

Genie

Tuning effort

Marco Roda
mroda@liverpool.ac.uk

on behalf of GENIE collaboration



University of Liverpool

15 March 2021
New Directions in Neutrino-Nucleus Scattering

Neutrino MC generators: our vision



- Connect neutrino fluxes and observables
 - event topologies and kinematics
- Good generators
 - optimal coverage of physics processes
 - Uncertainty validation
 - Tune the *physics* models
- Specific requirements for *experiments*
 - Shareable configurations
 - Data agreement
 - ⇒ Simple models can be perfectly acceptable

We don't believe in a *perfect theory* approach

- There are always things that need to be derived from measurements
- ⇒ Dealing with errors is unavoidable
- ⇒ Errors are part of the analysis procedures
 - See reweight approach

GENIE - www.genie-mc.org

Core GENIE mission - from GENIE by-law

Framework "... provide a state-of-the-art neutrino MC generator for the world experimental neutrino community ..."

Universality "... simulate all processes for all neutrino species and nuclear targets, from MeV to PeV energy scales ..."

Global fit "... perform global fits to neutrino, charged-lepton and hadron scattering data and provide global neutrino interaction model tunes ..."

- Well established generator
 - main generator for LAr experiments at Fermilab
- Modelling effort
 - Discussed by Steven
- Tuning effort
 - Deal with quantities that cannot be predicted
 - Control transition regions - from a model to another
- Other tools: Reweight package, flux drivers, geometry drivers, etc.

GENIE v3.00.06 (latest version) - v3.02.00 (soon to be released)

- Version 3 release introduced new concepts
- “Comprehensive Model Configurations”
 - Self-consistent collections of primary process models
 - Unique string identifier
 - --tune **G18_02a_00_000**
- Some of these models have been tuned
 - Identified by datasets and parameters
 - Unique string identifier
 - --tune **G18_02a_02_11b**



Tuning procedure

- ① “Brute force” scan the observables to be used for tuning (**bins**)
 - Decoupling prediction generation and minimisation
- ② The bin’s behaviour is summarised by response function
 - Polynomial functions of desired order
 - Including all the correlation terms up to the order of the polynomial
 - Fitted against the brute force points

Parameterisation - P dimensional parameter space, M order

$$O^i(\theta) = \alpha_0^i + \sum_{n=1}^P \beta_n^i \theta_n + \sum_{n \leq m} \gamma_{nm}^i \theta_n \theta_m + \dots + \sum_{n_1 \leq \dots \leq n_M} \xi_{n_1 \dots n_M}^i \prod_{\ell=1}^M \theta_{n_\ell} \quad (1)$$

More complex scenarios can be imagined

- ③ The minimisation is performed using the response functions
 - The response functions are extracted using Professor [1, 2]
 - The minimisation code is developed internally and it's based on Minuit
 - Takes into account datasets correlations
 - Can add nuisance parameters

The technique

Pros and Cons



Pros and Cons



- All parameters can be tuned
 - Not only reweightable
- It does not work on an event-by-event basis

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- Re-usability
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- It does not work on an event-by-event basis
- Changes in the analysis might require a re-run of the full chain
 - New bins
 - changes in the parameter ranges

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Some details

- Proved to be successful up ~ 20 parameters
 - Limit due to disk/CPU ⇒ It can be overcome
- 10 parameters, 5 order polynomial ⇒ about 4.5 k scan points
- 20 parameters, 5 order polynomial ⇒ about 23 k scan points

Response functions and reweight

- Response functions and reweight are separated (GENIE v3)
 - No reweighting for the tuned parameters
 - Unless a dedicated rweighting tool is already in place
 - We aim to find the best parameter values ⇒ best data description
 - We understand the release of tune results can be frustrating
- ⇒ We encourage experiments to develop their own analyses
 - Based on their own response functions
 - In this way they will be able to use our statistical results as their priors

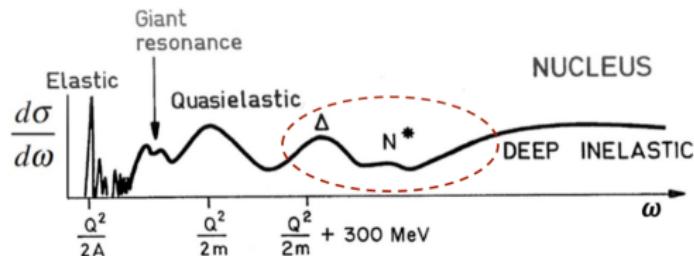
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- Reweight updated to include response function concepts (GENIE v4)
 - Response functions could be created in every parameter space
 - ⇒ Including each event generator parameter space
 - Reweight will be configurable with response functions produced by users
 - Overcome the intrinsic limitation of “writing a dedicated module”
 - Still the reweight limitation will apply
 - Hard to reweight things like binding energies, masses, thresholds, etc
 - ⇒ Experiments will produce their own response functions for their analysis bins
 - Optimised for their flux, detector composition, etc.

Tuning details

The physics

Resonant to Deep Inelastic Scattering transition



$$\frac{d^2 \sigma_{\text{inel}}}{dQ^2 dW} = \begin{cases} \frac{d^2 \sigma_{\text{RES}}}{dQ^2 dW} + \frac{d^2 \sigma_{\text{SIS}}}{dQ^2 dW} & \text{for } W < W_{\text{cut}} \\ \frac{d^2 \sigma_{\text{DIS}}}{dQ^2 dW} & \text{for } W \geq W_{\text{cut}} \end{cases}$$

$$\frac{d^2 \sigma_{\text{SIS}}}{dQ^2 dW} = \frac{d^2 \bar{\sigma}_{\text{DIS}}}{dQ^2 dW} \Theta(W_{\text{cut}} - W) \sum_m f_m(Q^2, W)$$

$$f_m(Q^2, W) = R_m \cdot P_m^{\text{had}}(Q^2, W)$$

Parameter	GENIE parameter name	Default value	Min value	Max value	Prior
W_{cut} (GeV/c ²)	Wcut	1.7	1.5	2.3	
M_A^{QE} (GeV/c ²)	QUE-Ma	0.999	0.75	1.10	1.014 ± 0.014 [3]
M_A^{RES} (GeV/c ²)	RES-Ma	1.12	0.8	1.3	1.12 ± 0.03 [4]
$R_{v/p}^{\text{CC1}\pi}$	DIS-HMultWgt-vp-CC-m2	0.10	0.0	0.4	
$R_{v/p}^{\text{CC2}\pi}$	DIS-HMultWgt-vp-CC-m3	1.00	0.0	2.0	
$R_{v/n}^{\text{CC1}\pi}$	DIS-HMultWgt-vn-CC-m2	0.30	0.0	0.35	
$R_{v/n}^{\text{CC2}\pi}$	DIS-HMultWgt-vn-CC-m3	1.00	0.8	3.0	
S_{RES}	RES-CC-XSecScale	1.0	0.6	1.2	
S_{DIS}	DIS-CC-XSecScale	1.032	0.9	1.15	1 ± 0.05

Datasets

Observables and data

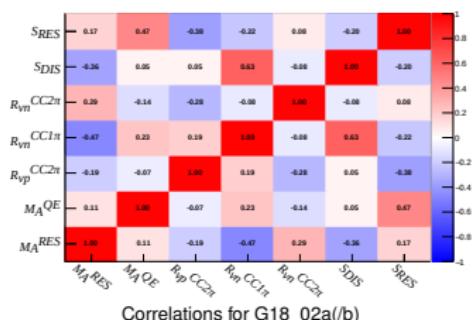
- Bubble chambers data
 - Muon (anti)neutrinos CC interactions on hydrogen and deuterium targets
 - ANL 12FT, BNL 7FT, FNAL 15FT and BEBC
 - Integrated cross sections as a function of the neutrino energy
 - ν_μ and $\bar{\nu}_\mu$ CC inclusive scattering [5–32]
 - ν_μ and $\bar{\nu}_\mu$ CC quasi-elastic scattering [5, 15, 29, 33–41]
 - ν_μ and $\bar{\nu}_\mu$ CC single-pion production [34, 42–50]
 - $\nu_\mu + n \rightarrow \mu^- + n + \pi^+$ and $\nu_\mu + p \rightarrow \mu^- + p + \pi^+$
 - $\nu_\mu + n \rightarrow \mu^- + p + \pi^0$
 - $\bar{\nu}_\mu + p \rightarrow \mu^+ + p + \pi^-$ and $\bar{\nu}_\mu + n \rightarrow \mu^+ + n + \pi^-$
 - ν_μ CC two-pion production [51]
 - $\nu_\mu + p \rightarrow \mu^- + n + 2\pi^+$ and $\nu_\mu + p \rightarrow \mu^- + n + \pi^+ + \pi^-$
 - $\nu_\mu + p \rightarrow \mu^- + p + \pi^+ + \pi^0$
- Not all datasets are used:
 - Only bins with $E_\nu > 0.5$ GeV
 - Only latest available version from each experiment
- Our predictions take into account experimental cuts
- Correlations between data from the same experiments
 - ⇒ Taken into account with experiment related nuisance parameters

Results

Best fit parameters

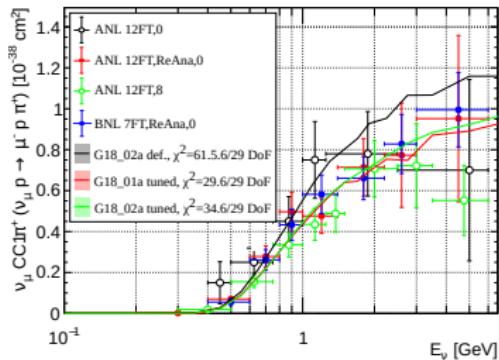
- Two fits: one for each CMC
 - Both behaving similarly
⇒ confidence in the procedure
- Results as expected from previous iterations
 - RES interaction suppressed ~ 15 %
 - 1 π very suppressed
 - at the limit of the region
 - 2 π enhanced
- Reasonable goodness of fit
 - These numbers take into account the nuisance parameters

Parameter	G18_01a(/b)	G18_02a(/b)
W_{cut}	1.94	1.81
M_A^{QE}	1.00 ± 0.01	1.00 ± 0.013
M_A^{RES}	1.09 ± 0.02	1.09 ± 0.014
$R_{\nu p}^{\text{CC1}\pi}$	0.06 ± 0.03	0.008
$R_{\nu p}^{\text{CC2}\pi}$	1.1 ± 0.2	0.94 ± 0.075
$R_{\nu n}^{\text{CC1}\pi}$	0.14 ± 0.03	0.03 ± 0.010
$R_{\nu n}^{\text{CC2}\pi}$	2.8 ± 0.4	2.3 ± 0.12
S_{RES}	0.89 ± 0.04	0.84 ± 0.028
S_{DIS}	1.03 ± 0.02	1.06 ± 0.01
$\chi^2/157 \text{ DoF}$	1.84	1.64

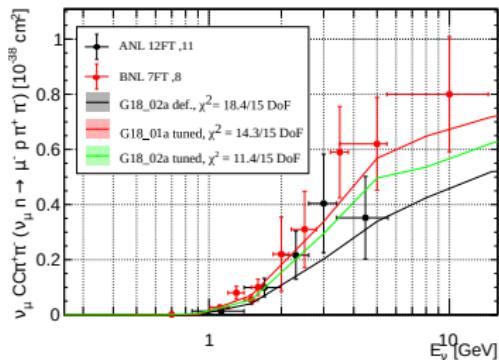


Results

Best fit predictions



- The agreement improves for every observable
 - True also for inclusive datasets despite observed tensions with exclusive channels
- The predictions of the two CMC after tuning are in strong agreement



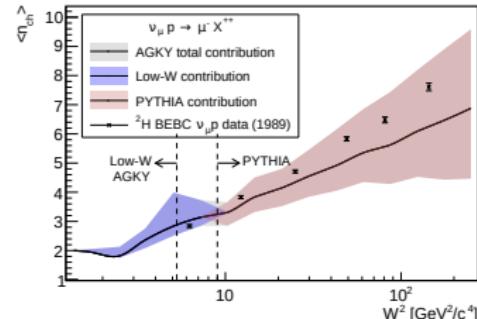
The modeling

Hadronisation modeling

- Another aspect of the inelastic modeling
 - Independent to cross section modeling
 - Focusing on $\mathcal{P}(n|W)$ instead of $\mathcal{P}(W)$
 - Empirically

- $\langle n_{\text{ch}} \rangle(W) = \alpha_{\text{ch}} + \beta_{\text{ch}} \ln \left(\frac{W^2}{\text{GeV}^2/c^4} \right)$
- $\langle n_{\text{neut}} \rangle = \frac{1}{2} \langle n_{\text{ch}} \rangle$

- Two different models inside GENIE
 - Pythia - for high W
 - based on Lund fragmentation model
 - Low W AGKY
 - Empirical distribution based on Lévy distribution
 - $\langle n \rangle P \left(\frac{n}{\langle n \rangle} \mid c \right) = 2e^{-c} c^{\frac{n}{\langle n \rangle} + 1} \left[\Gamma \left(c \frac{n}{\langle n \rangle} + 1 \right) \right]^{-1}$
 - Joint linearly
 - from $W_{\min}^{\text{tr}} = 2.3 \text{ GeV}$ to $W_{\max}^{\text{tr}} = 3.0 \text{ GeV}$



Default GENIE v3 low-W AGKY parameters

Parameter	$\nu_\mu p$	$\nu_\mu n$	$\bar{\nu}_\mu p$	$\bar{\nu}_\mu n$
α_{ch}	0.40	-0.20	0.02	0.80
β_{ch}	1.42	1.42	1.28	0.95
c	7.93	5.22	5.22	7.93

Default GENIE v3 PYTHIA parameters

Parameter	Name in PYTHIA	Value
$P_{S\bar{S}}$	PARJ(2)	0.30
$\langle p_\perp^2 \rangle [\text{GeV}^2]$	PARJ(21)	0.44
$E_{\text{CutOff}} [\text{GeV}]$	PARJ(33)	0.20
Lund a	PARJ(41)	0.30
Lund b [GeV^{-2}]	PARJ(42)	0.58

Datasets

Datasets and previous attempts

- Data from FNAL 15 ft and BEBC
- Data in the form of $\langle n_{\text{ch}} \rangle$ vs W
 - from muon (anti)neutrinos on H , 2H , heavier targets
 - We only used hydrogen and deuterium data
 - there are other observables but not used this time
 - data sensitive to c parameters
 - to take into account those observables priors have been used on E_{CutOff} and $\langle p_{\perp}^2 \rangle$
- Data samples divided by the nucleon hit in the interaction
 - Selection based on the topology
- Every experiment tried to make their own estimation of the linear coefficients
 - Not enough for Pythia contribution
 - Not satisfactory as a global fit
 - Tensions between hydrogen and deuterium data

$\nu_{\mu} + p \rightarrow \mu^- X^{++}$	
FNAL 15 ft (1976)	H
BEBC (1983)	H
BEBC (1990)	H
BEBC (1992)	H
FNAL 15 ft (1983)	2H
BEBC (1989)	2H

$\nu_{\mu} + n \rightarrow \mu^- X^+$	
FNAL 15 ft (1983)	2H
BEBC (1984)	2H
BEBC (1989)	2H

$\bar{\nu}_{\mu} + p \rightarrow \mu^+ X^0$	
FNAL 15 ft (1981)	H
BEBC (1983)	H
BEBC (1990)	H
BEBC (1992)	H
BEBC (1982)	2H
BEBC (1989)	2H

$\bar{\nu}_{\mu} + n \rightarrow \mu^+ X^-$	
BEBC (1982)	2H
BEBC (1989)	2H

Results

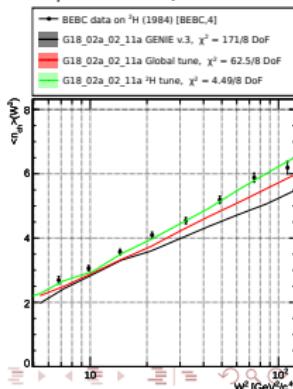
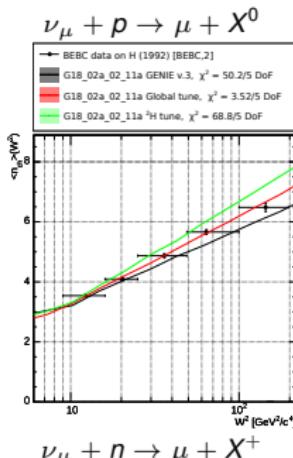
Best fit parameters

Parameter	GENIE parameter name	Nominal value	Allowed range	Global Fit	${}^2\text{H}$ Fit
Low-W empirical model					
$\alpha_{\nu p}$	KNO-Alpha-vp	0.40	[−1.0, 2.0]	1.1 ± 0.3	1.2 ± 0.4
$\alpha_{\nu n}$	KNO-Alpha-vn	-0.20	[−1.0, 2.0]	$1.75^{+0.14}_{-0.11}$	-0.58 ± 0.07
$\alpha_{\bar{\nu} p}$	KNO-Alpha-vbp	0.02	[−1.0, 2.0]	$1.32^{+0.16}_{-0.14}$	1.9 ± 0.08
$\alpha_{\bar{\nu} n}$	KNO-Alpha-vbn	0.80	[−1.0, 2.0]	1.11 ± 0.09	1.07 ± 0.3
$\beta_{\nu p}$	KNO-Beta-vp	1.42	[0.0, 2.5]	0.79 ± 0.15	0.9 ± 0.3
$\beta_{\nu n}$	KNO-Beta-vn	1.42	[0.0, 2.5]	0.5 ± 0.1	1.9 ± 0.3
$\beta_{\bar{\nu} p}$	KNO-Beta-vbp	1.28	[0.0, 2.5]	0.8 ± 0.1	0.3 ± 0.1
$\beta_{\bar{\nu} n}$	KNO-Beta-vbn	0.95	[0.0, 2.5]	$0.88^{+0.09}_{-0.08}$	0.9 ± 0.2
PYTHIA					
$P_{s\bar{s}}$	PYTHIA-SSBarSuppression	0.30	[0.0, 1.0]	0.27 ± 0.04	0.29 ± 0.05
$P_{(p_\perp^2)} [\text{GeV}^2/c^2]$	PYTHIA-GaussianPt2	0.44	[0.1, 0.7]	0.43 ± 0.05	0.43 ± 0.04
$E_{\text{CutOff}} [\text{GeV}]$	PYTHIA-RemainingEnergyCutoff	0.20	[0.0, 1.0]	0.30 ± 0.04	0.24 ± 0.05
<i>Lund a</i>	PYTHIA-Lunda	0.30	[0.0, 2.0]	1.53 ± 0.13	1.85 ± 0.15
<i>Lund b</i> [GeV/c^2]	PYTHIA-Lundb	0.58	[0.0, 1.5]	1.16 ± 0.09	1.0 ± 0.2
$\chi^2 =$					
$87.9/62 \text{ DoF}$					
$29.5/32 \text{ DoF}$					

- Generally good goodness of fit
- Mostly symmetrical χ^2 profiles \Rightarrow Gaussian behaviour
- A few strong (anti)correlation observed: due to transition region

Best fit predictions

- The data prefers a higher multiplicity at high W
 - Low-W is bound by energy conservation
 - In general better agreement with data
 - Very strong tension identified
 - Global fit p-value $4 \cdot 10^{-12}$
 - Deuterium fit p-value 0.94
 - Previously observed [61]
 - Not completely understood
 - The problem gets complicated when other observables are included
 - Already tension of these data vs others
 - e.g. π^0 multiplicity, ratio of multiplicity variance over average multiplicity
 - Interference with the free nucleon tune
 - A proper tuning procedure should go in inverse order
 - Possible systematic errors in the reconstruction of W



Summary

- Overview of the tuning effort within GENIE
 - Procedures
 - Expected plan
- Satisfactory free nucleon tune
 - Better understood data
 - Analyses from different groups have been going on for a long time
 - Paper ready to submission
- Hadronisation tune is new
 - Data is more tricky
 - First time a global tune of this scale has been attempted
 - Much more to be learned
 - Paper in internal review

Backup

Collaboration

Luis Alvarez-Ruso [9], Costas Andreopoulos [5,7], Adi Ashkenazi [4], Christopher Barry [5], Steve Dennis [5], Steve Dytman [6], Hugh Gallagher [8], Steven Gardiner [3], Walter Giele [3], Robert Hatcher [3], Or Hen [4], Libo Jiang [6], Rhiannon Jones [5], Igor Kakorin [2], Konstantin Kuzmin [2], Anselmo Meregaglia [1], Donna Naples [6], Vadim Naumov [2], Afroditis Papadopoulou [4], Gabriel Perdue [3], Marco Roda [5], Vladislav Syrotenko [8], Jeremy Wolcott [8], Júlia Tena Vidal [5], Julia Yarba [3]

1. **CENBG, Université de Bordeaux, CNRS/IN2P3**
33175 Gradignan, France
2. **Joint Institute for Nuclear Research (JINR)**
Dubna, Moscow region, 141980, Russia
3. **Fermi National Accelerator Laboratory**
Batavia, Illinois 60510, USA
4. **Massachusetts Institute of Technology (MIT) Dept. of Physics**
Cambridge, MA 02139, USA
5. **University of Liverpool, Dept. of Physics**
Liverpool L69 7ZE, UK
6. **University of Pittsburgh, Dept. of Physics and Astronomy**
Pittsburgh PA 15260, USA
7. **UK Research and Innovation, Science and Technology Facilities Council**
Rutherford Appleton Laboratory, Particle Physics Dept.
Harwell Oxford Campus, Oxfordshire OX11 0QX, UK
8. **Tufts University, Dept. of Physics and Astronomy**
Medford MA 02155, USA
9. **University of Valencia**
Valencia, Spain

Roles of generators in oscillation physics

- Compare data and models
 - Reliability and validity region
⇒ You cannot study oscillations without fully understood models
- Compare dataset against dataset
 - Data quality and data sources are increasing ⇒ **tensions**
 - ⇒ joint analyses
 - ⇒ comparing results from different experiments
- **Global fits**
 - A generator is the ideal place for global fits
 - Controls the model implementation
 - Finding the best parameters
 - Cross Section priors based on data
- Feedback for experiments
 - Drive the format of cross section releases
 - Hint toward key measurements

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