



The Science
and Design of
the LBNF
Beamline

Mary Bishai
(on behalf of
LBNF/DUNE)
Brookhaven
National Lab

DUNE Science
Requirements

Beamline
overview

Primary Beam
Neutrino Beam
Decay Pipe
Absorber

Beam Flux
Systematics

Beam Flux
Stability

Future Beam
Upgrades

Summary

The Science and Design of the LBNF Beamline

NuFact 2018, Aug 12-18, 2018, Blacksburg, VA

**Mary Bishai (on behalf of LBNF/DUNE)
Brookhaven National Lab**

August 14, 2018





Outline

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- Decay Pipe
- Absorber

3 Beam Flux Systematics

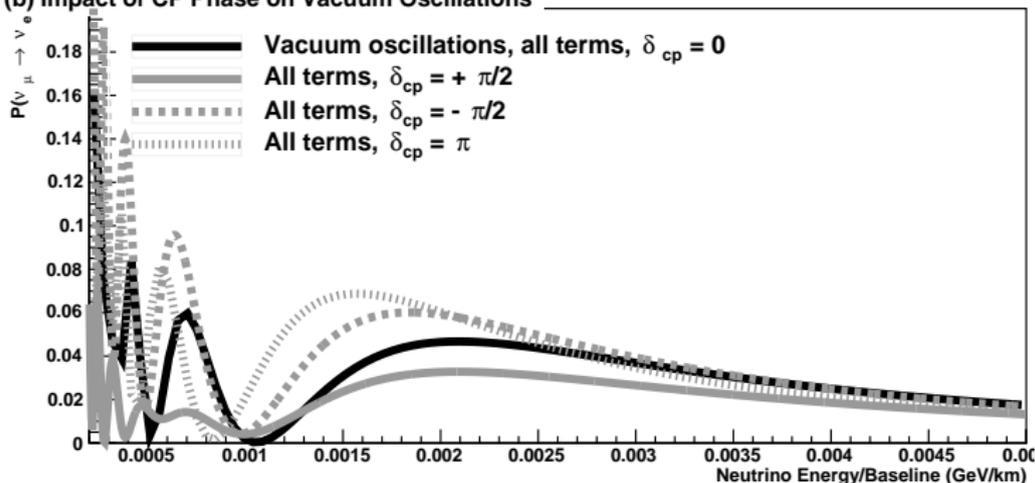
4 Beam Flux Stability

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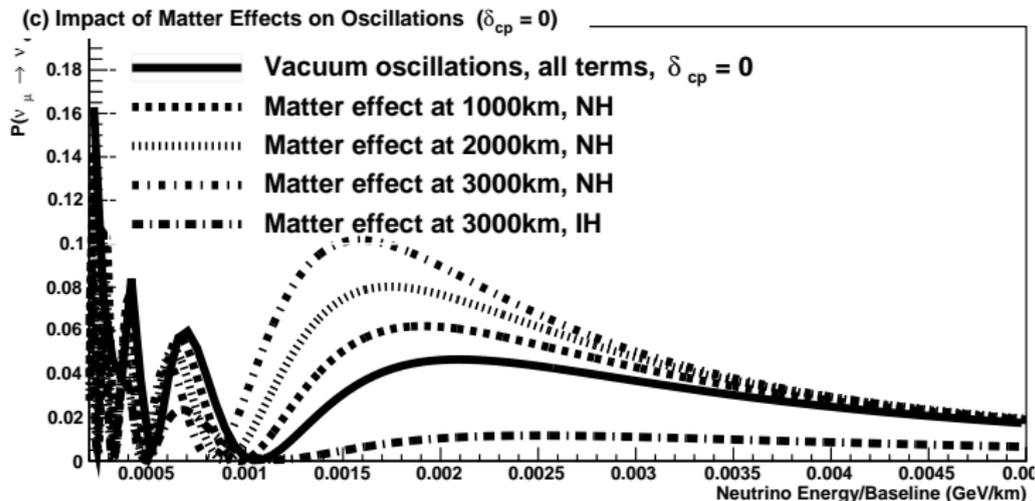
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Science Requirements for the LBNF Beamline

(b) Impact of CP Phase on Vacuum Oscillations

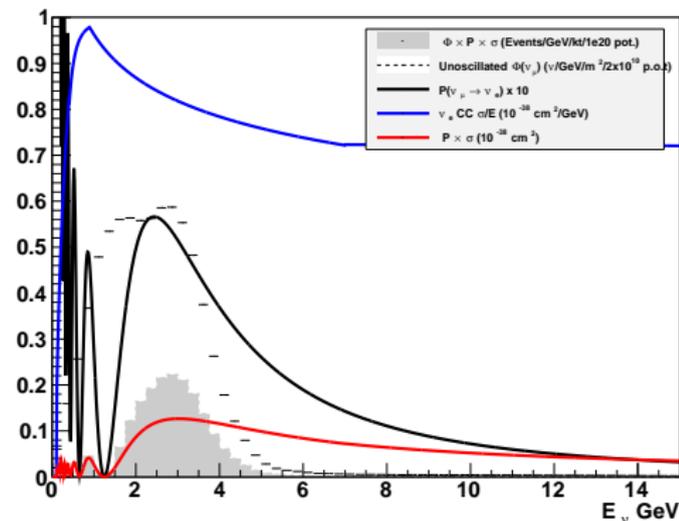


Science Requirements for the LBNF Beamline



Science Requirements for the LBNF Beamline

$\nu_{\mu} \rightarrow \nu_e$ Appearance at 1300 km





Global Science Requirements of the Beam

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From DUNE/LBNF's formal requirement document:

Glo-Sci-13: The neutrino beam spectrum shall cover the energy region of the first two oscillation maxima affected by muon-neutrino conversion from the atmospheric parameters.

Glo-Sci-60: The neutrino beam spectrum shall extend beyond the first maximum to higher energies, while maintaining a high signal to background ratio to obtain the maximum number of charged current signal events.

Glo-Sci-14: The neutrino beam spectrum shall be tunable so that beam with both lower peak energy (below the first oscillation node) and higher peak energy (significantly higher than the first oscillation node) can be achieved without substantial downtime that reduces the overall exposure.



Global Science Requirements of the Beam

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From DUNE/LBNF's formal requirement document:

Glo-sci-16: The beam shall be sign-selected to provide separate neutrino and anti-neutrino beams with high purity to enable measurement of CP violation mass hierarchy, and precision oscillation measurement.

Glo-sci-17: The beam shall be aimed at the far detector with an angular accuracy that allows the determination of the far detector spectrum using the near detector measurements. The angular accuracy shall not be the dominant factor in the determination of oscillation parameters.

Glo-sci-22: The beam monitoring systems shall have sufficient energy and spatial and temporal resolution that when combined with the detailed knowledge of the beam line geometry, a timely (few hours) feedback of beam performance, stability, as well as a data-driven estimate of the neutrino flux will be obtained.



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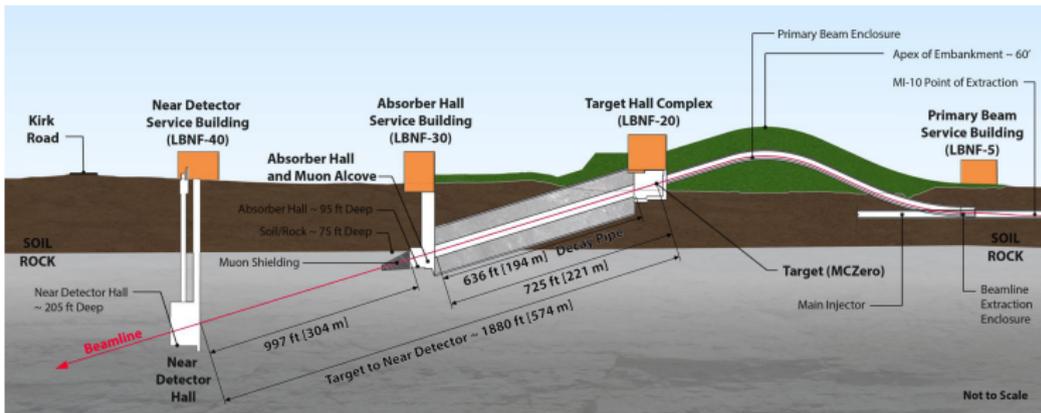
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LBNF Beamline Design and Optimization

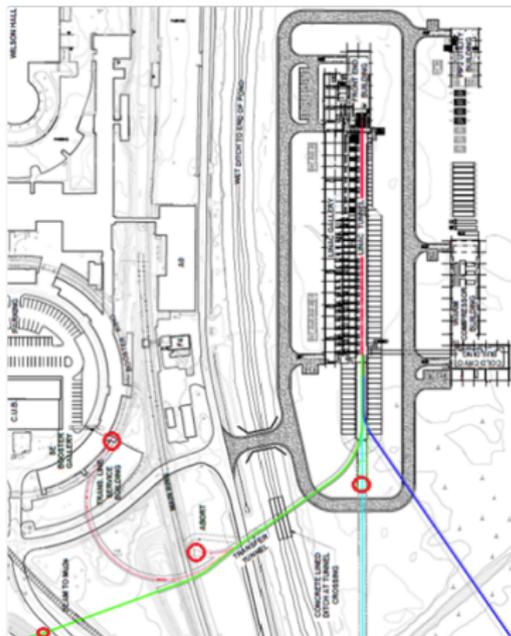


- **Primary proton beam 60-120 GeV**
- **Initial 1.2 MW beam power, upgradable to 2.4 MW**
- **Embankment allows target complex to be at grade**
- **Wide-band beam (on-axis) optimized for CP violation sensitivity**
- **Decay pipe: 194m x 4m diameter, He filled**

ND default: 574 m from target, FD: 1300 km

PIP-II Components

- 800 MeV linac
 - Warm Front End
 - SRF section
- Linac-to-Booster transfer line
 - 3-way beam split to: (1) Beam dump, (2) Booster & (3) Mu2e-II
- Upgraded Booster
 - 20 Hz, 800 MeV injection
 - New injection area
 - Resonant Magnet Upgrade
- Upgraded Recycler & MI
 - RF in both rings
- Conventional Facilities
 - Includes 2 empty slots at the linac end ($L \approx 23$ m)
 - Up to 1 GeV
- Cryogenic Plant



6 12/12/2017 Paul Derwent | PIP-II Conceptual Design

Received DOE CD1* approval Aug '18. Operations start ~ 2027

* DOE Critical Decision 1: Approve alternative selection and cost range. Conceptual design released

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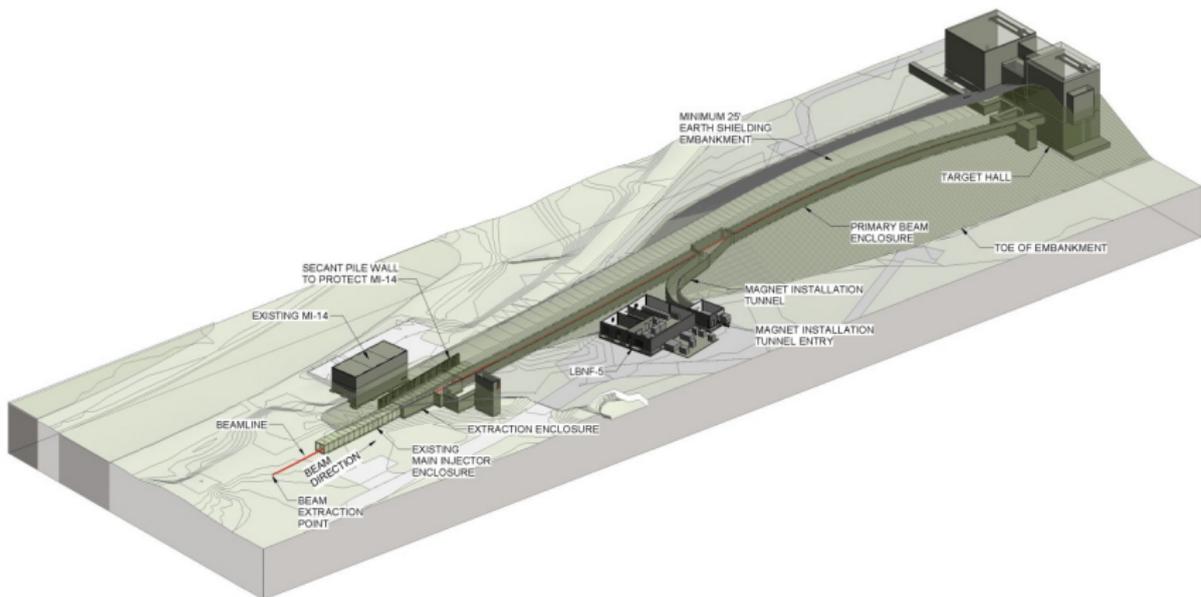
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Primary Beamline Design

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Parameter	Protons per cycle	Cycle Time (sec)	Beam Power (MW)
≤ 1.2 MW Operation - Current Maximum Value for LBNF			
Proton Beam Energy (GeV):			
60	7.5E+13	0.7	1.03
80	7.5E+13	0.9	1.07
120	7.5E+13	1.2	1.20
≤ 2.4 MW Operation - Planned Maximum Value for LBNF 2nd Phase			
Proton Beam Energy (GeV):			
60	1.5E+14	0.7	2.06
80	1.5E+14	0.9	2.14
120	1.5E+14	1.2	2.40

Beam optics point to 79 conventional magnets: 25 dipoles, 21 quadrupoles, 23 correctors, 6 kickers, 3 Lambertsons, 1 C Magnet

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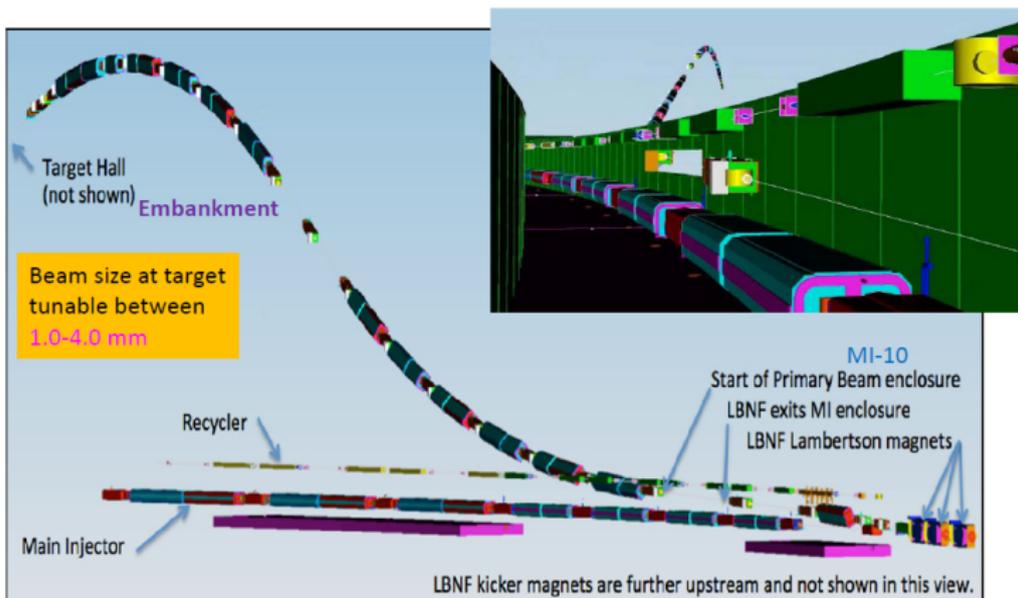
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Primary Beamline Design

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Table 6-3: Summary of primary-beam magnet specifications

Magnet	Description	Steel Length	Nom. Strength at 120 GeV	Count
RKB type Kicker	LBNF extraction	~1.3 m	0.058 T	6
ILA	MI Lambertson	2.800 m	0.532 / 1.000 T	3
ICA	MI C Magnet	3.353 m	1.003 T	1
IDA	MI Dipole 6 m	6.100 m	1.003 - 1.604 T	13
IDD	MI Dipole 4 m	4.067 m	1.003 - 1.604 T	12
3Q120	120 inch quadrupole	3.048 m	9.189 - 16.546 T/m	17
3Q60	60 inch quadrupole	1.524 m	11.135 - 17.082 T/m	4
IDS	LBNF trim dipoles	0.305 m	Up to 0.365 T	23

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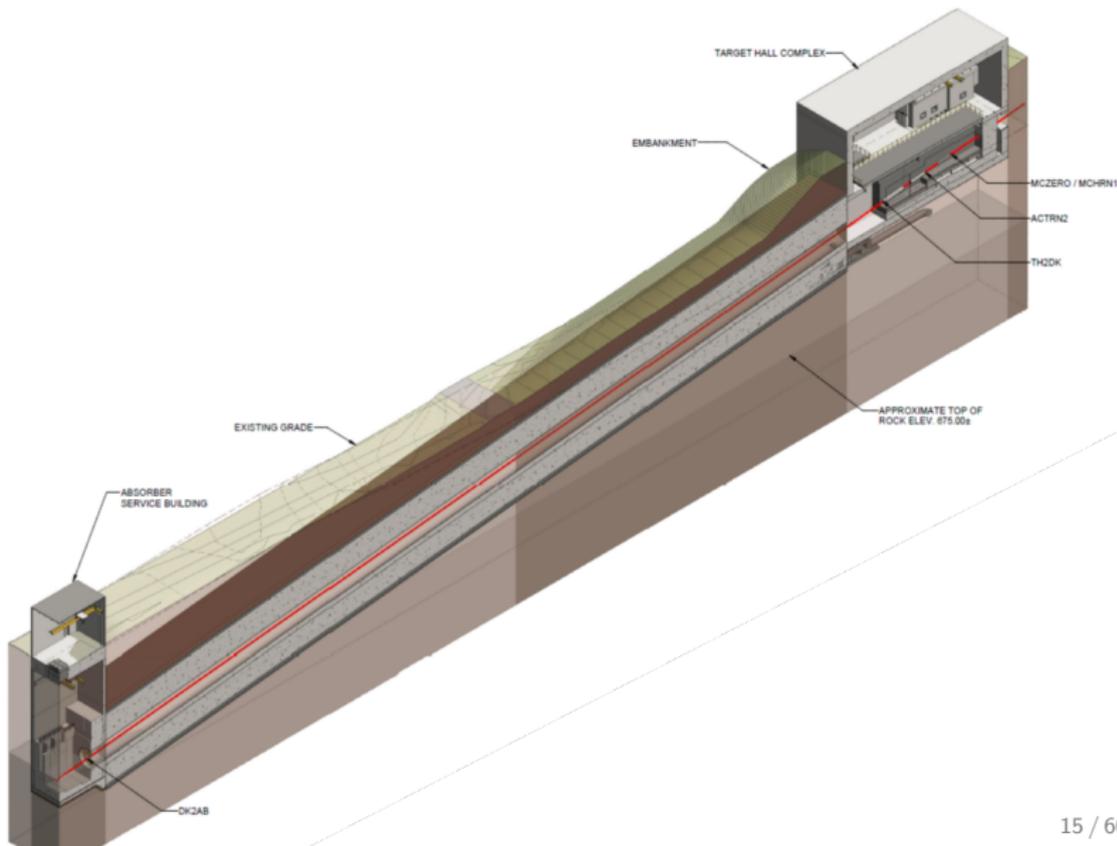
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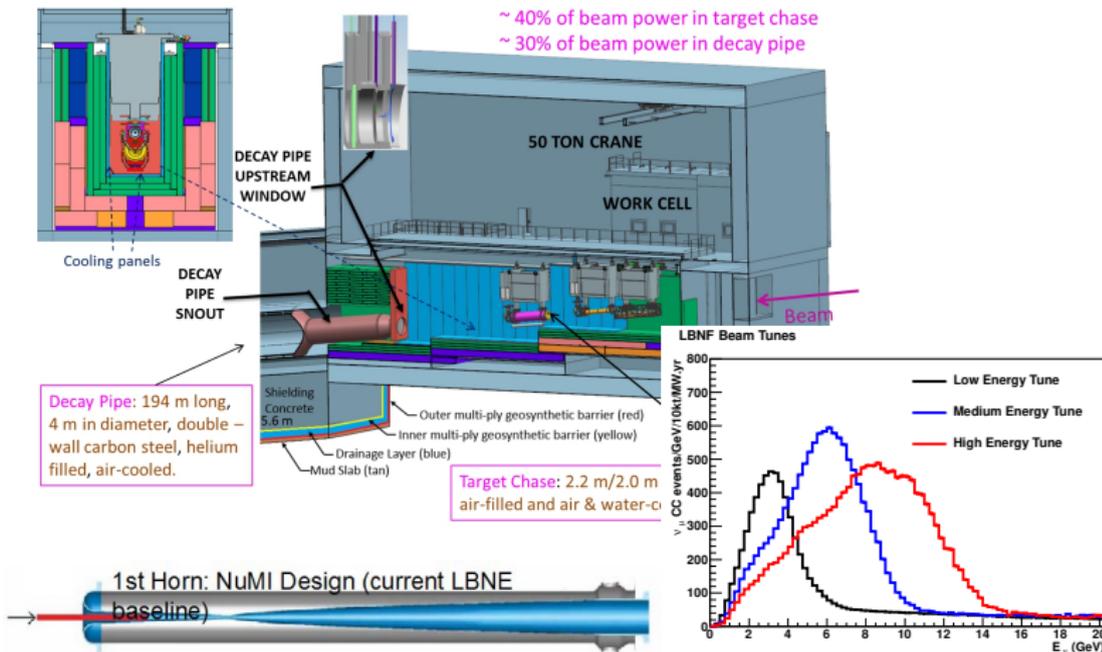
Future Beam Upgrades

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The LBNF Reference Design (up to 2015)

Initial conceptual design was of a *tunable wide-band* NuMI-style focusing:



LBNF has switched to CPV optimized focusing design with 3 horns

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Optimization of flux for Physics: CP Violation

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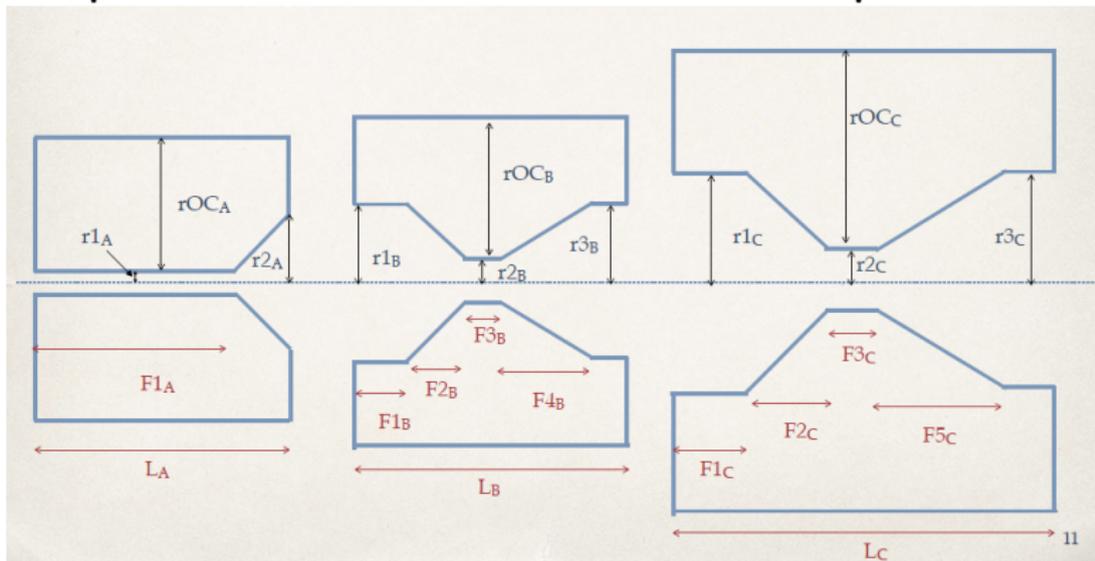
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- The 2015 CD1R reference design for LBNF/DUNE is a NuMI-like movable target (segmented rectangular graphite fins with water cooling $\approx 1\text{m}$ long) and 2 modified NuMI horns 6.6m apart
- In Sep 2017 LBNF adopted a focusing design with 3 horns optimized using a *genetic algorithm* with the physics parameter to be measured (CPV sensitivity) used to gauge fitness.
- Target geometry is optimized at the same time, as well as proton beam energy with realistic Main Injector power profile (1.03 MW at 60 GeV to 1.2 MW at 120 GeV).
- Limits on horn current, diameter and length are imposed based on experience with T2K and NuMI horn manufacturing
- Limits on horn separation imposed based on size of target chase.

Horn parameters used in GEANT4 simulation for GA optimization:

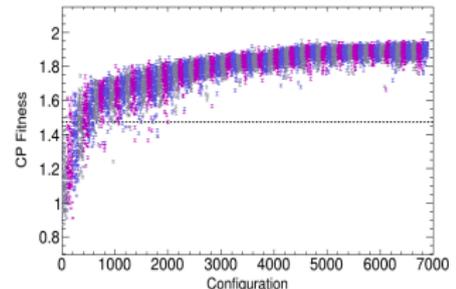


Schematic of the Genetic Algorithm:



CP Fitness = minimum significance with which 75% of δ_{cp} can be determined $\neq 0$ or π for a given exposure

Fast CP Fitness estimator was determined by calculating the change in CP sensitivity given some fixed change in a single energy bin.



CP Fitness vs configuration

Optimized horn design with 297kA current :



Parameter	Value	Parameter	Value
Horn A Length (mm)	2218	Horn A F1 (% of length)	53
Horn A R1 (mm)	43	Horn A OC Radius (mm)	369
Horn A R2 (mm)	33		
Horn B Length (mm)	3932	Horn C Length (mm)	2184
Horn B R1 (mm)	159	Horn C R1 (mm)	284
Horn B R2 (mm)	81	Horn C R2 (mm)	131
Horn B R3 (mm)	225	Horn C R3 (mm)	362
Horn B F1 (% of length)	31	Horn C F1 (% of length)	20
Horn B F2 (% of length)	22	Horn C F2 (% of length)	9
Horn B F3 (% of length)	2	Horn C F3 (% of length)	7
Horn B F4 (% of length)	16	Horn C F4 (% of length)	35
Horn B OC Radius (mm)	634	Horn C OC Radius (mm)	634
Horn B Position (mm)	2956	Horn C Position (mm)	17806

Optimized target is 4λ (2m C) with $\sigma_{\text{beam}} = 2.7\text{mm}$, $E_p \sim 110\text{ GeV}$



Optimization of flux for Physics: CP Violation

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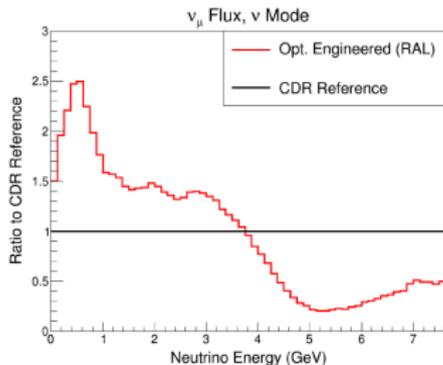
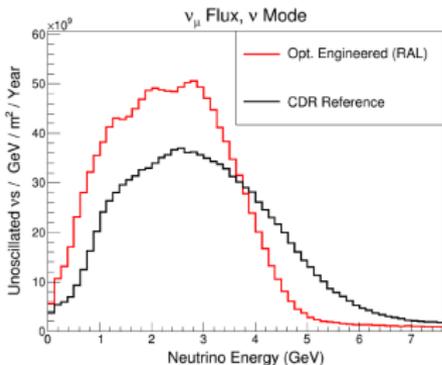
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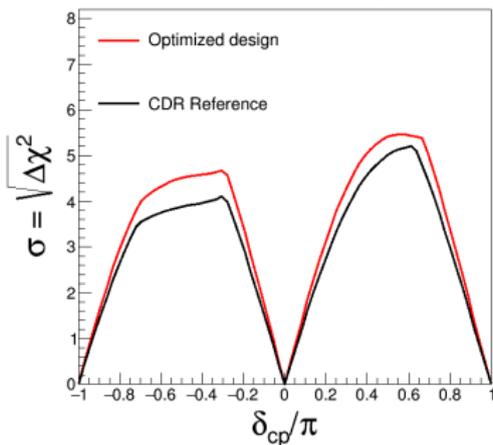
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CP violation sensitivity



Computationally advanced optimization techniques = significant gain in flux and CPV sensitivity from many small changes

Gain in sensitivity \equiv 70% increase in FD mass



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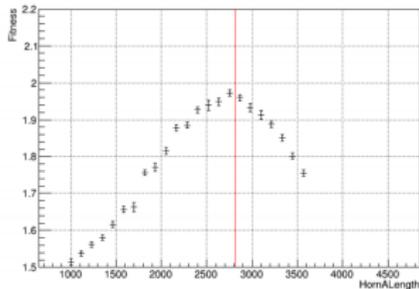
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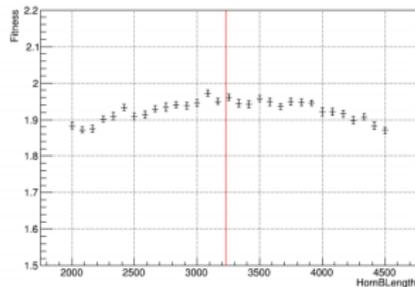
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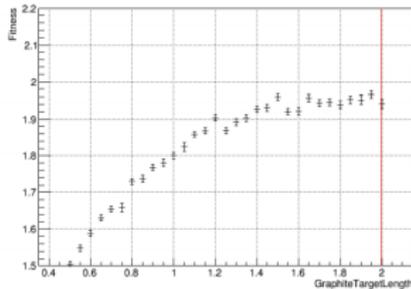
Scan over some sample optimization parameters:



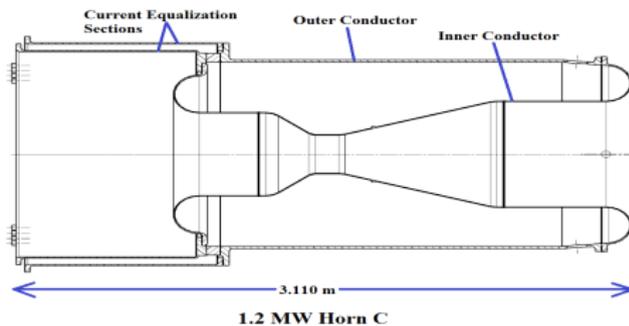
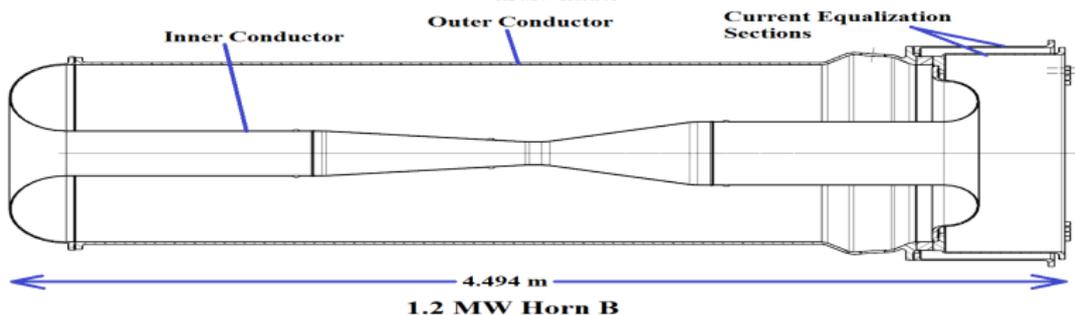
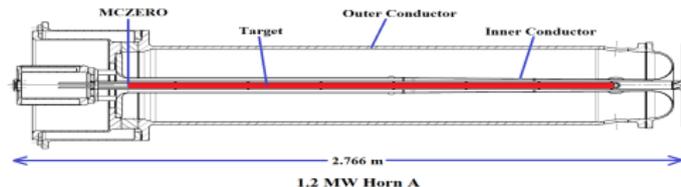
Horn A length



Horn B length



Target length



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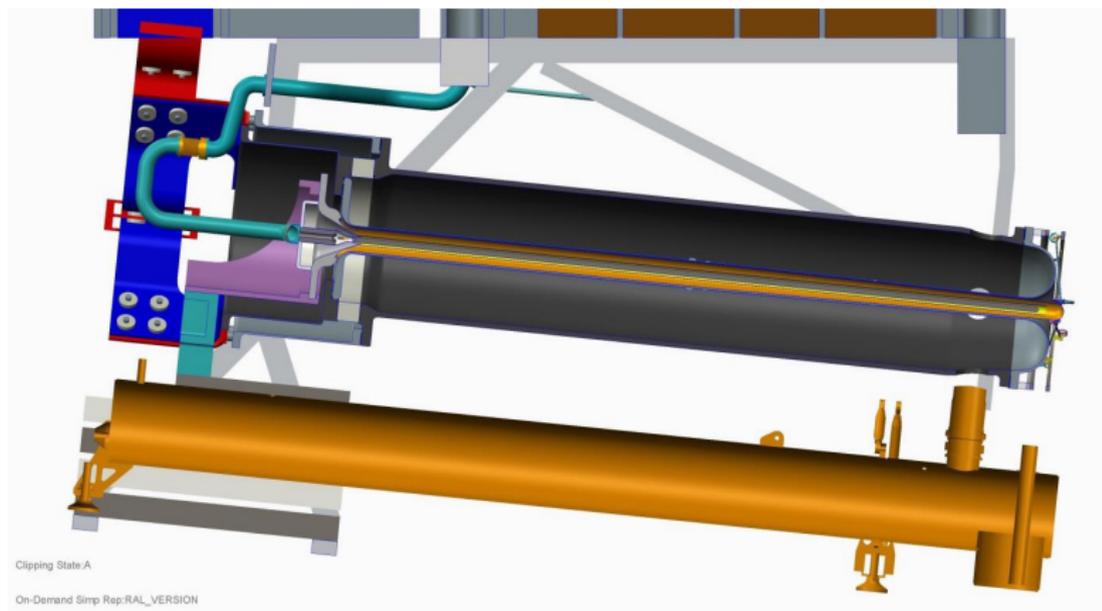
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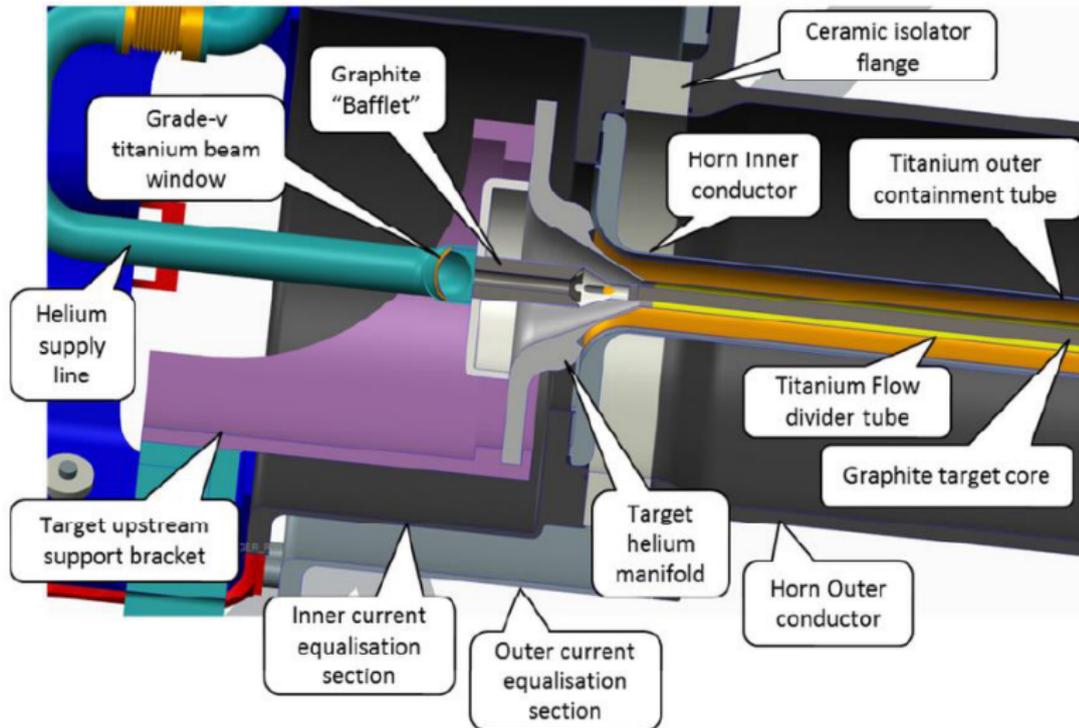
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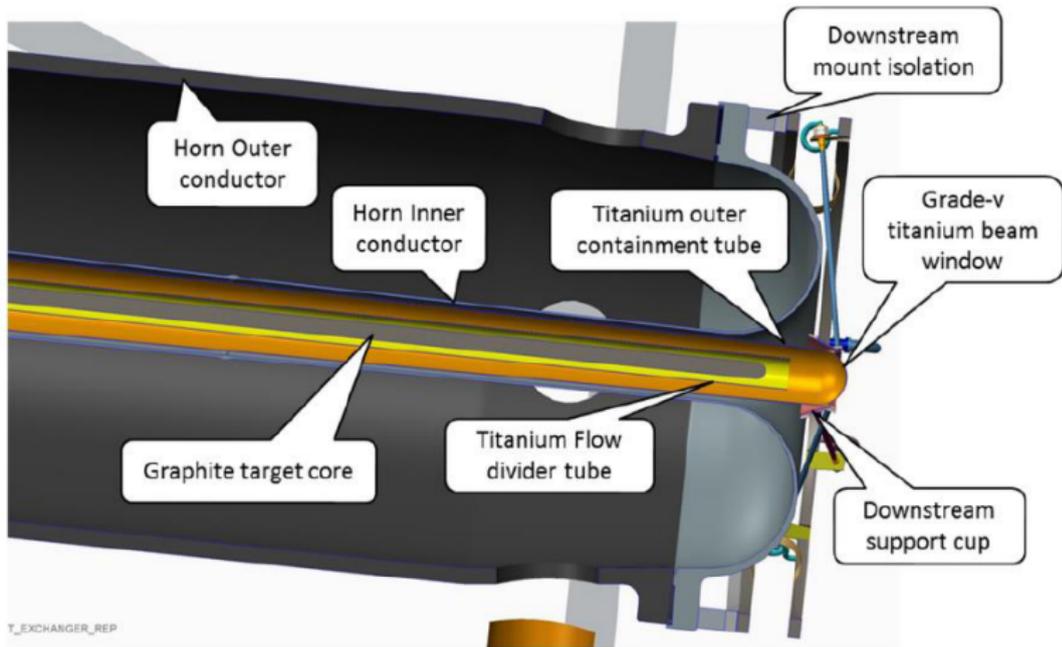
Horn A and 2m target design



Upstream section of target in Horn A (RAL design)



Downstream section of target in Horn A (RAL design)



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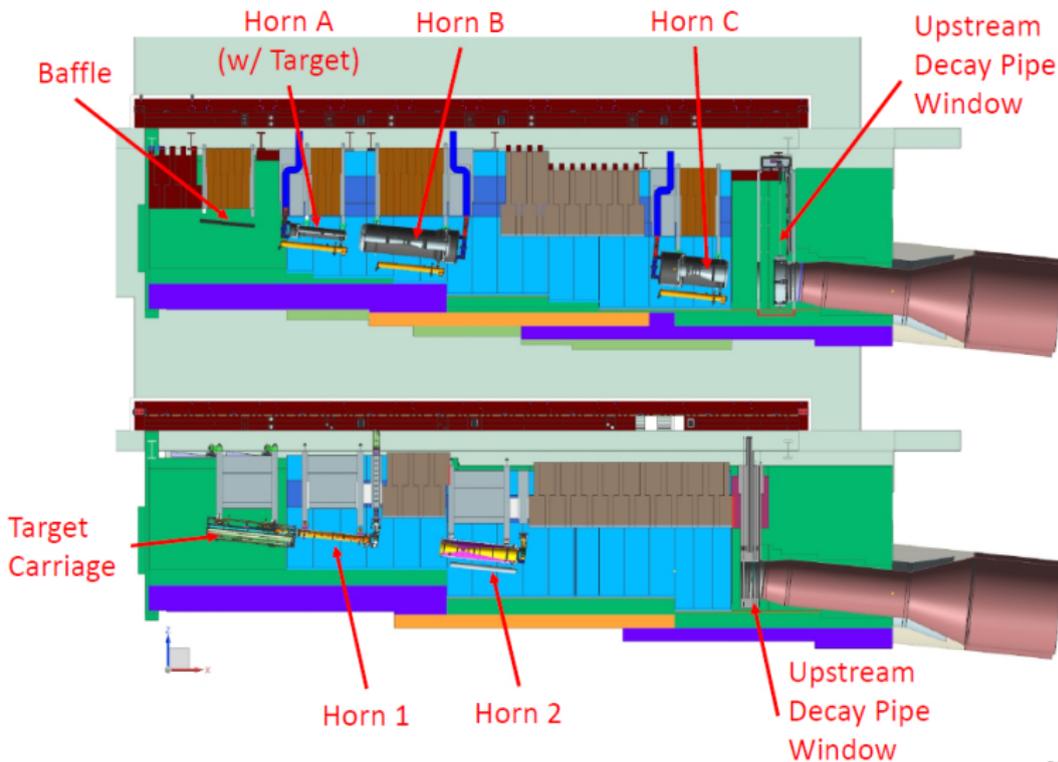
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Target hall layout 3-horn optimized vs CD1R reference designs:



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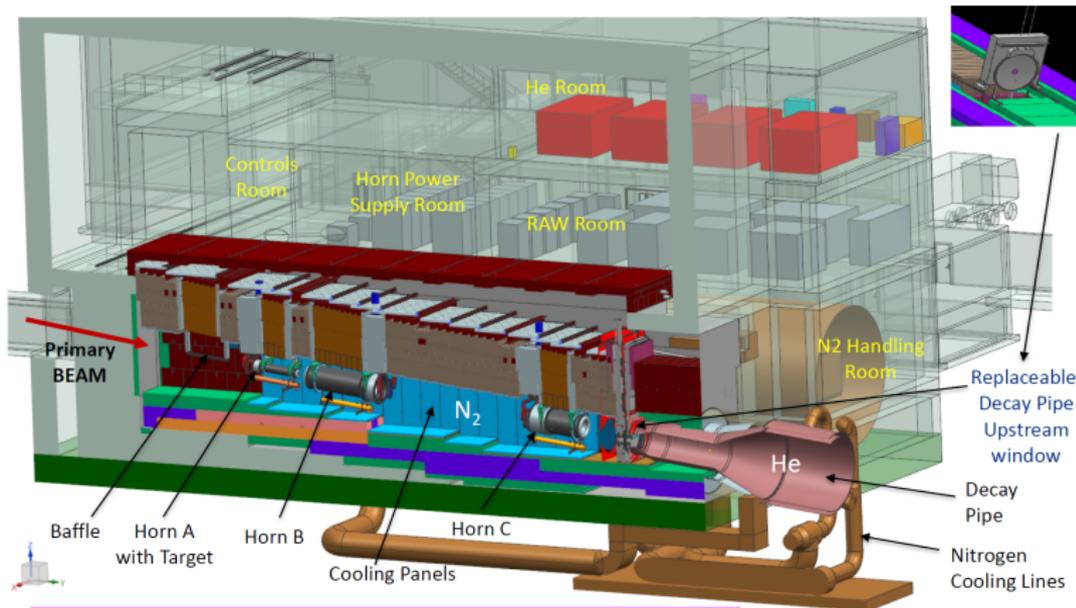
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Target Chase: 2.2 m/2.0 m wide, 34.3 m long nitrogen-filled and nitrogen plus water-cooled (replaceable cooling panels).

Decay Pipe Design

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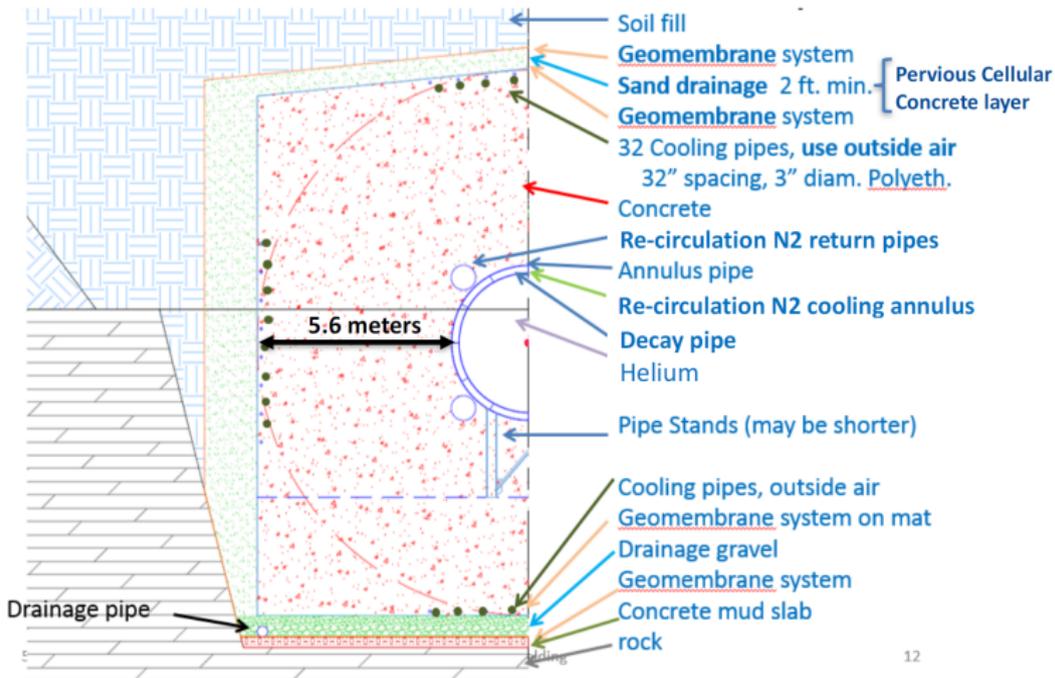
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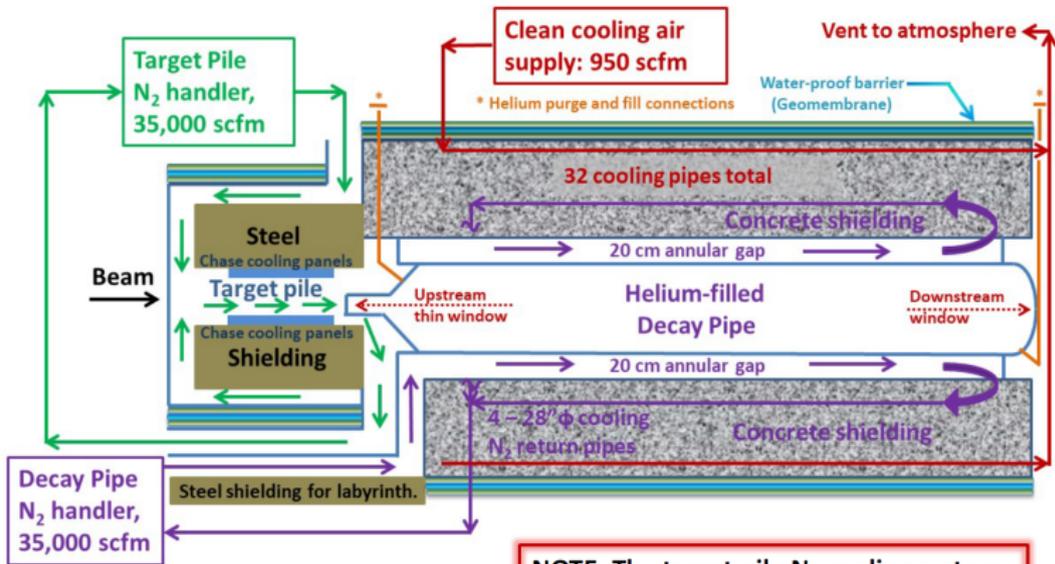
Future Beam Upgrades

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Decay pipe is 194m long and 4m in diameter, He filled



NOTE: The target pile N₂ cooling system and the decay pipe N₂ cooling system are two separate systems.

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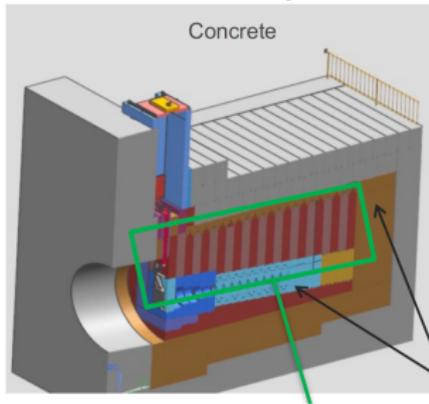
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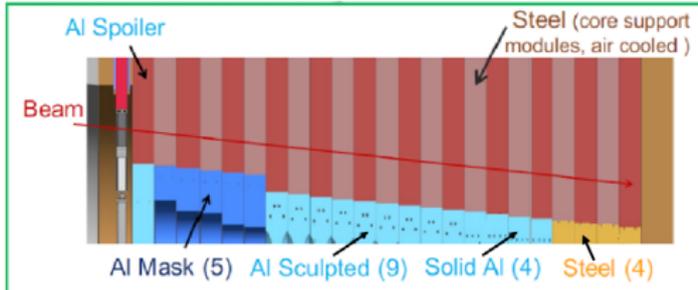
Absorber Core (water cooled)



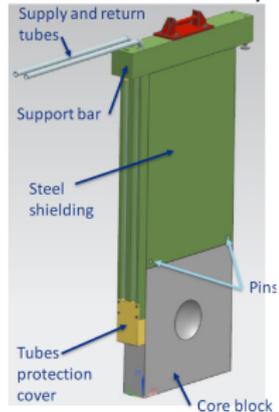
- Made of removable, water-cooled Al and Steel "T-Blocks" (based on NuMI T-Block design)
- Gun drilled water channels in core blocks.
- Flexible, modular design that can be repaired/replaced or upgraded in the future.

Steel
(air cooled)

Absorber Core – Water Cooled



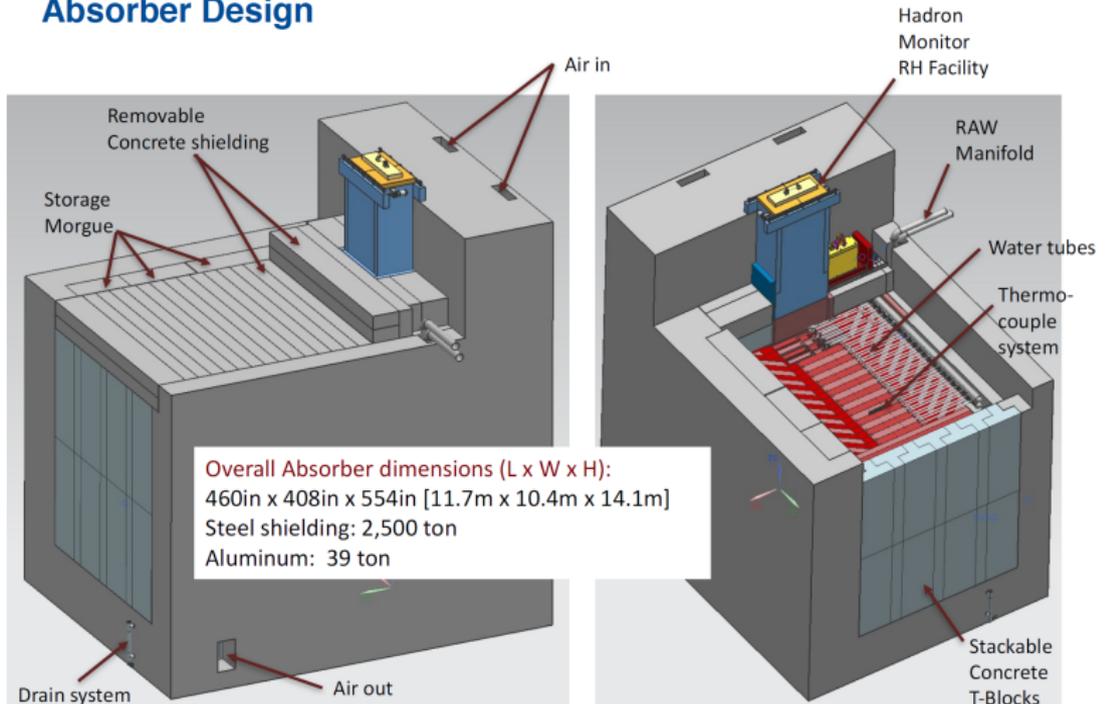
Absorber Core Module Assy.



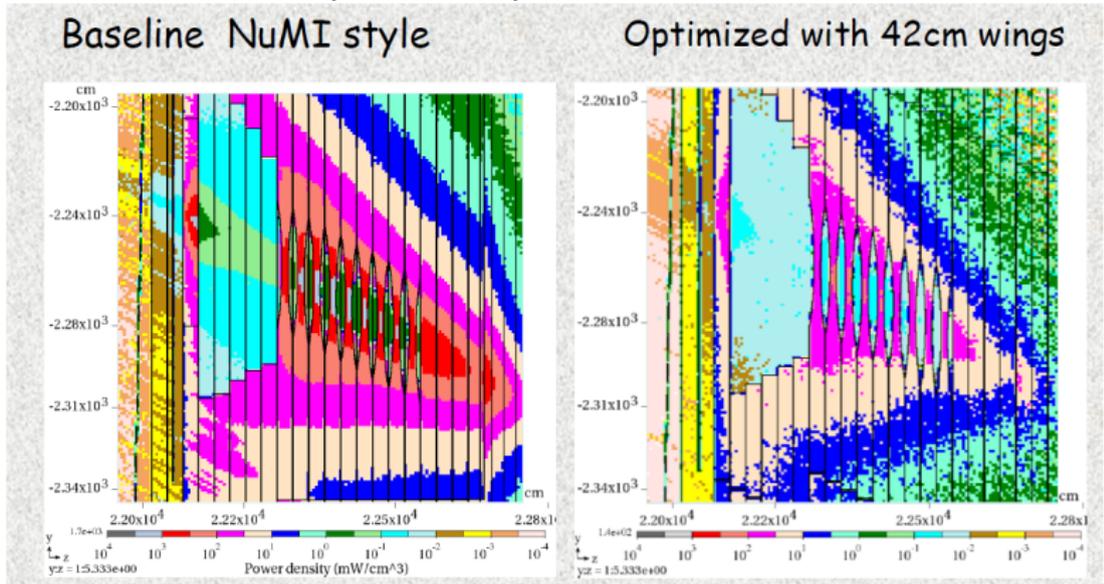
NuMI T-Block



Absorber Design



MARS15 simulation (N. Mokhov):

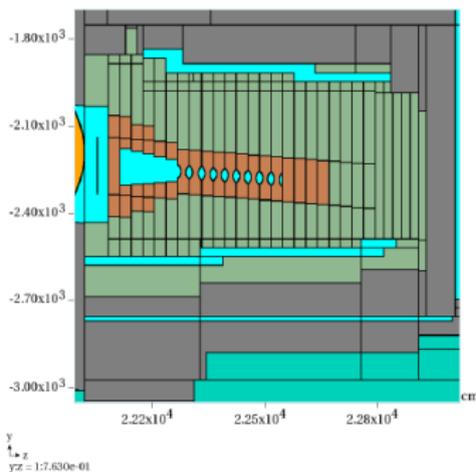


Optimized focusing design and wings added to target could reduce peak energy deposition in absorber core 8-12 ×.

Absorber design at CD1R (2015) was not optimized - non-uniformity makes monitoring of tertiary beam muons difficult. Uniform absorber design developed in 2017-2018:

Ref. Hadron Absorber (RHA)

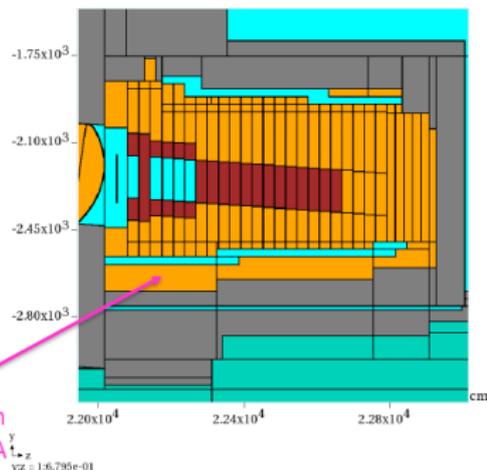
Non-uniformity, sculpting needed for RD



CFD simulation points to temperature reduction by 40% (89°C) in the optimized design with UHA

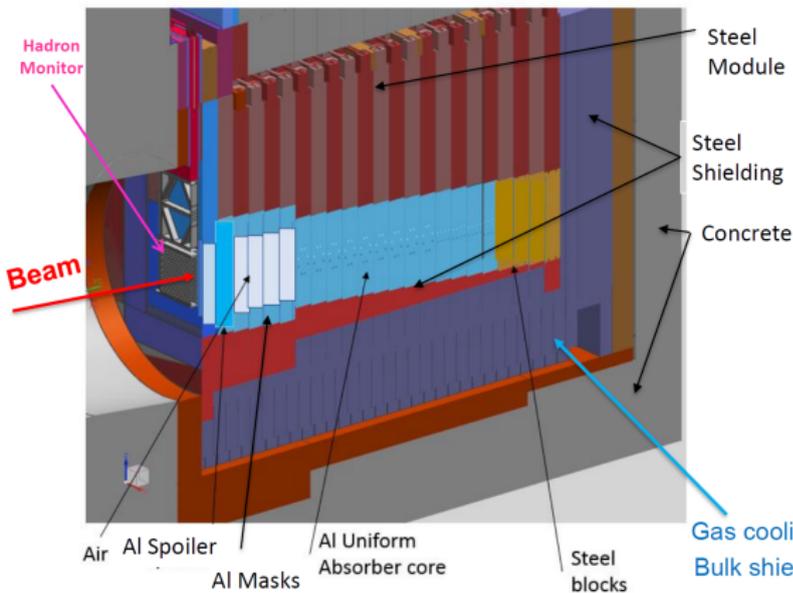
Uniform Hadron Abs (UHA)

No sculpting, larger uniform masks, larger core blocks (60"→67"), 1/16" windows on mask blocks – better for muon measurements

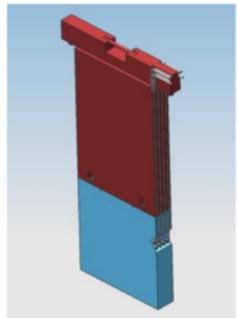


Uniform absorber is now LBNF baseline design

Large diameter core with uniform density to enable good muon monitoring



Flexible, modular hanger design of water cooled core



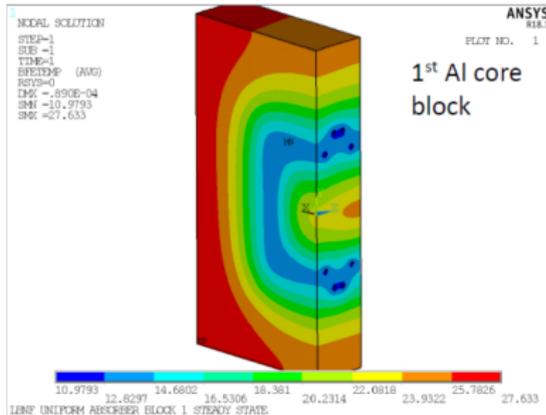
1 ft thick core blocks are easily replaceable via Remote Handling

Gas cooling of Bulk shielding

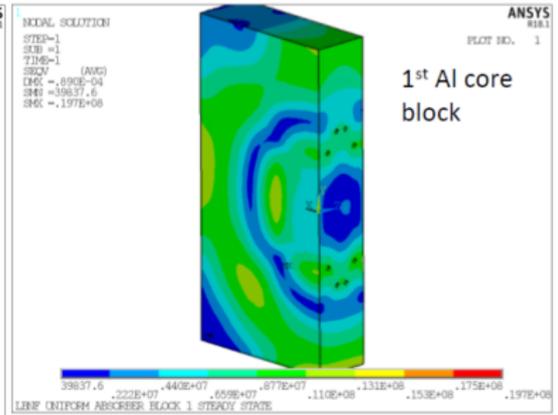
MARS15-ANSYS analysis of the uniform absorber core in the optimized beam completed for normal case. Spoiler and steel blocks are cool and show low stress:

UHA – 28 C M_{ax} (RHA 88 C M_{ax})

UHA – 20 MP_a M_{ax} (RHA 103 MP_a M_{ax})



Temperature



Pressure

Accident case has been analyzed and is fine



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Beam Flux Systematics with Optimized Design



Neutrino Beam Flux Comparisons

(For 50 kton.years* of exposure. No detector effects)

Experiment	Baseline	$\nu_\mu \rightarrow \nu_\mu$	$\nu_\mu \rightarrow \nu_\tau$	$\nu_\mu \rightarrow \nu_e$ δ_{CP} range
Current Experiments				
T2K 30 GeV, 750 kW 9×10^{20} POT/year	295km (off-axis)	900	< 1	40 - 70
MINOS LE 120 GeV, 700 kW 6×10^{20} POT/year	735km	11,000	115	230-340
NO ν A 120 GeV, 700 kW 6×10^{20} POT/year	810km (off-axis)	1500	10	120 - 200
Future Experiments				
LBNF/DUNE LE † 80 GeV, 1.1MW 1.5×10^{21} POT/year	1,300km	4300	160	350 - 600
LBNF/DUNE ME † 120 GeV, 1.2MW 1.1×10^{21} POT/year	1,300km	12,000	690	290 - 430
T2HK 30 GeV, 1.3 MW		(see T2K numbers)		

* Facility duty factor taken into consideration

† 2012 LBNE CDR Reference Design with NuMI style focusing

Future LB experiments expect $\mathcal{O}(1000)$ $\nu_\mu \rightarrow \nu_e$ CC events

Need to control *total* systematics to a few %

The Science and Design of the LBNF Beamline

Mary Bishai (on behalf of LBNF/DUNE) Brookhaven National Lab

DUNE Science Requirements

Beamline overview

Primary Beam Neutrino Beam Decay Pipe Absorber

Beam Flux Systematics

Beam Flux Stability

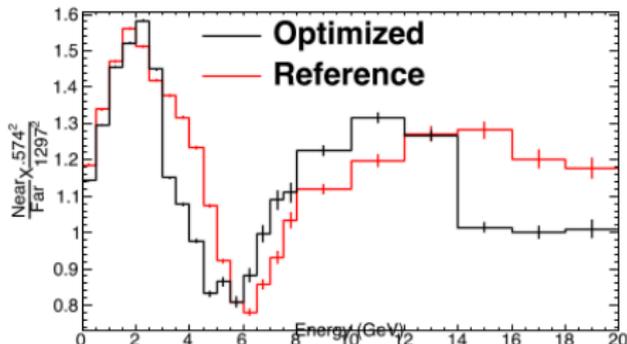
Future Beam Upgrades

Summary

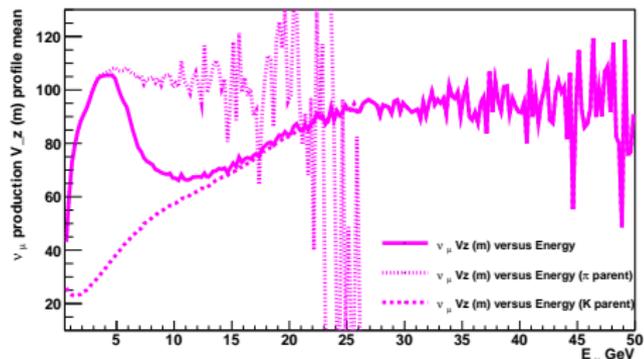
Simple ratio of Near/Far spectrum:

Neutrino parent decay location in decay pipe:

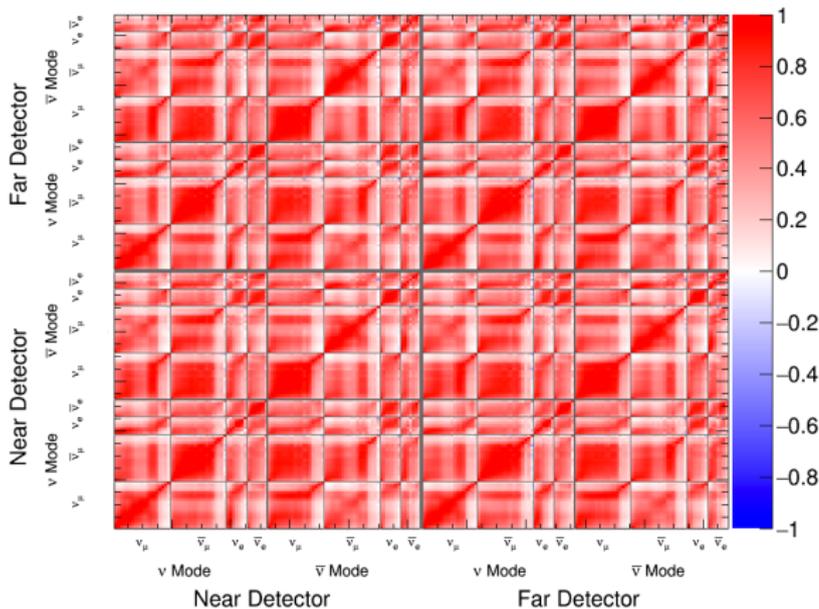
π/K decay kinematics and decay channel geometry are primary reason for strange shape of N/F ratio



ν_{μ} events at FD (1300km)



To correctly relate near to far fluxes - need to use a correlation matrix:



Flux correlation matrix comes from simulation and is highly correlated

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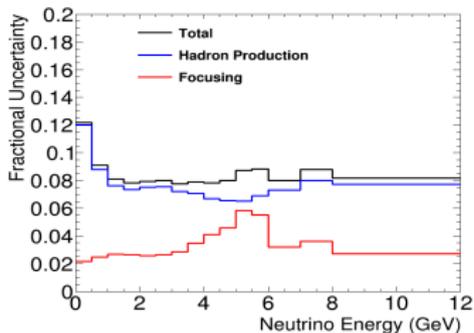
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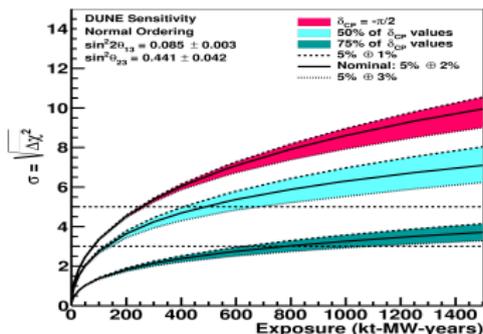
Future Beam Upgrades

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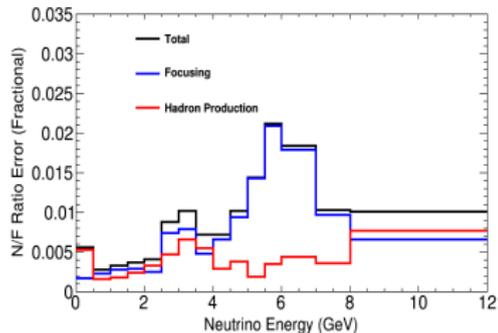
Uncertainty on ND flux prediction



CP Violation Sensitivity



Residual uncertainty at FD



How well do we actually trust the simulation to correctly estimate the uncertainties on near \rightarrow far extrapolation?

Proposal by Laura Fields at Fermilab:

The **LBNF Spectrometer** is a concept for a thick-target hadron production measurement **after the focusing horns**. It would involve a **replica of the LBNF target and horns** in an external beamline at Fermilab. In addition to hadron production in the target, the spectrometer would also **measure hadron production and absorption in the horns and the effects of the magnetic fields** in the horns.



Detector technology is always challenging

Need to get more people interested and involved to succeed



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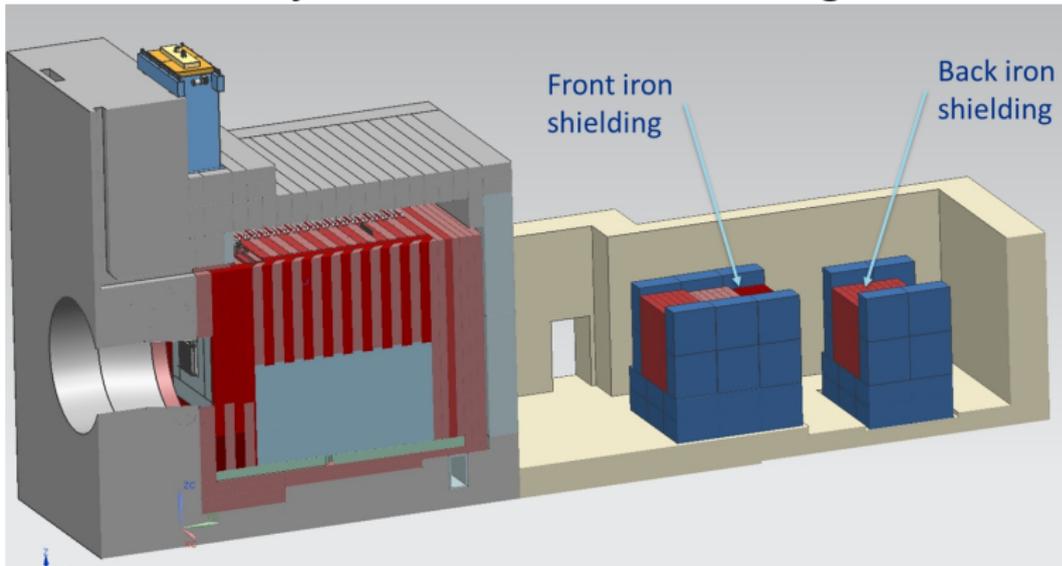
**Beam Flux
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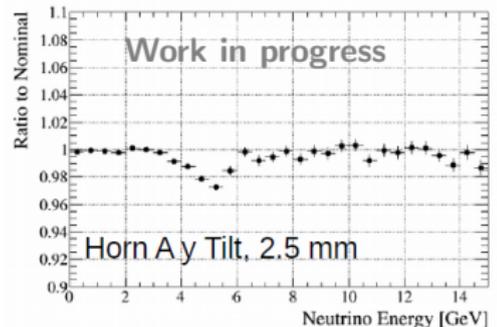
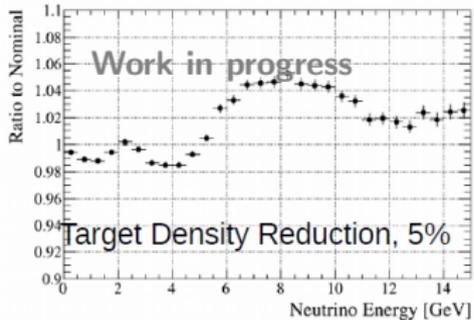
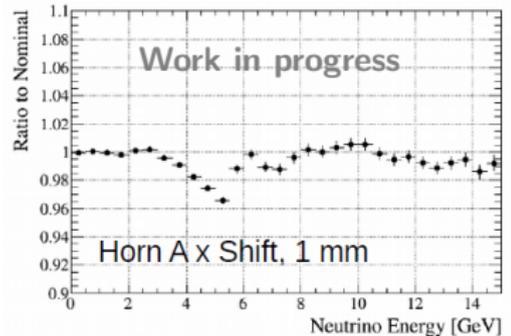
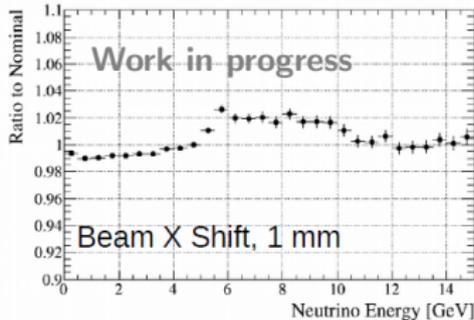
Monitoring the Neutrino Flux Stability

Layout of Muon Alcove and Shielding



High intensity makes it difficult to measure μ spectrum accurately. With a 2.4 MW beam, the absorber thickness is too large to sample the lower energy muons. But these systems play an essential role in monitoring *flux stability*

ν Spectrum Changes



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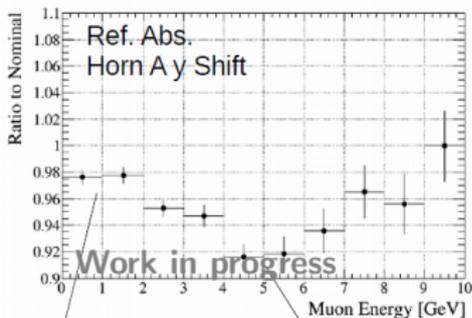
Beam Flux Systematics

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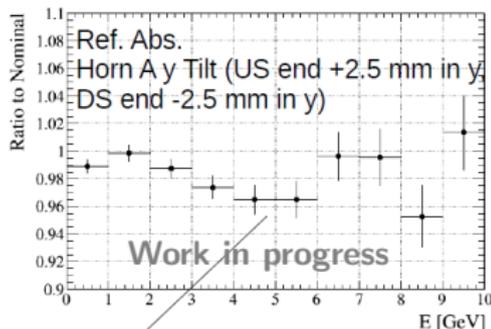
Summary

μ Spectrum Changes



Reduction in total flux

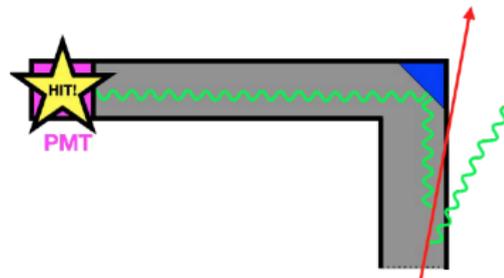
Shape changes at max near 5 GeV



Changes are v. small - need novel detector concepts

- **Array of ionization detectors:** Measures muon beam center and intensity. Spill by spill monitoring of beam stability. Both diamond and silicon under study
- **Threshold gas Cherenkov detector (R&D):** Uses signal intensity at different gas pressure and angles to extract rough muon spectrum.
- **Stopped muon counters (R&D):** separate stations with steel shielding in between could measure muon flux at several energies. Better measurement of beam flux spectrum and composition.

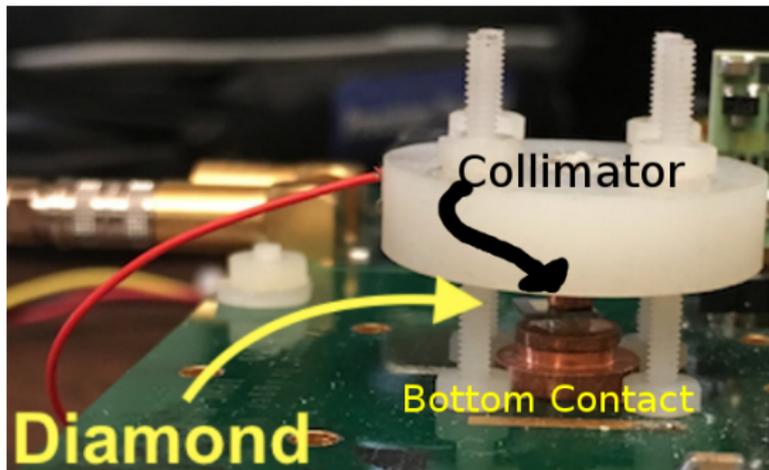
Gas Cherenkov counter concept:



Prototype in NuMI beamline:



Currently only ionization detectors included in the beam design.



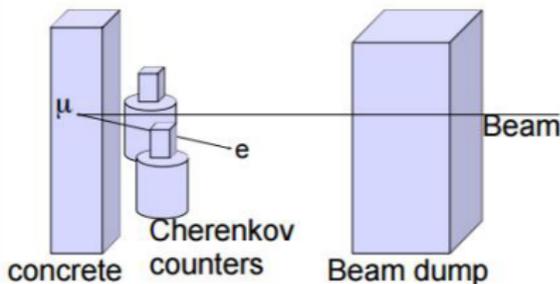
Use polycrystal chemical vapor deposition (pCVD) diamond - detects ionizing radiation when a large voltage potential (1V per μm of thickness) is applied across two sides of the diamond. Diamond is radiation hard.

pCVD detector prototype installed in NuMI during 2018 shutdown.



From K. Hiraide, *Muon monitor using the decay electrons*, NBI2003 Workshop

Strategy

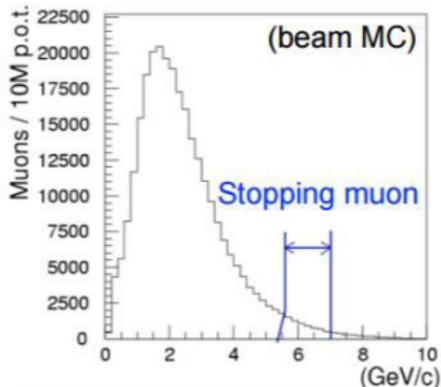


- Counting the decay electrons from muons stopping at the wall of μ -pit
- Measuring spatial and time distributions of events

- Energy loss of muons in the beam dump
- Range of electrons in the concrete



We can measure muons of
5.2~7.0 GeV/c
 by counting the decay electrons



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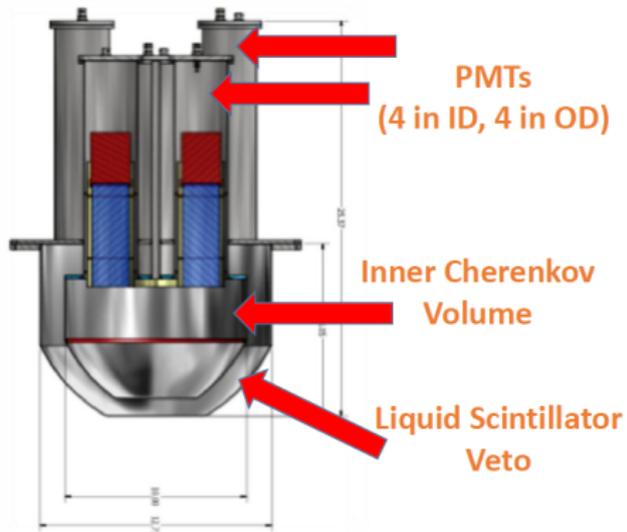
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Prototypes being commissioned with cosmics.



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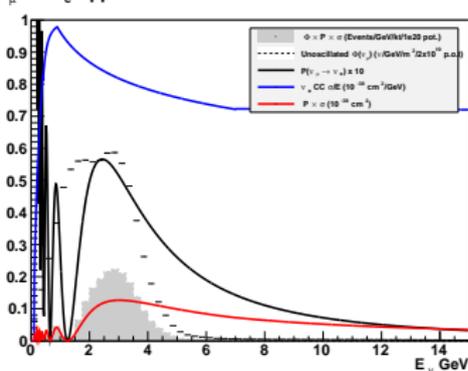
Summary

Possible Future Upgrades to the LBNF Beamline to enhance the DUNE Physics Program

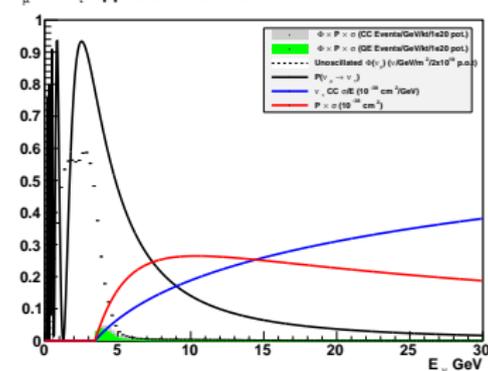
NuMI-like reference design could be tuned to higher energy to observe $\nu_\mu \rightarrow \nu_\tau$ with high statistics.

2015 two horn optimized design $E_p = 66$ GeV:

$\nu_\mu \rightarrow \nu_e$ Appearance at 1300 km



$\nu_\mu \rightarrow \nu_\tau$ Appearance at 1300 km



$\nu_\mu \rightarrow \nu_e$ 290 events

$\nu_\mu \rightarrow \nu_\tau$ 60 events

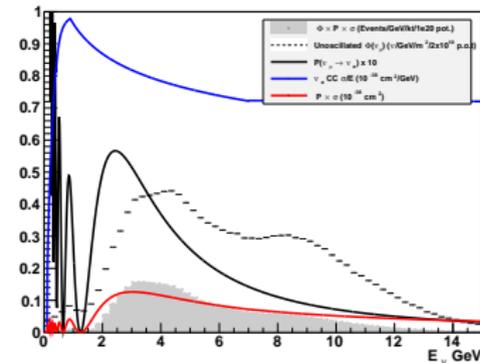
in 40 ktons, 1 year at 1.2 MW

Tunability of beam and ν_τ Flux (v. preliminary)

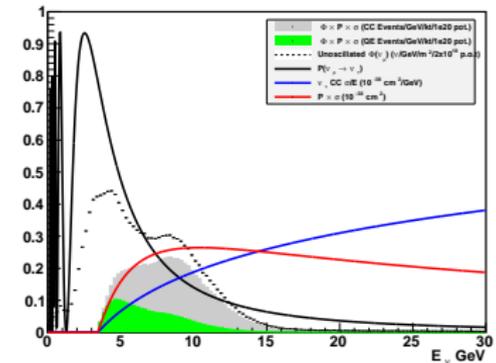
NuMI-like reference design could be tuned to higher energy to observe $\nu_\mu \rightarrow \nu_\tau$ with high statistics.

LBNF target -2m from horn 1, NuMI focusing 230 kA, horns 17m apart

$\nu_\mu \rightarrow \nu_e$ Appearance at 1300 km



$\nu_\mu \rightarrow \nu_\tau$ Appearance at 1300 km



$\nu_\mu \rightarrow \nu_e$ 330 events

$\nu_\mu \rightarrow \nu_\tau$ 700 events

in 40 ktons, 1 year at 1.2 MW

Increase ν_τ appearance 10x!!!

Increase high energy ν_e appearance - good for NSI/Sterile searches

Masud, M. Bishai

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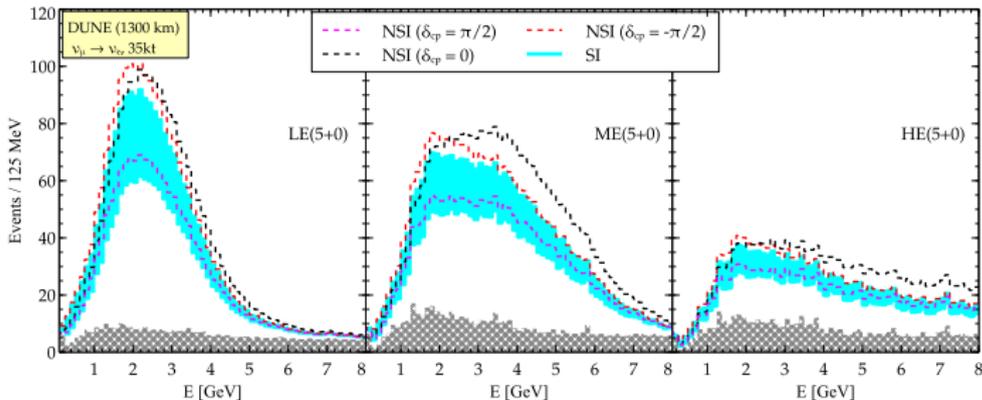
Future Beam Upgrades

Summary

Study NSI sensitivity with GLoBeS using $\nu_\mu \rightarrow \nu_{\mu,e}$ and 3 sample LBNF-like beam tunes : LE, ME and HE*.

NSI parameters used:

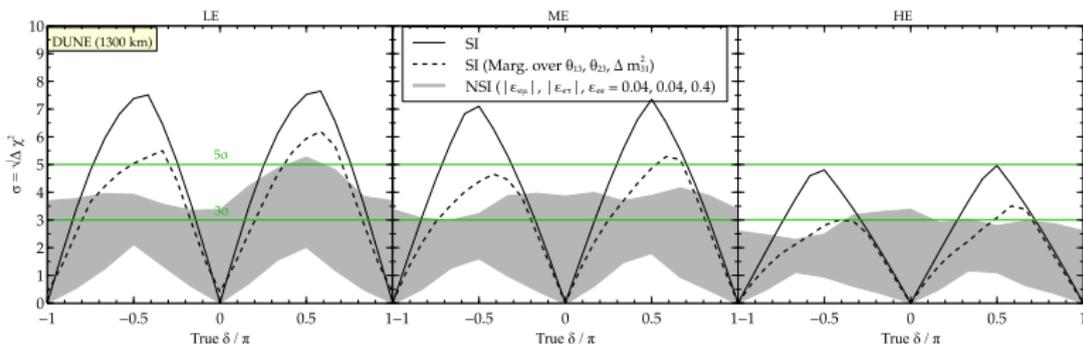
$$|\epsilon_{e\mu}| = 0.04, |\epsilon_{e\tau}| = 0.04, \epsilon_{ee} = 0.4, \phi_{e\mu} = 0, \phi_{e\tau}$$



NSI effects in $\nu_\mu \rightarrow \nu_e$ are larger at higher energy

* 2 NuMI horns, 230kA, 6.6m apart and horns were not moved for higher energy beam tunes (non-optimal beams). Decay pipe was assumed to be 250m.

M. Masud, P. Mehta, M. Bishai arXiv:1704.08650



SI/NSI separation metric:

(the abbreviations *tr* and *ts* indicate *true* and *test* respectively.)

Thus for the NSI-SI separation,

$$\chi^2(\delta_{tr}) = \min_{\delta_{ts}} \sum_{i=1}^x \sum_j^2 \frac{\left[N_{NSI}^{i,j}(\delta_{tr}, |\epsilon|, \varphi) - N_{SI}^{i,j}(\delta_{ts} \in [-\pi, \pi]) \right]^2}{N_{NSI}^{i,j}(\delta_{tr}, |\epsilon|, \varphi)} \quad (2)$$

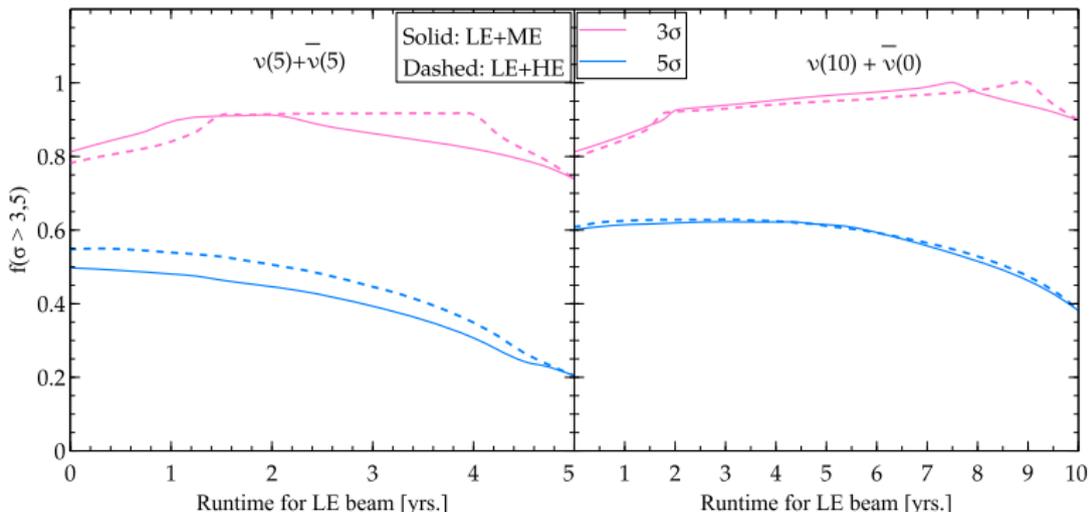
where, we have marginalised over the standard CP phase δ in the test dataset. This χ^2 was calculated using a set of conservative values of the NSI parameters (see Table 2).



SI/NSI Separation with Different Beams

M. Masud, P. Mehta, M. Bishai arXiv:1704.08650

Fraction of SI δ_{cp} for which SI/NSI can be separated at the $3/5\sigma$ level:



Can achieve 3σ separation for $> 80\%$ of true δ_{cp}

No beam optimization attempted yet!

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Summary and Conclusions



Summary and Conclusions

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Summary

- The LBNF beam design is advancing towards a preliminary design for CD2 (expected end of 2019)
- The CPV optimized neutrino beamline design with 3 horns is now the baseline design
- The challenging design of Horn A with a 4λ graphite target is proceeding
- The simplified hadron absorber redesign for the optimized beam is well advanced.
- A “real-time” beam monitoring system is a requirement for the LBNF beamline, and R&D for a muon monitoring system is proceeding but is still very challenging.
- New physics capabilities with high energy beams produced from the 2015 CD1R reference design are being explored.
- A full evaluation of the flux uncertainties from the optimized beam has been carried out. Uncertainties are at the 1% level in the oscillation region.