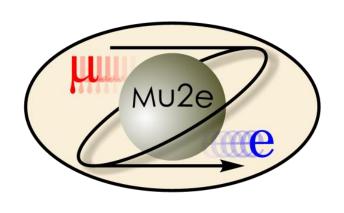
The Mu2e Experiment at Fermilab



STEVEN BOI, UNIVERSITY OF VIRGINIA

ON BEHALF OF THE MU2E COLLABORATION

NUFACT 2018

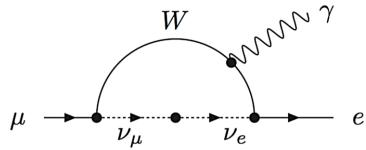


Charged Lepton Flavor Violation and the SM

In an extension of the standard model, one which includes neutrino oscillations, lepton flavor is only approximately conserved, with some charged lepton flavor violation (CLFV) occurring due to neutrino oscillations.

Such a process is exceedingly rare, a rate too small to be observed.

Bonus: since we will never observe this, it is not a background to CLFV searches.



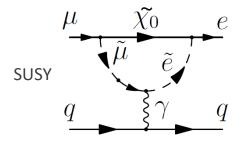
BR(
$$\mu \to e\gamma$$
) = $\frac{3\alpha}{32\pi} \left| \sum_{i=2,3} U_{\mu i}^* U_{ei} \frac{\Delta m_{1i}^2}{M_W^2} \right|^2 < 10^{-5\alpha}$



CLFV beyond the SM

While CLFV within the SM is effectively unobservable, the rate of $\mu + N \rightarrow e + N$ is uniquely sensitive to a host of BSM physics.

- Observation of $\mu \rightarrow e$ conversion is an unambiguous sign of new physics.
- Easy to produce high intensity muon beams.
- BSM models predict effects that would be measurable by nextgeneration experiments.



★★★
Large effects
★★
Small effects
★
No effects

Flavor physics effects for the most interesting observables in a selection of SUSY and non-SUSY models.

	4						
	AC	RVV2	AKM	δ LL	FBMSSM	LHT	RS
$D^0 - \bar{D}^0$	***	*	*	*	*	***	?
ϵ_K	*	***	***	*	*	**	***
$S_{\psi\phi}$	***	***	***	*	*	***	***
$S_{\phi K_S}$	***	**	*	***	***	*	?
$A_{\mathrm{CP}}\left(B \to X_{s}\gamma\right)$	*	*	*	***	***	*	?
$A_{7,8}(B\to K^*\mu^+\mu^-)$	*	*	*	***	***	**	?
$A_9(B o K^\star\mu^+\mu^-)$	*	*	*	*	*	*	?
$B \to K^{(\star)} \nu \bar{\nu}$	*	*	*	*	*	*	*
$B_s \to \mu^+ \mu^-$	***	***	***	***	***	*	*
$K^+ \to \pi^+ \nu \bar{\nu}$	*	*	*	*	*	***	***
$K_L o \pi^0 \nu \bar{\nu}$	*	*	*	*	*	***	***
$\mu \to e \gamma$	***	***	***	***	***	***	***
$\tau \to \mu \gamma$	***	***	*	***	***	***	***
$\mu + N \rightarrow e + N$	***	***	***	***	***	***	***
d_n	***	***	***	**	***	*	***
d_e	***	***	**	*	***	*	***
		***	**	***	***	*	?

A. de Gouvêa , et al

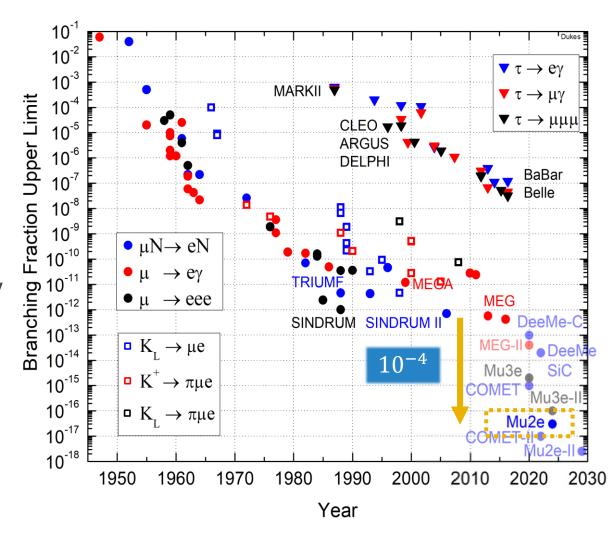


Historical CLFV Searches

Since the 1940s, searches for CLFV have been made, but so far only tighter constraints have been placed.

Next-generation experiments will explore unconstrained phase space favored by many BSM models.

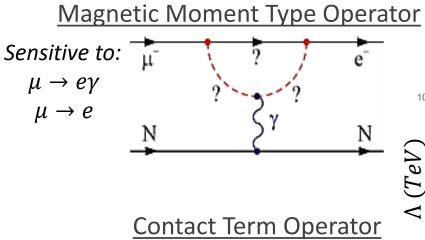
Mu2e intends to improve the sensitivity by four orders of magnitude over the present limit.

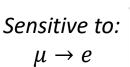


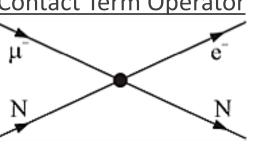
$\mu + N \rightarrow e + N$

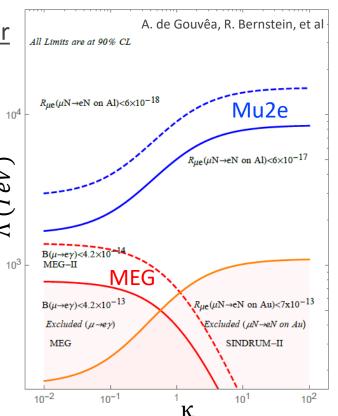
A simplified CLFV Lagrangian:

$$\mathcal{L}_{CLFV} = \frac{m_{\mu}}{(1+\kappa)\Lambda^2} \bar{\mu}_R \sigma_{\mu\nu} e_L F^{\mu\nu} + \frac{\kappa}{(1+\kappa)\Lambda^2} \bar{\mu}_L \gamma_{\mu} e_L \sum_{q=u,d} \bar{q}_L \gamma_{\mu} q_L$$









Λ: Mass Scale

Indirect probing of mass scales up to 10^4 TeV.

κ: relative strength of magnetic moment type and contact terms

doi:10.1146/annurev.nucl.58.110707.171126

de Gouvea and Vogel, arXiv:1303.4097v2

$\mu + N \rightarrow e + N$

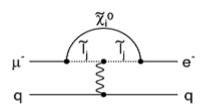
A simplified CLFV Lagrangian:

$$\mathcal{L}_{CLFV} = \frac{m_{\mu}}{(1+\kappa)\Lambda^2} \bar{\mu}_R \sigma_{\mu\nu} e_L F^{\mu\nu}$$

MAGNETIC MOMENT TYPE **OPERATOR**

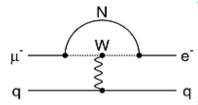
$$\mathcal{L}_{CLFV} = \frac{m_{\mu}}{(1+\kappa)\Lambda^2} \bar{\mu}_R \sigma_{\mu\nu} e_L F^{\mu\nu} + \frac{\kappa}{(1+\kappa)\Lambda^2} \bar{\mu}_L \gamma_{\mu} e_L \sum_{q=u,d} \bar{q}_L \gamma_{\mu} q_L$$

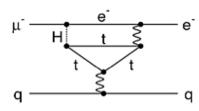
CONTACT TERM OPERATOR



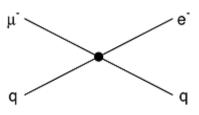
Supersymmetry $R_{\mu\rho} \sim 10^{-15}$

Heavy Neutrinos $\left|U_{\mu N}U_{eN}\right| \sim 8 \times 10^{-13}$





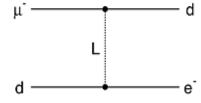
Two Higgs Doublet $g(H_{ue}) \sim 10^{-4} g(H_{uu})$

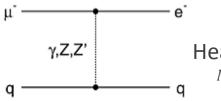


Compositeness $\Lambda_c \sim 3000 \, TeV$

Leptoquarks

$$M_{LQ} = 3000 \left(\lambda_{\mu d} \lambda_{ed} \right)^{\frac{1}{2}}$$





Heavy Z' $M_{7'} = 3000 \, TeV/c^2$

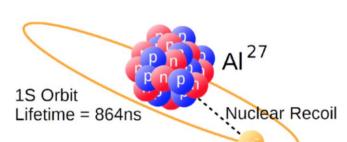
Mu2e will search for the coherent, neutrino-less $\mu \rightarrow e$ conversion in the presence of an aluminum nucleus.

$$\mu^{-} + (A, Z) \rightarrow e^{-} + (A, Z)$$

The signal is a monoenergetic electron:

•
$$E_e \approx m_{\mu} - E_b - E_{recoil} \approx 104.97 \; MeV$$

Signal can be distinguished from background



The ratio of $\mu + N \rightarrow e + N$ conversions to the number of muon captures by the aluminum nuclei will be measured:

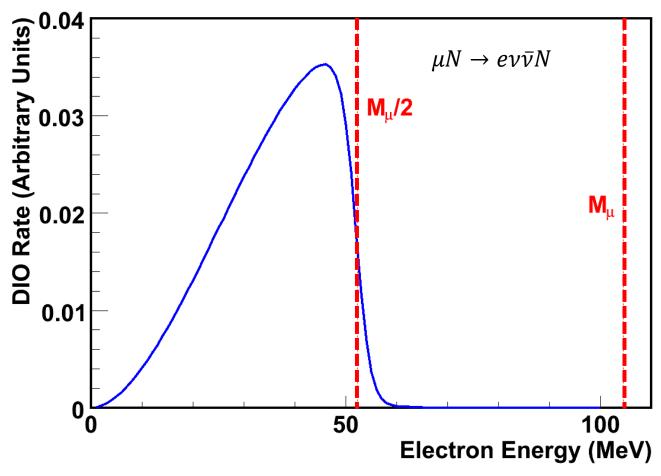
$$R_{\mu e} = \frac{\Gamma(\mu^{-} + (A, Z) \to e^{-} + (A, Z))}{\Gamma(\mu^{-} + (A, Z) \to \nu_{\mu} + (A, Z - 1))}$$

Mu2e intends to surpass the current limit in sensitivity by 4 orders of magnitude.

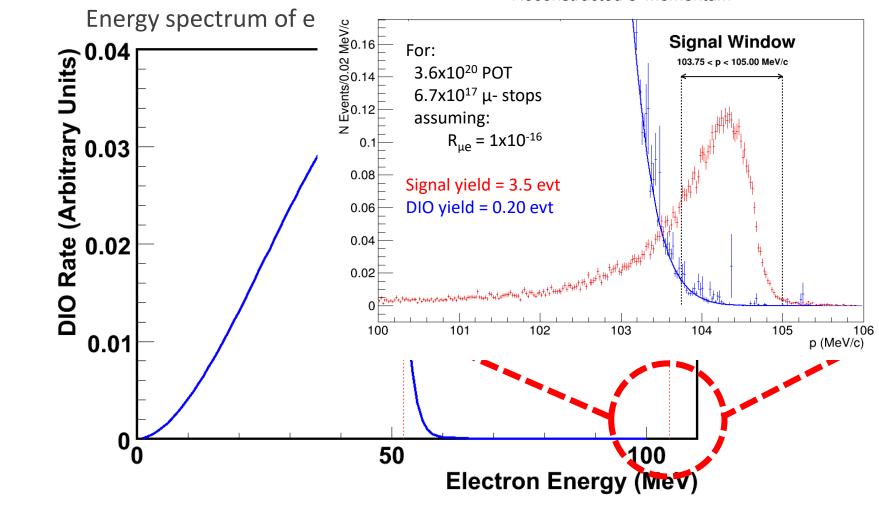
• Indirect probing of mass scales $\sigma(10^4 \text{ TeV})$

A single event sensitivity (SES) of 3×10^{-17} is required.

Energy spectrum of electron emitted by muon decay in orbit



Reconstructed e Momentum



Intent:

- Four order of magnitude improvement in sensitivity over the current limit
 - Set by SINDRUM II for $\mu + N \rightarrow e + N$
 - Probing of mass scales up to 10⁴ TeV

Achievements:

- World's highest intensity muon beam
 - Graded magnetic fields for greater muon collection
- Sub-event level management of backgrounds
 - Pulsed beam structure
 - High-efficiency Cosmic Ray Veto detector
 - Tracker blind to >99% muon decay in orbit spectrum
 - Fast timing Calorimeter with good energy resolution



The Mu2e Experiment

(AT FERMILAB)

Mu2e Collaboration

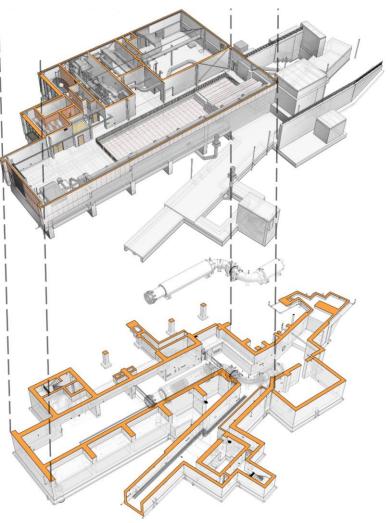
Comprised of 229 members from 37 Institutions.



Mu2e at Fermilab







Mu2e Beamline

Using the existing accelerator infrastructure, Mu2e will be the second building in the "Muon Campus" at Fermilab.

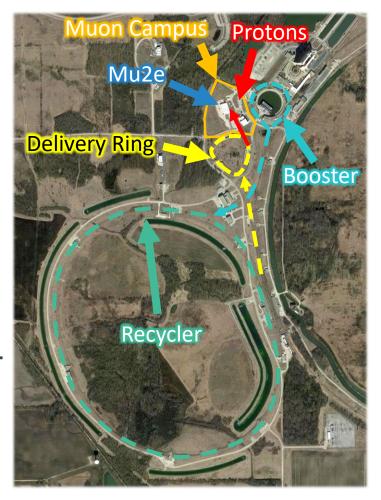
Booster provides 8 GeV protons to the Recycler.

Recycler stacks protons into 4 bunches.

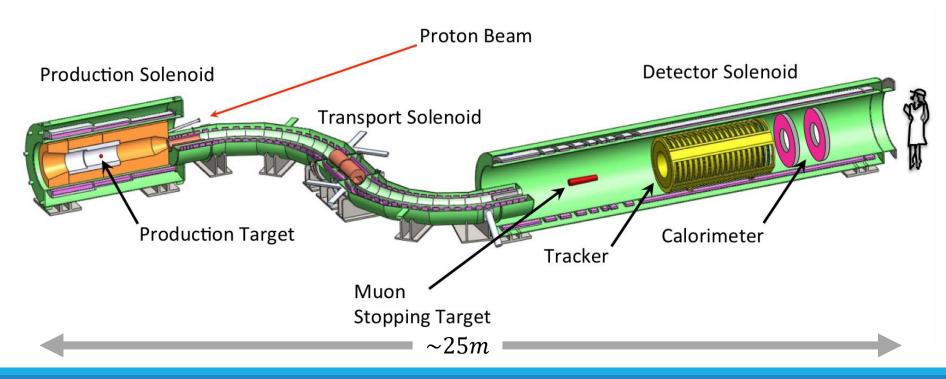
Delivery Ring takes 1 out of every 4 bunches from the Recycler.

Mu2e slow-extracts **Protons** every 1695ns.

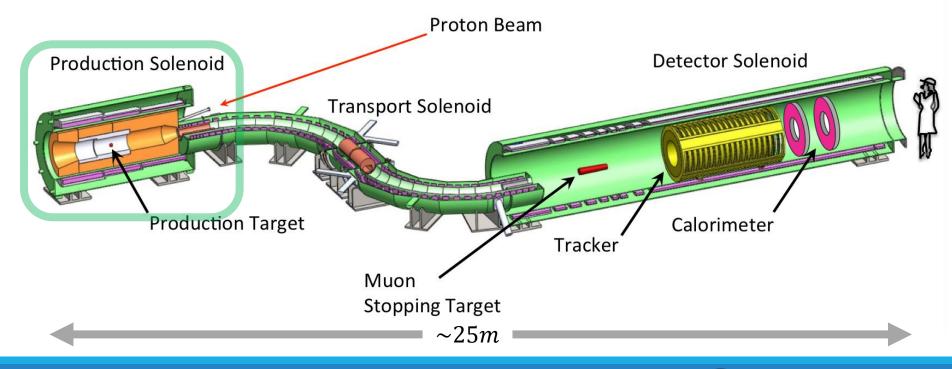
Runs with minimal impact on NOvA.



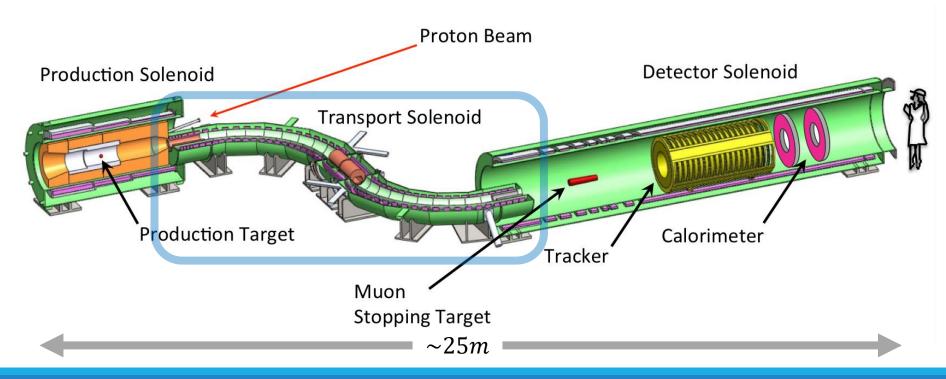
- Production Solenoid
- Transport Solenoid
- Detector Solenoid



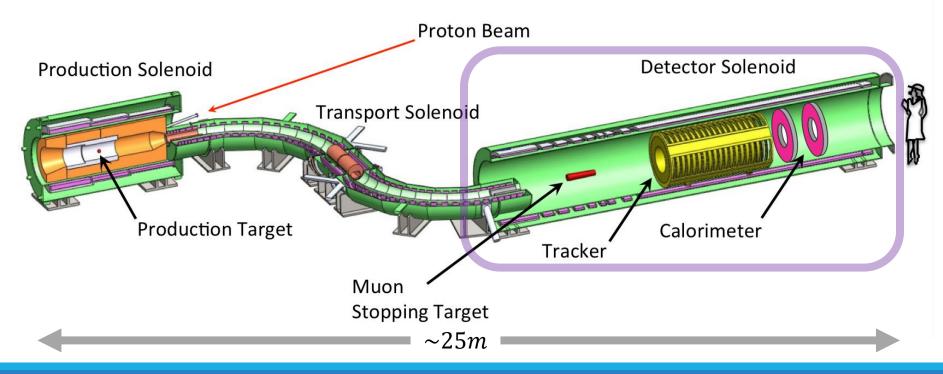
- Production Solenoid
- Transport Solenoid
- Detector Solenoid



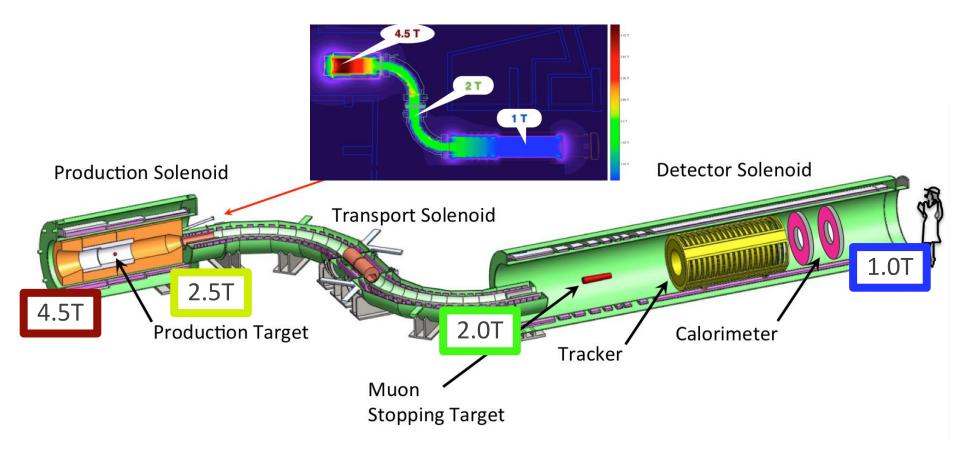
- Production Solenoid
- Transport Solenoid
- Detector Solenoid



- Production Solenoid
- Transport Solenoid
- Detector Solenoid



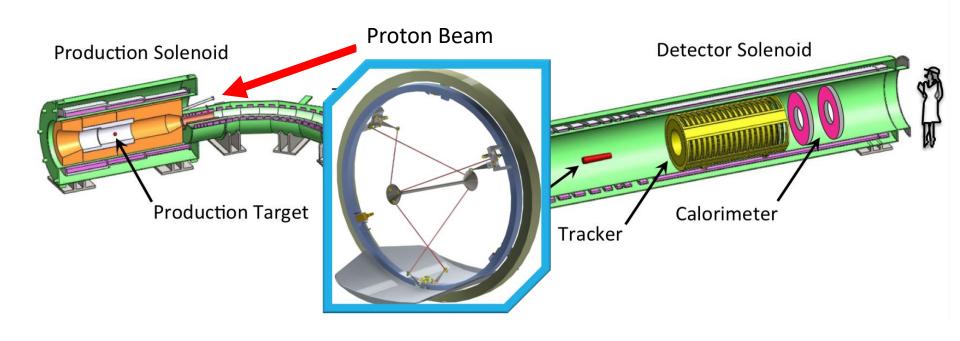
Graded magnetic fields sweep particles into the detector solenoid, creating the highest intensity muon beam to date.



Mu2e Apparatus Production Solenoid

A proton beam incident on a tungsten production target produces π^- .

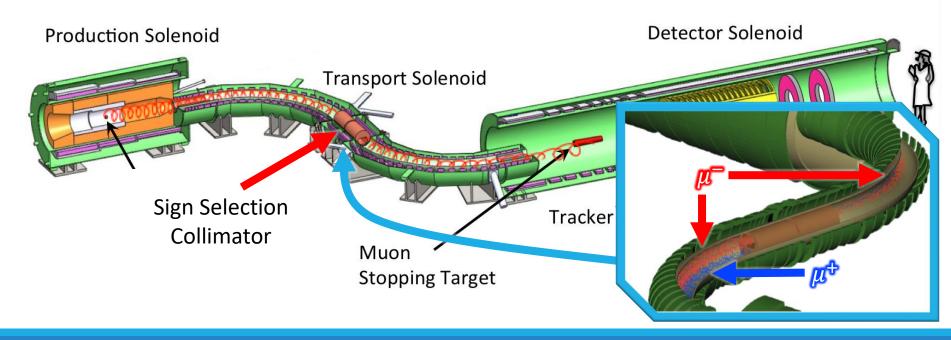
- \circ 8 GeV beam with 1695ns between pulses, well timed to $au^{Al}=$ 864ns
- 6×10^{12} protons/second $\rightarrow 10^{10}$ stopped μ^- /second
- π^- move toward the transport solenoid, decaying into μ^- .



Mu2e Apparatus Transport Solenoid

 π^- and μ^- beam propagate through the transport solenoid (TS) to the detector solenoid.

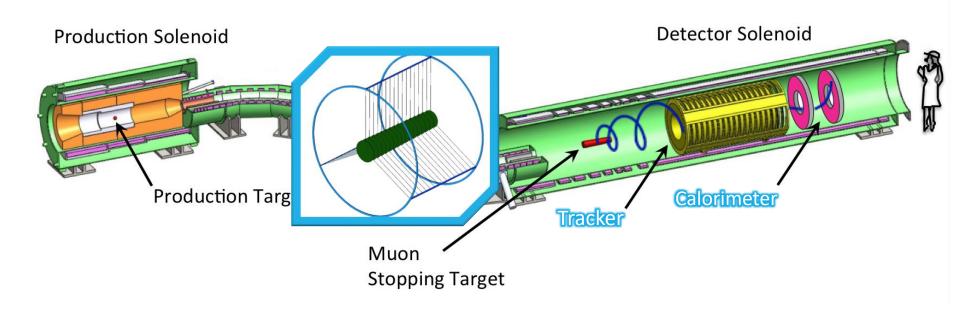
- Momentum selection
- Sign selecting collimator in the middle of the TS.



Mu2e Apparatus Detector Solenoid

 μ^- stop in an Aluminum target comprised of 37 thin 100 μ m foils.

- With a rate of 6 \times 10¹² protons/second incident on the production target, $1.1 \times 10^{10}~\mu^-$ /second reach the stopping target.
- The momentum and energy of e^- coming out of the stopping target is measured by the tracker and calorimeter.



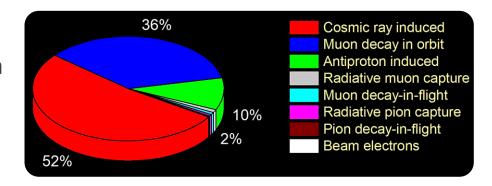
The single event sensitivity (SES) is expected to be 3×10^{-17} , which is around 40 $\mu \rightarrow e$ conversion events over the lifetime of the experiment (3 years) at $R_{\mu e} = 10^{-15}$.

For Mu2e to be a high-sensitivity experiment, the backgrounds which can mimic $\mu + N \rightarrow e + N$ conversion must be well-understood and kept at a sub-event level.

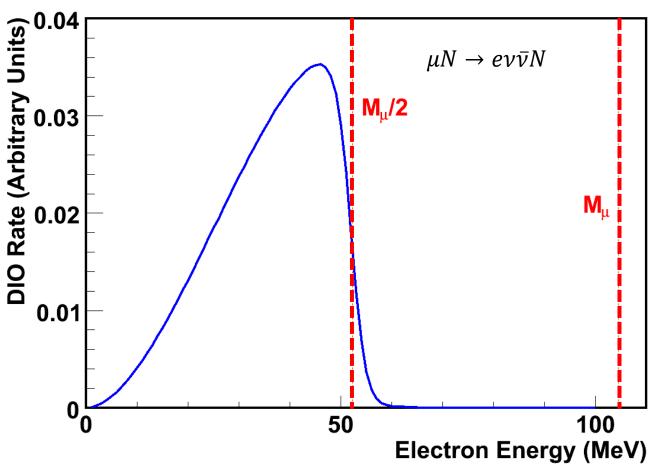
0.4 events over the lifetime of the experiment.

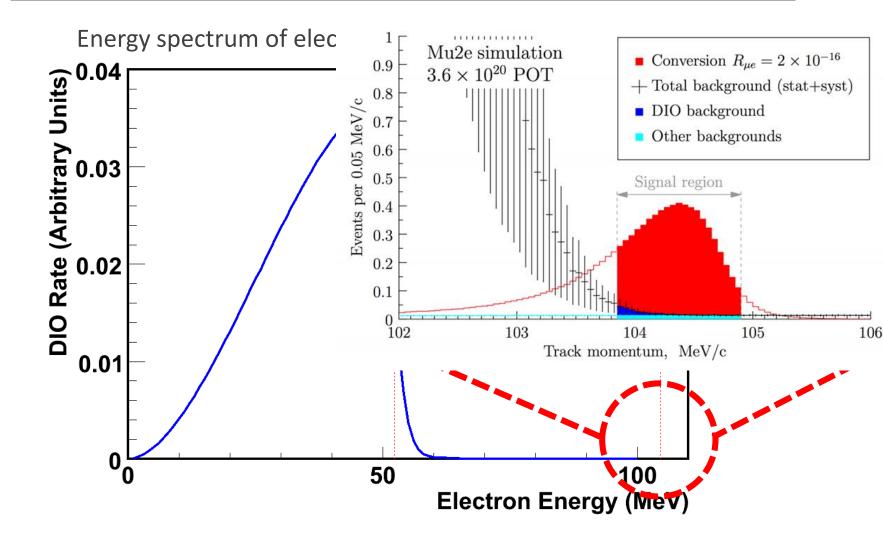
The three most prominent backgrounds for the experiment are the products of decay in orbit (DIO), anti-proton processes, with the largest background being the production of conversion-like electrons due to cosmic rays.

- Tracker and Calorimeter handles
- Cosmic Ray Veto



Energy spectrum of electron emitted by muon decay in orbit





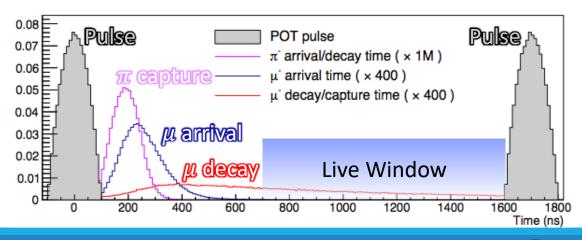
Process	Expected Number
Cosmic Ray Muons	0.25 ± 0.026
DIO	0.14 ± 0.09
RPC	0.025 ± 0.003
Antiprotons	0.047 ± 0.024
Muon DIF	< 0.003
Pion DIF	$0.001 \pm < 0.001$
Beam Electrons	$< 5 \times 10^{-4}$
Total	0.46 ± 0.10

The live window is delayed by 700ns relative to the proton pulse.

- $^{\circ}$ π reaching and stopping in the stopping target undergo radiative pion capture (RPC). Since the live window is delayed, emission of a conversion-like electron caused by RPC is mitigated.
- Beam flash is prompt, but can blind detector components.

Protons arriving out of time with respect to the pulses must be kept to a minimum.

- Can generate additional π , μ which can fake $\mu + N \rightarrow e + N$
- Require 10^{-10} out-of-pulse/in-pulse protons
- Measured and monitored throughout the experiment.

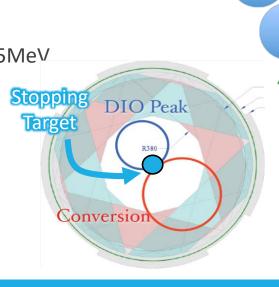


High-precision measurements of e^- momentum is made using a low-mass, straw drift tube, annular tracker, transverse to the beam axis.

Due to being un-instrumented close to the beam axis, sthe tracker has the distinct advantage of being blind to beam flash and next to all of the DIO spectrum.

The tracker is built out of panels, 6 panels per plane, 2 planes per station, with 18 stations total in the tracker detector.

- Total 216 panels, ~21,000 straws
- Momentum Resolution <200KeV/c @ 105MeV
- 30º rotation for stereo reconstruction
- 5 mm diameter straws
- \circ 12 μ m Mylar walls
- 25 μ m tungsten wires
- Filled with Ar/CO₂

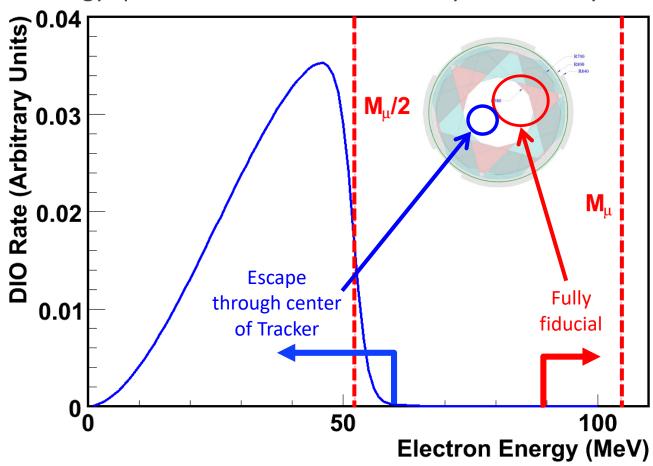




3270.0 mm

28

Energy spectrum of electron emitted by muon decay in orbit



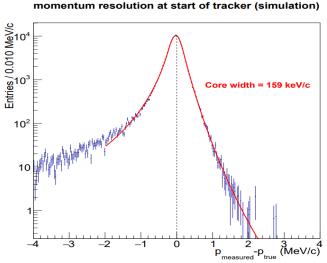
Several tracker panels have been built and tested using cosmic rays and a Fe⁵⁵ source.

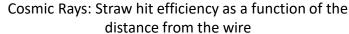
Full plane assembled at Fermilab.

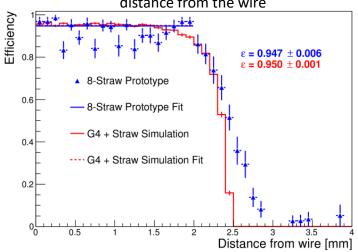
Panel assembly at the University of Minnesota (with Fermilab and U. Houston).

QA of panel components @ CUNY, Duke and LBL/UC Berkeley.

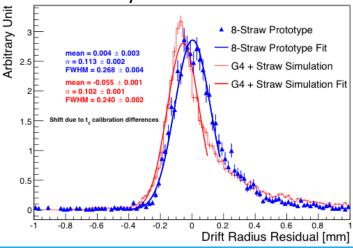








Cosmic Rays: Transverse Resolution





A high-granularity crystal-based calorimeter will measure the energy of e^- coming out of the stopping target.

racker

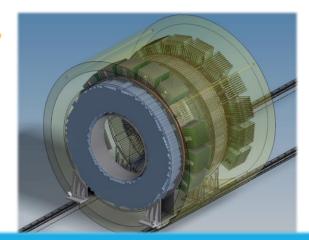
 This measurement is complementary to the Tracker's momentum measurement.

It provides:

- Precise timing $\sigma_t \sim 100 \text{ps}$
- Particle identification
- Seeds for the Tracker
 - Protection against pattern recognition errors in Tracker
- μ rejection x200 with 96% e^- efficiency

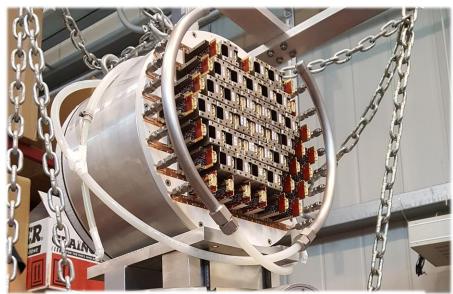
Comprised of 2 annuli of CsI crystal scintillators.

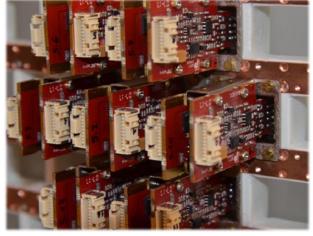
- Radiation Hard
- $^{\circ}$ Good energy resolution $^{\sigma_E}/_E \approx 5\%$ @ 105MeV

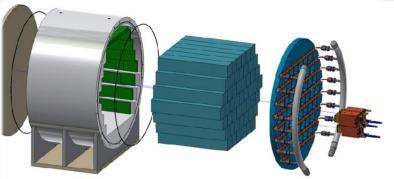


1400 CsI Crystals 3×3×20 cm³ 2800 SiPMs 2 SiPMs per crystal Disk Radii 37-66cm

A fully-instrumented prototype calorimeter was built and tested.

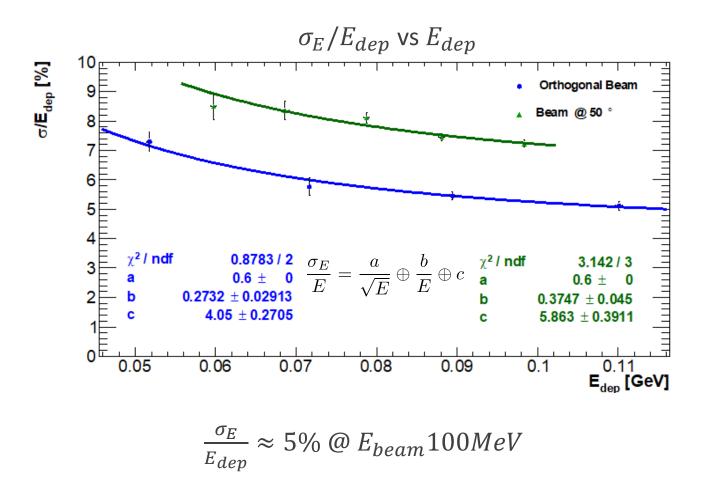


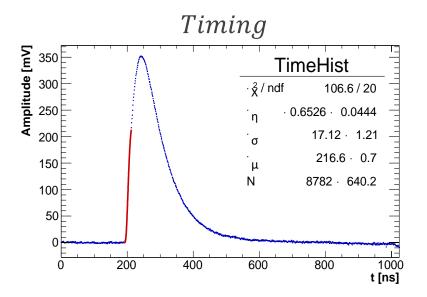








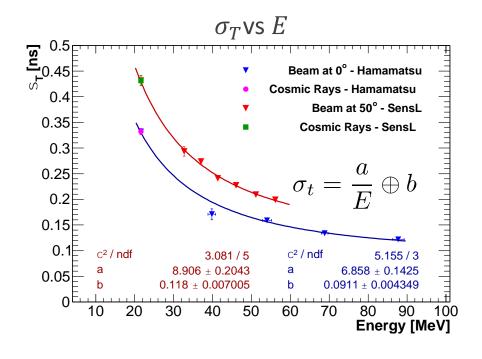




$$\frac{\sigma_T(T_1-T_2)}{\sqrt{2}}{\sim}132ps \ @ \ E_{beam}=100 MeV$$

Log Normal fit on leading edge.

Constant Fraction method used CF = 5%.

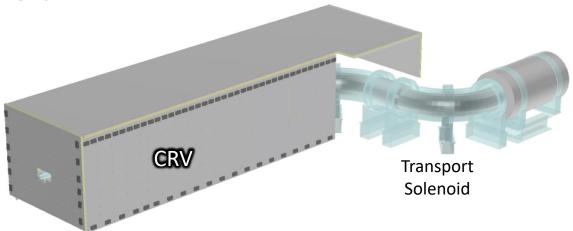




Cosmic rays infiltrating the experiment have the potential to create electrons and positrons through in-flight decays, as well as secondary interactions and delta-ray production in materials within the experiment's detector.

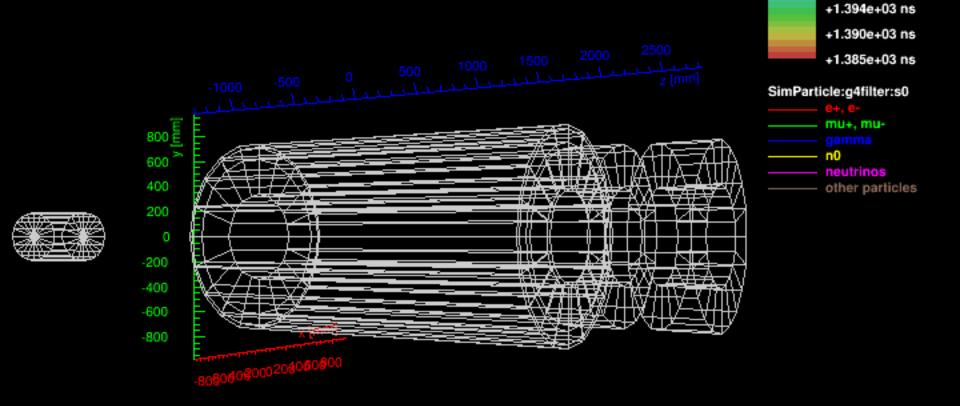
It is expected that one conversion-like event will occur per day, which can be suppressed by implementation of the CRV.

The CRV must have an overall efficiency of 99.99%, which limits the number of background events to 0.25 over the run time of the experiment.



Production Solenoid





Run #: 16467478 Sub Run #: 240

Conversion-like e^- from Cosmic Rays

Cosmic rays entering the detector have the ability to create conversion-like electrons, faking muon-conversion signal.

The CRV will be active-shielding against the entrance of cosmic rays.

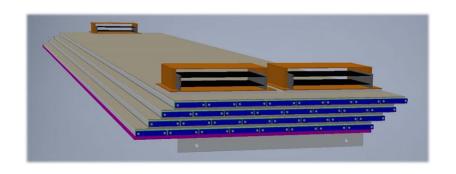


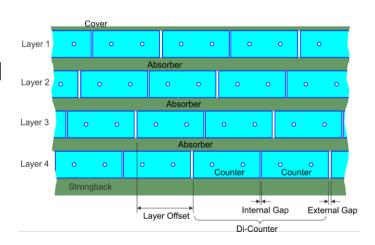
+1.399e+03 ns

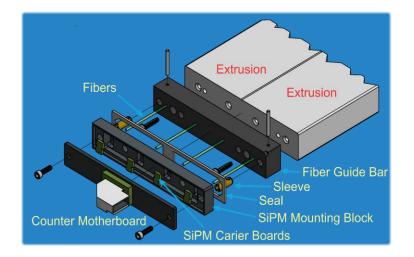


Four layers of extruded scintillator, read out by wavelength shifting fibers (WLS) and SiPMs, cover the entire detector solenoid and part of the transport solenoid.

- Total coverage area of 391 m²
- Total mass of \sim 68,500 kg \cong 75.5 tons
- 5,504 scintillator extrusions
- 52.7 km of WLS fibers
- 19,808 SiPMs
- 310 Front End Boards



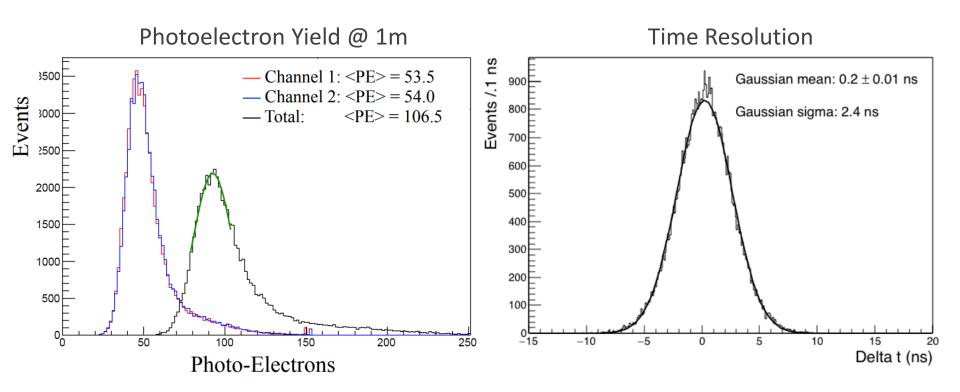






Three prototype, 4.5m-long detector modules and detector components have been built and tested.







Currently being fabricated at the University of Virginia.

There are 5 CRV posters being presented in the poster session.

• Check them out!

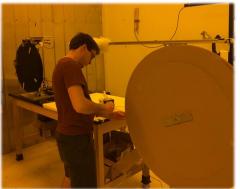












Status of Mu2e

Successfully passed DOE reviews.

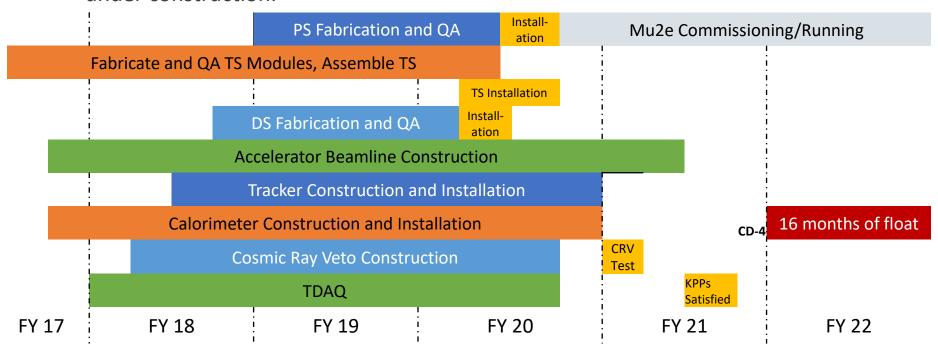
Solenoid installation will begin in 2020.

The detector hall construction has finished.

Experiment commissioning will begin in 2020.

The various detectors are currently under construction.

On track to begin taking data in 2022.



Summary

Mu2e has excellent discovery potential and the ability to reveal BSM physics.

- Improvement in sensitivity by 4 orders of magnitude
- Probing of higher mass scales $\sigma(10^4 \ TeV)$
- Sensitive to a wide range of BSM physics models
- Complementary to searches conducted by other experiments

The experiment is under construction and on schedule to begin commissioning in 2020.

Technical Design Report:

http://arXiv.org/abs/1501.05241

Web site:

http://mu2e.fnal.gov

