

# Results and Prospects from NOvA

*Jianming Bian*

*For the NOvA Collaboration*

*University of California, Irvine*

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*NUFACT2018, Blacksburg, Virginia*

# NOvA Physics Goals

- $\nu_e$  appearance +  $\nu_\mu$  disappearance
  - **Mass hierarchy**:  $m_3 > m_{1,2}$  or  $m_{1,2} > m_3$ ? Implications for absolute neutrino masses, unified theories and neutrino-less double beta decay searches
  - **CP phase  $\delta_{CP}$** : whether neutrinos and antineutrinos behave the same way in oscillation? Implications for matter-antimatter asymmetry
  - **Octant of  $\theta_{23}$** : Is  $\theta_{23}$  exactly  $45^\circ$ ? Is  $\nu_3$  more strongly coupled to  $\nu_\tau$  or  $\nu_\mu$ ?

(This talk, technical details see Talk 182, Erica Smith, 08/16/2018)

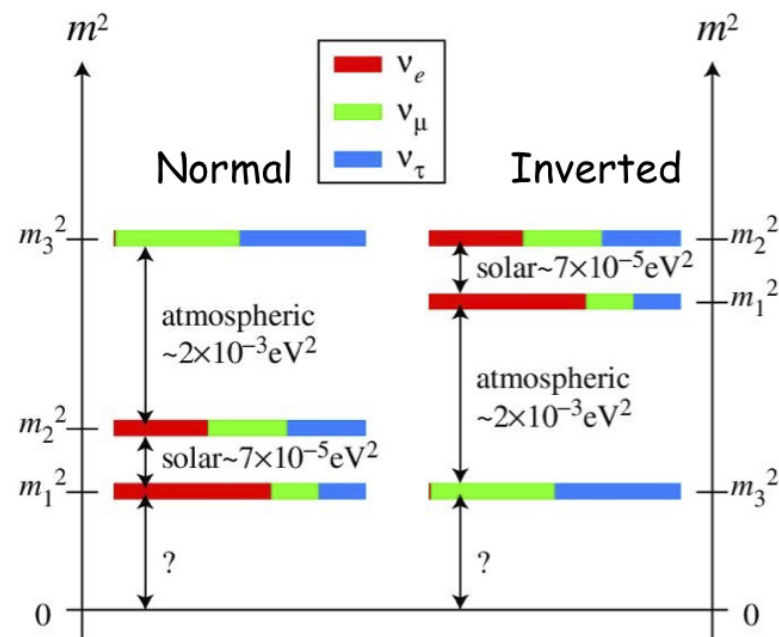
- **NC disappearance**
  - **Sterile neutrino search**: are there other neutrinos beyond the three known active flavors?

(See Talk 190, Michael Wallbank, 08/17/2018)

- Also, cross sections, exotic phenomena and non-beam physics

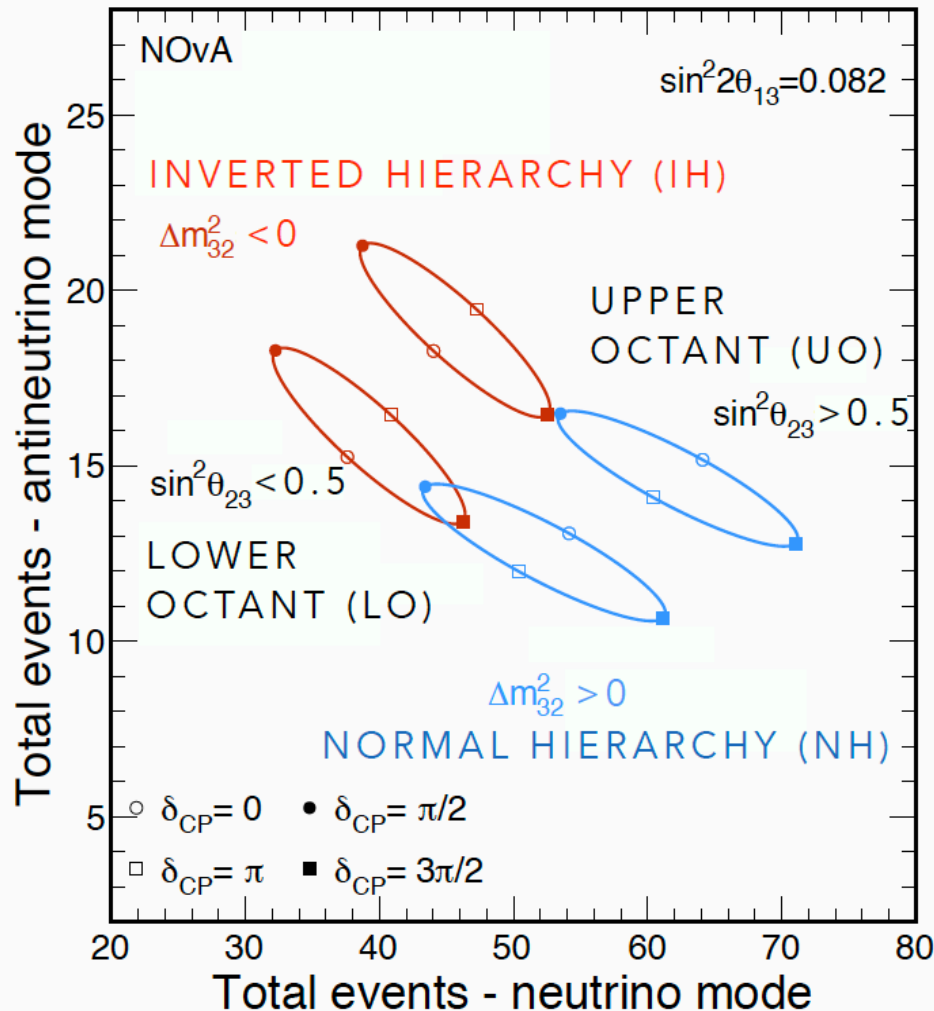
(See Talk 71, Matt Judah, 08/13/2018)

This talk: New  $\nu_e$  and  $\nu_\mu$  oscillation results with NOvA's first antineutrino data



# Appearance and Disappearance at NOvA

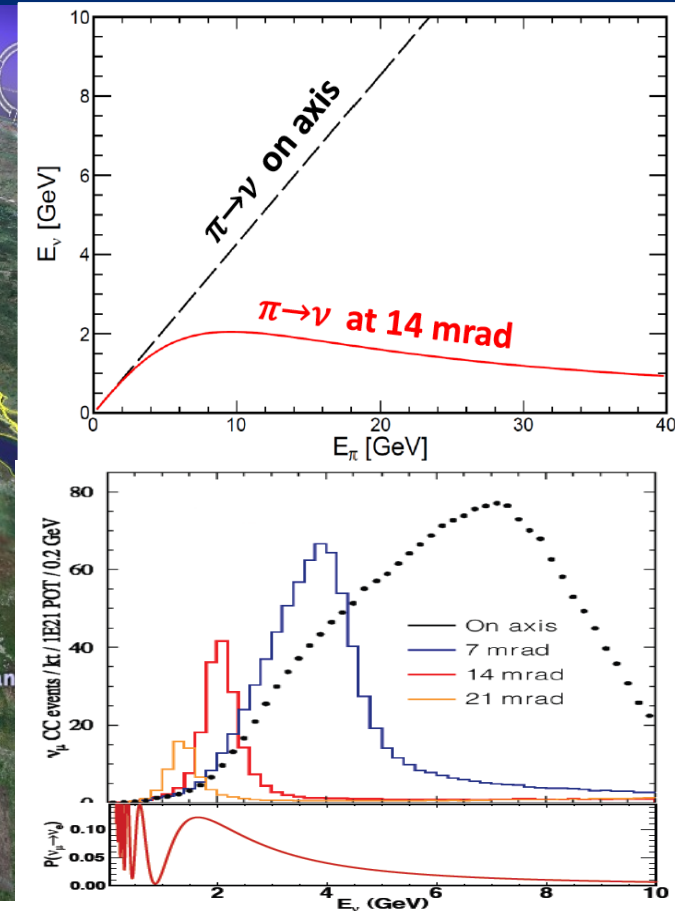
$\nu_e/\bar{\nu}_e$  Appearance event counts



- Measuring  $\nu_e$  and  $\bar{\nu}_e$  appearance probabilities with  $\nu_\mu$  and  $\bar{\nu}_\mu$  beam
- When other parameters fixed,  $\nu_\mu \rightarrow \nu_e$  and  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  appearance probabilities depend on
  - sign of  $\Delta m_{32}^2$
  - $\delta_{CP}$
  - octant of  $\theta_{23}$
- $\nu_\mu$  and  $\bar{\nu}_\mu$  disappearance provides high precision  $\Delta m_{32}^2$  and  $\sin^2 2\theta_{23}$ , constrain  $\theta_{23}$  octant



# NuMI Off-Axis $\nu_e$ Appearance Experiment

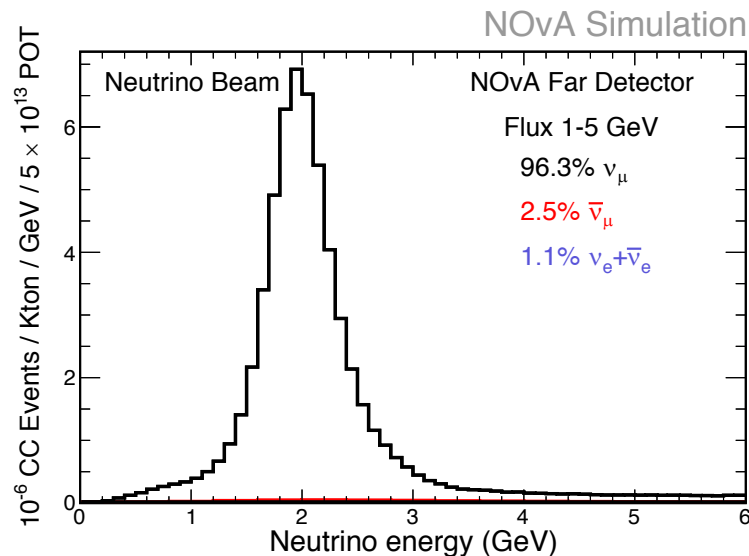


- Upgraded NuMI muon neutrino beam at Fermilab (700 kW design goal achieved)
- Longest baseline in operation (810 km), large matter effect ( $\pm 30\%$ ), sensitive to mass hierarchy
- Far/Near detector sited 14 mrad off-axis, narrow-band beam around oscillation maximum, small wrong sign components

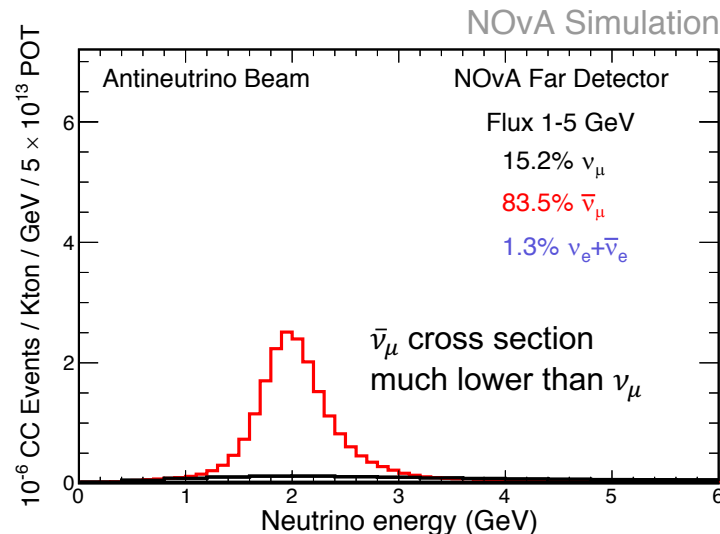


# Beam Performance

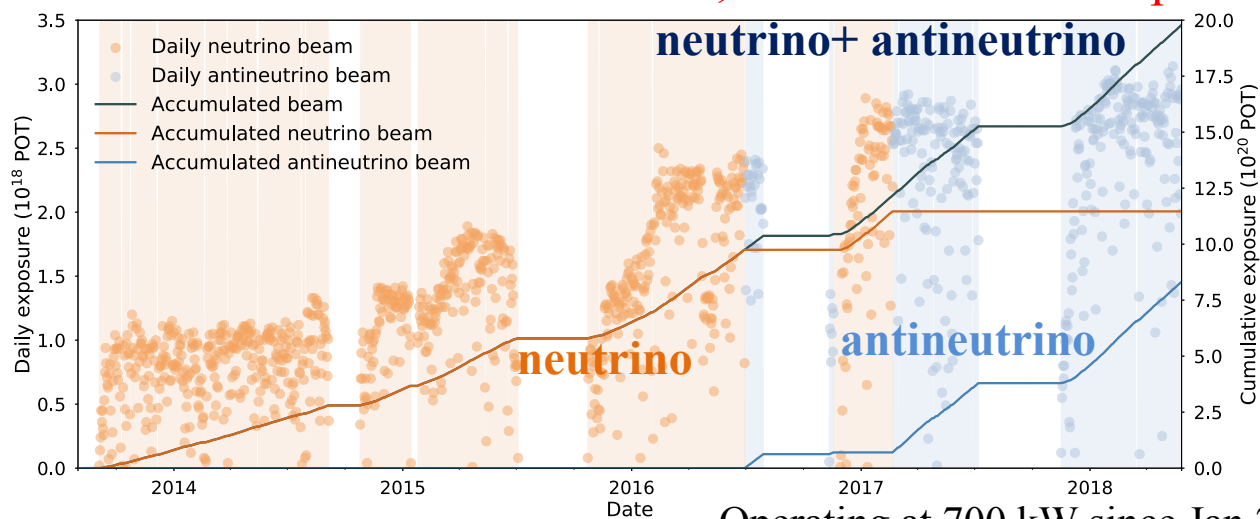
## CC event rates at FD in neutrino beam



## CC event rates at FD in antineutrino beam



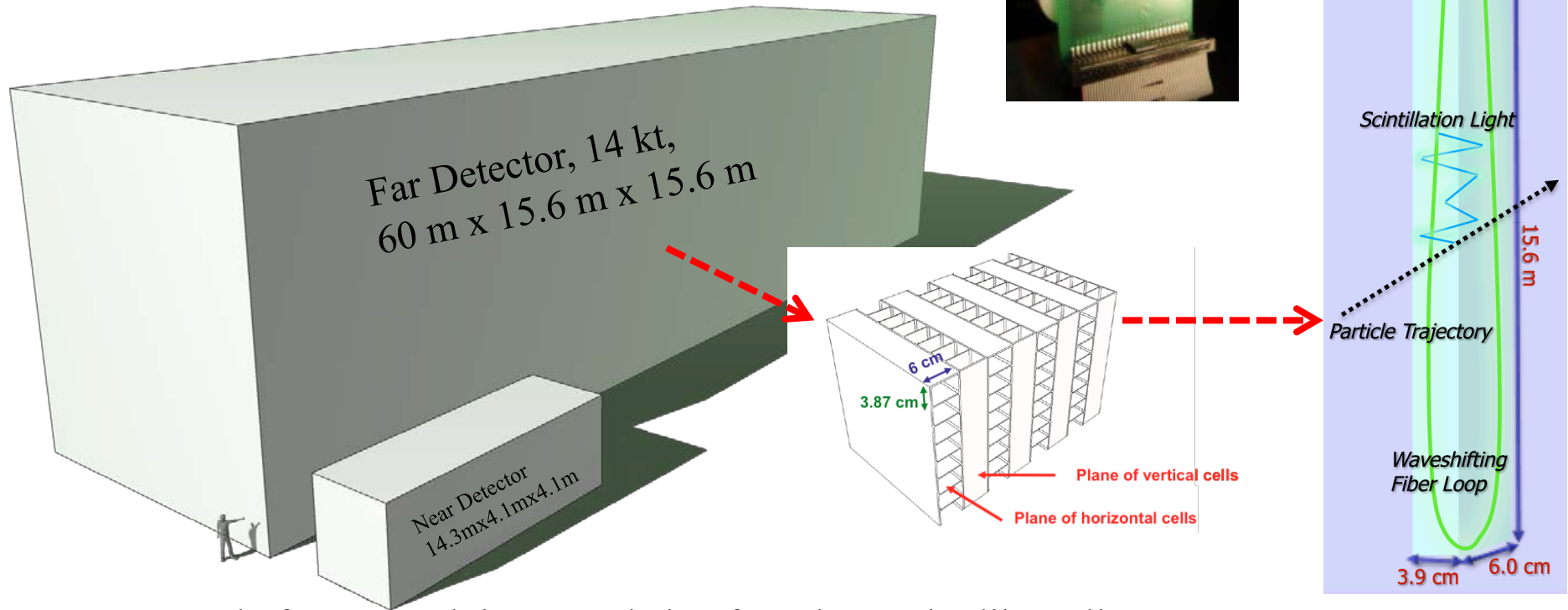
- **Neutrino beam data:**  $8.85 \times 10^{20}$  POT, taken Feb 2014 - Feb 2017
- **First antineutrino data:**  $6.9 \times 10^{20}$  POT, taken Feb 2017 - April 2018



Operating at 700 kW since Jan 2017

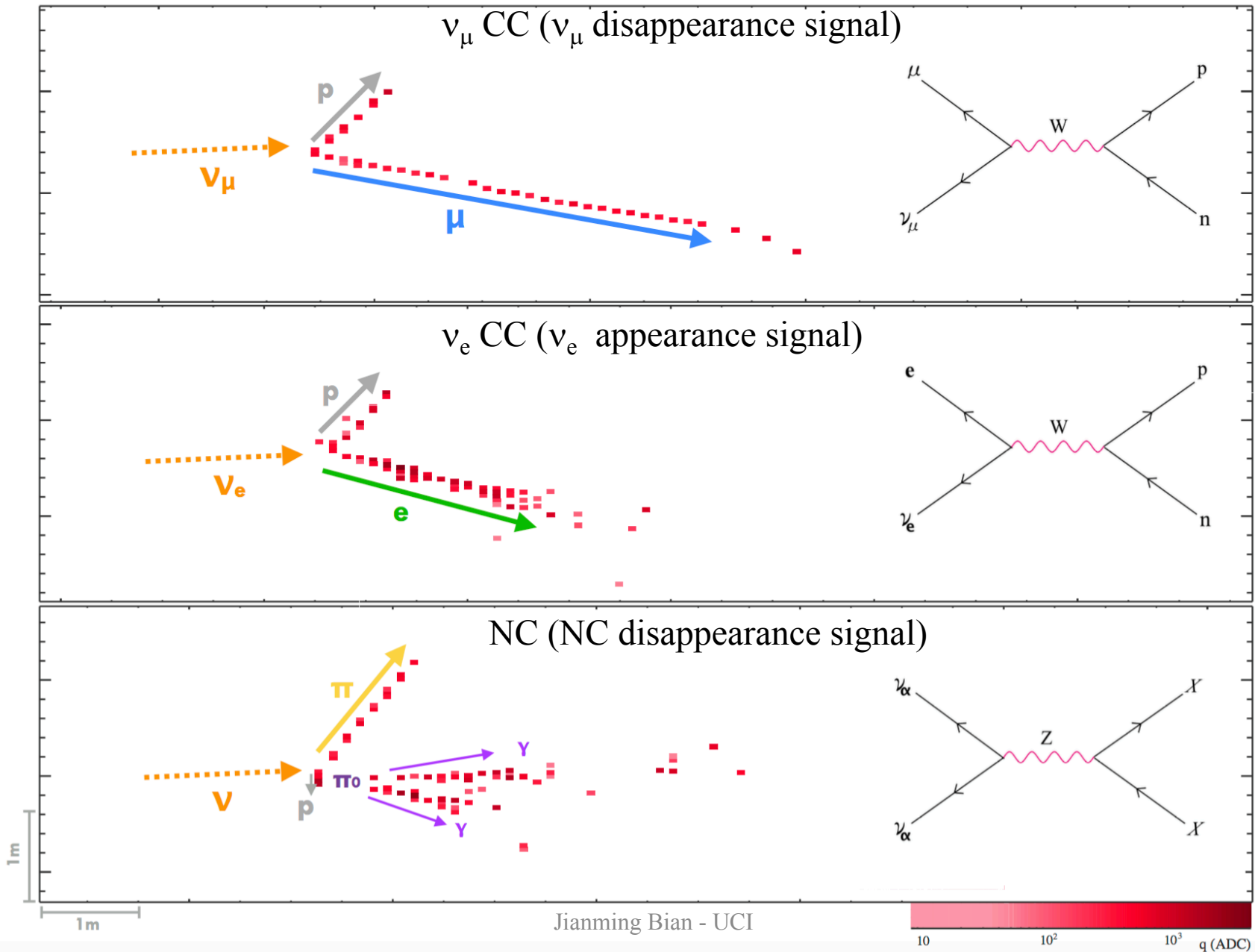
# The NOvA Detectors

- 14-kton Far Detector
- 344,064 detector cells
- 0.3-kton functionally identical Near Detector
- 18,432 cells



- Composed of PVC modules extruded to form long tube-like cells
- Each cell: filled with liquid scintillator, has wavelength-shifting fiber (WLS) routed to Avalanche Photodiode (APD)
- Cells arranged in planes, assembled in alternating vertical and horizontal directions
- Low-Z and low-density, each plane just 0.15  $X_0$ , great for  $e^-$  vs  $\pi^0$  separation

# NOvA Event Topologies

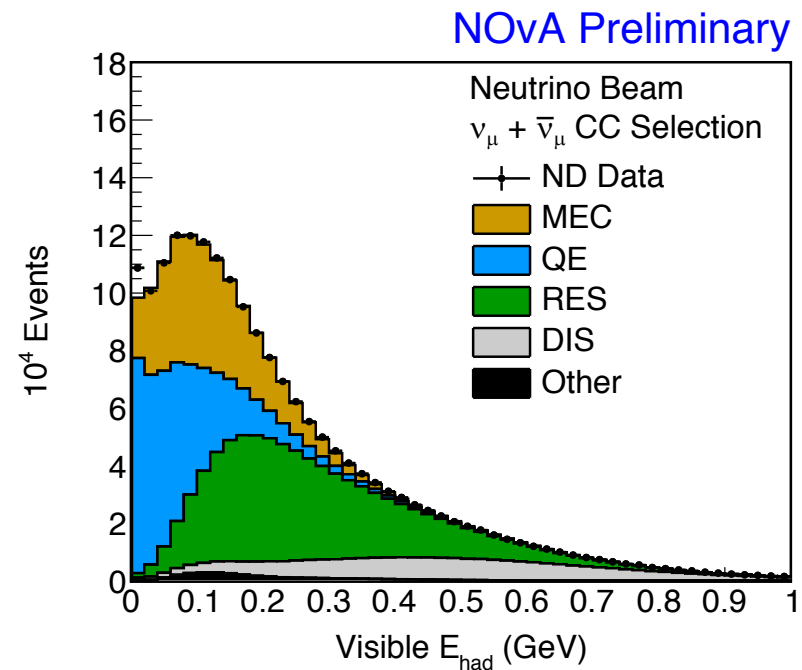
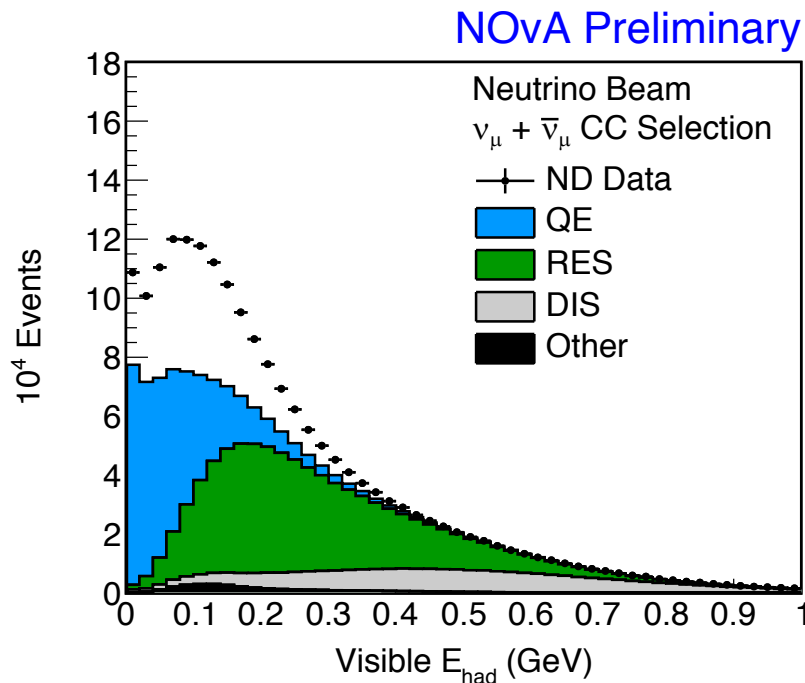




# Neutrino Interaction Tuning

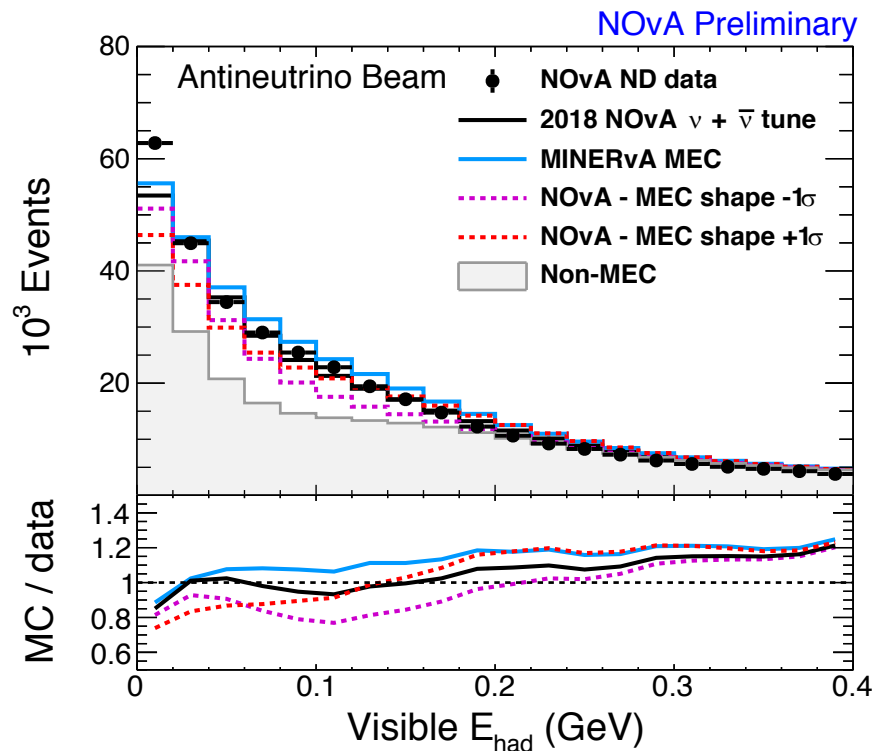
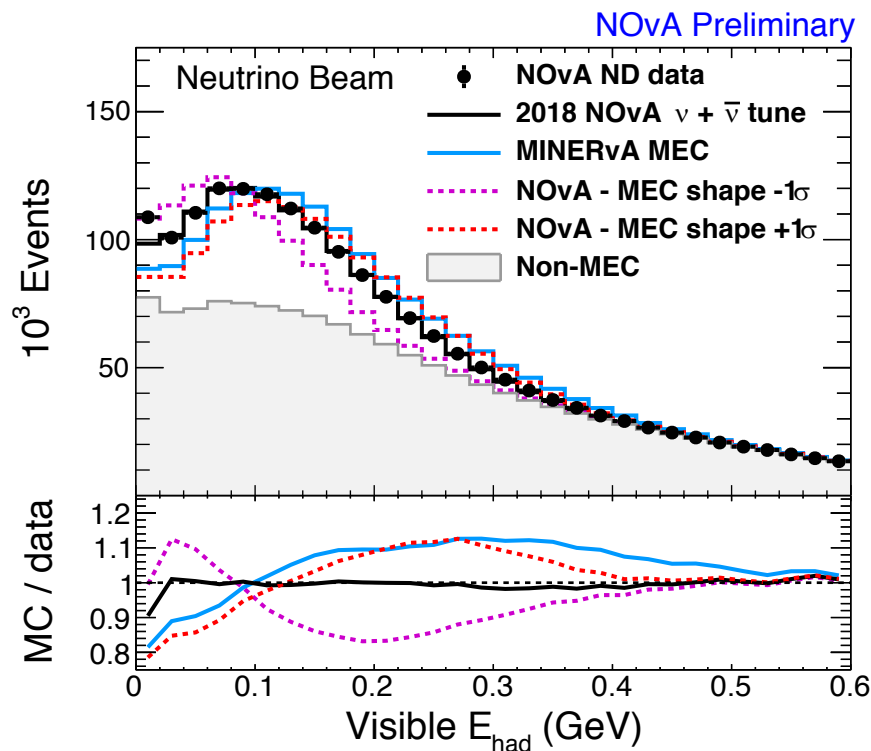
- QE, RES tuned to consider long-range nuclear correlations using València model via work of R. Gran (MINERvA) [<https://arxiv.org/abs/1705.02932>]
- DIS at high invariant mass ( $W > 1.7 \text{ GeV}/c^2$ ) weighted up 10% based on NOvA data
- Empirical MEC (Meson Exchange Current) model for Multi-nucleon ejection (2p2h)  
[T. Katori, AIP Conf. Proc. 1663, 030001 (2015)], amount tuned in 2D 3-momentum and energy transfers ( $q_0 = E_\nu - E_\mu$ ,  $|\mathbf{q}| = |\mathbf{p}_\nu - \mathbf{p}_\mu|$ ) space to match ND data

(See Talk 134, Jeremy Wolcott, 08/17/2018)

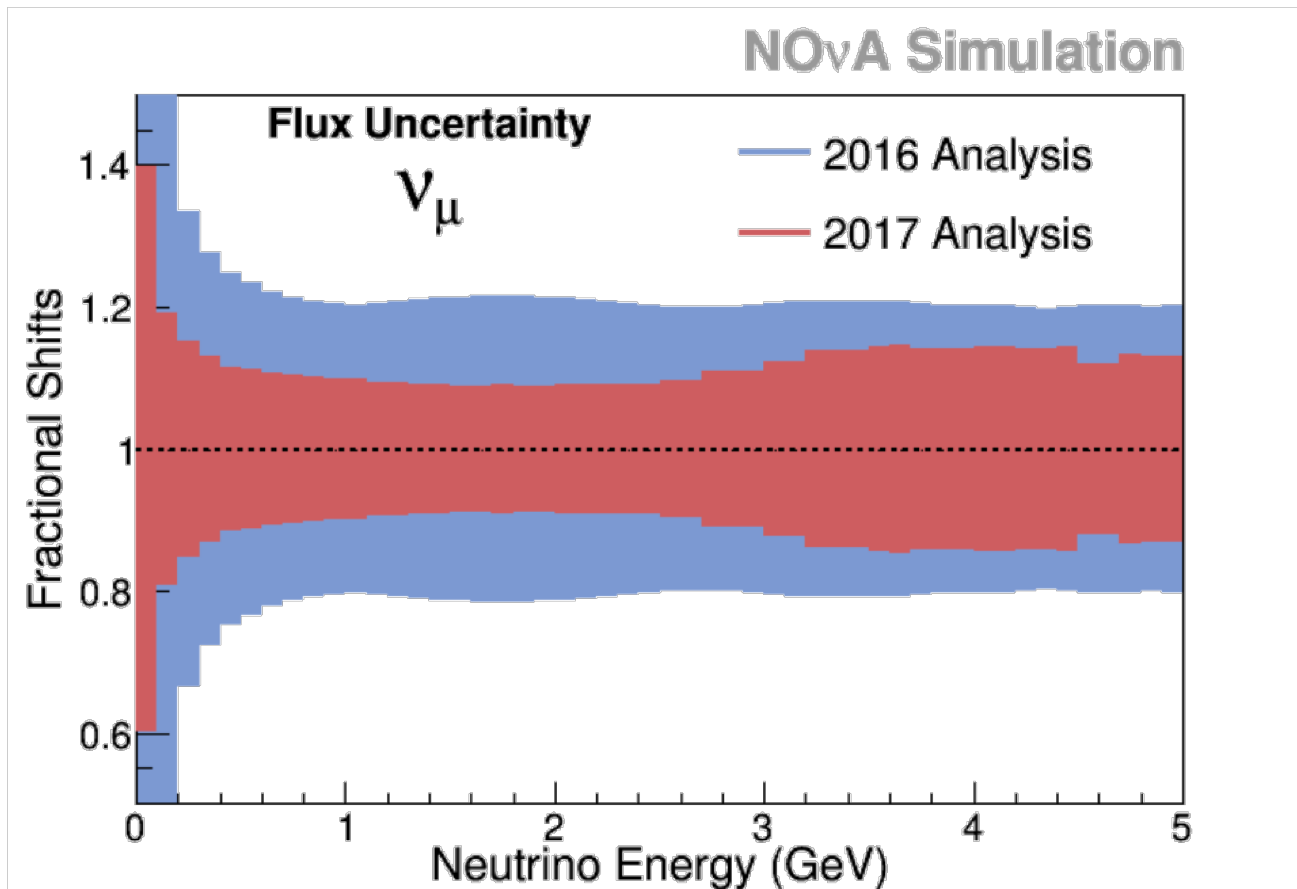


# Neutrino Interaction Tuning

- Empirical MEC (Meson Exchange Current) model for Multi-nucleon ejection (2p2h) [T. Katori, AIP Conf. Proc. 1663, 030001 (2015)], amount tuned in 2D 3-momentum and energy transfers ( $q_0 = E_\nu - E_\mu$ ,  $|\mathbf{q}| = |\mathbf{p}_\nu - \mathbf{p}_\mu|$ ) space to match ND data
- MEC shape systematic estimated by re-fitting using models with QE and RES related systematic shifts



# Improved Flux Model



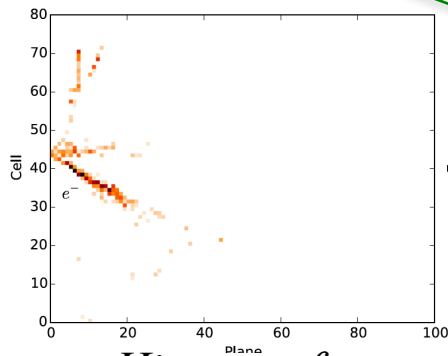
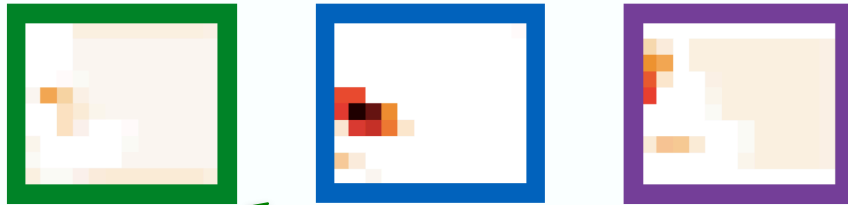
- Package to Predict the Flux (PPFX) from MINERvA (Phys. Rev. D 94, 092005. 2016).
  - Based on thin target hadron production data from NA49 and MIPP.
- Significantly reduced systematic uncertainties.
  - Central values also changed within prior systematics, but not shown here.



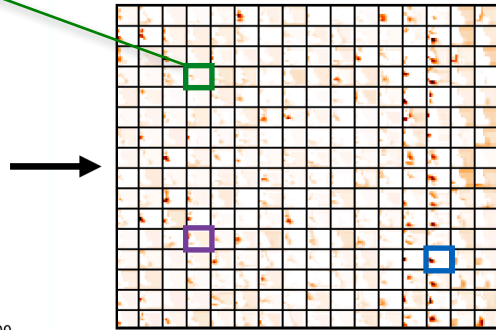
# Deep-Learning based PID for $\nu_e$ and $\nu_\mu$ Analyses

- CVN: a convolutional neural network (CNN), based on modern image recognition technology
- Introduce convolution filters to extract features from the hit map for the training of the neural net
- Statistical power equivalent to 30% more exposure than previous  $\nu_e$  PIDs
- $\nu_e$ ,  $\nu_\mu$  and NC analyses all use CVN as event selector

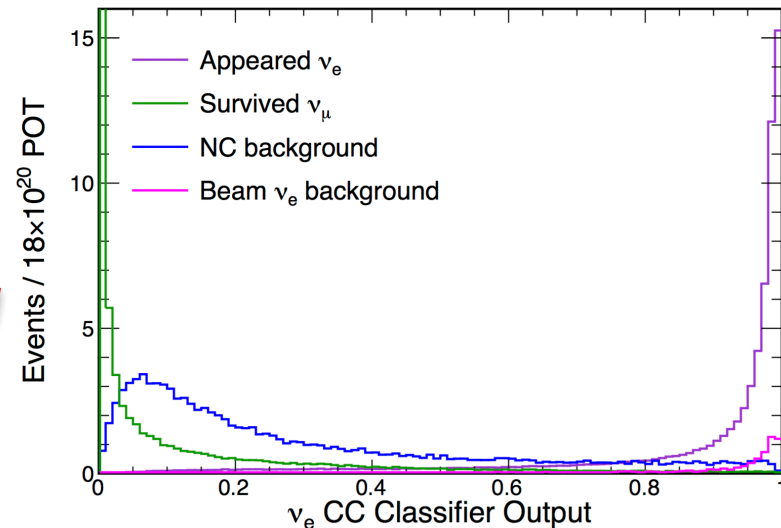
*Outputs of convolutional filters (features)*



*Hit map of  
a  $\nu_e$  CC event*



*CVN output in the far detector MC*

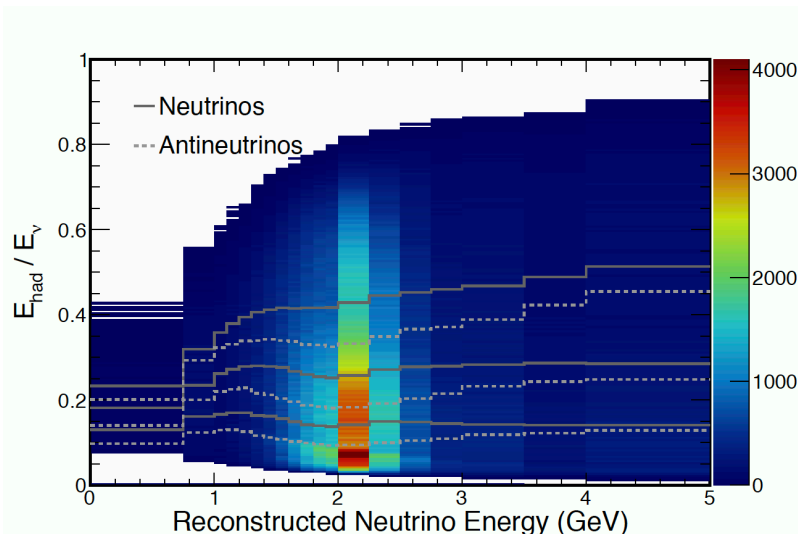


(See Poster 206, Fernanda Psihas)

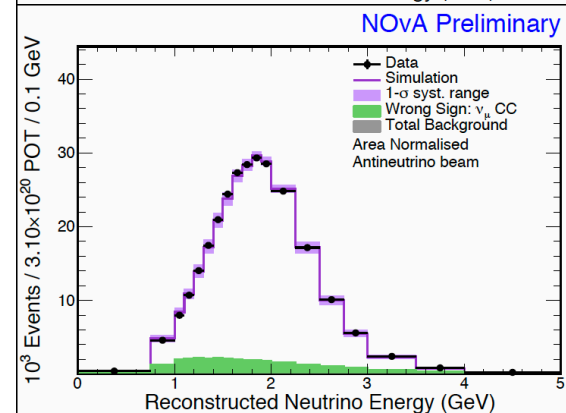
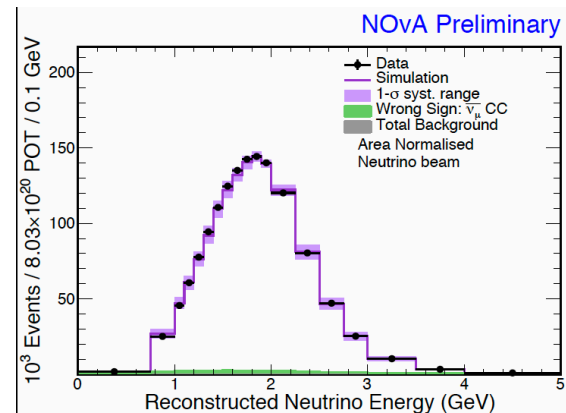
At NOvA, CVN has been extended to single particle ID, energy reconstruction (for future analyses), etc

# Near Detector Spectrum ( $\nu_\mu$ disappearance)

- Select  $\nu_\mu$  ( $\bar{\nu}_\mu$ ) CC in ND from neutrino (antineutrino) beam, wrong sign contamination 3% (11%)
- $E_\nu = E_\mu + E_{\text{had}}$ , data split in 4 equal energy quantiles based on  $E_{\text{had}}/E_\nu$ , resolution varies from 5.8% (5.5%) to 11.7% (10.8%) for neutrino (antineutrino) beam.
- Normalize ND MC to data in each  $E_\nu$  bin, then extrapolate the 4 quantiles to FD



Reco  $\nu_\mu$  ( $\bar{\nu}_\mu$ ) energy, all Quartiles



Area-normalized, **shape-only systematics**

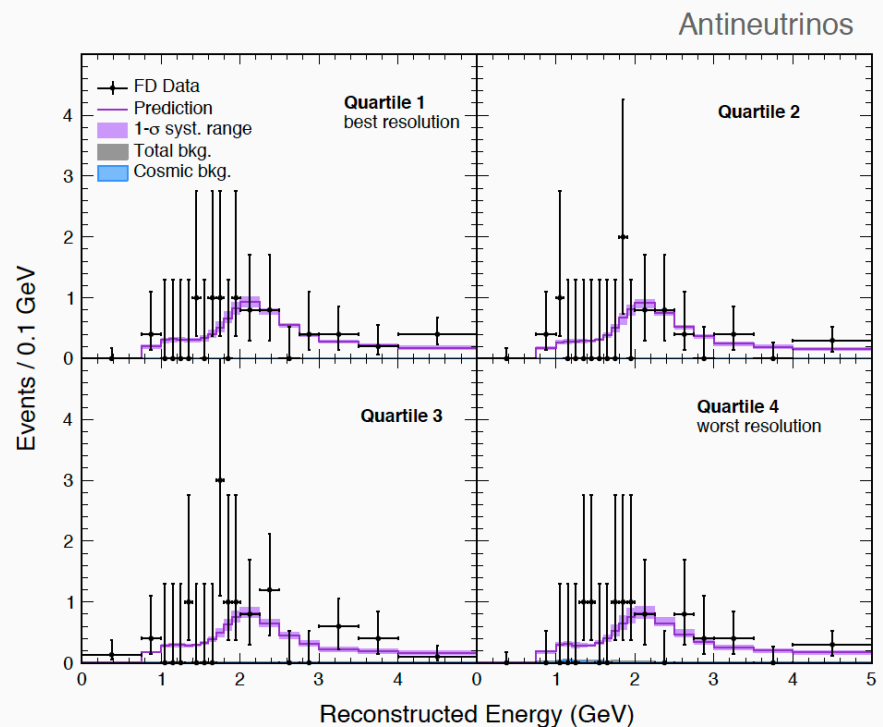
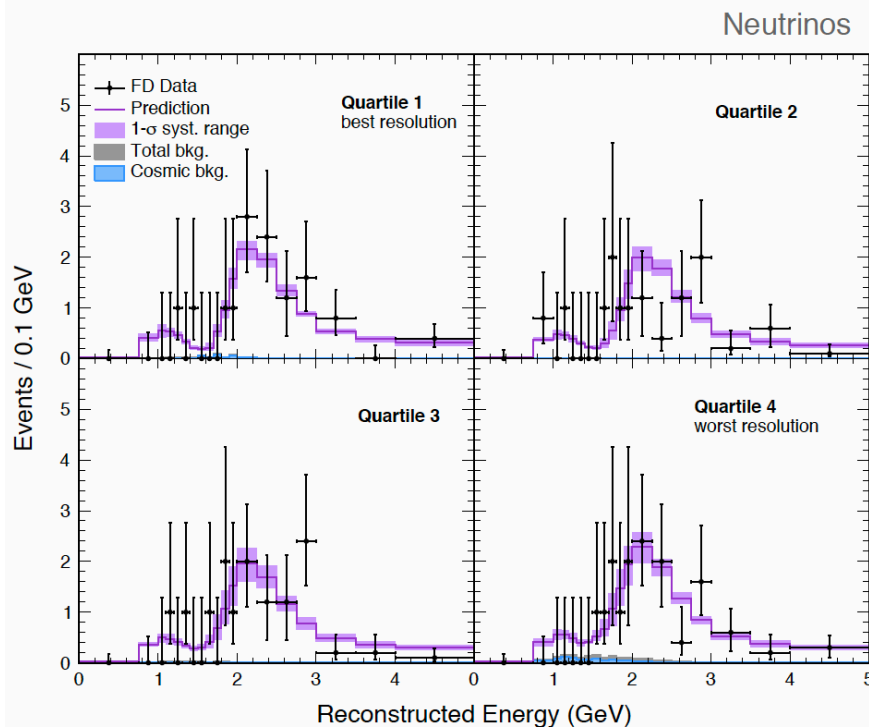
Data/MC normalization difference:

1.3% and 0.5% for  $\nu_\mu$  and  $\bar{\nu}_\mu$

# $\nu_\mu$ Data at Far Detector

- FD selection:
  - Additional BDT to reduce cosmic backgrounds
  - Estimate cosmic background rate from timing sidebands of the NuMI beam triggers and cosmic trigger data

$\nu_\mu$  events in 4 quartiles, each quartile extrapolated independently

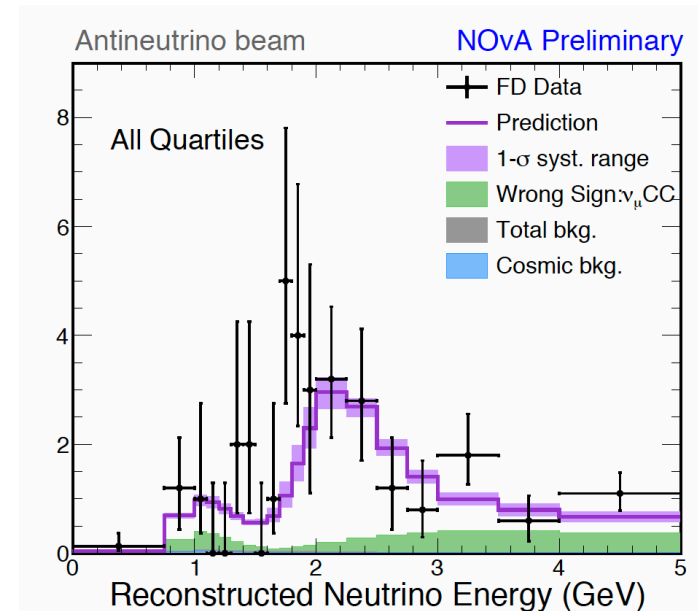
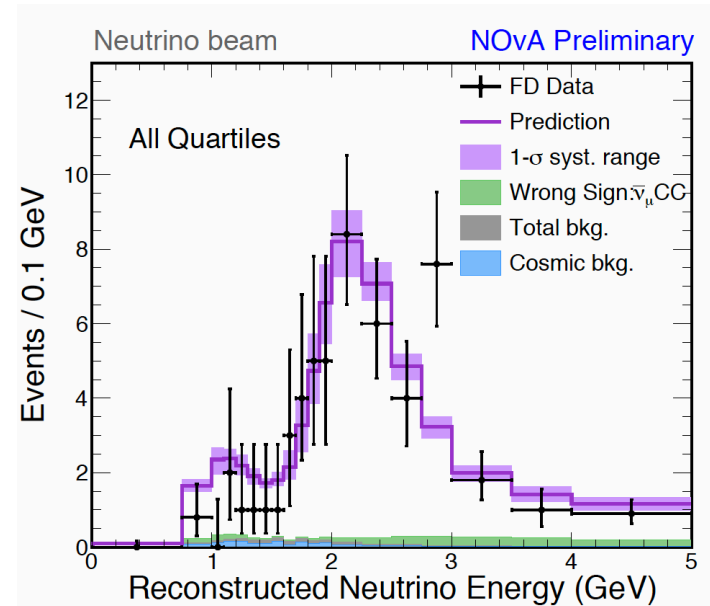




# $\nu_\mu$ Data at Far Detector

- FD selection:
  - Additional BDT to reduce cosmic backgrounds
  - Estimate cosmic background rate from timing sidebands of the NuMI beam triggers and cosmic trigger data
- Neutrino beam:
  - Observe 113 events
  - Expect  $730 +38/-49(\text{syst.})$  w/o oscillations
- Antineutrino beam:
  - Observe 65 events
  - Expect  $266 +12/-14(\text{syst.})$  w/o oscillations

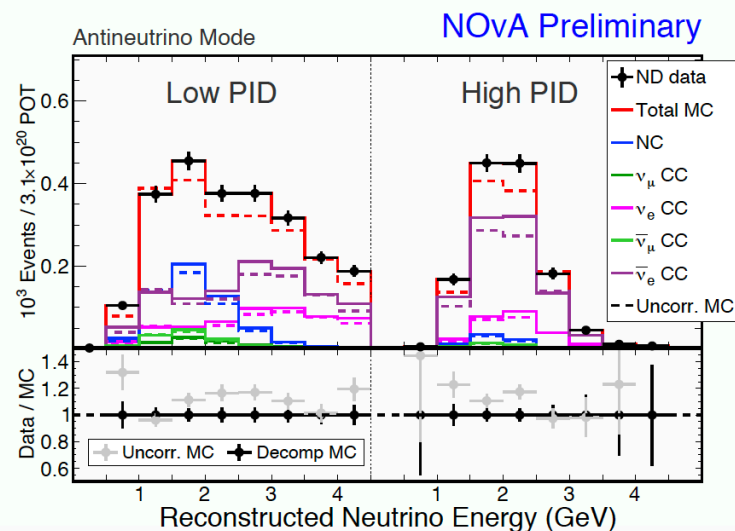
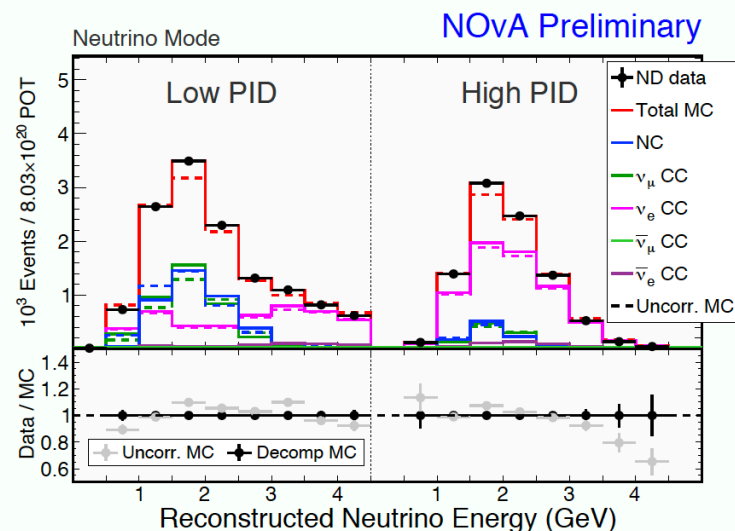
4 quartiles combined



# Near Detector Spectrum ( $\nu_e$ appearance)

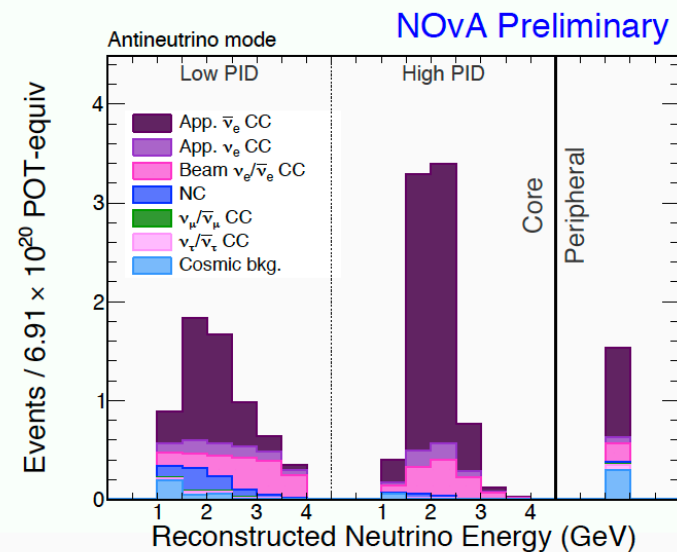
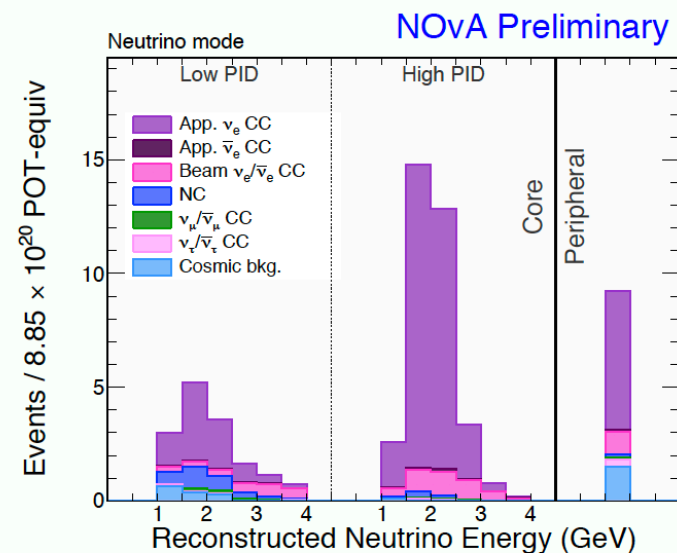
- Select  $\nu_e$  ( $\bar{\nu}_e$ ) CC in ND from neutrino (antineutrino) beam
- $E_\nu = f(E_e, E_{had})$ , data split into low and high particle ID (purity) range
- For neutrino beam:
  - Contained and uncontained  $\nu_\mu$  events constrain the  $\pi/K$  contributions to the beam  $\nu_e$ 's.
  - Michel electrons constrains NC/ $\nu_\mu$  CC balance in each  $E_\nu$  bin
- For antineutrino beam, scale all components evenly to match data
- ND $\rightarrow$ FD extrapolation: Each component propagated independently in energy and PID bins

(See Talk 182, Erica Smith, 08/16/2018)



# $\nu_e$ Far Detector Prediction

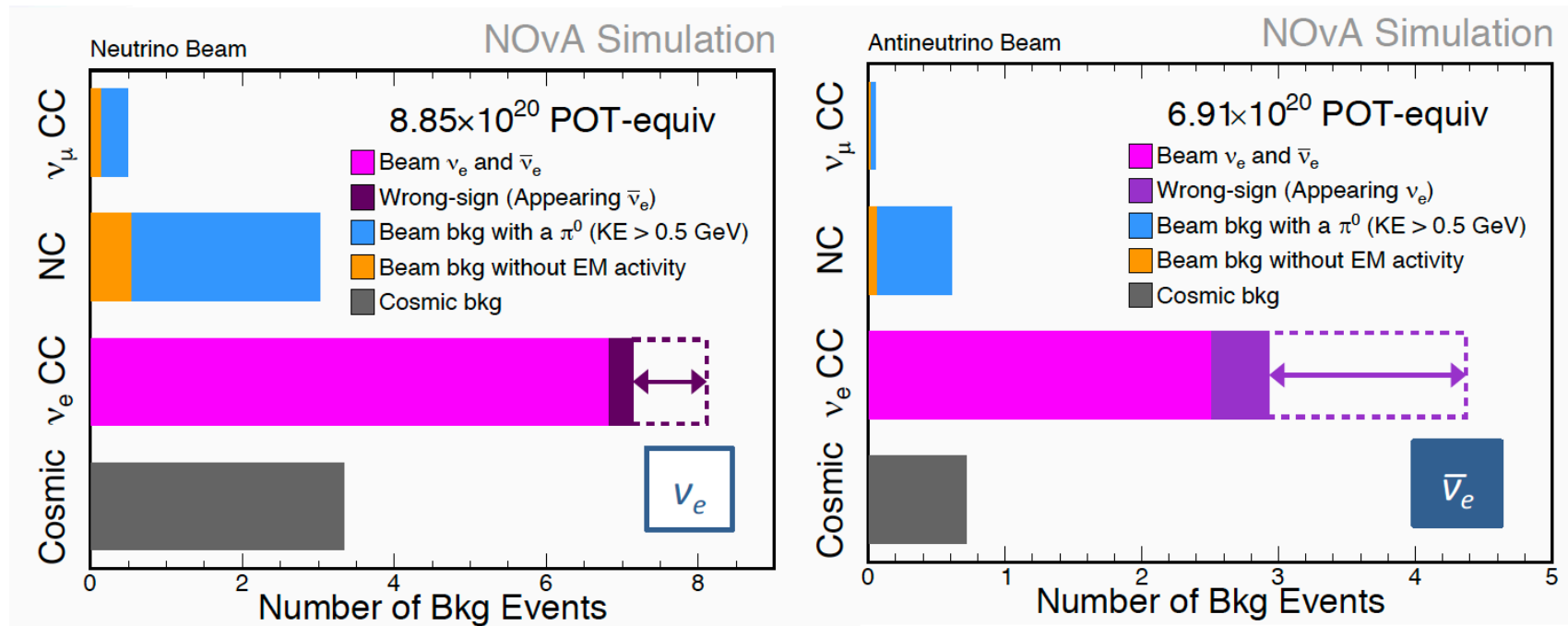
- FD selection:
  - Add a one-bin peripheral with less stringent containment selection to include more signal
  - Use location dependent BDT and tight PID cuts to recover signal events in this peripheral bin
- ND→FD extrapolation: Each component propagated independently in energy and PID bins
- Neutrino beam:
  - Background: 11 beam, 3 cosmic and < 1 wrong sign
- Antineutrino beam:
  - Background events : 3.5 beam, <1 cosmic and 1 wrong sign





# $\nu_e$ Far Detector Backgrounds

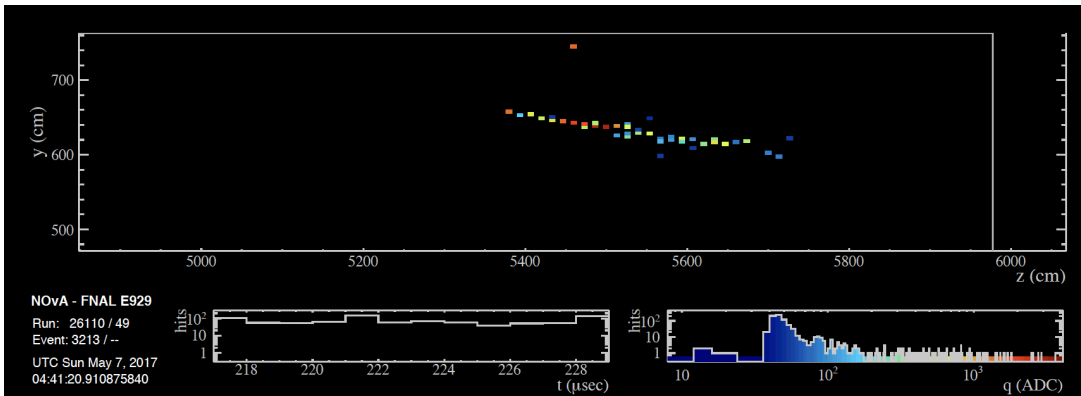
- Neutrino beam:
  - Background: 11 beam, 3 cosmic and  $< 1$  wrong sign
- Antineutrino beam:
  - Background events : 3.5 beam,  $< 1$  cosmic and 1 wrong sign



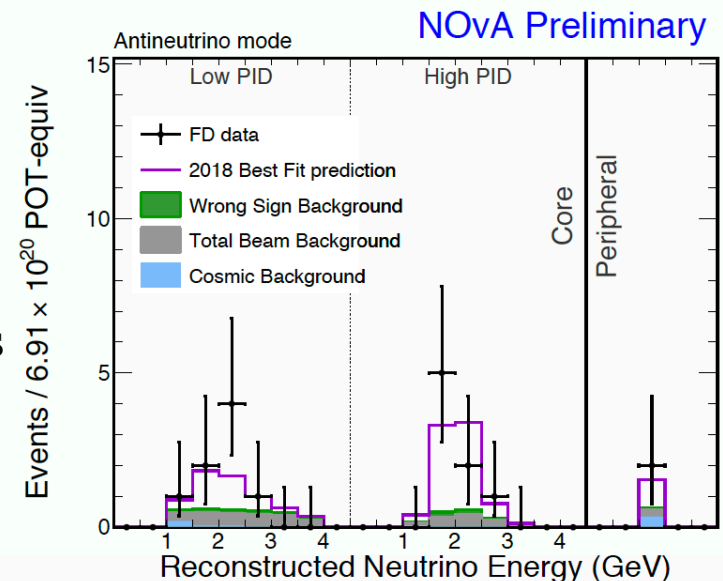
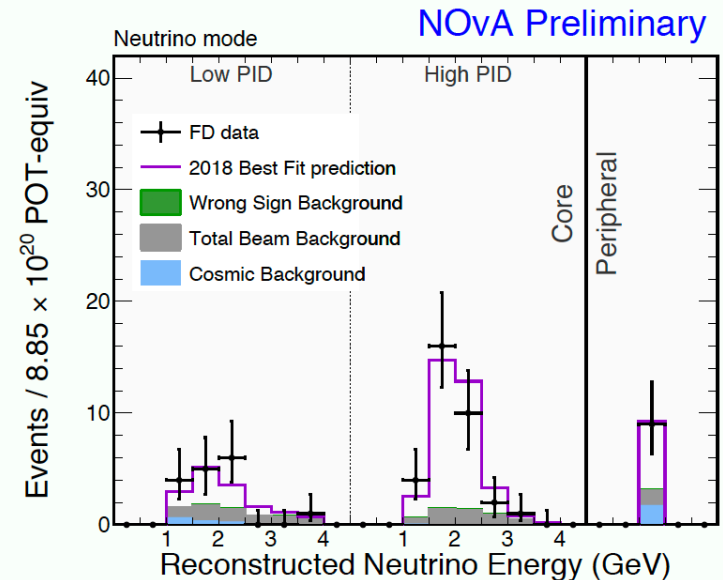
- Major backgrounds from beam  $\nu_e$
- Wrong sign background depends on oscillation

# $\nu_e$ Data at Far Detector

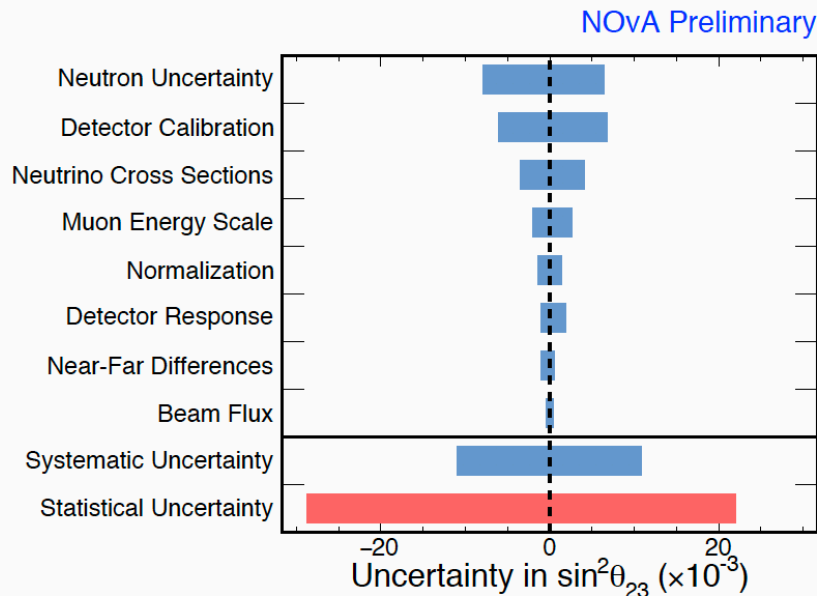
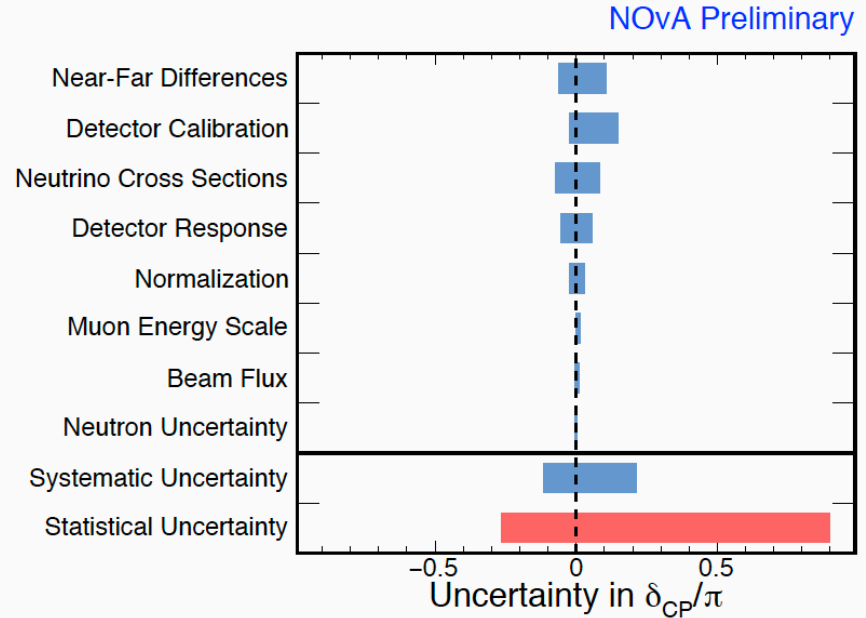
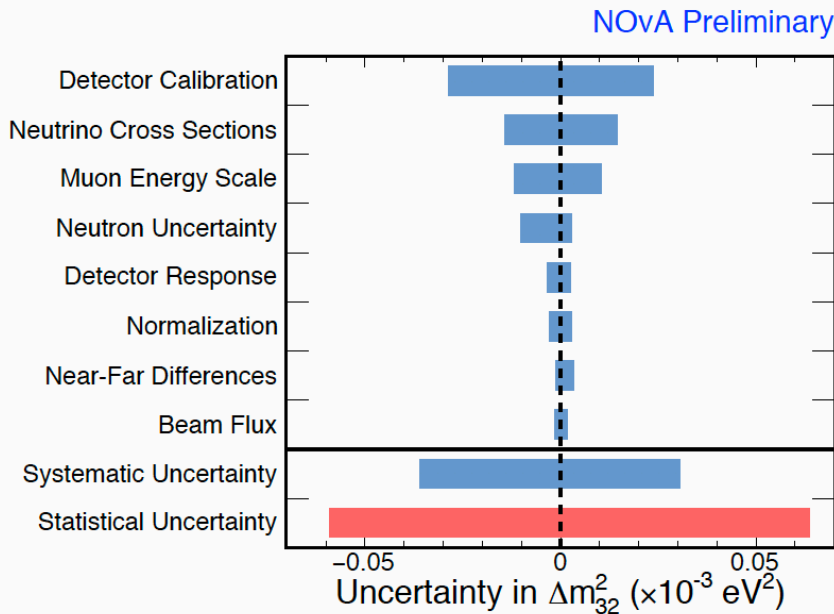
## Selected $\nu_e$ candidate in FD Data



- Neutrino beam:
  - Observe 58 events, expect 15 background events
- Antineutrino beam:
  - Observe 18 events, expect 5.3 background events
- $> 4\sigma$   $\bar{\nu}_e$  appearance



# Systematic Uncertainties (Joint fit)

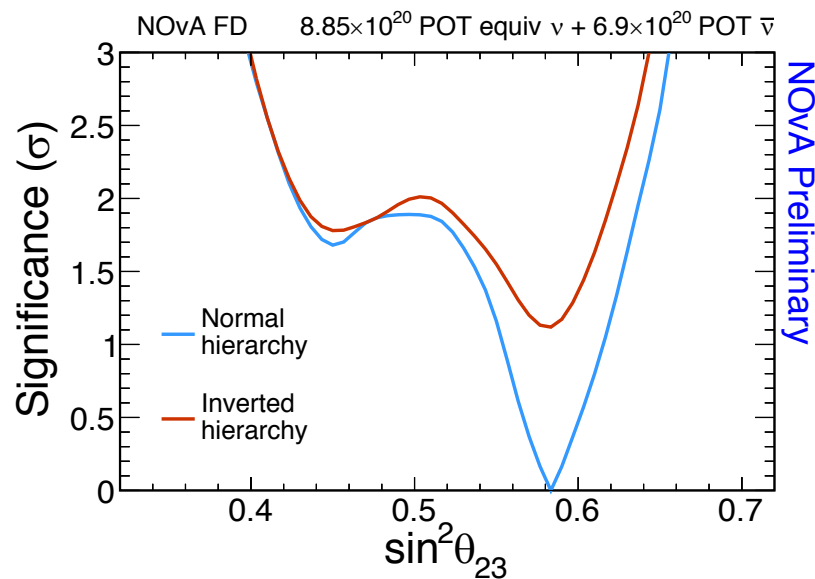
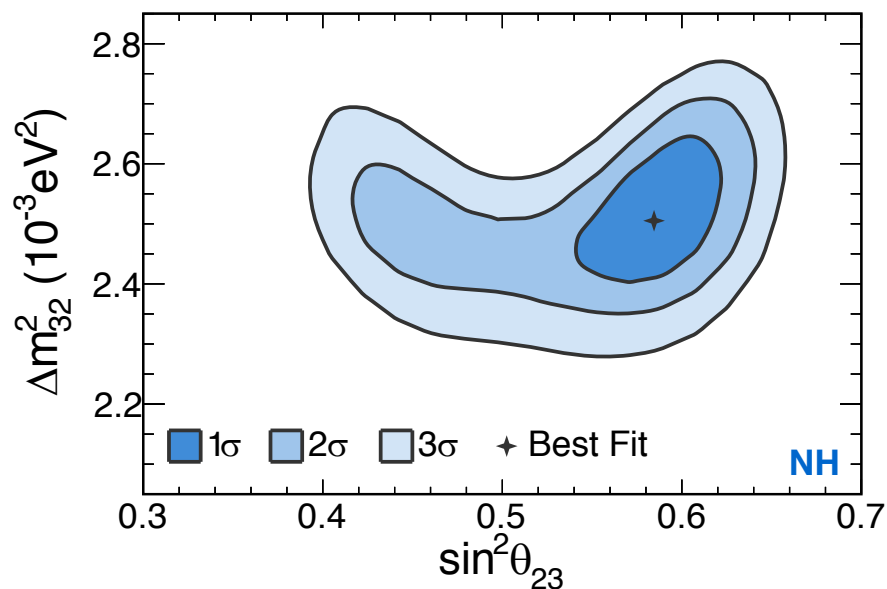


- Largest systematics for  $\nu_\mu$  and  $\nu_e$  are **calibration** and **cross-sections**.
- Both analyses are statistically limited.
- Upcoming NOvA test beam program will address calibration and detector response uncertainties

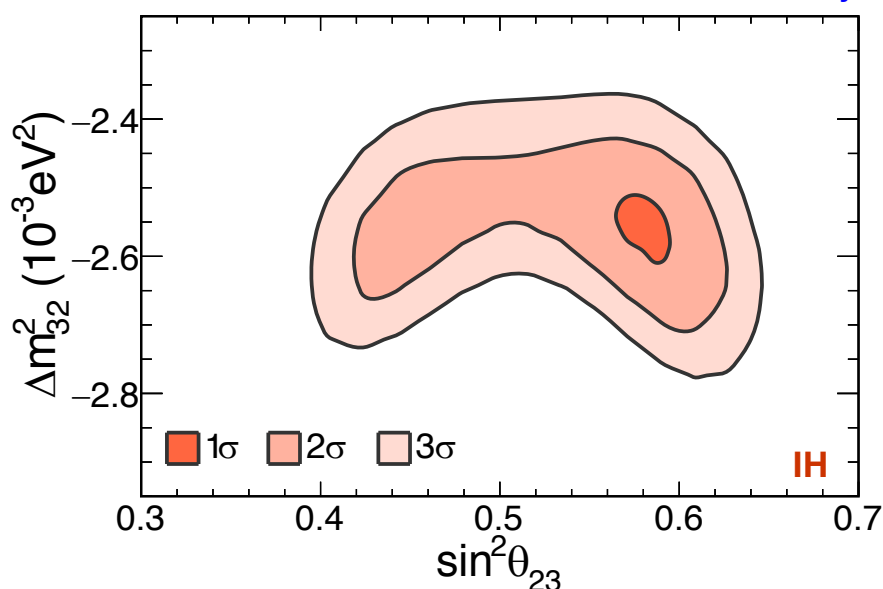
(See Poster 205, Andrew Sutton)

# Joint Appearance and Disappearance

NOvA Preliminary



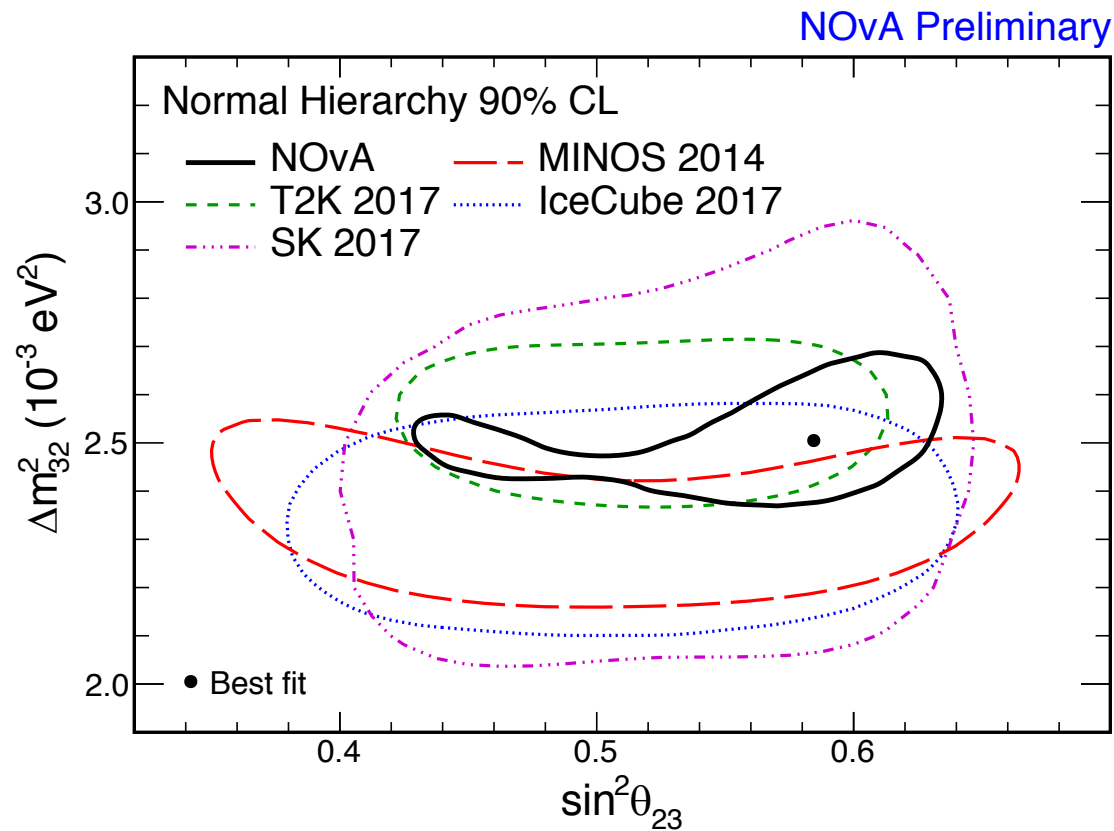
NOvA Preliminary



- Statistically limited, largest systematics for  $\nu_\mu$  and  $\nu_e$  are calibration and cross-sections.
- Best fit:
  - Normal Hierarchy
  - $\sin^2 \theta_{23} = 0.58 \pm 0.03$  (UO)
  - $\Delta m^2_{32} = (2.51 + 0.12 - 0.08) \cdot 10^{-3} \text{ eV}^2$
- Prefer non-maximal at 1.8 $\sigma$ , favor upper octant at similar level

# Joint Appearance and Disappearance

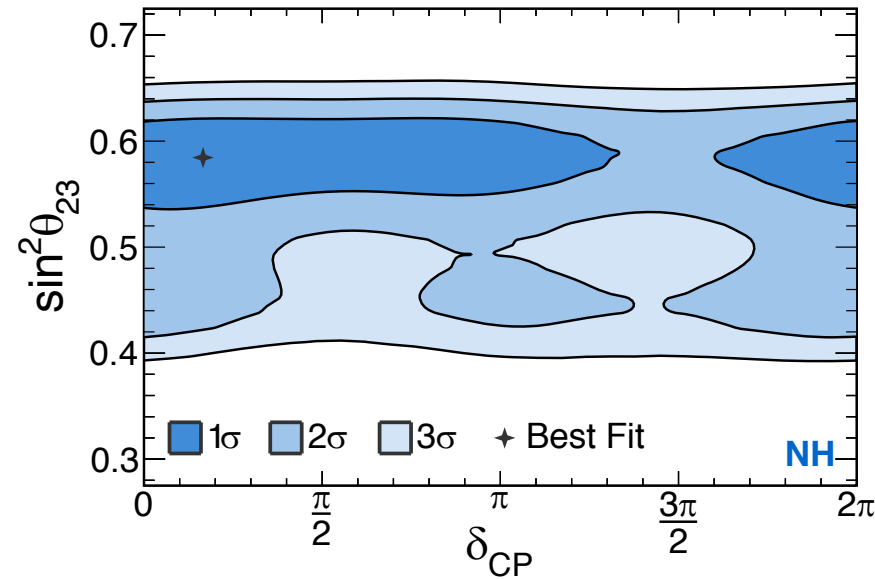
NOvA's allowed 90% C.L. regions are compatible to other experiments



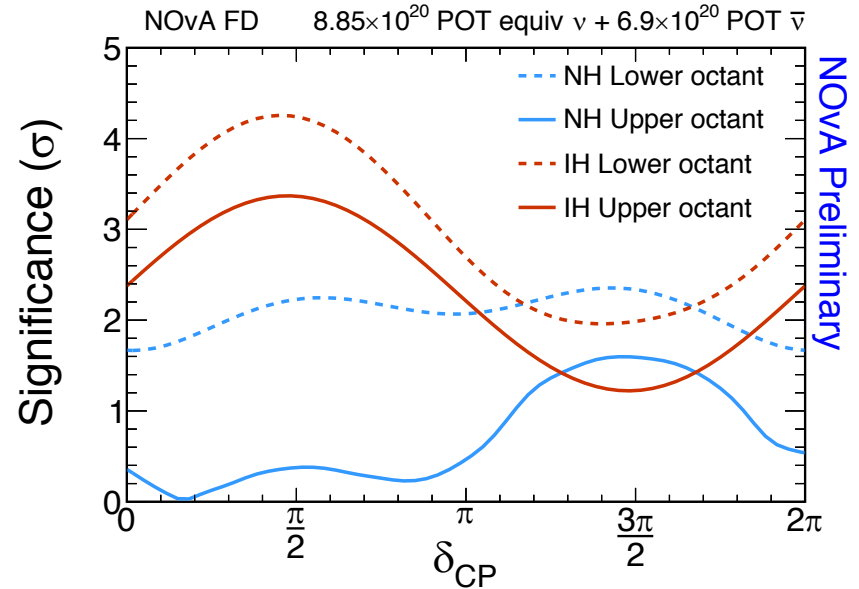
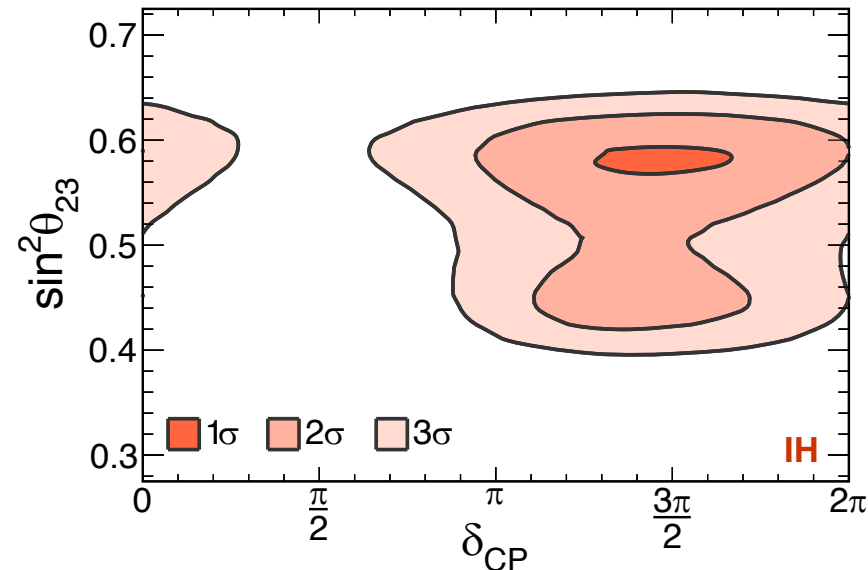


# Joint Appearance and Disappearance

NOvA Preliminary



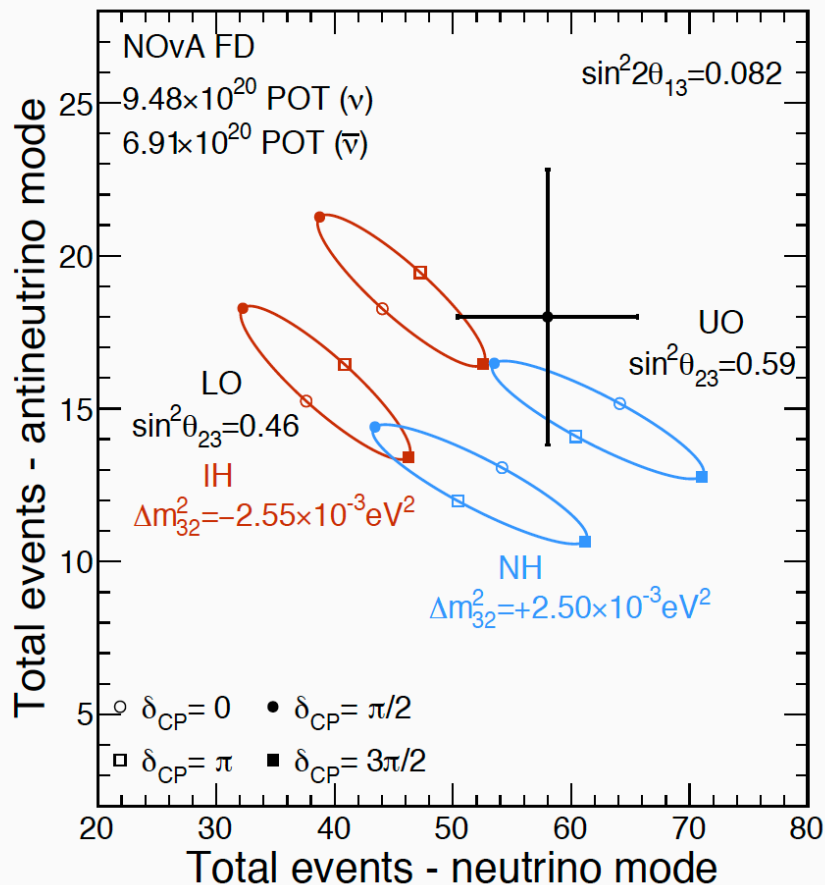
NOvA Preliminary



- Statistically limited, largest systematics for  $\nu_\mu$  and  $\nu_e$  are calibration and cross-sections
- Best fit:
  - Normal Hierarchy
  - $\delta_{CP} = 0.17\pi$
  - $\sin^2 \theta_{23} = 0.58 \pm 0.03$  (UO)
  - $\Delta m_{32}^2 = (2.51 + 0.12 - 0.08) \cdot 10^{-3} \text{ eV}^2$
- Consistent with all  $\delta_{CP}$  values in NH at  $< 1.6\sigma$
- Exclude  $\delta = \pi/2$  in IH at  $> 3\sigma$
- Prefer NH at  $1.8\sigma$

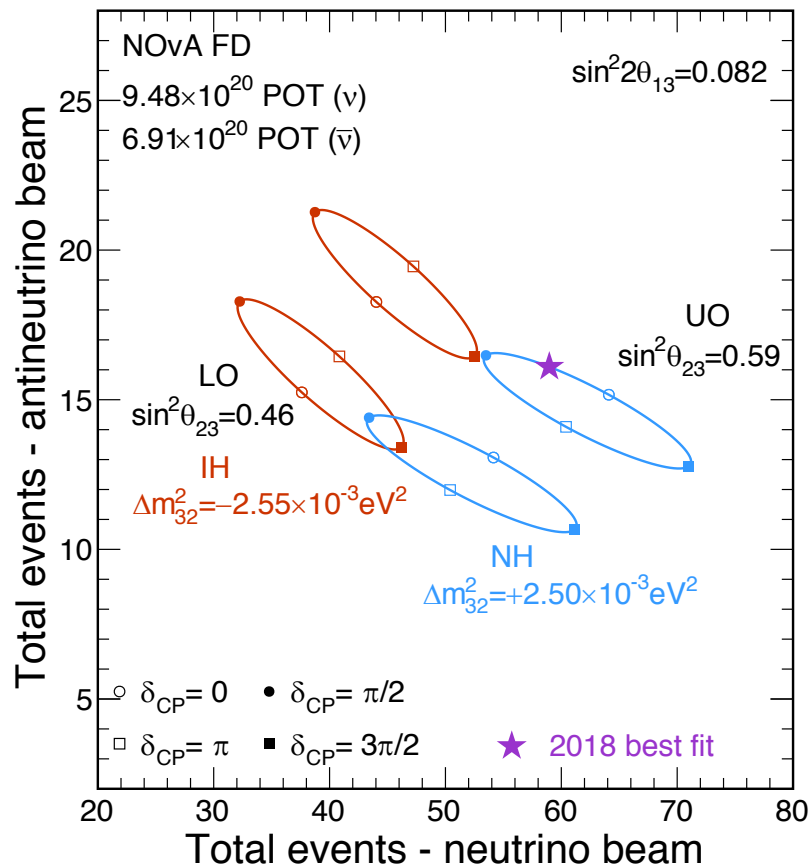
# Joint Appearance and Disappearance

$\nu_e/\bar{\nu}_e$  appearance event counts



Best fit from  $\nu_e/\bar{\nu}_e + \nu_\mu/\bar{\nu}_\mu$  combined analysis

NOvA Simulation



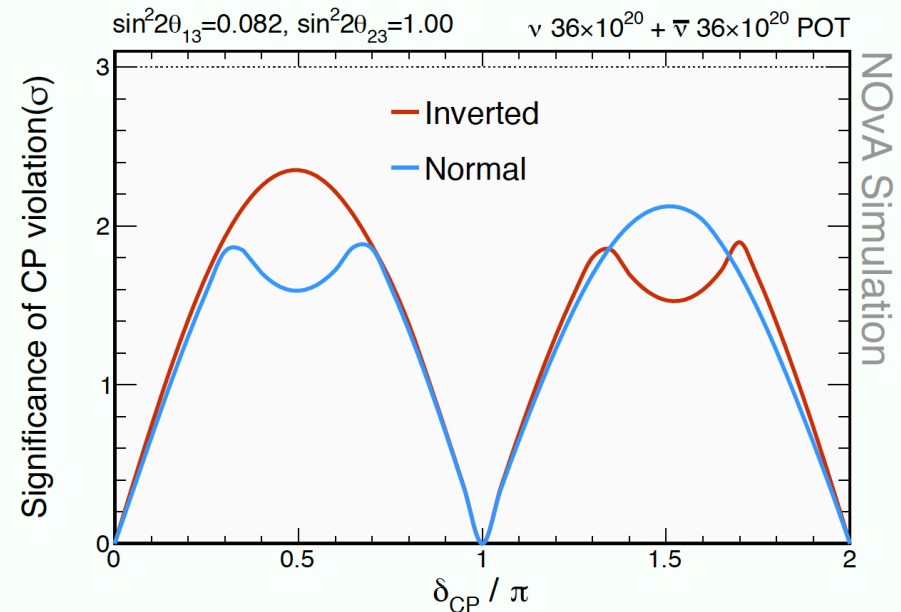
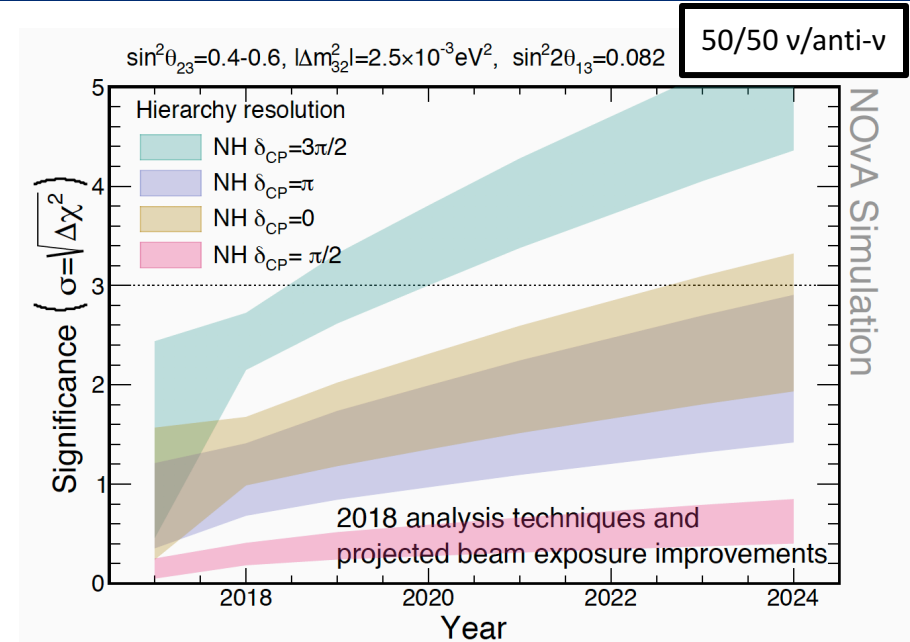
Error bars represent counting uncertainty of  $\nu_e/\bar{\nu}_e$  appearance, full power from joint fit to  $\nu_e/\bar{\nu}_e + \nu_\mu/\bar{\nu}_\mu$  energy/PID spectra

- Prefer non-maximal at  $1.8\sigma$ , favor upper octant
- Consistent with all  $\delta_{CP}$  values in NH at  $< 1.6\sigma$
- Exclude  $\delta = \pi/2$  in IH at  $> 3\sigma$
- **Prefer NH at  $1.8\sigma$**

# Looking Forward

- Taking antineutrino data since 2017, switch back to neutrinos in 2019, run 50% neutrino, 50% anti-neutrino
- Extended running through 2024, test beam program and potential accelerator improvement to enhance ultimate reach
- If  $\delta_{CP}=3\pi/2$ , 3  $\sigma$  sensitivity to MH by 2020,  $\sim 5 \sigma$  by 2024
- 3  $\sigma$  to MH for 30-50% (depending on octant) of  $\delta_{CP}$  range by 2024
- 2+  $\sigma$  to CP at  $\delta_{CP}=3\pi/2$  or  $\delta_{CP}=\pi/2$  by 2024

*Thank you!*

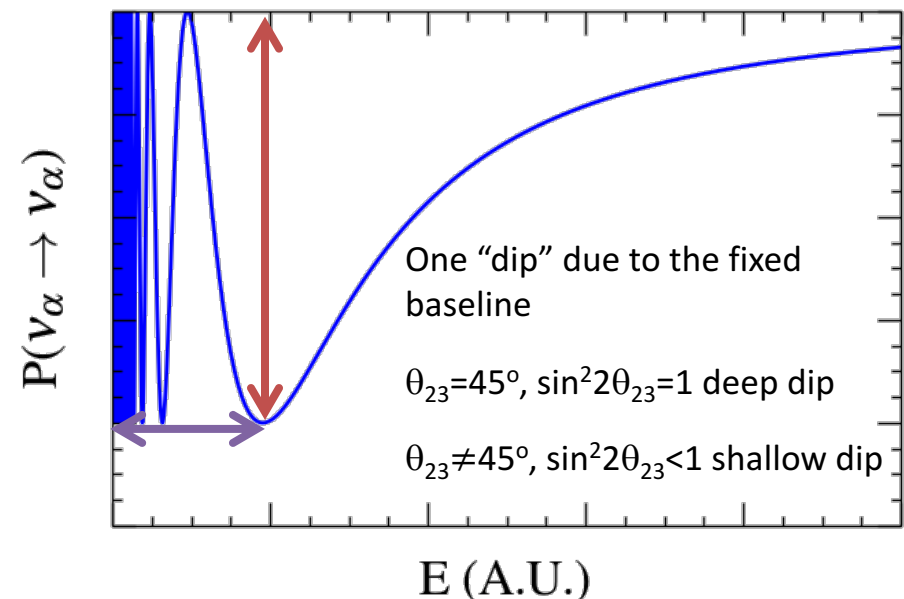
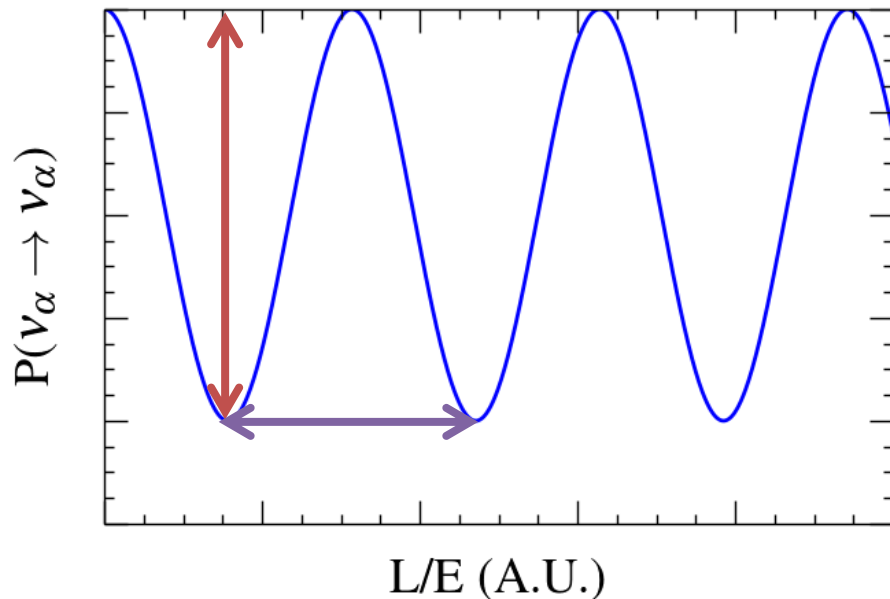


# *Backup*

# $\nu_\mu$ disappearance

$$P(\mu\mu) \approx 1 - \sin^2 2\theta_{23} \sin^2 \left( \frac{\Delta m_{32}^2 L}{4E} \right)$$

$\nu_\mu$  disappearance: High precision  $\Delta m_{32}$  and  $\sin^2 2\theta_{23}$ , constrain octant



Nunokawa, Parke, Valle, in "CP Violation and Neutrino Oscillations",  
 Prog.Part.Nucl.Phys. 60 (2008) 338-402.



# $\nu_e$ appearance

$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) \approx & \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2(A-1)\Delta}{(A-1)^2} \\
 & + 2\alpha \sin \theta_{13} \cos \delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{(A-1)} \cos \Delta \\
 & - 2\alpha \sin \theta_{13} \sin \delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{(A-1)} \sin \Delta
 \end{aligned}$$

$$\alpha = \frac{\Delta m_{21}^2}{\Delta m_{31}^2}$$

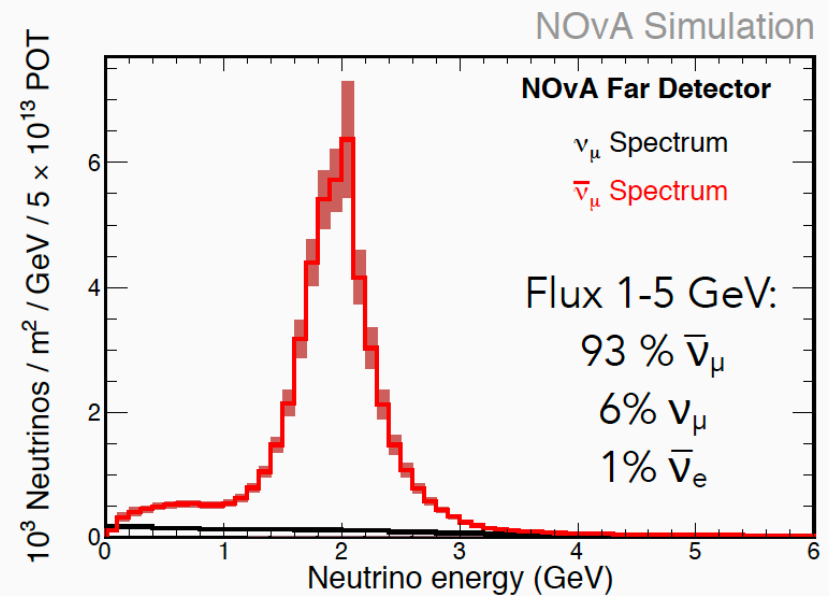
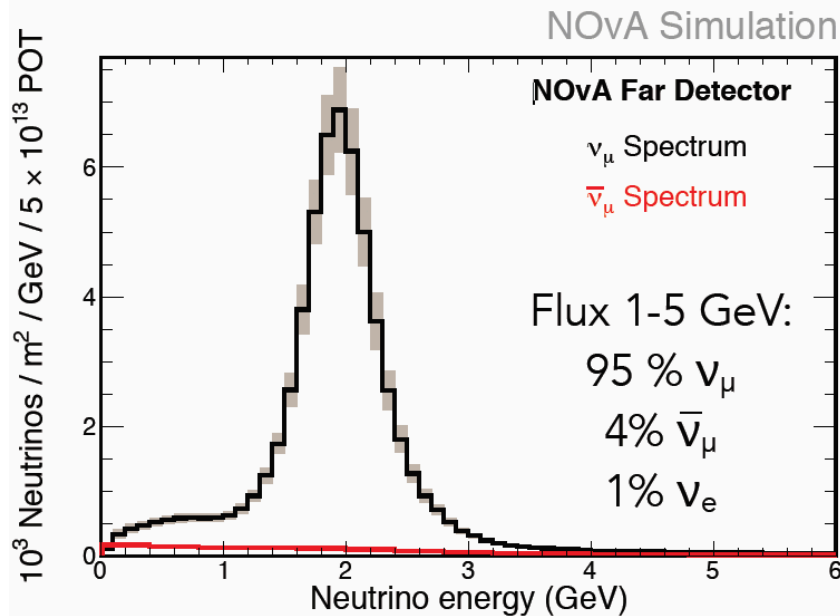
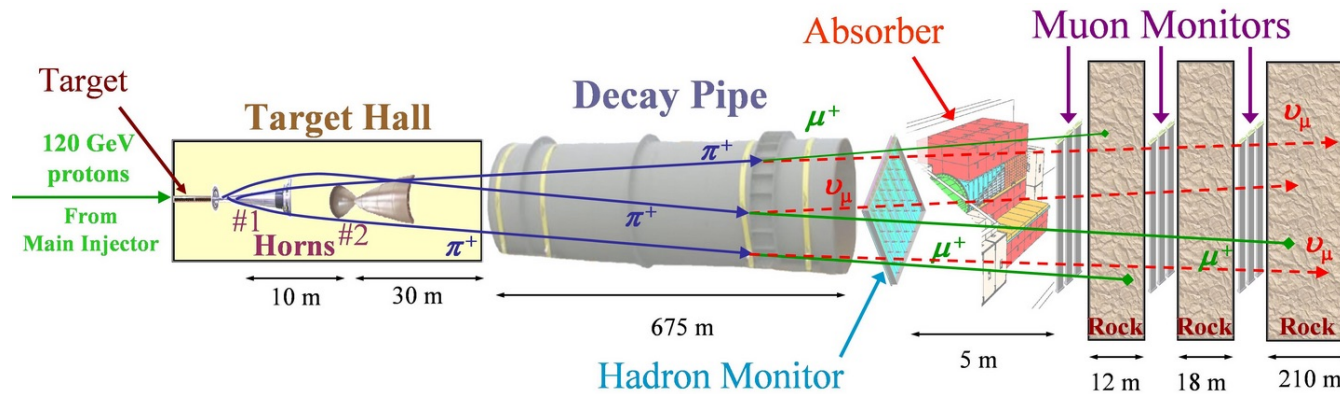
$$\Delta = \frac{\Delta m_{31}^2 L}{4E}$$

Matter effect

$$A = +G_f N_e \frac{L}{\sqrt{2}\Delta}$$

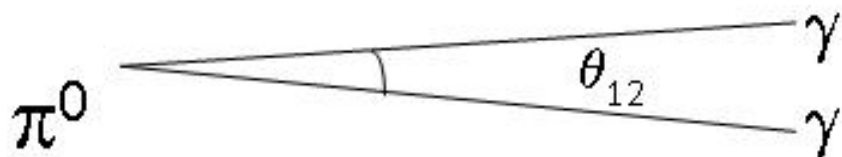
- Measuring mass hierarchy (sign of  $\Delta$  value),  $\delta_{CP}$  and octant of  $\theta_{23}$  with  $\nu_e$  appearance,
- $P(\nu_\mu \rightarrow \nu_e)$  difference between  $\Delta > 0$  and  $\Delta < 0$  enlarged by matter effect  $A$  ( $\propto L$  when fix  $L/E$  to oscillation maximum)

# NuMI Off-Axis $\nu_e$ Appearance Experiment



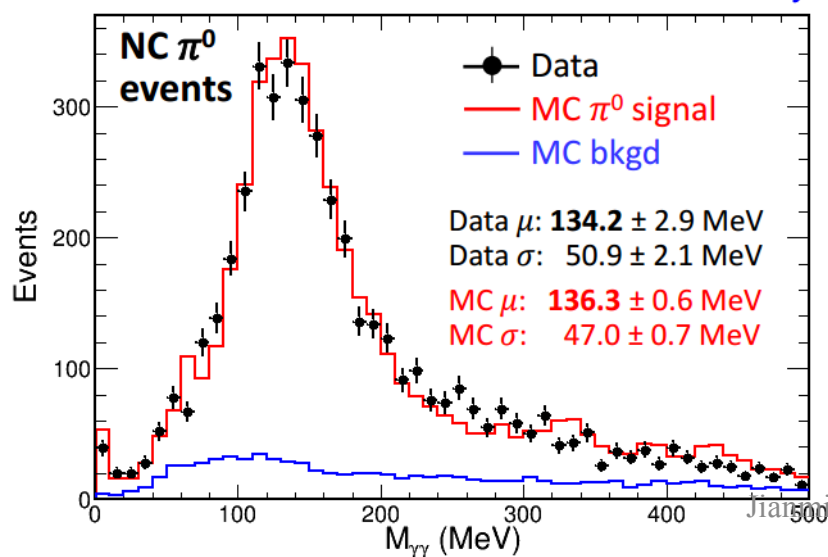
# Systematic Error in Calibration

- Our calibration is built on dE/dx from stopping cosmic muons.
- Control samples for calibration uncertainty
  - $\pi^0$  mass peak in ND
  - Michel electrons in ND and FD

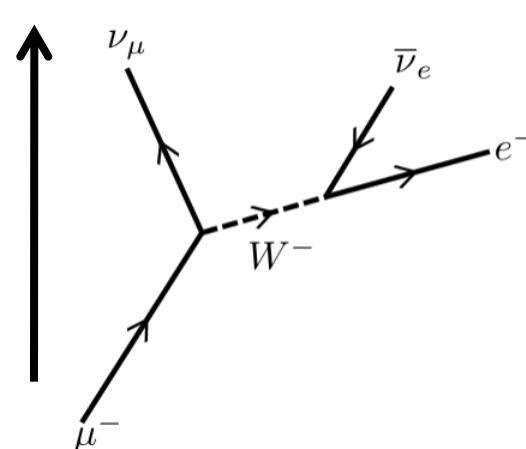


$$m_{\pi^0}^2 = 2E_{\gamma 1} E_{\gamma 2} (1 - \cos \theta_{12})$$

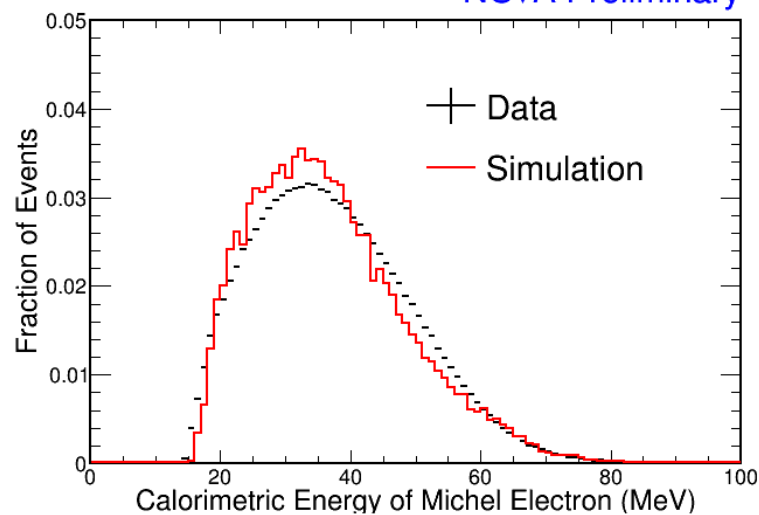
NOvA Preliminary



Michel electrons  
from muon decays

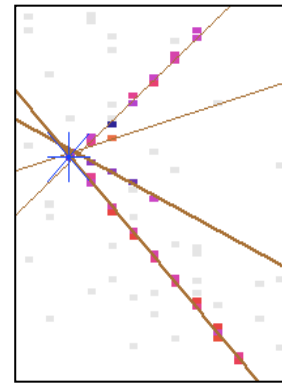
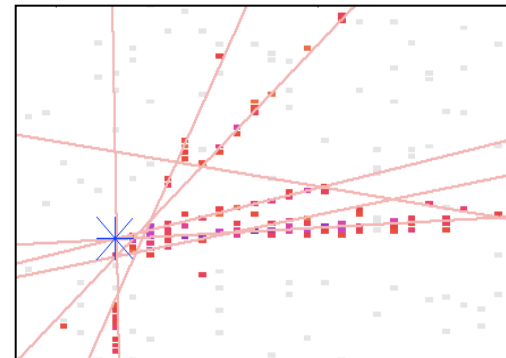
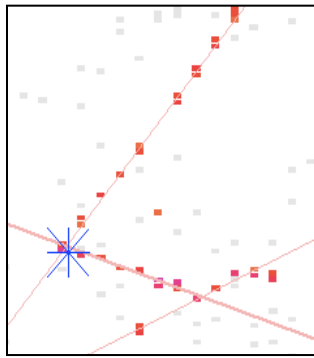


NOvA Preliminary

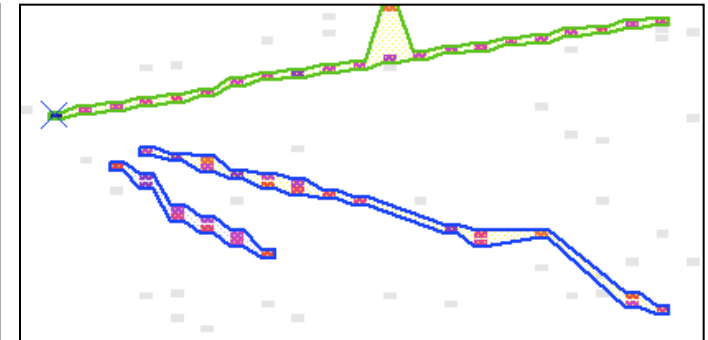
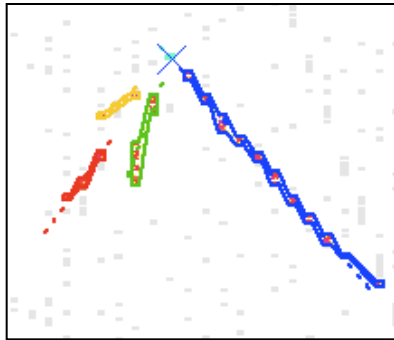


# Prong/track Reconstruction

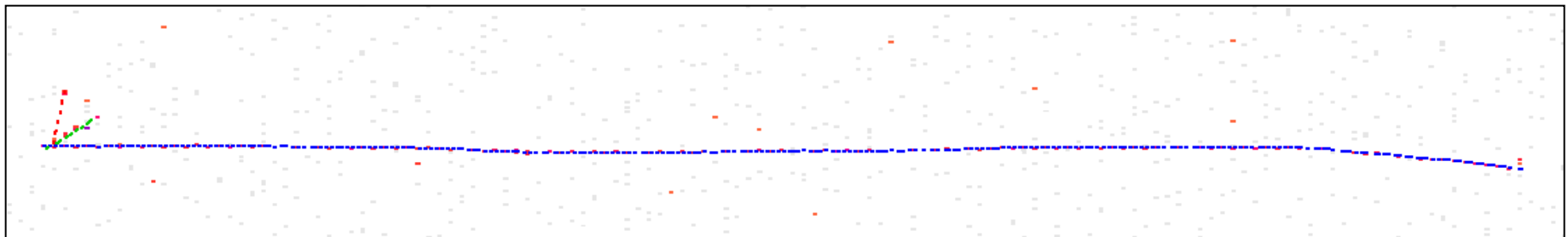
**Vertexing:** Find lines of energy depositions with Hough transform. Then determine the vertex that all lines converge to



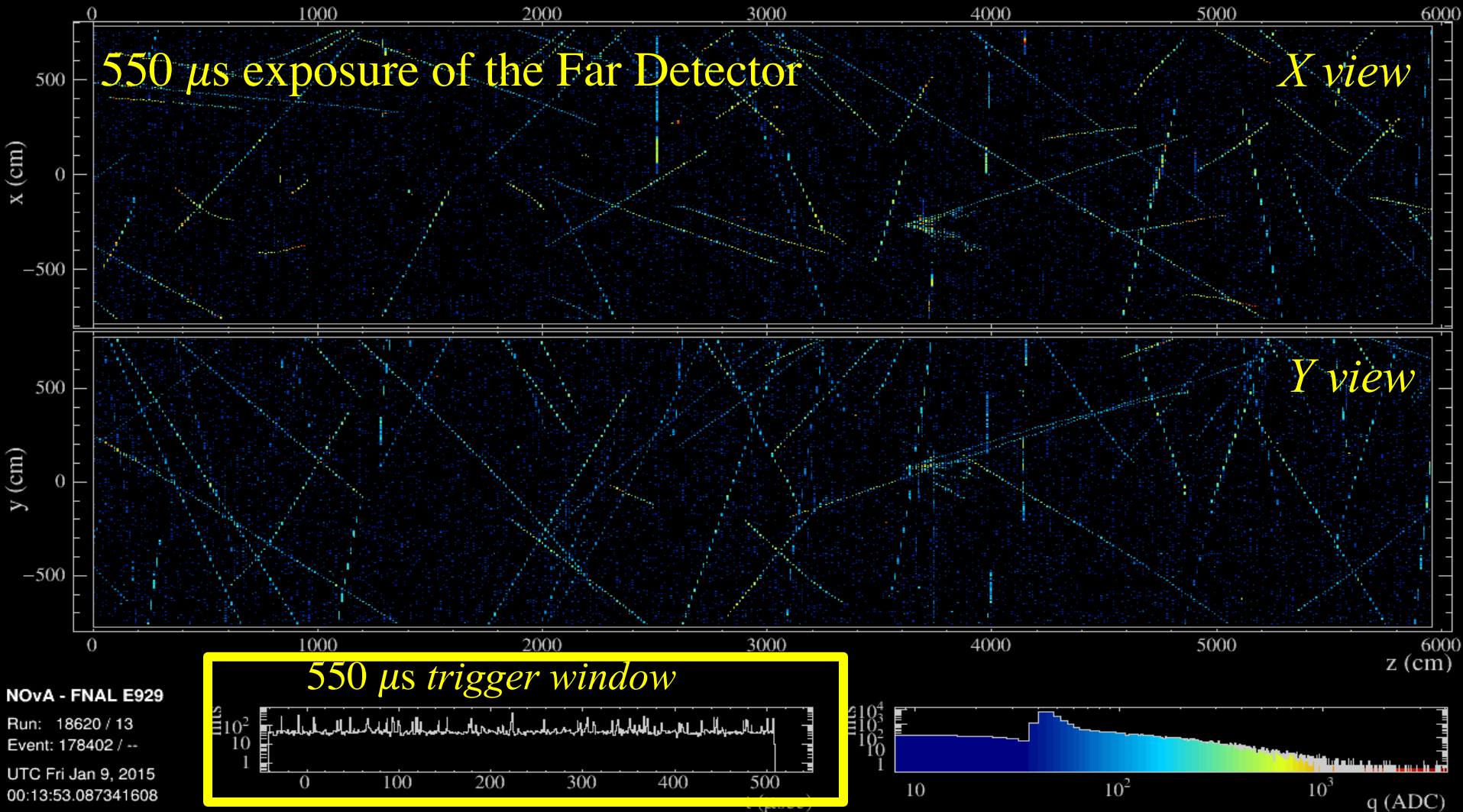
**Shower Clustering:** Based on the vertex and the lines, showers are reconstructed by angular clustering



**Tracking:** Trace particle trajectories with **Kalman filter** tracker (below). Also have a **cosmic ray tracker** that reconstructs cosmic tracks with high speed.



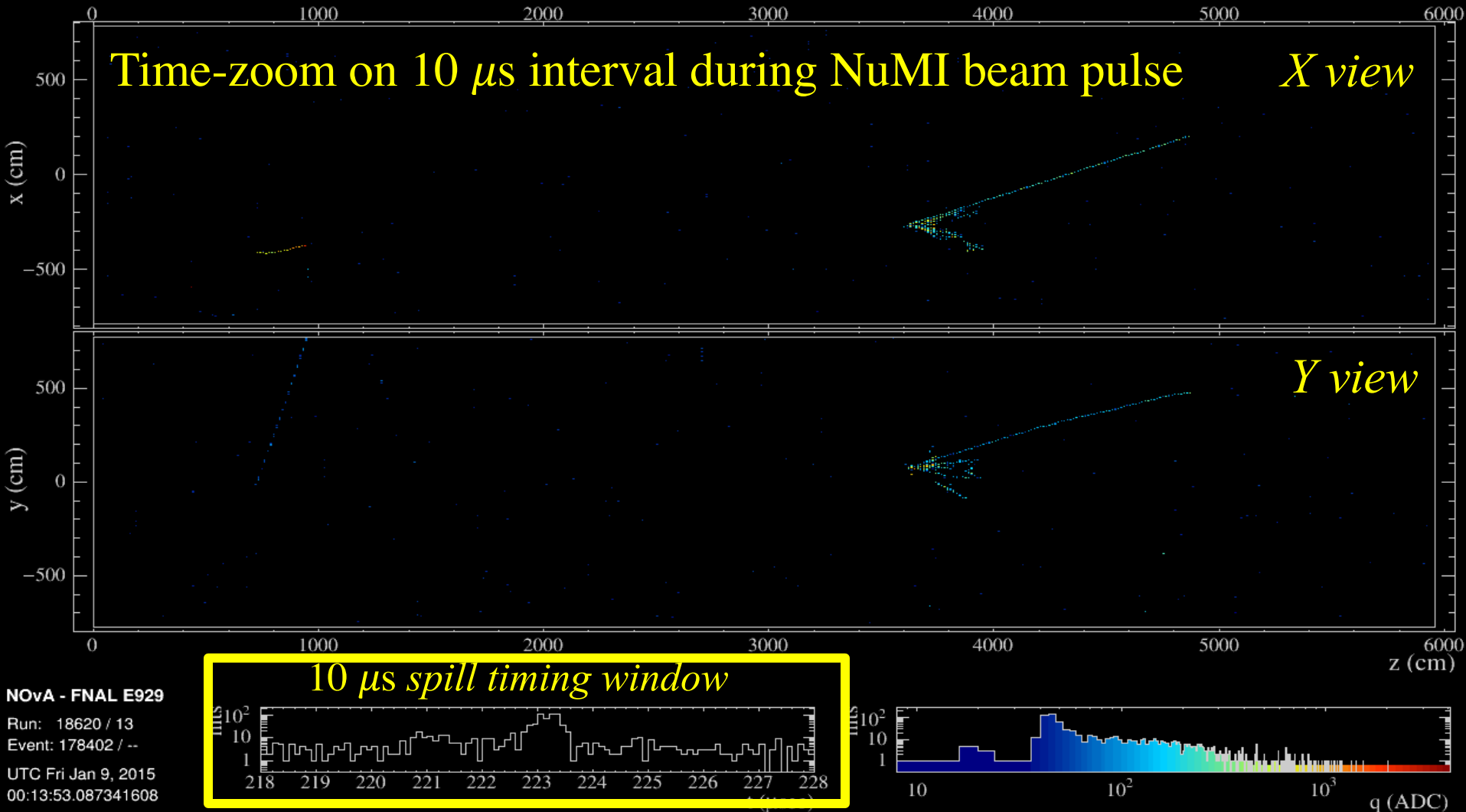
# Event clustering



- Because NOvA is on surface, hits in a trigger window are a combination of cosmic and beam events.
- First step in reconstruction is to cluster hits by space-time coincidence to separate neutrino hits and cosmic hits.

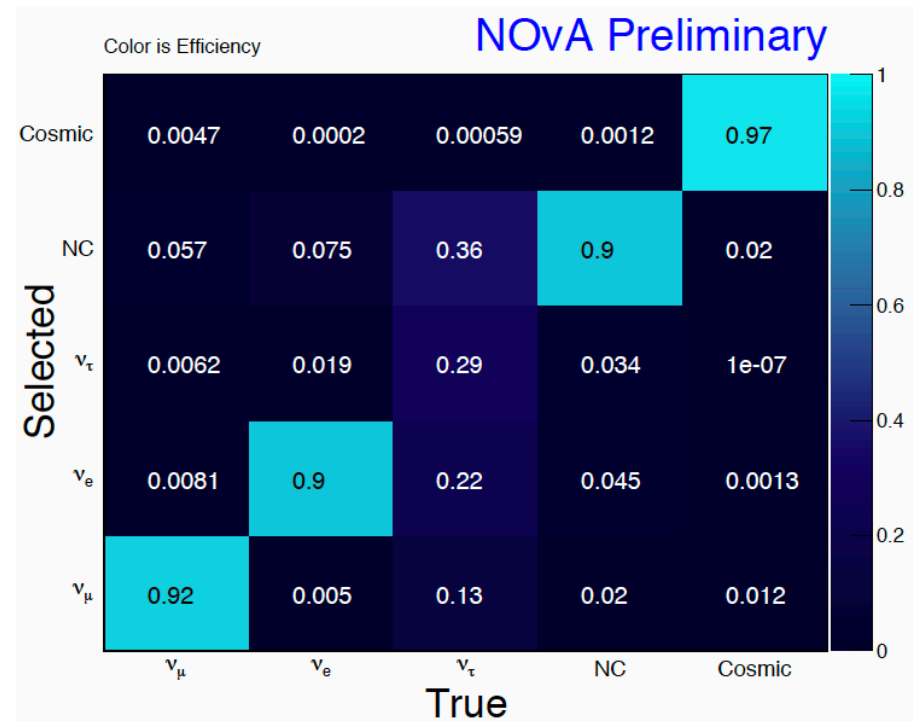
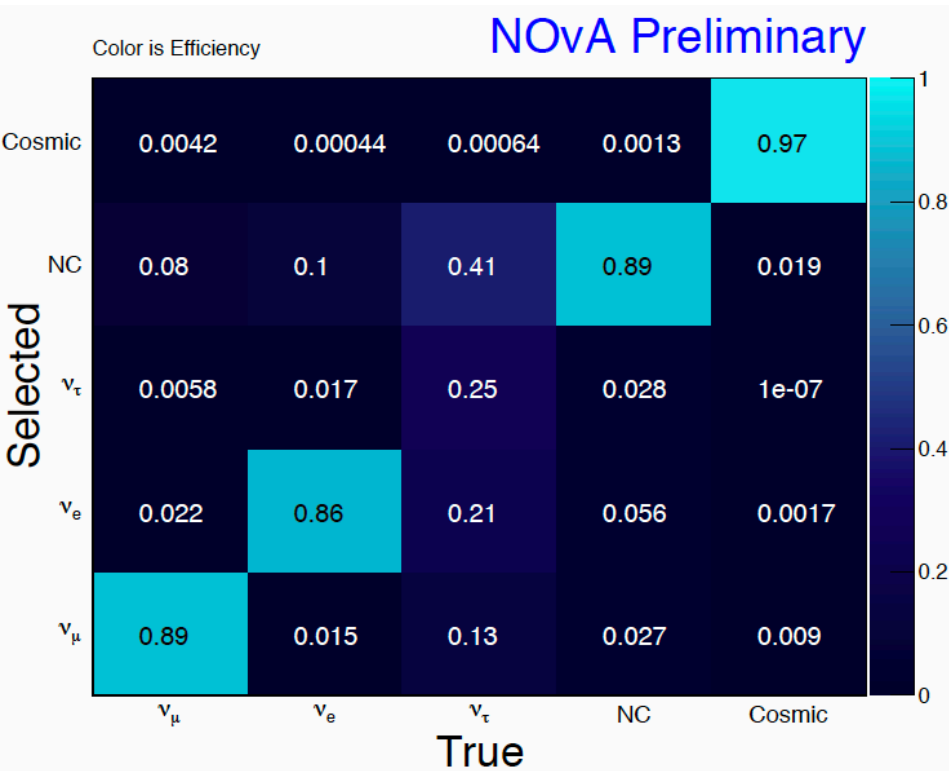


# Event clustering



Event clusters that contain neutrino interactions can be correctly selected in the neutrino spill timing window

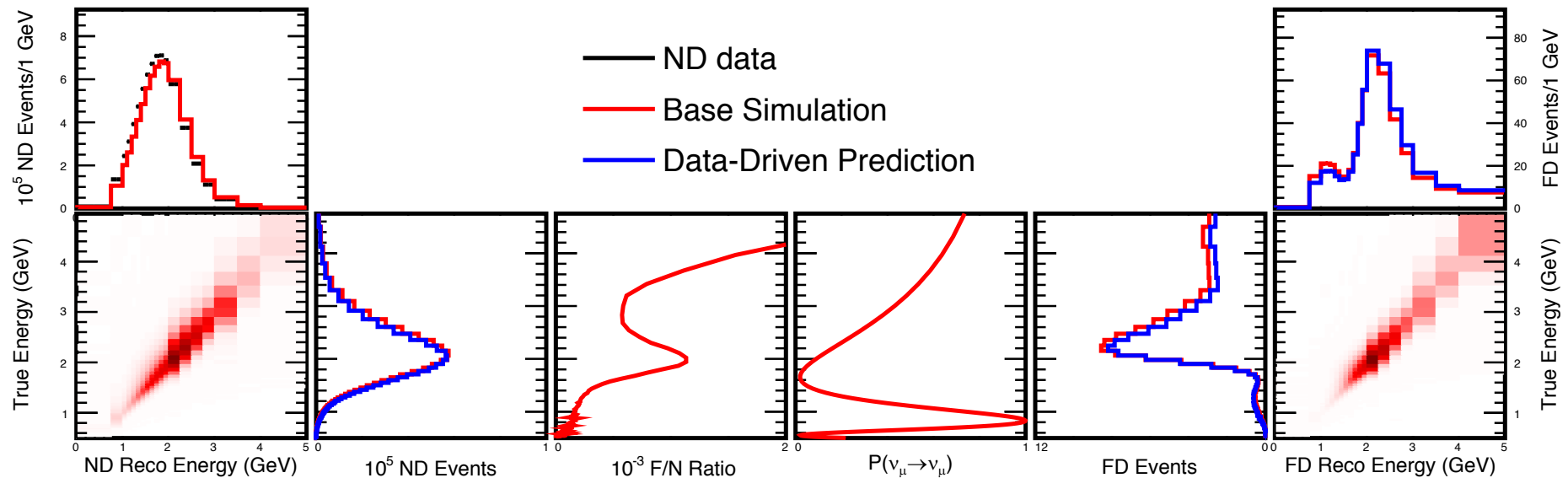
# PID efficiencies



# Analysis Strategy

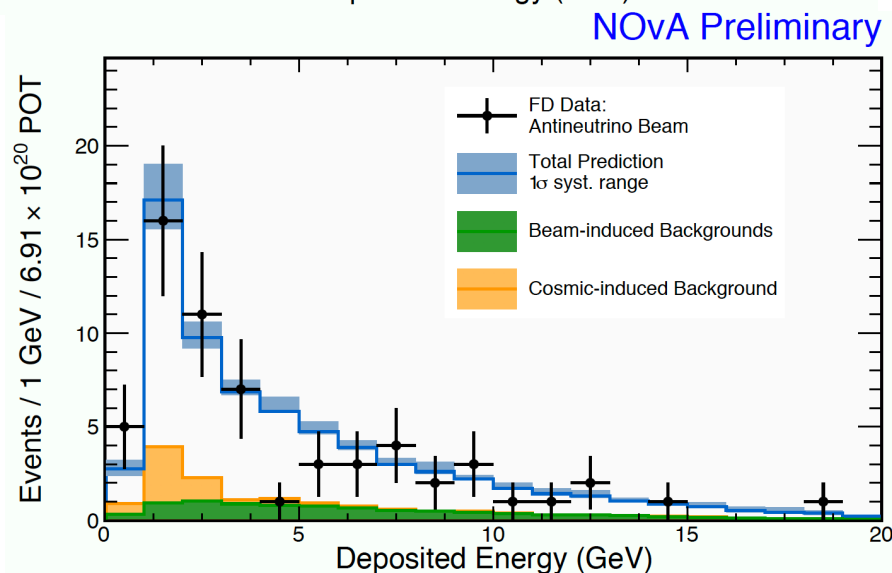
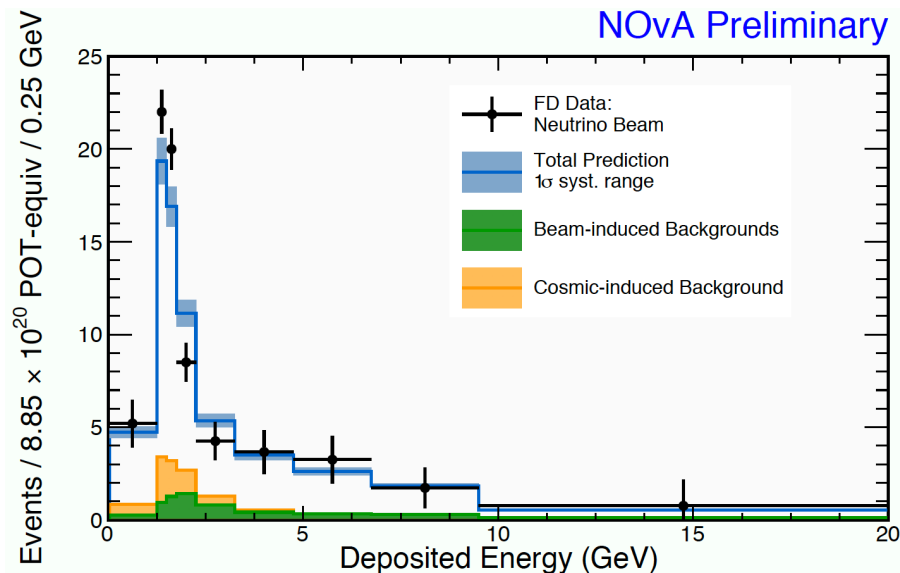
- Separate  $\nu_\mu/\nu_e$ /NC signal from beam backgrounds
- Extrapolate observed ND spectrum to FD, reject cosmic rays in FD, make FD unoscillated prediction
- Measure shapes and yields of signal events in energy/PID bins in FD to determine oscillation parameters

ND $\rightarrow$ FD extrapolation for  $\nu_\mu$  disappearance



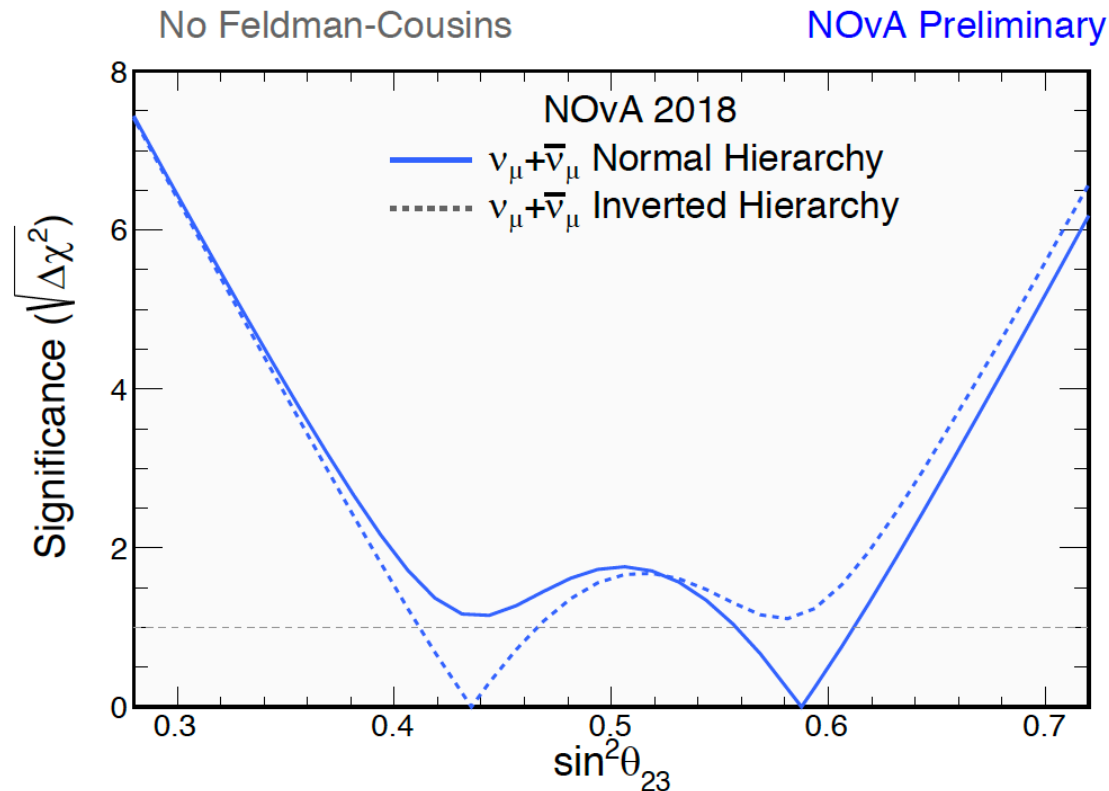
# Observed NC events in Far Detector

- FD selection:
  - NC CVN selection applied
  - Additional Deep-learning based cosmic rejection
- Neutrino beam:
  - Observe 201 events, predict  $188 \pm 13$  (syst.) events (38 bkg.)
- Antineutrino beam:
  - Observe 61 events, predict  $69 \pm 8$  (syst.) events (16 bkg.)
- No significant suppression for NC observed, consistent with 3-flavor oscillation



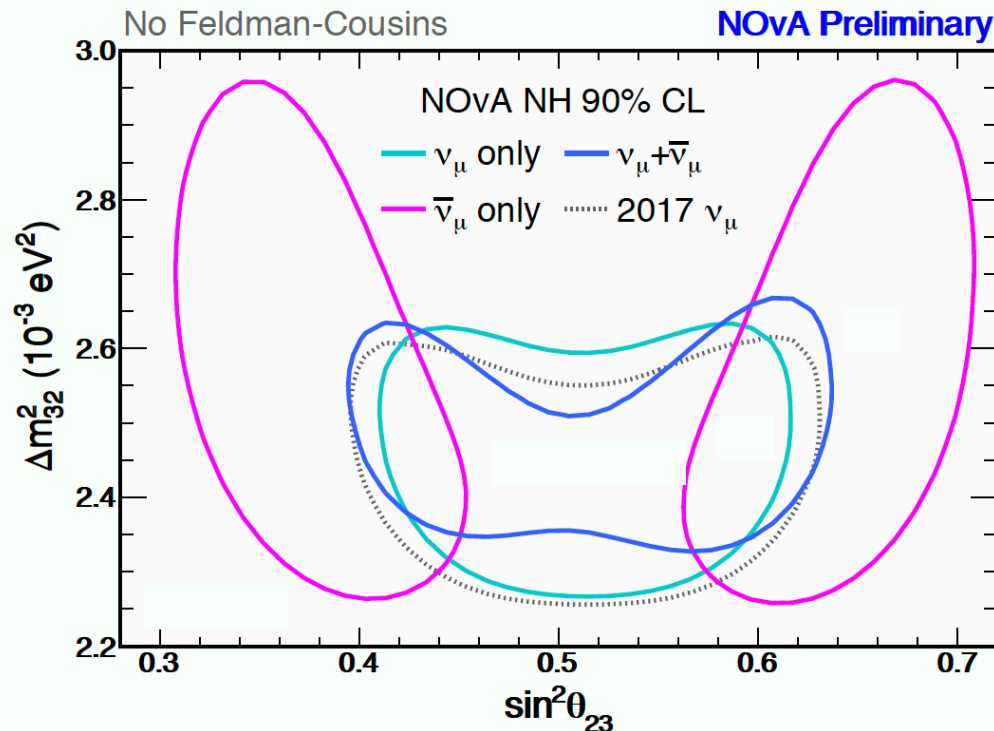
# $\nu_\mu$ appearance fit

- Combined data of neutrino and antineutrino beams fitted assuming CPT invariance
- If fit separately,  $\bar{\nu}_\mu$  data prefers non-maximal while  $\nu_\mu$  prefers maximal
- $\chi^2$ s consistent with combined fit oscillation parameters with  $p > 4\%$
- Matter effects introduce small asymmetry in the point of maximal disappearance,  $\sim 1\sigma$  prefers Upper (Lower) Octant in NH (IH)



# $\nu_\mu$ appearance fit

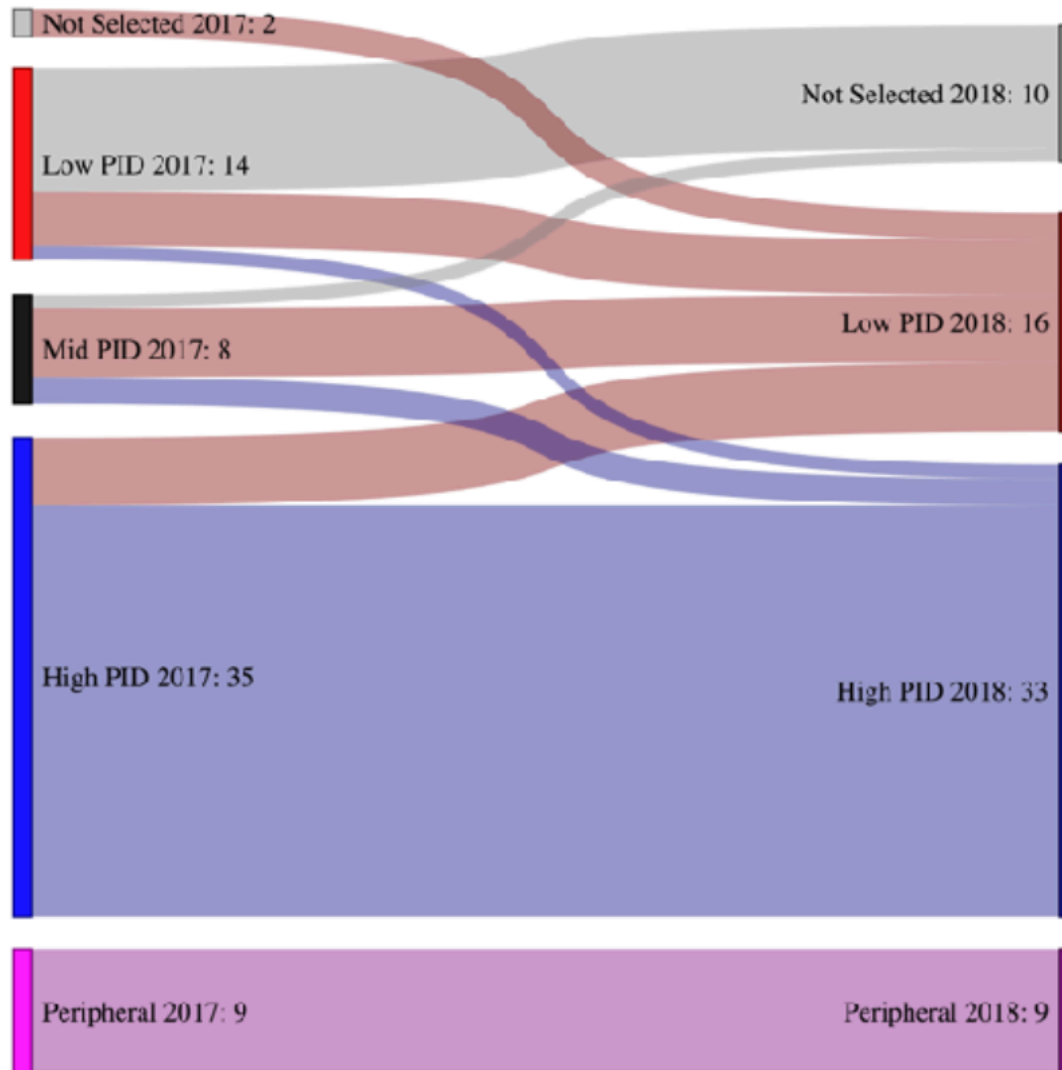
- Combined data of neutrino and antineutrino beams fitted assuming CPT invariance
- If fit separately,  $\bar{\nu}_\mu$  data prefers non-maximal while  $\nu_\mu$  prefers maximal
- $\chi^2$ s consistent with combined fit oscillation parameters with  $p > 4\%$
- Matter effects introduce small asymmetry in the point of maximal disappearance





# 2017/2018 RHC $\nu_e$ FD Data

66 FD data  
events in 2017  
analysis



58 FD data  
events in 2017  
analysis

Change in data events after retraining of PID, new training improved bkg rejection

# Systematic Uncertainties (Joint Fit)

Source of Uncertainty	$\sin^2\theta_{23} (\times 10^{-3})$	$\delta_{CP}/\pi$	$\Delta m_{32}^2 (\times 10^{-3} \text{ eV}^2)$
Beam Flux	+0.42 / -0.48	+0.0088 / -0.0048	+0.0016 / -0.0015
Detector Calibration	+6.9 / -6.1	+0.15 / -0.023	+0.024 / -0.029
Detector Response	+1.9 / -0.99	+0.055 / -0.054	+0.0027 / -0.0034
Muon Energy Scale	+2.6 / -2.1	+0.015 / -0.0026	+0.01 / -0.012
Near-Far Differences	+0.56 / -1.1	+0.11 / -0.064	+0.0033 / -0.0013
Neutrino Cross Sections	+4.2 / -3.5	+0.085 / -0.072	+0.015 / -0.014
Neutron Uncertainty	+6.4 / -7.9	+0.002 / -0.0052	+0.0028 / -0.01
Normalization	+1.4 / -1.5	+0.031 / -0.024	+0.0029 / -0.0027
Systematic Uncertainty	+9.6 / -11	+0.21 / -0.11	+0.032 / -0.035
Statistical Uncertainty	+22 / -29	+0.9 / -0.27	+0.064 / -0.059