

# Standard and non-standard neutrino physics at reactor experiments

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NUFACT

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# Outline

## 1 Preliminaries

- Results

## 2 The pheno approach to the NSI

- Motivation

## 3 CC-Like NSI at $\sim 1\text{km}$ reactors

- What are the current limits?

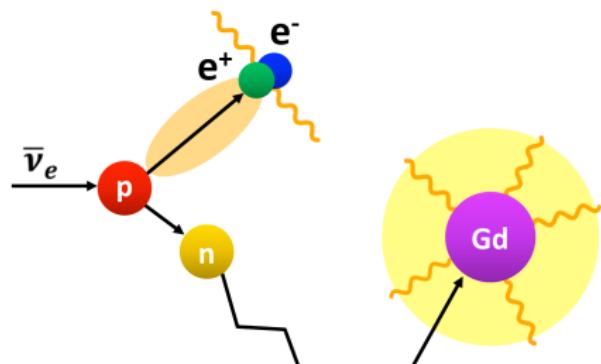
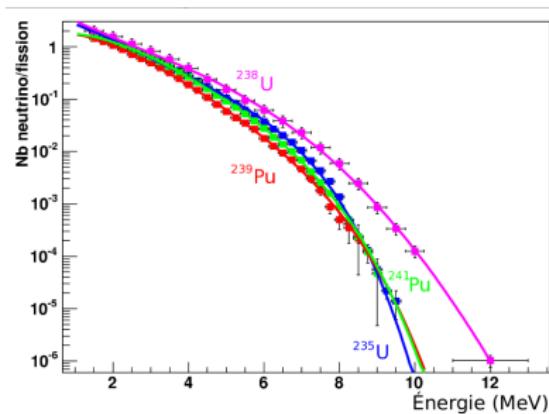
## 4 NSI ‘preliminary’ results

## 5 Summary and Conclusions

# $\bar{\nu}_e$ production and detection @ current reactor exps.

**Production:**  $\beta$  decay of  
 $k = {}^{235}\text{U}$ ,  ${}^{239}\text{Pu}$ ,  ${}^{241}\text{Pu}$  and  ${}^{238}\text{U}$

**Detection:** Inverse  $\beta$  decay,  
 $\bar{\nu}_e + p \rightarrow n + e^+$



Flux parametrizations:  $\Phi_k(E)$

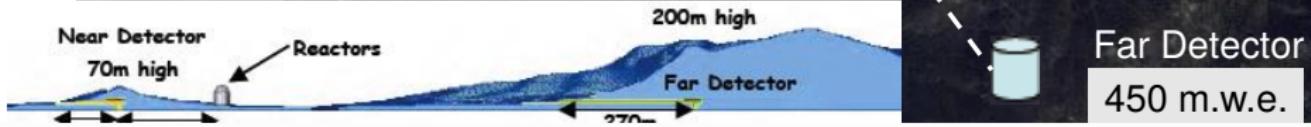
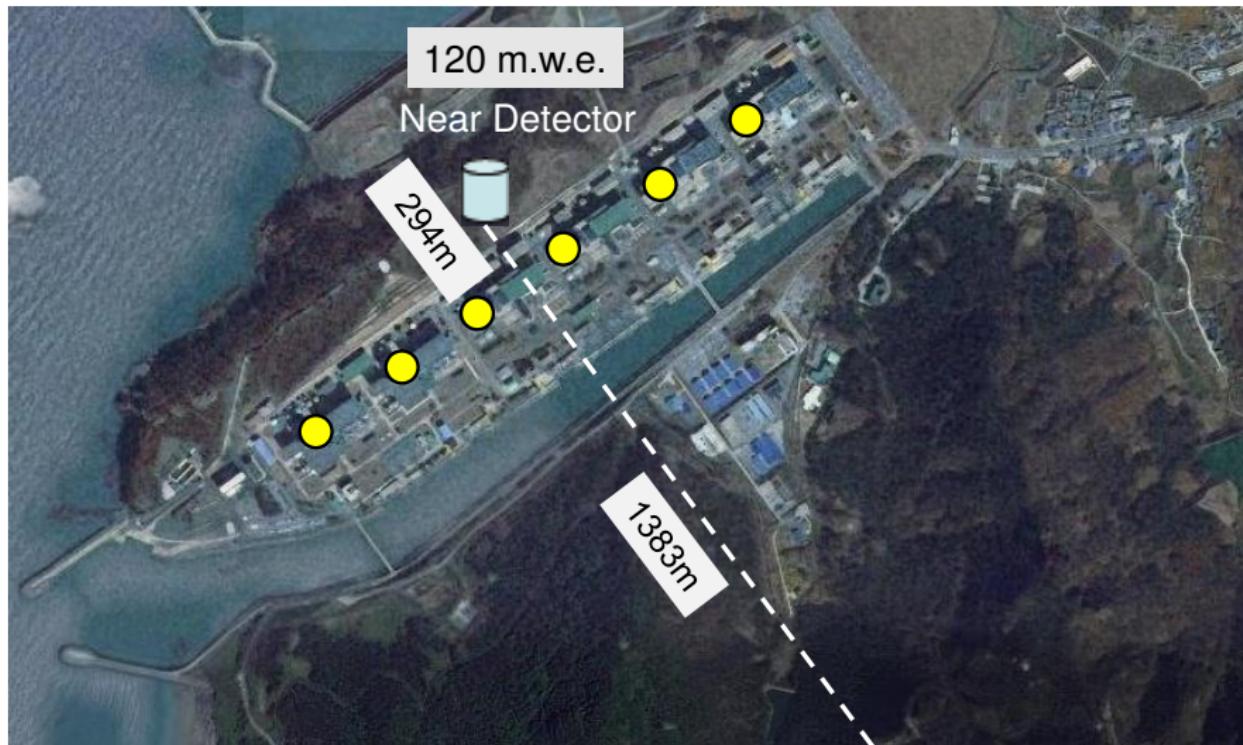
P. Huber (PRC 84 (2011))

T. Mueller *et al.* (PRC 83 (2011))

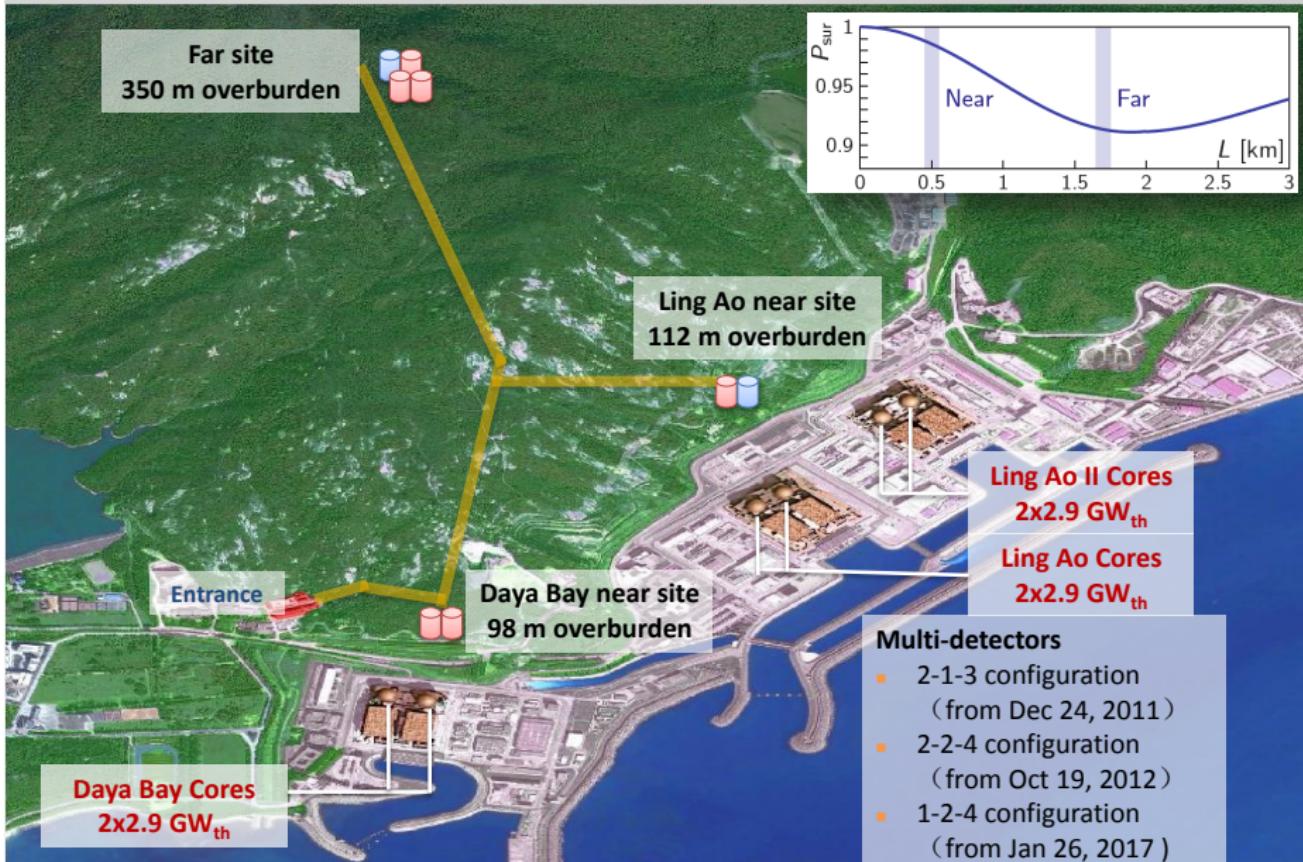
Coincidence signals:  
annihilation and delayed  $n$ -capture.  
Prompt  $e^+$ -  
( $\sim 30\mu\text{s}$ )

For  $\sim 1\text{km}$  baseline,  $\bar{\nu}_e$ s propagate to FD practically in Vacuum!

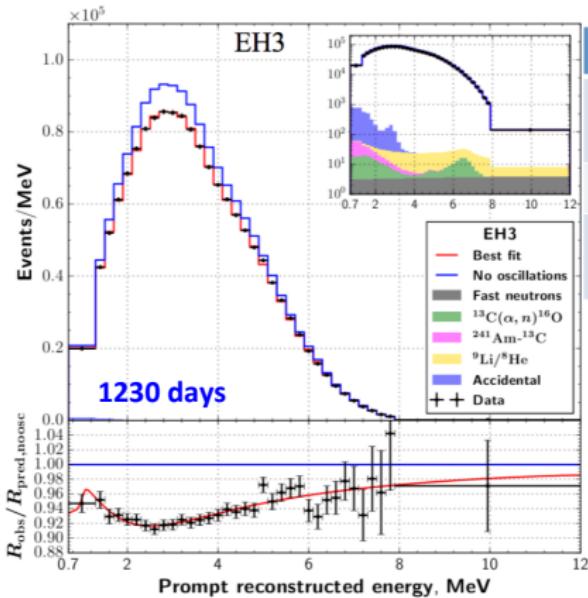
# RENO Experimental Set-up



# The Daya Bay Experiment



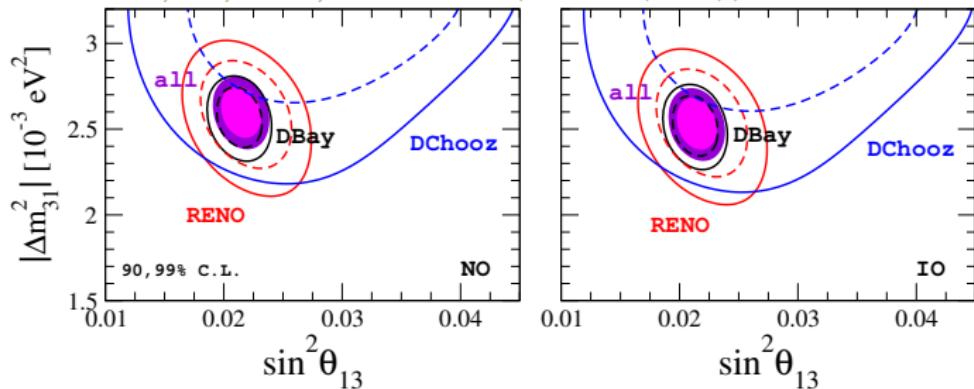
# Summary of IBD candidates



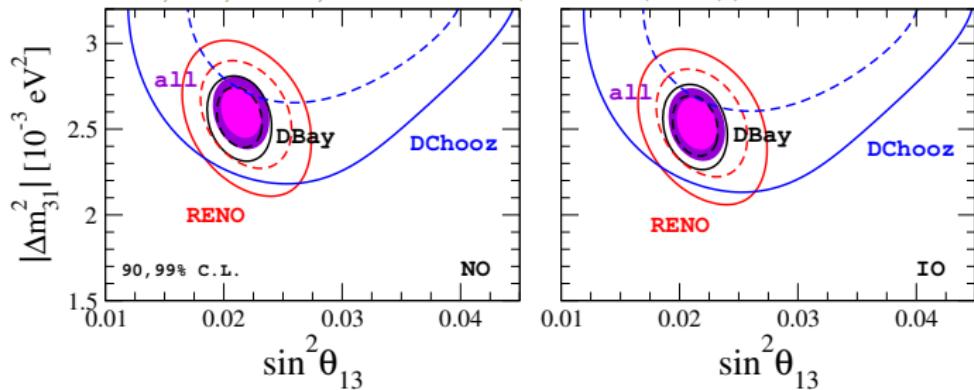
6-AD: 217 days (Dec/2011 – Jul/2012)  
8-AD: 1013 days (Oct/2012 – Jul/2015)

	EH1	EH2	EH3
IBD candidates	1,203,969	1,033,209	308,150
B/S ratio	$1.8 \pm 0.2\%$	$1.5 \pm 0.2\%$	$2.0 \pm 0.2\%$
IBD rate (day <sup>-1</sup> )	1058.5	998.2	285.2

- Over 2.5M (300K) IBD candidates in total (the far site)!
- $\leq 2\%$  backgrounds
- $^9\text{Li}/^8\text{He}$  has the largest uncertainty on B/S ratio:  $0.1\% \sim 0.15\%$

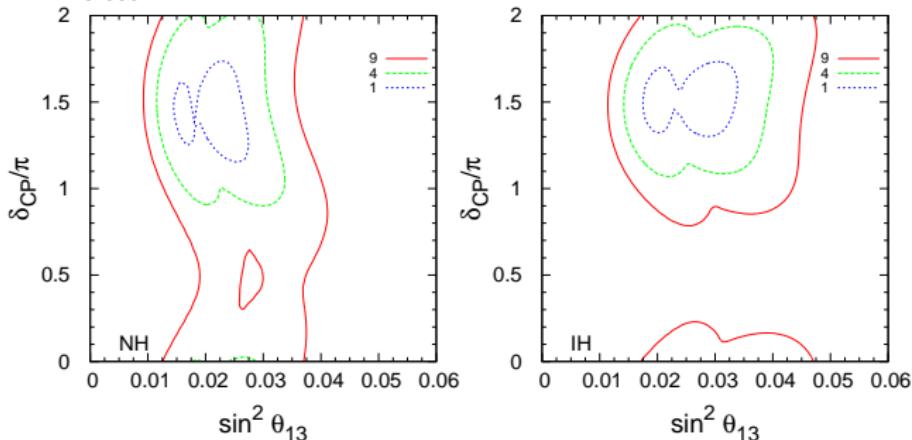


$$\sin^2 \theta_{13} = (2.160^{+0.083}_{-0.069}) \times 10^{-2} \text{ (global)}$$



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LBL-only



More info in <https://globalfit.astroparticles.es/>

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## Motivating the NSI

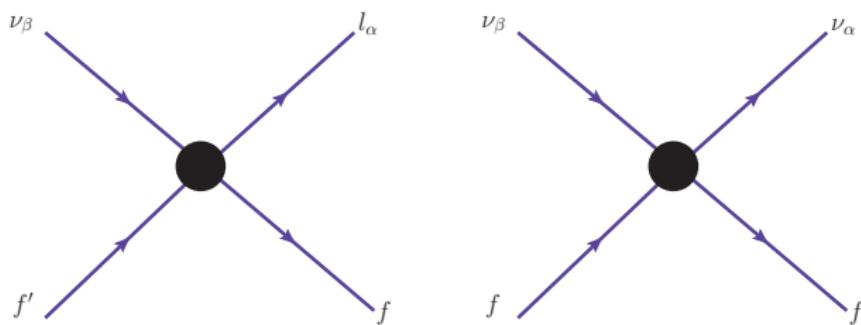
- Neutrino masses imply physics BSM.
- New neutrino interactions are expected in SM extensions
- The exchange of new heavy particles could leave a low energy ‘fingerprint’ in the form of NSI.
- NSI provide a model independent framework to include ‘NP’.
- It is worthy to quantify the ‘amount’ of NSI allowed by the neutrino oscillation data.

# The pheno approach to the NSI

Effective four-fermion operators

L. Wolfenstein (PRD **17**(1978)), J.W.F Valle (PLB **199**(1987))

M.M Guzzo *et al.* (PLB **260**(1991)), E. Roulet (PRD **44**(1991))



$$\begin{aligned}\mathcal{L}_{V\pm A} = & \frac{G_F}{\sqrt{2}} \sum_{f,f'} \tilde{\varepsilon}_{\alpha\beta}^{S(D), f, f', V\pm A} \left[ \bar{\nu}_\beta \gamma^\rho (1 - \gamma^5) \ell_\alpha \right] \left[ \bar{f}' \gamma_\rho (1 \pm \gamma^5) f \right] \\ & + \frac{G_F}{\sqrt{2}} \sum_f \tilde{\varepsilon}_{\alpha\beta}^{m, f, V\pm A} \left[ \bar{\nu}_\alpha \gamma^\rho (1 - \gamma^5) \nu_\beta \right] \left[ \bar{f} \gamma_\rho (1 \pm \gamma^5) f \right] + \text{h.c.},\end{aligned}$$

Also, see J. Kopp *et al.* (PRD **77**(2008)) for a discussion of other Lorentz structures for the NSI, not distinguishable with current reactor and accelerators.

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# Analysis framework

- $\tilde{\varepsilon}_{\alpha\beta}^{m,f,V\pm A} \rightarrow 0$  and  $\tilde{\varepsilon}_{e\beta}^{S(D),u,d,V\pm A} \rightarrow \varepsilon_{e\beta}^{S(D)}$
- $|\bar{\nu}_\alpha^s\rangle = |\bar{\nu}_\alpha\rangle + \sum_\gamma \varepsilon_{\alpha\gamma}^{s*} |\bar{\nu}_\gamma\rangle$  ( $I_\alpha$  produced).  $|\bar{\nu}_\beta^d\rangle = |\bar{\nu}_\beta\rangle + \sum_\gamma \varepsilon_{\gamma\beta}^{d*} |\bar{\nu}_\gamma\rangle$  ( $I_\beta$  detected)
- $\varepsilon^s$  &  $\varepsilon^d$  are general complex matrices!
- detection (inv.  $\beta$ -decay) Vs production ( $\beta$ -decay)  $\varepsilon_{e\alpha}^s = \varepsilon_{\alpha e}^{d*} \equiv \varepsilon_\alpha = |\varepsilon_\alpha| e^{i\phi_\alpha}$
- The **effective** oscillation probability is given by:

$$P_{\bar{\nu}_e^s \rightarrow \bar{\nu}_e^d}^{\text{eff.}} \simeq 1 + \overbrace{4|\varepsilon_e| \cos \phi_e}^{\text{'zero distance term'}} - 4 [\sin \theta_{13} + s_{23} |\varepsilon_\mu| \cos(\delta - \phi_\mu) + c_{23} |\varepsilon_\tau| \cos(\delta - \phi_\tau)]^2 \sin^2 \Delta_{31} + \mathcal{O}(\varepsilon)^2$$

Also, notice the ' $\delta$ ', ' $\theta_{23}$ ' dependence!

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Also, notice the ' $\delta$ ', ' $\theta_{23}$ ' dependence!

Analysis details: One parameter at a time,  $\theta_{23} = \pi/4$ , and NH.

- I.  $\varepsilon_e \neq 0$ , II.  $\varepsilon_{\mu,\tau} \neq 0$ , and Flavour Universal, III.  $\varepsilon_{e,\mu,\tau} \equiv \varepsilon$ .
- $S_d (1 + a_{\text{norm}} + \sum_r \omega_r^d \alpha_r + \xi_d) + (1 + \beta_d) B_d$  &  $\chi^2 = \sum_k (\eta_k / \sigma_k)^2$  for each  $\eta_k$ .

see Agarwalla, Bagchi, DVF, Tórtola (JHEP 1517 (2015)) for more analysis details

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Here, 'preliminary' results for **shape analysis** of 1230 days of Daya Bay data.

# Current bounds

CC-like NSI

Biggio, et al. (JHEP 090 (2009))

Bounds extracted from:

- $V^{ud}$  determination: From Kaon decays  $\rightarrow V^{us}$  (and assuming CKM unitarity) compared with the derivation from beta decays (affected by NSI).
- Universality tests: Ratios  $\pi \rightarrow e(\mu)\nu$  and  $\tau \rightarrow \pi\nu$  decay rates modified by quark CC-like NSI.
- Non-observation of flavor change at NOMAD ('zero distance effect').  
Channels  $\nu_\mu \rightarrow \nu_e$  ( $|\varepsilon_{\mu e}^{ud A}|$ ,  $|\varepsilon_{e\mu}^{ud L(R)}|$ ),  $\nu_e \rightarrow \nu_\tau$  ( $|\varepsilon_{e\tau}^{ud}|$ ), and  $\nu_\mu \rightarrow \nu_\tau$  ( $|\varepsilon_{\mu\tau}^{ud A}|$ ,  $|\varepsilon_{\tau\mu}^{ud L(R)}|$ ).

Assuming only one parameter at a time (90% C.L. for 1 d.o.f):

$$\mathcal{X} = \begin{bmatrix} V & L(R) & V \\ A & A & A \\ L(R) & L(R) & A \end{bmatrix}, |\varepsilon_{\alpha\beta}^{ud} \mathcal{X}_{ij}| < \begin{bmatrix} 0.041 & 0.026(0.037) & 0.041 \\ 0.026 & 0.078 & 0.013 \\ 0.087(0.12) & 0.013(0.018) & 0.13 \end{bmatrix}$$

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We improved the limit on  $|\varepsilon_{ee}^{ud}| < 0.02$  NSI coupling, total-rate analysis of Daya Bay

see Agarwalla, Bagchi, DVF, Tórtola (JHEP 1517 (2015)) for details

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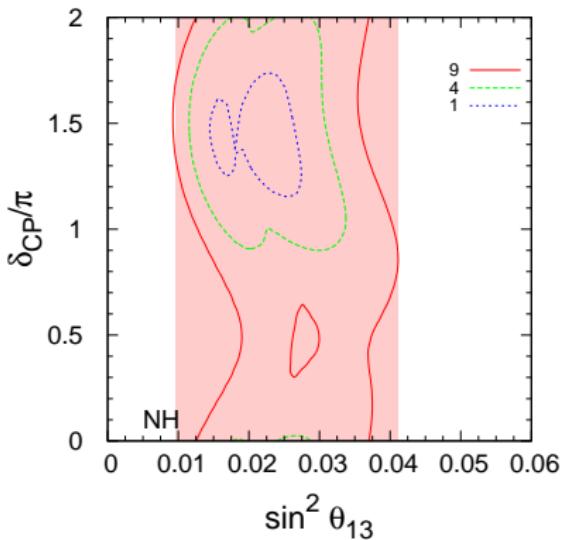
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# Results for the ' $\varepsilon_{ee}$ ' coupling

Showing  $s_{13}^2$  from LBL-only in vertical bands



$$0.0096 \leq \sin^2 \theta_{13} \leq 0.0411 \text{ (3}\sigma\text{)}$$

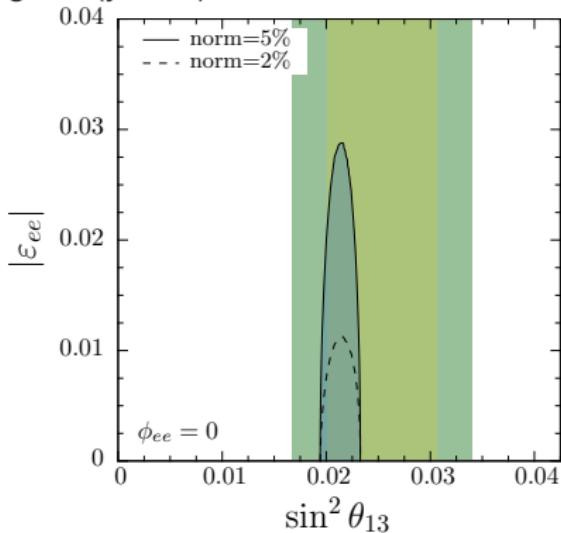
$$0.0201 \leq \sin^2 \theta_{13} \leq 0.0306 \text{ (1}\sigma\text{)}$$

# Results for the ' $\varepsilon_{ee}$ ' coupling

Showing  $s_{13}^2$  from LBL-only in vertical bands

Contours at 90% of C.L for 2 d.o.f.

Band: 90% ( $1\sigma$ ) of C.L for 1 d.o.f,  
green(yellow).

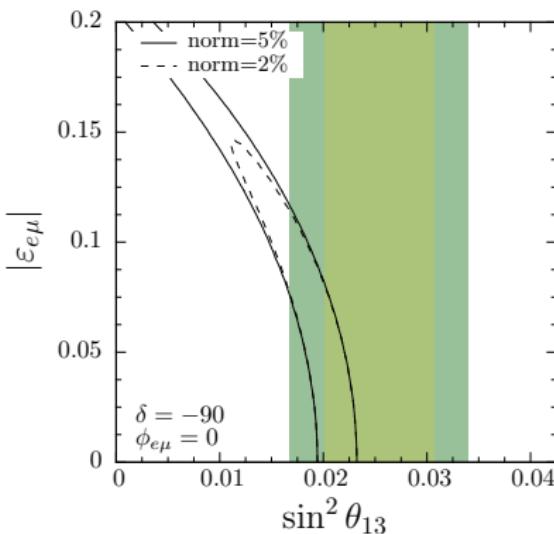
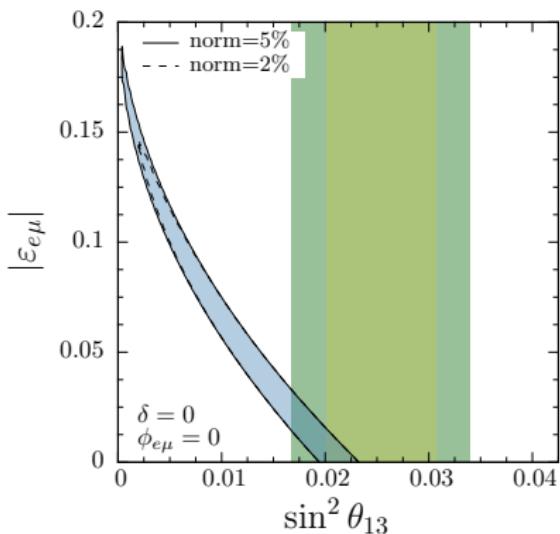


$$\tilde{s}_{13}^2 \approx s_{13}^2 - \frac{|\varepsilon_{ee}| \cos \phi_{ee}}{\sin^2 \Delta_{31}}$$

# Results for the ' $\varepsilon_{e(\mu,\tau)}$ ' case

Showing  $s_{13}^2$  from LBL-only in vertical bands

Contours at 90% of C.L for 2 d.o.f. Band: 90% ( $1\sigma$ ) of C.L for 1 d.o.f, green(yellow).



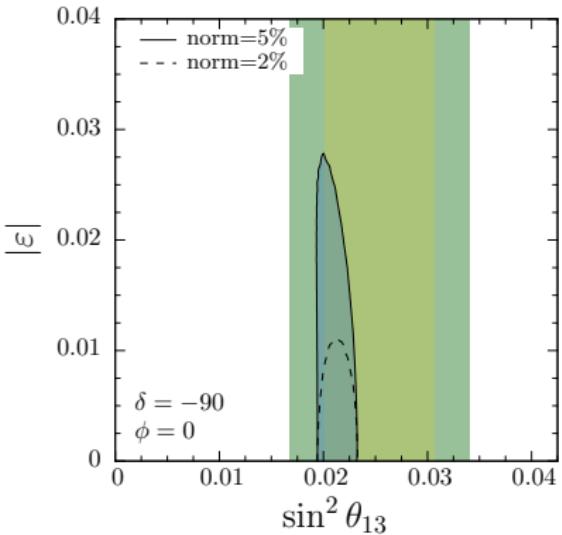
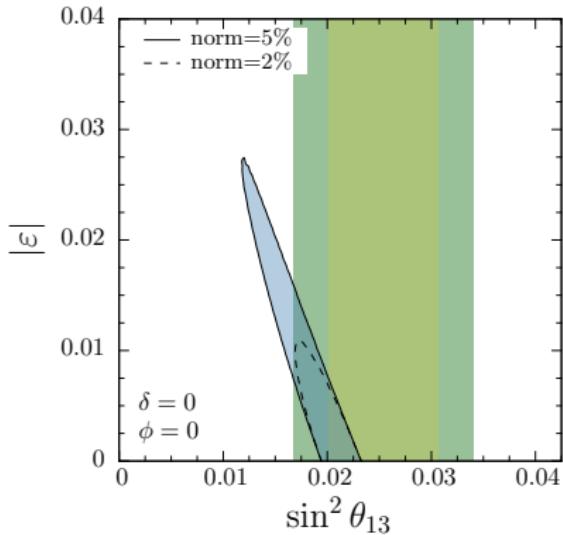
$$\tilde{s}_{13}^2 \approx s_{13}^2 + 2s_{13}s_{23}|\varepsilon_{e(\mu,\tau)}| \cos(\delta - \phi_{e(\mu,\tau)})$$

Cancellations when  $\cos(\delta - \phi_{e(\mu,\tau)}) = 0$ , higher order  $\varepsilon$  terms important!

# Results for the ‘FU’ case

Showing  $s_{13}^2$  from LBL-only in vertical bands

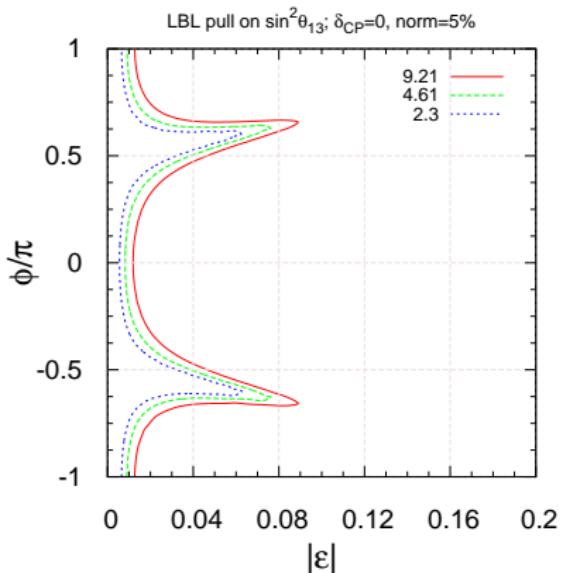
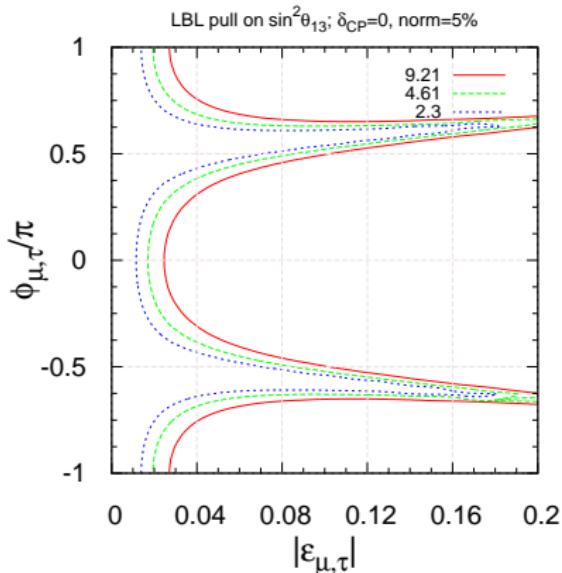
Contours at 90% of C.L for 2 d.o.f. Band: 90% ( $1\sigma$ ) of C.L for 1 d.o.f, green(yellow).



$$\tilde{s}_{13}^2 \approx s_{13}^2 - |\varepsilon| \left[ \frac{\cos \phi}{\sin^2 \Delta_{31}} - 4 s_{13} s_{23} \cos(\delta - \phi) \right]$$

# Including the LBL measurement of $\sin^2 \theta_{13}$

$$0.0201 \leq \sin^2 \theta_{13} \leq 0.0306 \text{ (1}\sigma\text{)}$$



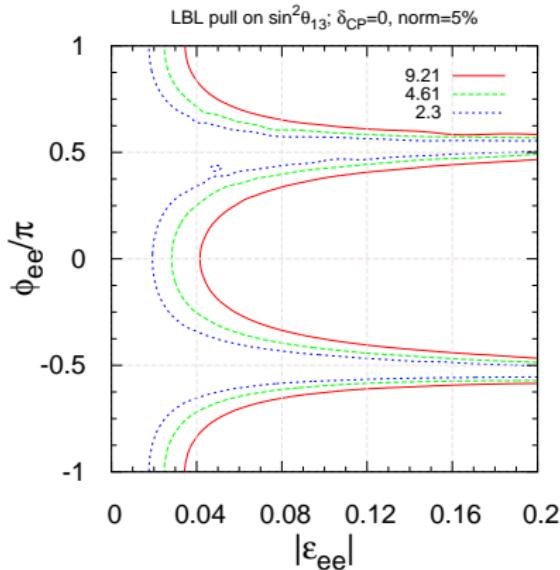
Cancellations when  $\cos(\delta - \phi_{\mu,\tau}) = 0$ , higher order  $\varepsilon$  terms important!

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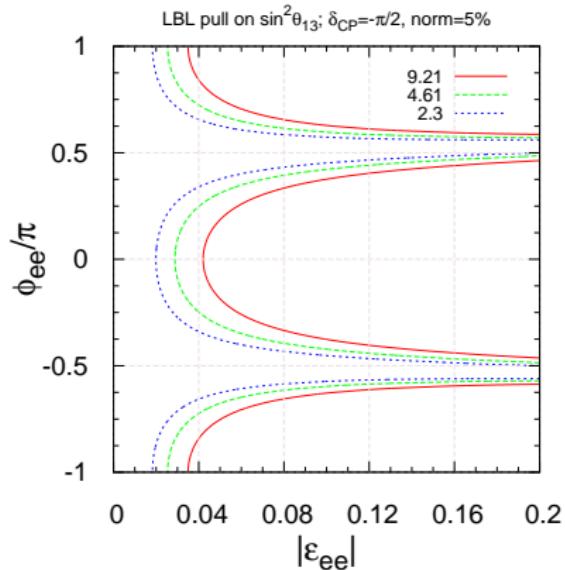
Daya Bay Vs RENO

$$0.0201 \leq \sin^2 \theta_{13} \leq 0.0306 \text{ (1}\sigma\text{)}$$

Daya Bay, 1230 days



RENO, 1500 days



Cancellations when  $\cos \phi_{ee} = 0$ , higher order  $\varepsilon$  terms important!

Work in progress: Updating RENO results with the latest data,  $\sim 2200$  days.

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## Summary

- The reactor mixing angle has been measured within a  $\sim 4\%$  precision mainly thanks to Daya Bay.
- Taking advantage of the Daya Bay data other beyond standard oscillation scenarios can be probed. In particular, CC-like NSI have been tested.
- Multidetector reactor neutrino experiments offer a clean probe of CC-like NSI. The  $\theta_{13}$  determination is in general NOT robust under CC-like NSI (due to the effect of the phases) while the value of the NSI constraints is limited by our current knowledge of the ‘absolute normalization of reactor neutrino fluxes’. **New physis might be ‘entangled’ with syst. errors.**
- By using the LBL result for the reactor mixing angle as an input, constraints on the ‘ $\varepsilon_{e(\mu,\tau)}$ ’ and ‘FU’ couplings substantially improved.

# THANK YOU

Thanks to FAPESP for financial support under 2017/01749-6

Thanks to the organizers for the invitation and for the partial financial aid.

Back up

## Complete second order NSI probability

Assuming  $\varepsilon_{e\alpha}^s = \varepsilon_{\alpha e}^{d*}$

$$\begin{aligned} P_{\bar{\nu}_e^s \rightarrow \bar{\nu}_e^d} &\simeq \underbrace{1 - \sin^2 2\theta_{13} \left( c_{12}^2 \sin^2 \Delta_{31} + s_{12}^2 \sin^2 \Delta_{32} \right) - c_{13}^4 \sin^2 2\theta_{12} \sin^2 \Delta_{21}}_{\text{Standard Model terms}} \\ &+ \underbrace{4|\varepsilon_e| \cos \phi_e + 2|\varepsilon_e|^2 (2 + \cos(2\phi_e)) + 2|\varepsilon_\mu|^2 + 2|\varepsilon_\tau|^2}_{\text{non-oscillatory NSI terms}} \\ &- \underbrace{4\{2s_{13}[s_{23}|\varepsilon_\mu| \cos(\delta - \phi_\mu) + c_{23}|\varepsilon_\tau| \cos(\delta - \phi_\tau)]\} \sin^2 \Delta_{31}}_{\text{oscillatory NSI terms}} \\ &- \underbrace{4\{s_{23}^2|\varepsilon_\mu|^2 + c_{23}^2|\varepsilon_\tau|^2 + 2s_{23}c_{23}|\varepsilon_\mu||\varepsilon_\tau| \cos(\phi_\mu - \phi_\tau)\} \sin^2 \Delta_{31}}_{\text{oscillatory NSI terms}} \end{aligned}$$

# Normalization of the NSI states

The new neutrino flavor states, in presence of NSI are normalized as:

$$|\nu_\alpha^s\rangle = \frac{1}{N_\alpha^s} \left( |\nu_\alpha\rangle + \sum_\gamma \varepsilon_{\alpha\gamma}^s |\nu_\gamma\rangle \right),$$

$$\langle\nu_\beta^d| = \frac{1}{N_\beta^d} \left( \langle\nu_\beta| + \sum_\eta \varepsilon_{\eta\beta}^d \langle\nu_\eta| \right),$$

where the normalization factors are given by:

$$N_\alpha^s = \sqrt{[(1 + \varepsilon^s)(1 + \varepsilon^{s\dagger})]_{\alpha\alpha}},$$

$$N_\beta^d = \sqrt{[(1 + \varepsilon^{d\dagger})(1 + \varepsilon^d)]_{\beta\beta}},$$

To see the normalization effect in the probability, we can compute the oscillation probability at zero distance:

$$P_{\nu_\alpha^s \rightarrow \nu_\beta^d}(L=0) = \left| \langle\nu_\beta^d| \nu_\alpha^s \rangle \right|^2 = \left| \frac{\delta_{\alpha\beta} + \varepsilon_{\alpha\beta}^s + \varepsilon_{\alpha\beta}^d + \sum_\gamma \varepsilon_{\alpha\gamma}^s \varepsilon_{\gamma\beta}^d}{N_\alpha^s N_\beta^d} \right|^2.$$

For  $\varepsilon_{\alpha\gamma}^s = \varepsilon_{\gamma\alpha}^{d*} = |\varepsilon_{\alpha\gamma}| \exp(i\phi_{\alpha\gamma})$  we find the oscillation amplitude is given by:

$$A_{\alpha\beta}(L=0) = \frac{\delta_{\alpha\beta} + 2\Re(|\varepsilon_{\alpha\beta}| \exp i\phi_{\alpha\beta}) + \sum_\gamma |\varepsilon_{\alpha\gamma}| |\varepsilon_{\beta\gamma}| \exp [i(\phi_{\alpha\gamma} - \phi_{\beta\gamma})]}{\sqrt{1 + 2\Re(|\varepsilon_{\alpha\alpha}| \exp i\phi_{\alpha\alpha}) + \sum_\gamma |\varepsilon_{\alpha\gamma}|^2} \sqrt{1 + 2\Re(|\varepsilon_{\beta\beta}| \exp i\phi_{\beta\beta}) + \sum_\gamma |\varepsilon_{\beta\gamma}|^2}} \neq 0.$$