FUTURE DUNE CONSTRAINTS ON EFT

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based on: A. Falkowski, GGdC, Z. Tabrizi, JHEP04(2018)101, arXiv:1802.08296

NuFACT 2018, Virginia Tech





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Consistent with the SM



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Consistent with the SM

New Physics?

Neutrino masses, dark matter, inflation, baryon asymmetry

Strong CP problem, flavour hierarchy, gauge coupling unification, naturalness



Main goals: measure the parameters governing neutrino oscillations (CP violating phase in the PMNS matrix and neutrino mass ordering), searches for proton decay and for neutrino from core-collapse supernovas



1) Constrain systematic uncertainties in neutrino flux and neutrino scattering cross section; 2) <u>neutrino - nucleus/electron scattering measurement and BSM</u> searches;

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$$W^{\mu+}\bar{\nu}_{a}\bar{\sigma}_{\mu}(1+\delta g_{L}^{We_{a}})e_{a} + \text{h.c.}]$$

$$\overline{+g_{Y}^{2}}Z^{\mu}e_{a}^{c}\sigma_{\mu}\left(-s_{\theta}^{2}Q_{f}+\delta g_{R}^{Ze_{a}}\right)\bar{e}_{a}^{c}$$

$$\overline{+g_{Y}^{2}}Z^{\mu}\sum_{f=e,\nu}\bar{f}_{a}\bar{\sigma}_{\mu}\left(T_{3}^{f}-s_{\theta}^{2}Q_{f}+\delta g_{L}^{Zf_{a}}\right)f_{a}$$

 $\mathcal{L}_{\text{wEFT}} \supset -\frac{2}{v^2} (\bar{\nu}_a \bar{\sigma}_\mu \nu_b) \left[g_{LL}^{abcd} (\bar{e}_c \bar{\sigma}_\mu e_d) + g_{LR}^{abcd} (e_c^c \sigma_\mu \bar{e}_d^c) \right]$

Formalism

Neutrino interactions with charged leptons



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 $\delta g_L^{Z\nu_a} = \delta g_L^{We_a} + \delta g_L^{Ze_a}$

Shifted Z couplings

 $\frac{1}{2} + s_{\theta}^{2} + \frac{1}{2} [c_{\ell\ell}]_{1111} + \frac{1}{2} [c_{\ell\ell}]_{1221} - \delta g_{L}^{W\mu} + 2s_{\theta}^{2} \left(\delta g_{L}^{We} + \delta g_{L}^{Ze}\right)$

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$$\delta g_L^{Z\nu_a} = \delta g_L^{We_a} + \delta g_L^{Ze_a}$$

Shifted Z couplings

$$1221 - \delta g_L^{W\mu} + \left(2s_\theta^2 \left(\delta g_L^{We} + \delta g_L^{Ze}\right)\right)$$

$$\frac{2}{9}\left(\delta g_L^{We} + \delta g_L^{Ze}\right)$$

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$$\delta g_L^{Z\nu_a} = \delta g_L^{We_a} + \delta g_L^{Ze_a}$$

Shifted Z couplings

$$\delta g_{LX}^{ii11} = \epsilon_{ii}^{eX}$$

Formalism

Neutrino interactions with quarks

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$$\overline{\left\{-g_Y^2 Z^{\mu} \sum_{q=u,d} \left[\bar{q}\bar{\sigma}_{\mu} \left(\left(T_3^q - s_{\theta}^2 Q_q\right) + \delta g_L^{Zq}\right)\right) q\right.\right\}}$$
$$\left\{-s_{\theta}^2 Q_q + \left(\delta g_R^{Zq}\right) \bar{q}^c\right]$$
$$\left\{\bar{u}\bar{\sigma}_{\mu} \left(V_{ud} + \left(\delta g_L^{Wq_1}\right) d + \text{h.c.}\right]\right\}$$

$$\frac{d}{d}(1+\bar{\epsilon}_L^{de_a})(\bar{e}_a\bar{\sigma}_\mu\nu_a)(\bar{u}\,\bar{\sigma}^\mu d)$$

Formalism

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Translation to NSI parameters

 $\tilde{V}_{ud}\bar{\epsilon}_L^{de_a} = \epsilon_{aa}^{udL}$ $\delta g_{LX}^{\nu_a q} = \epsilon_{aa}^{qX}$

 $- \frac{2}{v^2} (\bar{\nu}_a \bar{\sigma}_\mu \nu_a) \sum \left[g_{LL}^{\nu_a q} \bar{q} \bar{\sigma}^\mu q + g_{LR}^{\nu_a q} (q^c \sigma^\mu \bar{q}^c) \right]$ a=u.d

3 years neutrino mode + 3 years antineutrino mode

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 $N_{\text{events}} = \text{time} \times N_{\text{T}} \times \epsilon \times \int_{E_{\perp}}^{E_{f}} dE_{\nu} \frac{d\phi(E_{\nu})}{dE_{\nu}} \sigma(E_{\nu})$

3 years neutrino mode + 3 years antineutrino mode

 N_T for $I.I \times I0^{21}$ proton on target (POT) in (anti-)neutrino mode with a 120 GeV proton beam with 1.2 MW of power; Near Detector of 100 tonnes Argon mass

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3 years neutrino mode + 3 years antineutrino mode Efficiency: 85% for V_{μ} $(\overline{\nu_{\mu}})$, 80% for ν_{e} ($\overline{\nu_{e}}$)

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Cross section of the process

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Neutrino trident production

(see also Matheus' talk)

ν beam		$\overline{\nu}$ beam		
$\nu_{\mu} \rightarrow \nu_{\mu} \mu^{-} \mu^{+}$	357	$\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{\mu} \mu^{-} \mu^{+}$	305	
$\nu_e \rightarrow \nu_e \mu^- \mu^+$	1.27	$\overline{\nu}_e \rightarrow \overline{\nu}_e \mu^- \mu^+$	1.03	

$$\frac{\sigma(\nu_b \gamma^* \to \nu_a \ell_c^- \ell_d^+)}{\sigma_{\rm SM}(\nu_b \gamma^* \to \nu_a \ell_c^- \ell_d^+)} = \frac{\sigma(\bar{\nu}_a \gamma^* \to \nu_b \ell_c^- \ell_d^+)}{\sigma_{\rm SM}(\bar{\nu}_a \gamma^* \to \nu_b \ell_c^- \ell_d^+)}$$

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[Phys.Lett. B245, 271(1990)] $\sigma_{\mathrm{CHARMII}}/\sigma_{\mathrm{SM}} = 1.58 \pm 0.57$

[Phys.Rev.Lett. 66, 3117(1991)] $\sigma_{\rm CCFR}/\sigma_{\rm SM}=0.82\pm0.28$

 $\frac{1}{2} \approx 1 + 2 \frac{g_{LL,SM}^{abcd} \delta g_{LL}^{abcd} + g_{LR,SM}^{abcd} \delta g_{LR}^{abcd}}{(g_{LL,SM}^{abcd})^2 + (g_{LR,SM}^{abcd})^2}$

Neutrino trident production

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$\nu_e \rightarrow \nu_e \mu^- \mu^+$	1.27	$\overline{\nu}_e \rightarrow \overline{\nu}_e \mu^- \mu^+$	1.03

 $R_{\mu} \equiv \frac{\sigma(\nu_{\mu} \to \nu_{\mu} \mu^{-} \mu^{+}) + \sigma(\bar{\nu}_{\mu}}{\sigma(\nu_{\mu} \to \nu_{\mu} \mu^{-} \mu^{+})_{\rm SM} + \sigma(\bar{\nu}_{\mu})}$ $-0.039 < 2 \frac{g_{LL,SM}^{2222} \delta g_{LL}^{2222} + g}{(g_{LL,SM}^{2222})^2 + (g_{LL,SM}^{2222})^2 + (g_{LL,SM}^{22222})^2}$

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$$\frac{\to \bar{\nu}_{\mu} \mu^{-} \mu^{+})}{\to \bar{\nu}_{\mu} \mu^{-} \mu^{+})_{\rm SM}} = 1 \pm 0.039$$

$$\frac{g_{LR,\rm SM}^{2222} \delta g_{LR}^{2222}}{(g_{LR,\rm SM}^{2222})^2} < 0.039$$

$$\begin{split} \sigma_{\nu_{\mu}e} &= \frac{s}{2\pi v^4} \left[(g_{LL}^{2211})^2 + \frac{1}{3} (g_{LR}^{2211})^2 \right] \approx \frac{m_e E_{\nu}}{\pi v^4} \left[(g_{LL}^{2211})^2 + \frac{1}{3} (g_{LR}^{2211})^2 \right] \\ \sigma_{\bar{\nu}_{\mu}e} &= \frac{s}{2\pi v^4} \left[(g_{LR}^{2211})^2 + \frac{1}{3} (g_{LL}^{2211})^2 \right] \approx \frac{m_e E_{\nu}}{\pi v^4} \left[(g_{LR}^{2211})^2 + \frac{1}{3} (g_{LL}^{2211})^2 \right] \end{split}$$

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	$N_{\rm tot}^{\nu-e}$	$r^{\nu-e}_{\nu_{\mu}}$	$r^{\nu-e}_{\overline{\nu}_{\mu}}$	$r_{\nu_e}^{\nu-e}$	$r_{\overline{\nu}_e}^{\nu-e}$
$\nu ext{-mode}$	1.69×10^6	0.898	0.059	0.040	0.003
$\overline{\nu}$ -mode	1.29×10^6	0.103	0.867	0.013	0.017

$$R_{\nu e}^{i} \equiv \frac{x_{i}\sigma_{\nu_{\mu}e} + \bar{x}_{i}\sigma_{\bar{\nu}_{\mu}e}}{x_{i}\sigma_{\nu_{\mu}e}^{\mathrm{SM}} + \bar{x}_{i}\sigma_{\bar{\nu}_{\mu}e}^{\mathrm{SM}}}$$

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	$N_{\rm tot}^{\nu-e}$	$r^{\nu-e}_{\nu_{\mu}}$	$r^{\nu-e}_{\overline{\nu}_{\mu}}$	$r_{\nu_e}^{\nu-e}$	$r_{\overline{\nu}_e}^{\nu-e}$
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$$\begin{cases} x_{\nu} &= 0.9 \\ x_{\overline{\nu}} &= 0.1 \\ \bar{x}_i &= 1 - x_i \end{cases}$$

$$\delta R_{\nu e}^{i} = 2 \frac{(1+2x_{i}) \,\delta g_{LL}^{2211} g_{LL,\text{SM}}^{2211} + (3-2x_{i}) \,\delta g_{LR}^{2211} g_{LR,\text{SM}}^{2211}}{(1+2x_{i}) \,(g_{LL,\text{SM}}^{2211})^{2} + (3-2x_{i}) \,(g_{LR,\text{SM}}^{2211})^{2}}$$

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$-8.0 \times 10^{-4} < \delta R_{\nu e}^{\nu} < 8.0 \times 10^{-4} \qquad -9.1 \times 10^{-4} < \delta R_{\nu e}^{\bar{\nu}} < 9.1 \times 10^{-4}$

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[Falkowski, GGdC, Tabrizi, JHEP04(2018)101]

$$R_{\nu_a N} \equiv \frac{x \sigma_{\nu_a N \to \nu_a N} + \bar{x} \sigma_{\bar{\nu}_a N \to \bar{\nu}_a N}}{x \sigma_{\nu_a N \to e_a^- N} + \bar{x} \sigma_{\bar{\nu}_a N \to e_a^+ N}} = (g_L^{\nu_a})^2 + r^{-1} (g_R^{\nu_a})^2$$

$$r = \frac{x \sigma_{\nu_a N \to e_a^- N} + \bar{x} \sigma_{\bar{\nu}_a N \to e_a^+ N}}{x \sigma_{\nu_a N \to e_a^- N} + \bar{x} \sigma_{\bar{\nu}_a N \to e_a^+ N}} \qquad \begin{array}{c} \text{generalised} \\ \text{Llewellyn-Smith} \\ \text{formula} \end{array}$$

only dependence on the nuclear structure

$$r_{\nu} \sim 2.5$$

$$r_{\bar{\nu}} \sim 0.4$$

Cannot neglect systematics!

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	N_{tot}^{CC}	$r_{\nu_{\mu}}^{CC}$	$r_{\overline{\nu}_{\mu}}^{CC}$	$r_{\nu_e}^{CC}$	$r_{\overline{\nu}_e}^{CC}$
$\nu\text{-mode}$	$4.25 imes 10^8$	0.964	0.028	0.007	0.001
$\overline{\nu}$ -mode	1.74×10^8	0.201	0.790	0.004	0.005
	N_{tot}^{NC}	$r_{\nu_{\mu}}^{NC}$	$r_{\overline{\nu}_{\mu}}^{NC}$	$r_{\nu_e}^{NC}$	$r_{\overline{\nu}_e}^{NC}$
$\nu\text{-mode}$	1.48×10^8	0.956	0.037	0.006	0.001
	7				0 00 M

but

I) the ⁴⁰Ar target nuclei are not isoscalar and the LS formula has to be corrected 2) neglected admixture of electron neutrinos

$R^{i}_{\nu_{\mu}N} = R^{i}_{\nu_{\mu}N,\rm{SM}}(1 + \delta R^{i}_{\nu_{\mu}N})$

 $-9.6 \times 10^{-5} < \delta R^{\nu}_{\nu_{\mu}N} < 9.6 \times 10^{-5}$ $-1.4 \times 10^{-4} < \delta R^{\bar{\nu}}_{\nu_{\mu}N} < 1.4 \times 10^{-4}$

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 $\delta g_R^{\gamma\mu}$

$R^{i}_{\nu_{\mu}N} = R^{i}_{\nu_{\mu}N,SM} (1 + \delta R^{i}_{\nu_{\mu}N})$

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	[Falkowski et al. 1706.03783]	, [Falkowski,	[Falkowski, GGdC, Tabrizi, JHEP04(2018)101]				
Coefficient	$\Delta(\text{current})$	Δ (no sys.)	Δ (0.1% sys.)	Δ (1% sys.)			
δg_L^{We}	3.5	0.37	2.5	3.4			
$\delta g_L^{Z\mu}$	3.7	0.18	1.1	3.5			
δg_L^{Zu}	1.9	0.34	1.4	1.5			
δg_R^{Zu}	9.5	0.57	2.0	2.3			
δg_L^{Zd}	1.9	0.28	1.4	1.6			
δg_R^{Zd}	9.7	1.1	3.0	3.1			
$\delta g_R^{Wq_1}$	1.9	0.36	1.7	1.9			
$[c_{\ell\ell}]_{1122}$	28	2.6	2.6	21			
$[c_{\ell e}]_{2211}$	45	3.1	3.1	27			
$[c_{\ell\ell}]_{2222}$	2100	310	310	310			
$[c_{\ell e}]_{2222}$	6300	970	970	970			
$[c_{\ell q}^{(3)}]_{1111}$	1.9	0.36	1.7	1.9			
$[c_{\ell q}^{(3)}]_{2211}$	12	1.8	10	12			
$[c_{\ell q}]_{2211}$	210	3.0	30	180			
$[c_{\ell u}]_{2211}$	190	1.2	9.5	85			
$[c_{\ell d}]_{2211}$	370	2.4	19	170			

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Result

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[Falkowski, GGdC, Tabrizi, JHEP04(2018)101]

Conclusions

We investigated the precision reach in the determination of the wEFT and SMEFT Wilson coefficients relevant for the DUNE experiment.

We studied observables related to trident production, neutrino scattering off electrons and neutrino scattering off nuclei at the **DUNE** Near Detector.

The results show the importance of precision measurements in DUNE and the importance of the effort to reduce the experimental and theoretical source of systematic errors.

Systematic uncertainties

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systematic uncertainty

$$\delta R^{i}_{\nu_{\mu}N} \simeq 2 \frac{g^{\nu}_{L,\text{SM}} \delta g^{\nu_{\mu}}_{L}}{(g^{\nu}_{L,\text{SM}})^2}$$

$$g_{X,\mathrm{SM}}^{\nu} \delta g_X^{\nu_{\mu}} = \sum_{q=u,d} g_{LX,\mathrm{SM}}^{\nu_{\mu}q}$$

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 $\frac{+r_{i}^{-1}g_{R,\text{SM}}^{\nu}\delta g_{R}^{\nu_{\mu}}}{+r_{i}^{-1}(g_{R,\text{SM}}^{\nu})^{2}}$

 $\delta g_{LX}^{\nu_{\mu}q} - (g_{X,\mathrm{SM}}^{\nu})^2 \overline{\epsilon}_L^{\nu_{\mu}d}$