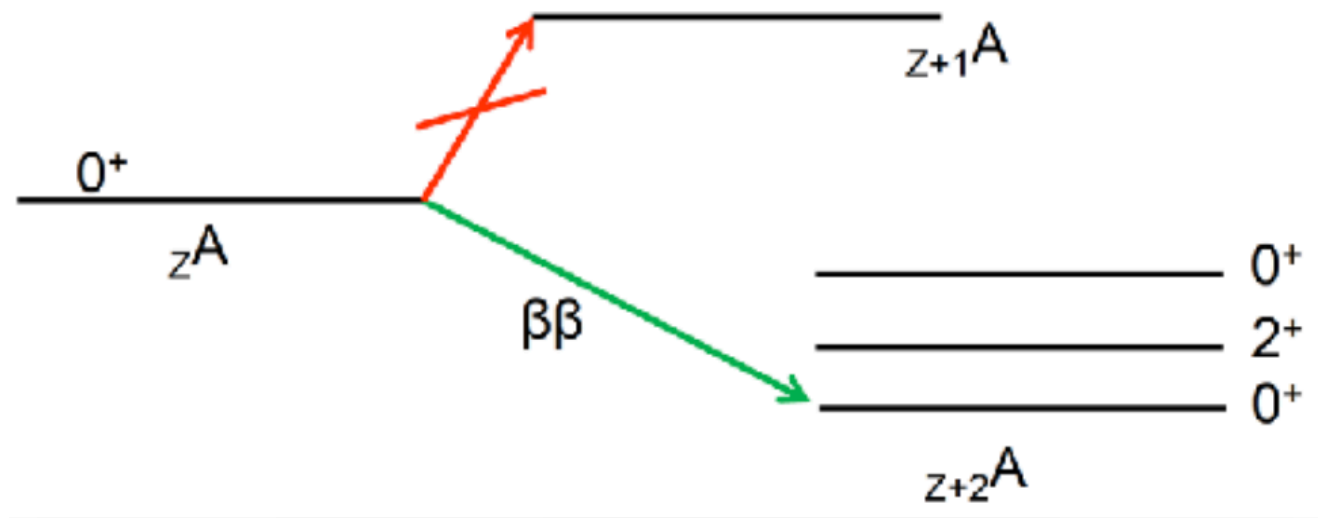


Processes within the Standard Model



Neutrinos are their own antiparticles
Lepton number is not conserved

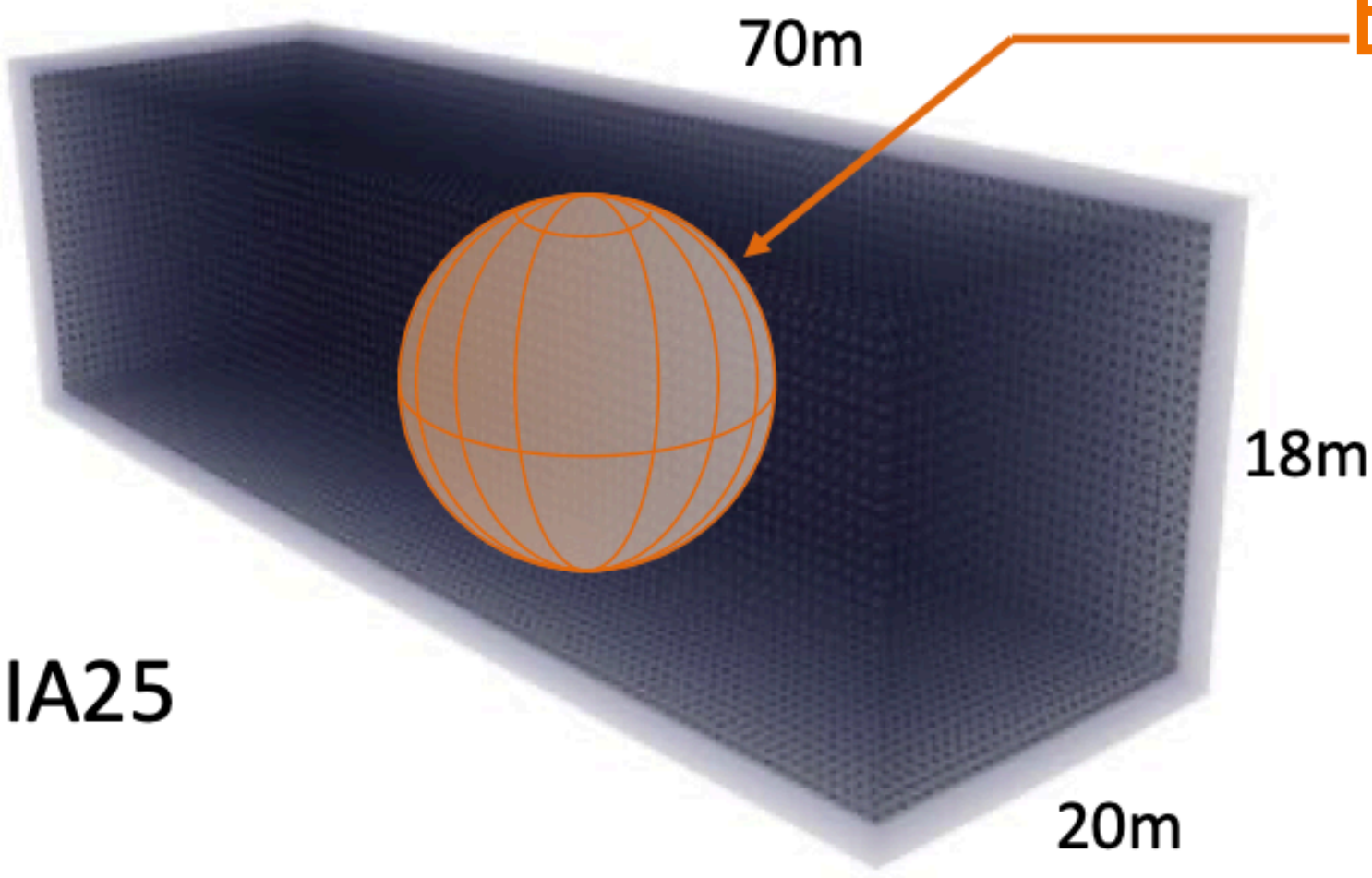
Elements for which normal beta decay is suppressed:
Germanium, Xenon, Tellurium



- violation of total **lepton number** conservation
- absolute neutrino masses
- mass ordering

Theia: Neutrinoless Double Beta Decay

Balloon for $0\nu\nu\beta$ isotope loaded liquid scintillator

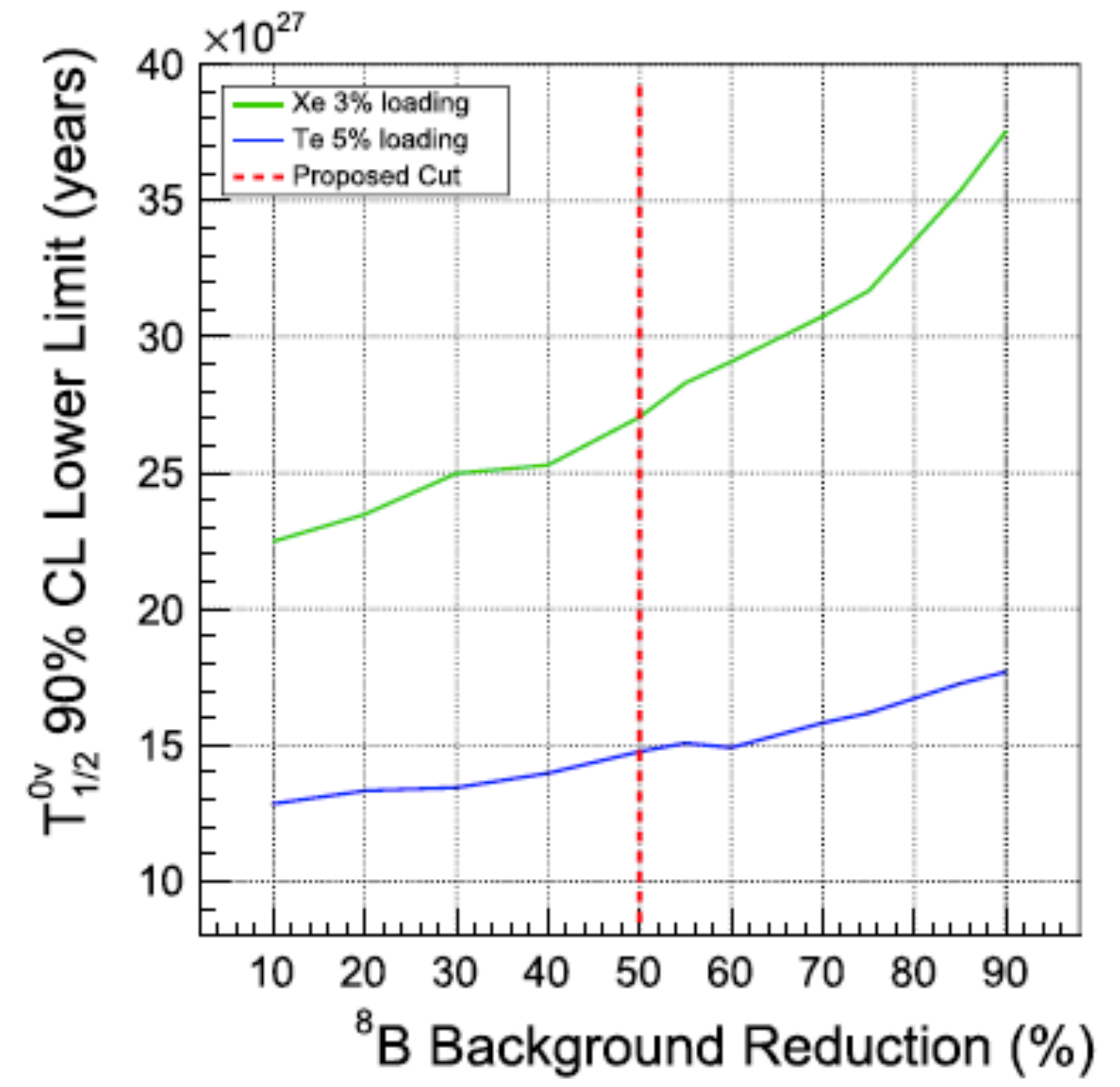
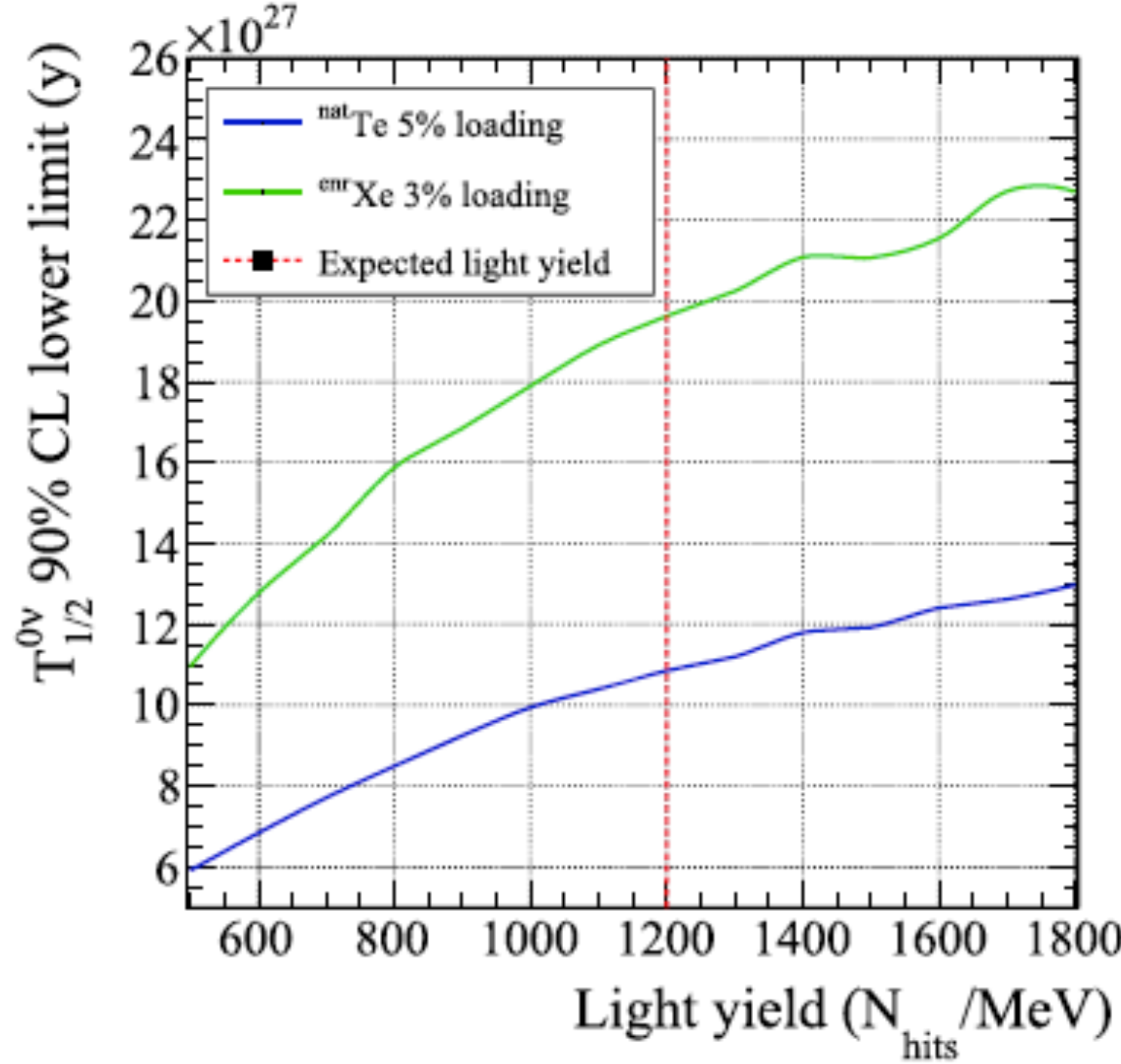
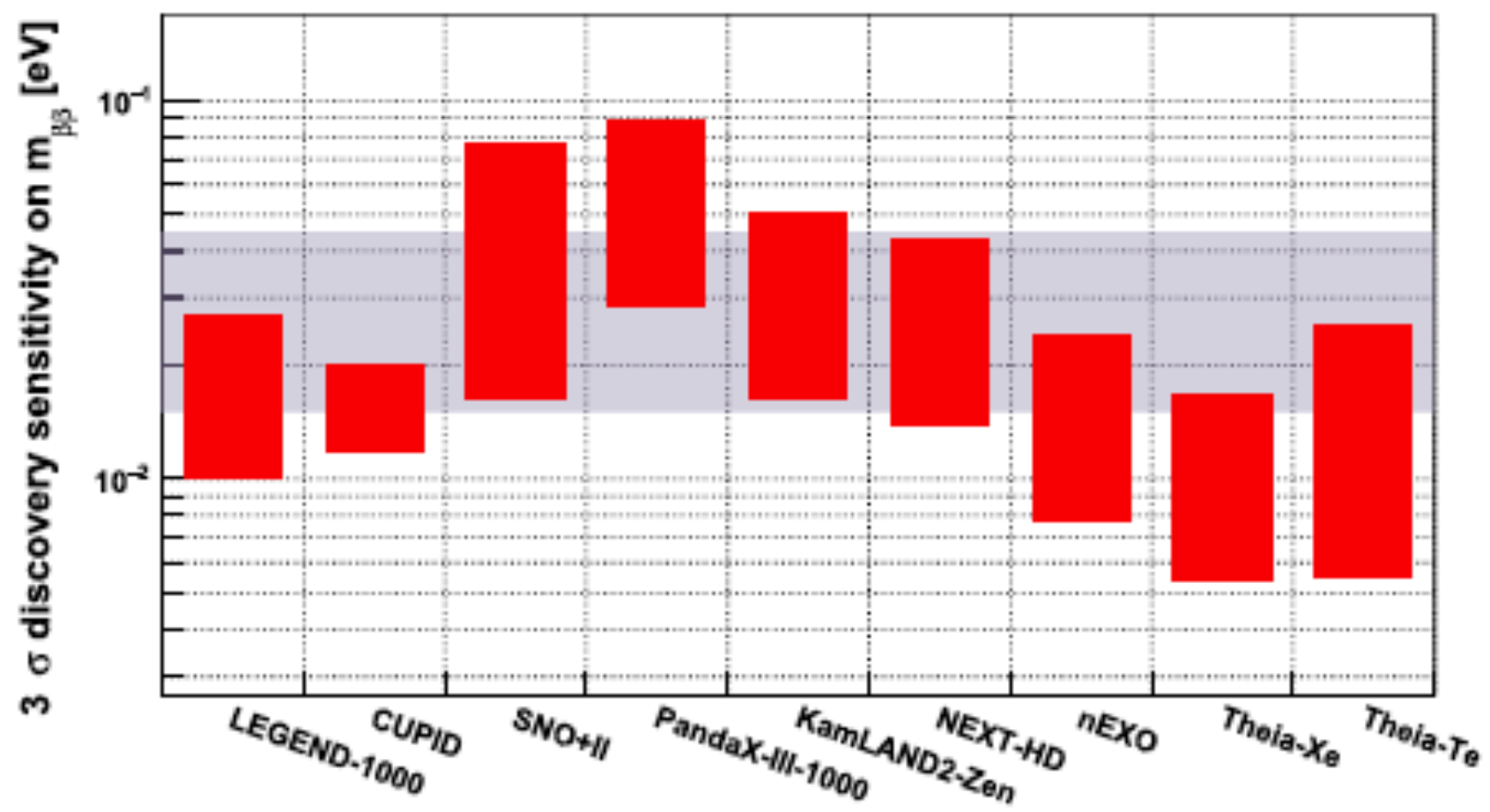


Isotopes in consideration: ^{136}Xe and ^{130}Te

Goal: Reach $T^{0\nu 2\beta}_{1/2} \sim 10^{28}$ years

Fiducialization and tagging techniques (triple coincidence, directionality, etc..) greatly reduce backgrounds

THEIA25



Theia: staged approach to physics goals

| | Primary physics goal | Reach | Exposure/assumptions |
|-----------|---------------------------------------|---|---|
| Phase I | Long-baseline oscillations | $>5\sigma$ for 30% of δ_{CP} | 524kt-MW-year |
| | Nucleon decay | $T > 3.8 \times 10^{34}$ year | 800 kt-year |
| | Supernova burst | $<1(2)^\circ$ pointing 20K(5K) events | 100(25)kt, 10kpc SN |
| | Diffuse Supernova Neutrino Background | 5σ | 125kt-year |
| Phase II | CNO neutrinos | $<5(10)\%$ | 300(62.5)kt-year |
| | Geoneutrinos | 2650 events | 100 kt-year |
| Phase III | $0\nu\nu\beta$ | $T_{1/2} < 1.1 \times 10^{28}$ year (90%C.L.) | 800 kt-year (Multi-tonne loaded LS in suspended vessel search) |

Conclusions

- Progress in the **novel target materials and photodetector technologies** opened the path for the **next-generation neutrinos experiments**
- **Theia** will employ the **advantages of these developments**

to **achieve**: *low energy threshold, good energy and position resolutions, directionality, large exposure*

and **to tackle a broad physics agenda**: *neutrino oscillations, solar, supernova neutrinos, and neutrinoless double beta decay*



Thank you for your attention!

QUESTIONS ARE WELCOME
now

or later

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<https://www.zarabagdasarian.com>

