

nEXO Double Beta Decay

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NATIONAL
ACCELERATOR
LABORATORY

5/27/16

HQL2016 - Brian Mong

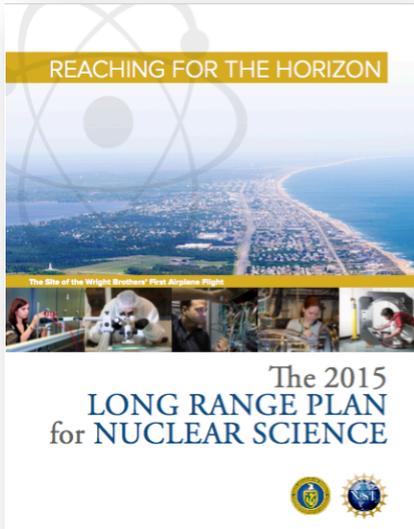


Neutrinos



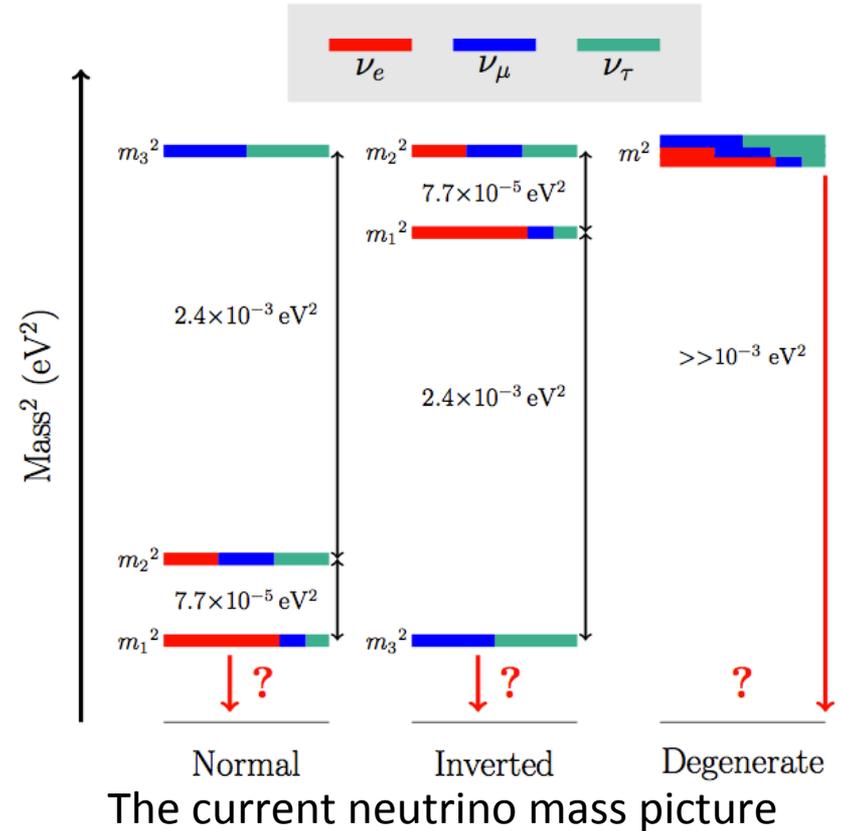
Neutrino oscillation experiments have demonstrated neutrinos have mass

- Matter/antimatter asymmetry and neutrino masses still a problem for SM
- Majorana neutrinos seem *likely* way to solve “neutrino problem” in SM

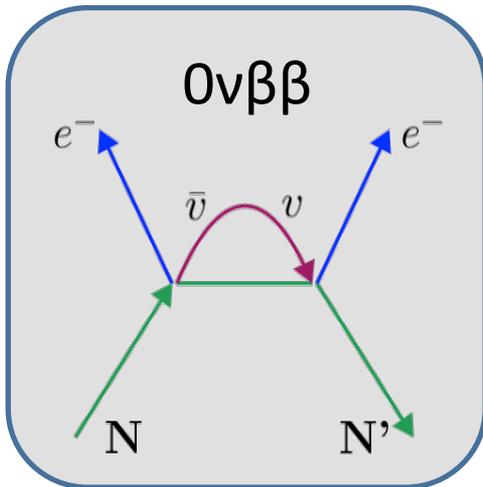
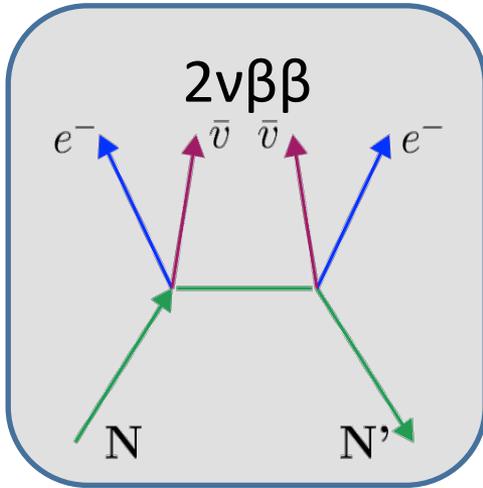


$0\nu\beta\beta$ most sensitive probe for Majorana nature of neutrinos

NSAC long range plan: US-led ton scale $0\nu\beta\beta$ experiment a priority



Neutrinoless double beta decay



- $0\nu\beta\beta$ decays require
 - Majorana nature of neutrino
 - Lepton number violation
- The half-life of $0\nu\beta\beta$ can be used to determine neutrino mass:

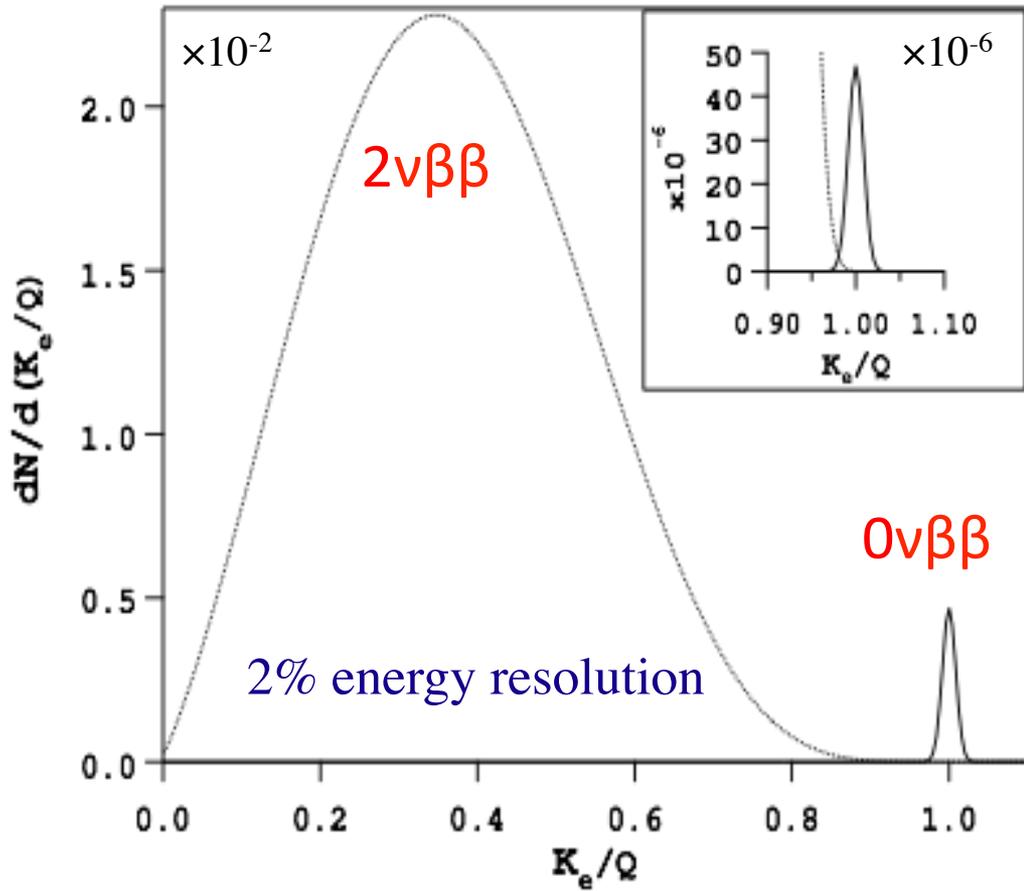
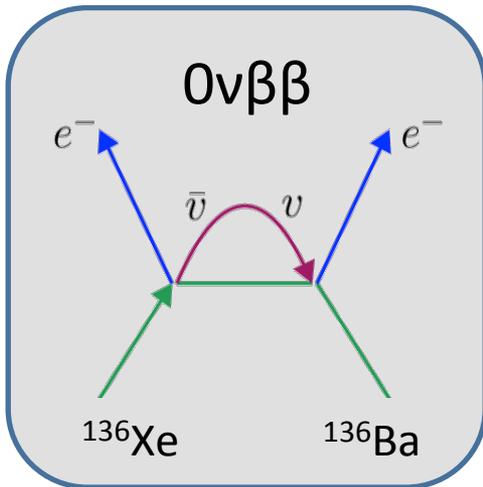
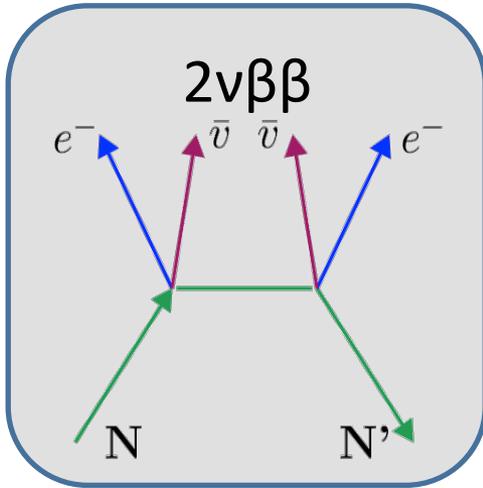
$$\langle m_\nu \rangle = \left(\overset{\text{Measure}}{T_{1/2}^{0\nu\beta\beta}} \overset{\text{Known}}{G^{0\nu\beta\beta}} (E_0, Z) \left| \overset{\text{Calculated (theory)}}{M_{GT}^{0\nu\beta\beta}} - \frac{g_V^2}{g_A^2} \overset{\text{Calculated (theory)}}{M_F^{0\nu\beta\beta}} \right|^2 \right)^{-1/2}$$

$$\langle m_\nu \rangle = \sum_{i=1}^3 |U_{\alpha i}|^2 m_i$$

$$\text{mass} \propto \frac{1}{\sqrt{T_{1/2}^{0\nu\beta\beta}}}$$

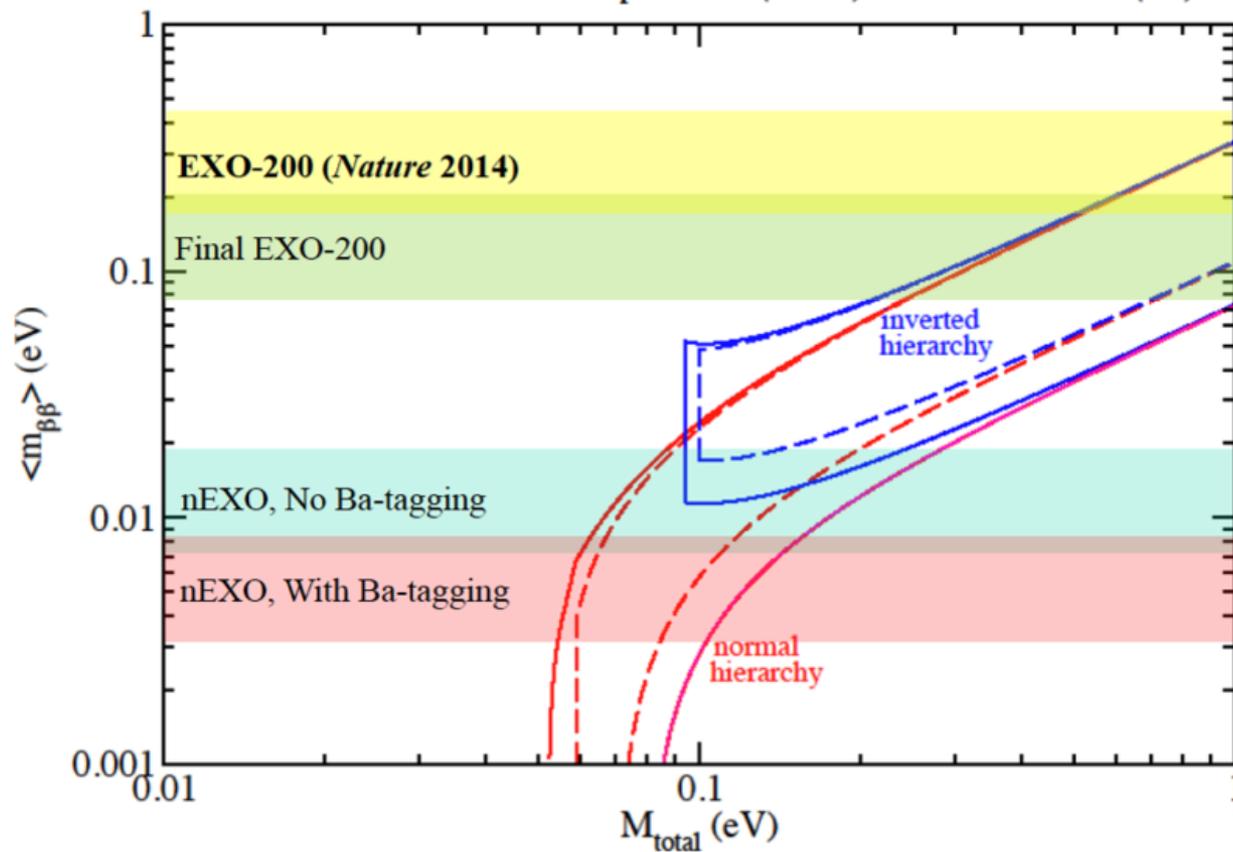
Neutrinoless double beta decay

[P. Vogel, arXiv:hep-ph/0611243]

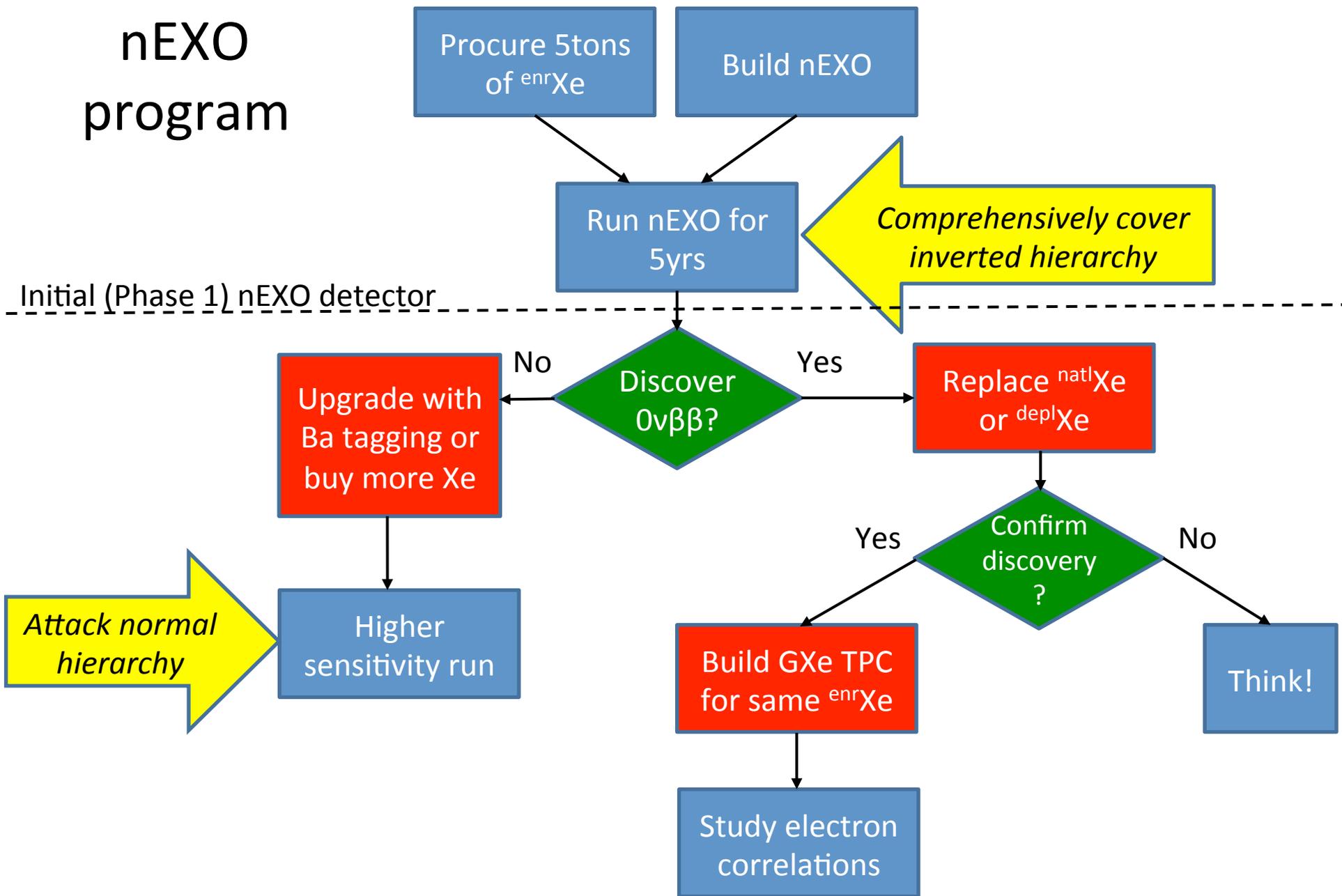


$$^{136}\text{Xe } Q_{\beta\beta} = 2457 \text{ keV}$$

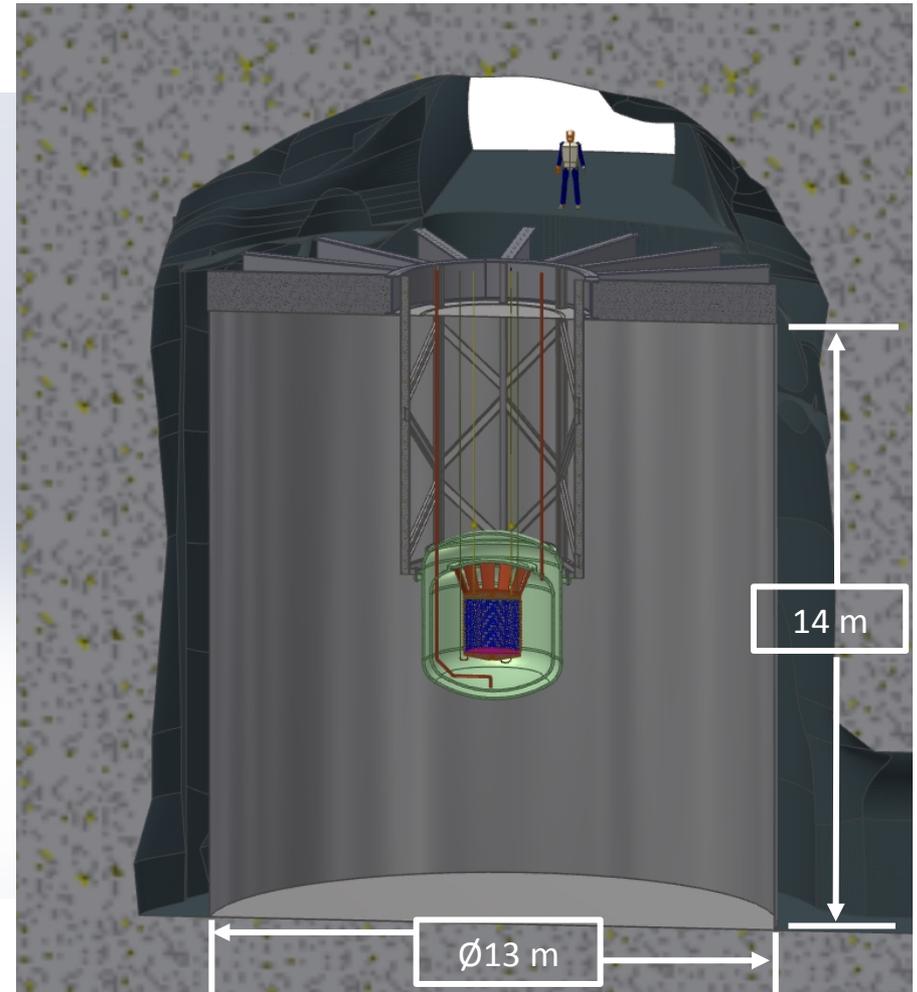
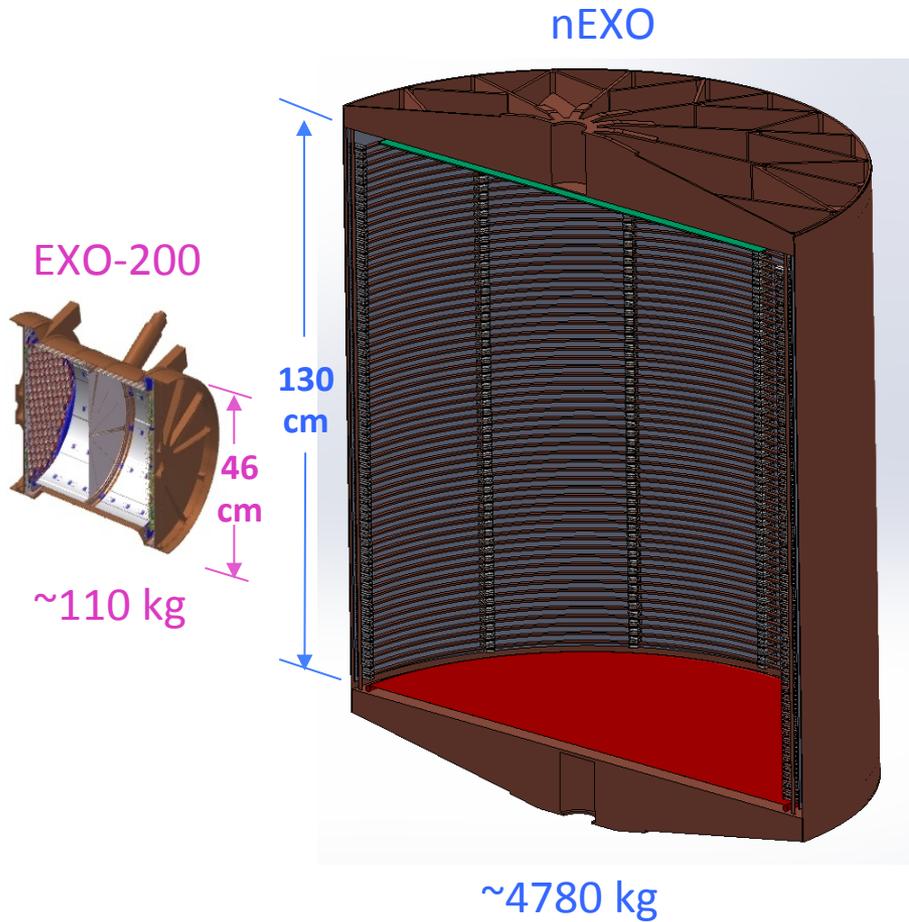
EXO-200 & nEXO projected sensitivities



nEXO program



nEXO – 5 ton detector



Building on EXO-200 successes

- EXO-200 started taking data in May 2011
 - “Observation of Two-Neutrino Double-Beta Decay in ^{136}Xe ...”, PRL (2011)
 - “Search for Neutrinoless Double-Beta Decay in ^{136}Xe with EXO-200”, PRL (2012)
 - “An improved measurement of the $2\nu\beta\beta$ half-life of Xe-136 with EXO-200”, PRC (2013)
 - “Search for Majorana neutrinos with the first two years of EXO-200...”, Nature (2014)
- EXO-200 demonstrated the power of the combination of *ionization density*, *energy resolution*, *event multiplicity*, and *event location*
- EXO-200 validated the choice of material and the model that was used to predict the detector background

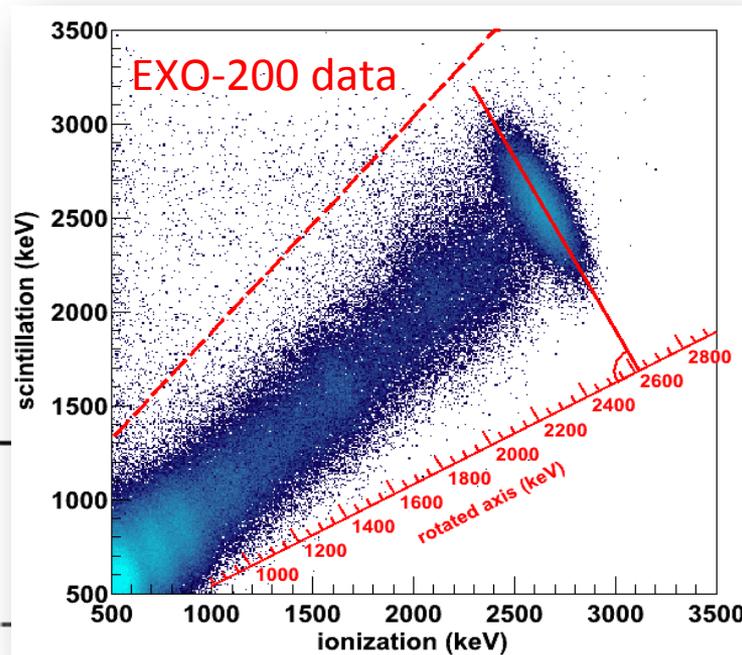
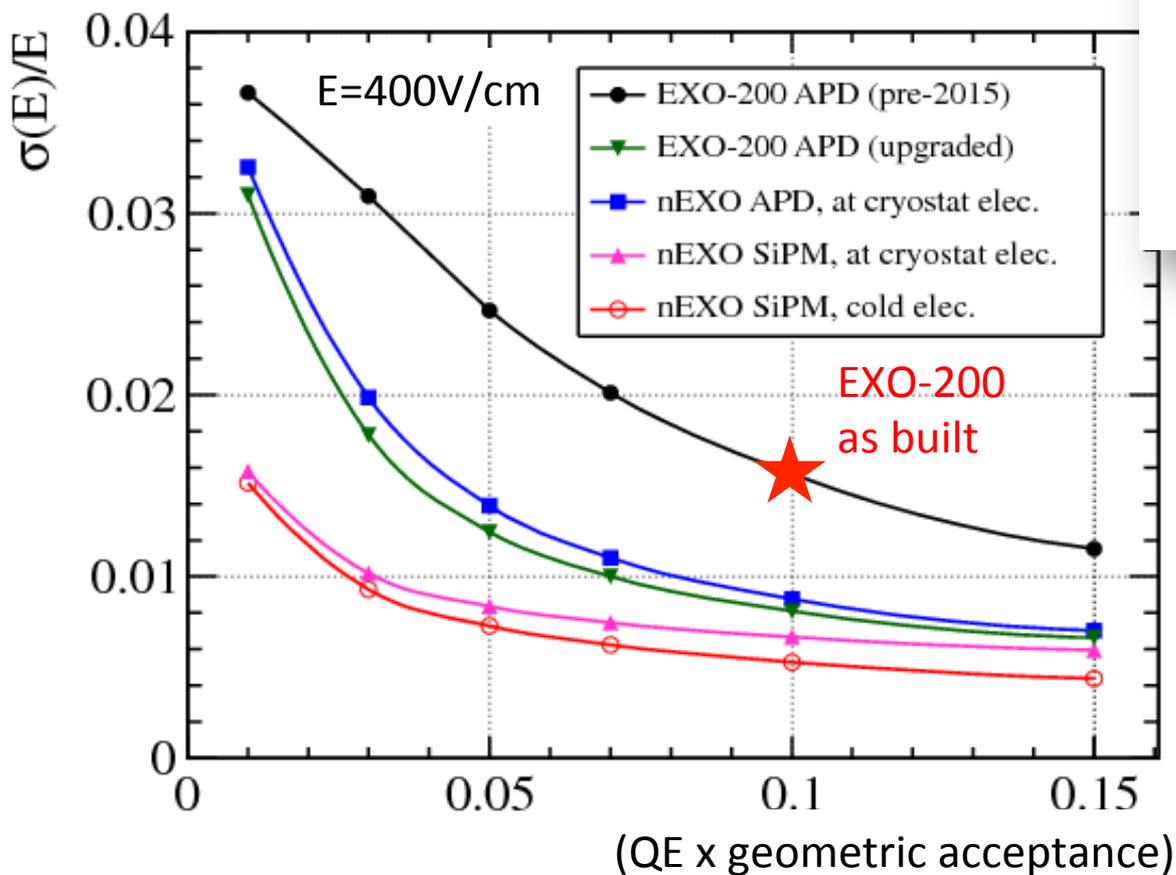
| Events in $\pm 2\sigma$ around Q | Radioactive bkgd prediction during construction | Radioactive bkgd prediction using present Monte Carlo | ^{137}Xe bkgd | Background from 0 ν analysis fit |
|----------------------------------|---|---|------------------------|--|
| 90%CL Upper | 48 | 22 | 7 | $31.1 \pm 1.8 \pm 3.3$ (39 events observed) |
| 90%CL Lower | 9.4 | 3.3 | | |

Optimization from EXO-200 to the nEXO scale

| What | Why |
|-------------------------------|---|
| ~30x volume/mass | To give sensitivity to the inverted hierarchy |
| No cathode in the middle | Larger low background volume/no ^{214}Bi in the middle |
| 6x HV for the same field | Larger detector and one drift cell |
| >3x electron lifetime | Larger detector and one drift cell |
| Better photodetector coverage | Energy resolution |
| SiPM instead of APDs | Higher gain, lower bias, lighter, E resolution |
| In LXe electronics | Lower noise, more stable, fewer cables/feedthroughs, E resolution, lower threshold for Compton ID |
| Lower outgassing components | Longer electron lifetime |
| Different calibration methods | Very “deep” detector (by design) |
| Deeper site | Less cosmogenic activation |
| Larger vessels | 5 ton detector and more shielding |

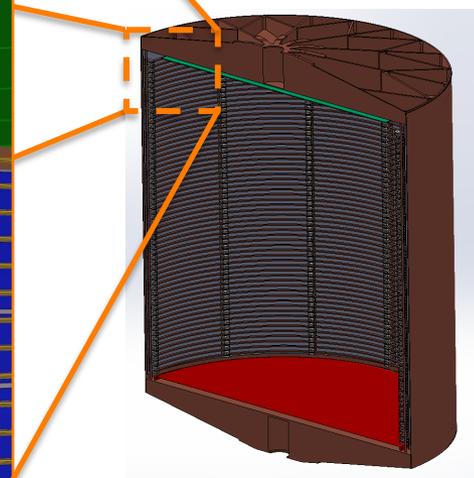
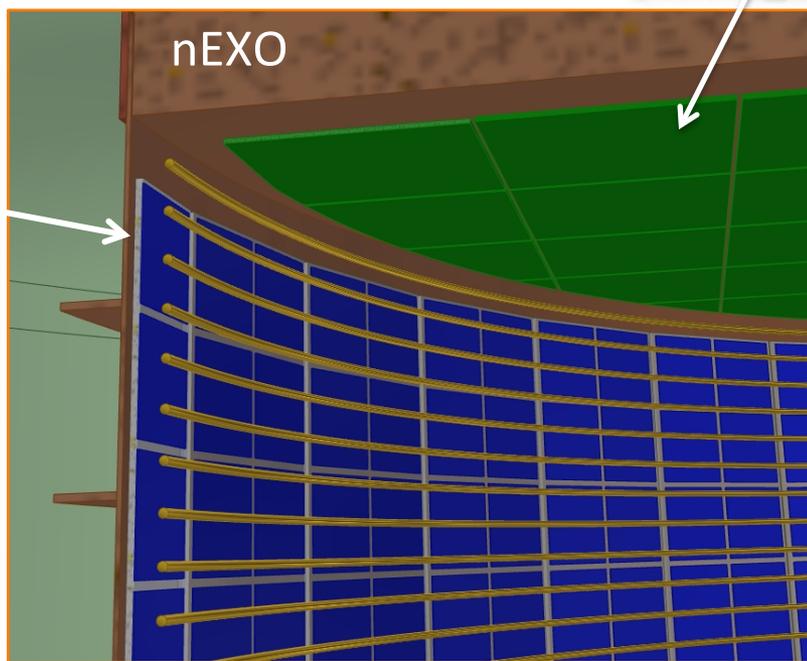
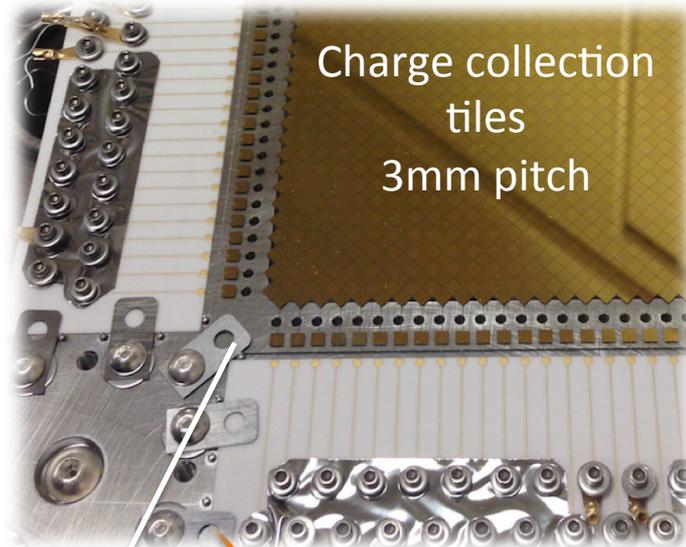
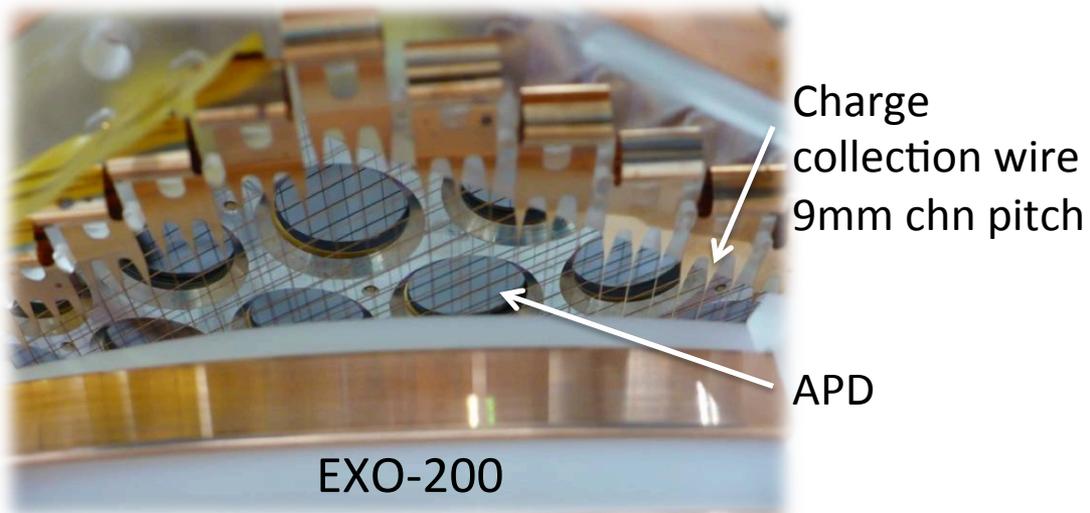
Energy resolution

Good energy resolution requires efficient readout of the scintillation and ionization signal



While we always assume 1% resolution for nEXO, a figure close to 0.5% is not out of the question

Energy Resolution: scintillation and charge collection R&D

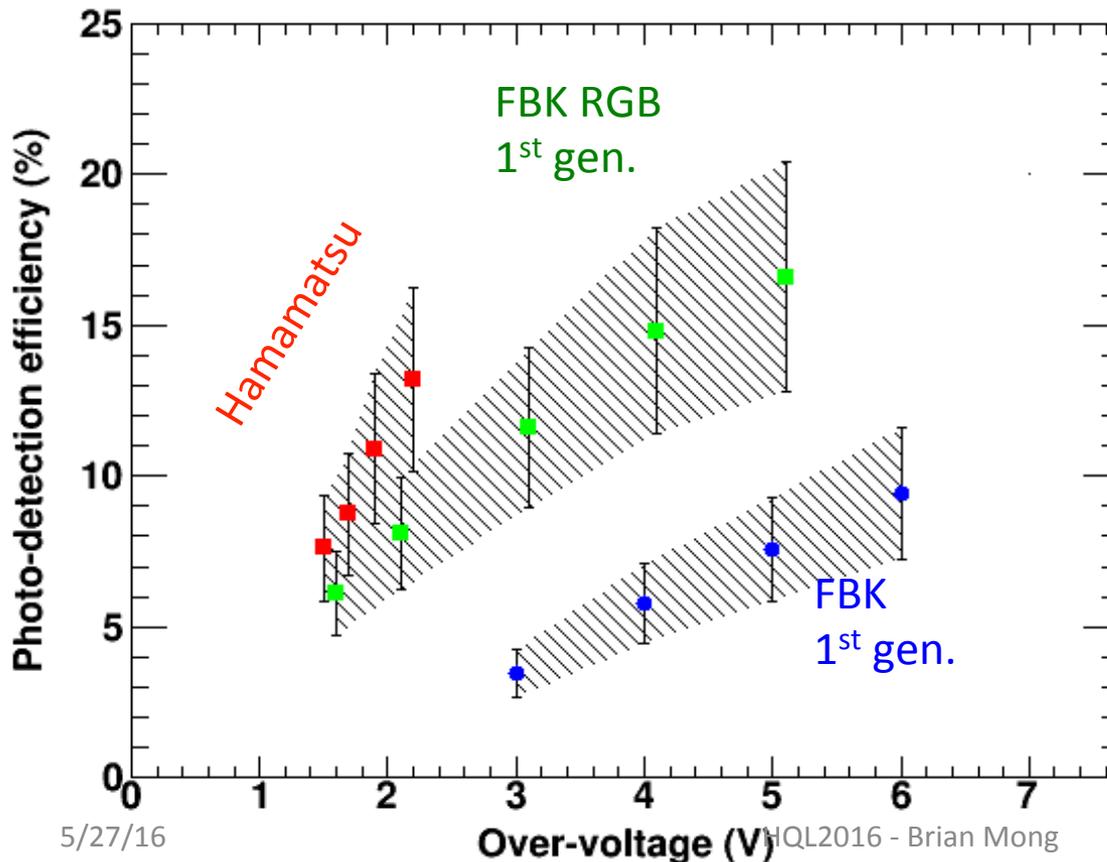


Energy Resolution: SiPM Photodetectors

Hamamatsu produces devices with QE= ~12% @ 175nm but until now they have refused to sell them un-encapsulated (hence are too radioactive)

First nEXO-specific run at FBK (Italy) provided ~10% QE
New “RGB” devices reach 15% QE with 7.7x7.7mm².

[I.Ostrovskiy et al. IEEE TNS 62 (2015) 1.]
DOI:[10.1109/TNS.2015.2453932](https://doi.org/10.1109/TNS.2015.2453932)



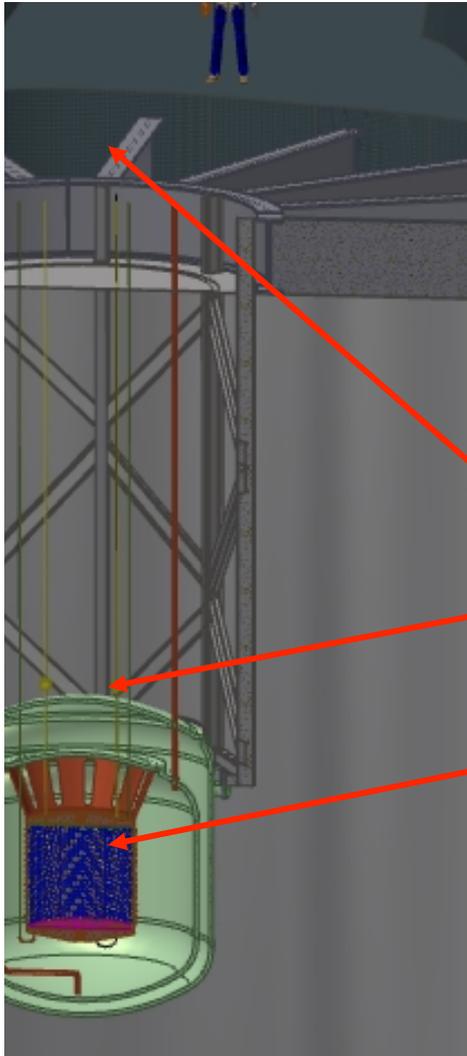
A new run at FBK will use a new technology (NUV-HD) with the promise to reach QE>15% @170nm, lower dark noise and crosstalk with 1x1cm² devices.

FBK devices are almost low activity enough (a Th anomaly needs to be investigated)

Energy resolution: cold electronics

Comparison for noise and threshold between front-end locations for the charge channel

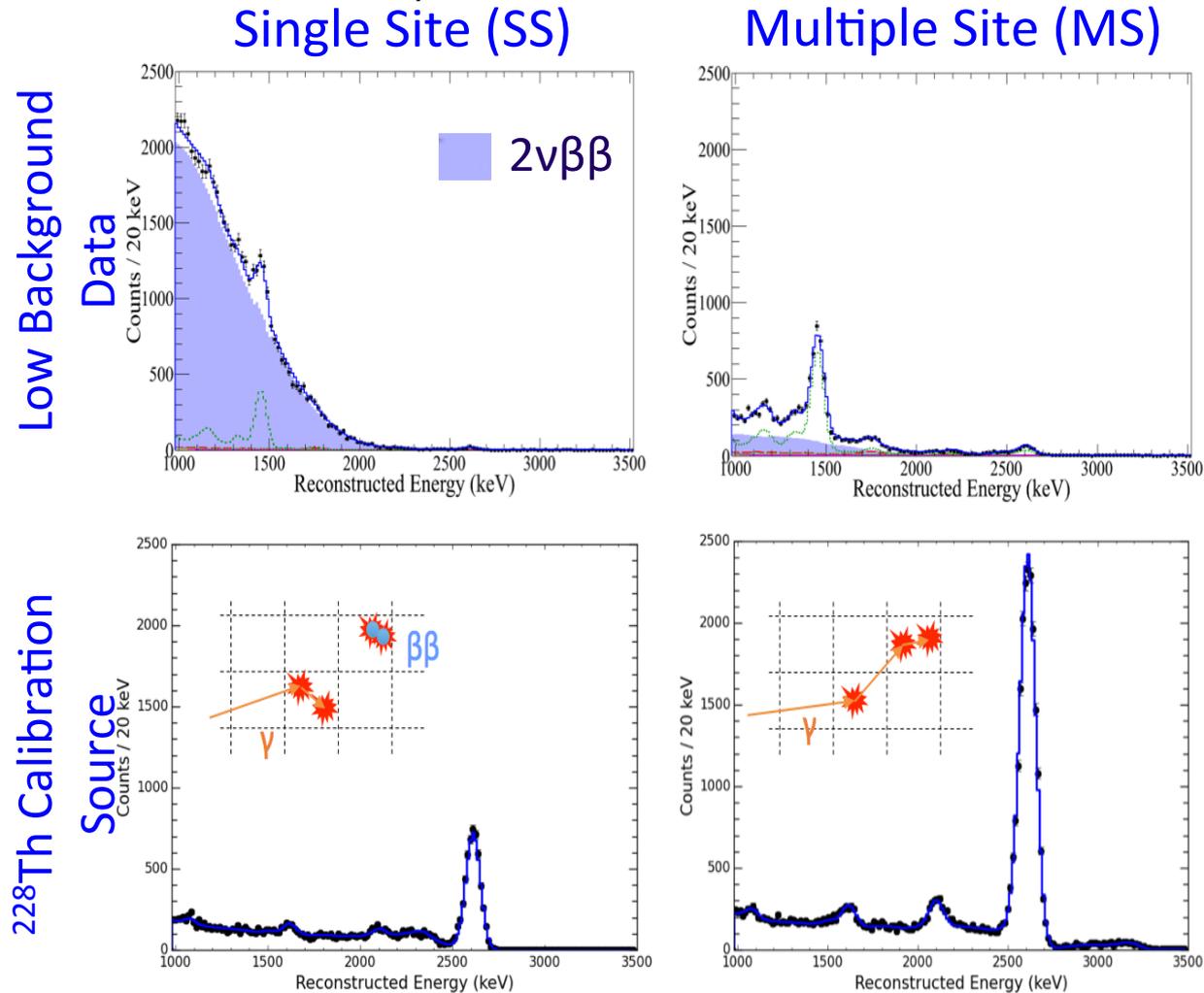
| Location | Cable length (m) | Total cap (pF) | Intrinsic RMS Noise (e) | RMS contribution to charge energy resolution | Charge cluster threshold (keV) |
|-------------------------|------------------|----------------|-------------------------|--|--------------------------------|
| In lab (warm) | 8 | 800 | 3200 | 2.5% | 600 keV |
| At cryostat (warm/cold) | 2 | 200 | 800 | 0.6% | 150 keV |
| Inside TPC (cold) | ~0 | <40 | <200 | 0.2% | 40 keV |



Assumes simple tile charge collection system with interleaved strips and EXO-200 style cables for the remote location cases

Event multiplicity and background discrimination

(example from EXO-200 data)



Improving charge threshold (near/cold electronics) and decreasing wire pitch (pads) will improve this discriminator

Event location: background identification and suppression

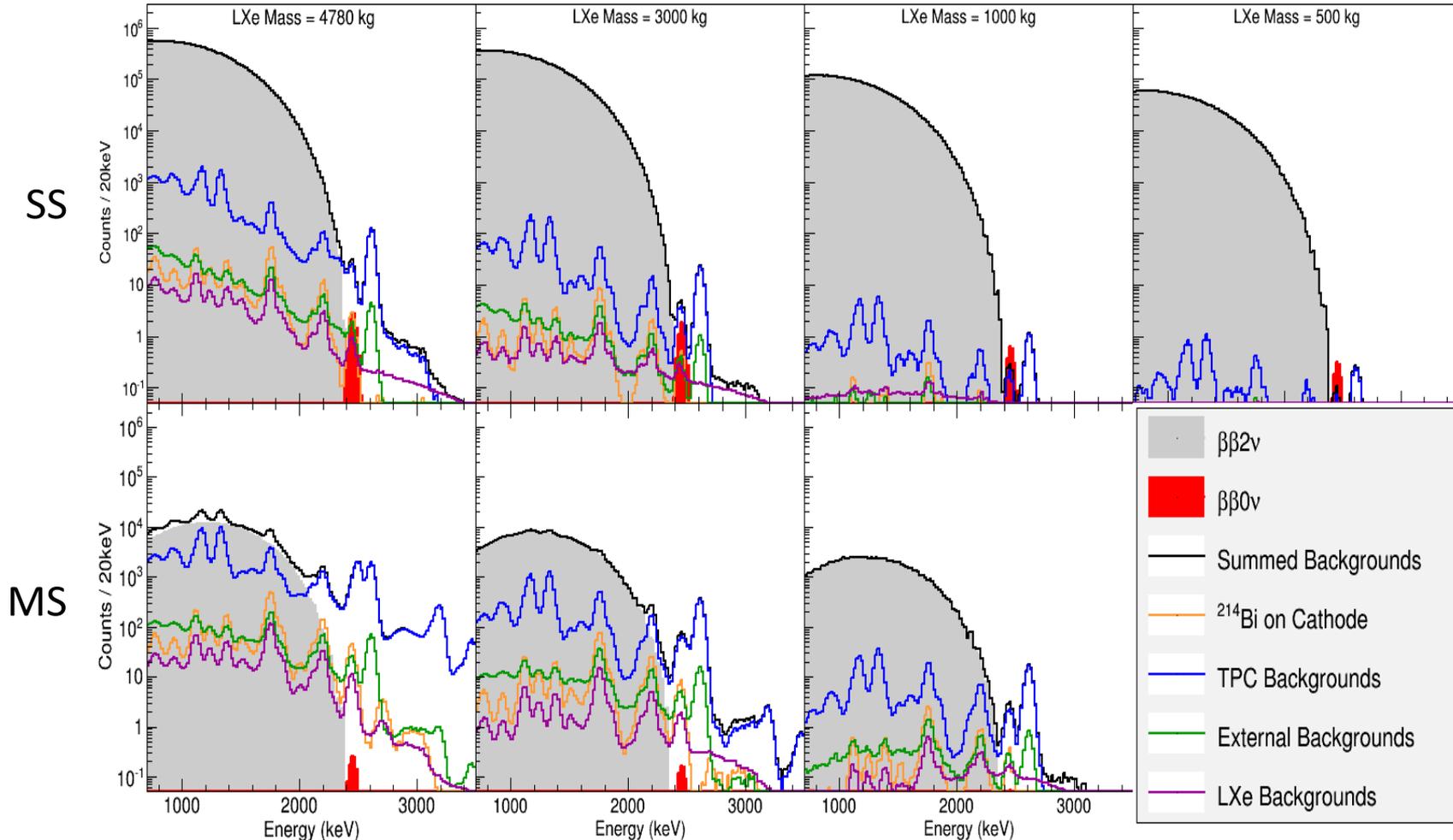
*Example: nEXO, 5 yr data, $0\nu\beta\beta$ @ $T_{1/2}=6.6\times 10^{27}$ yr,
projected backgrounds from subsets of the total volume*

Fid. LXe Mass = 4780kg

3000kg

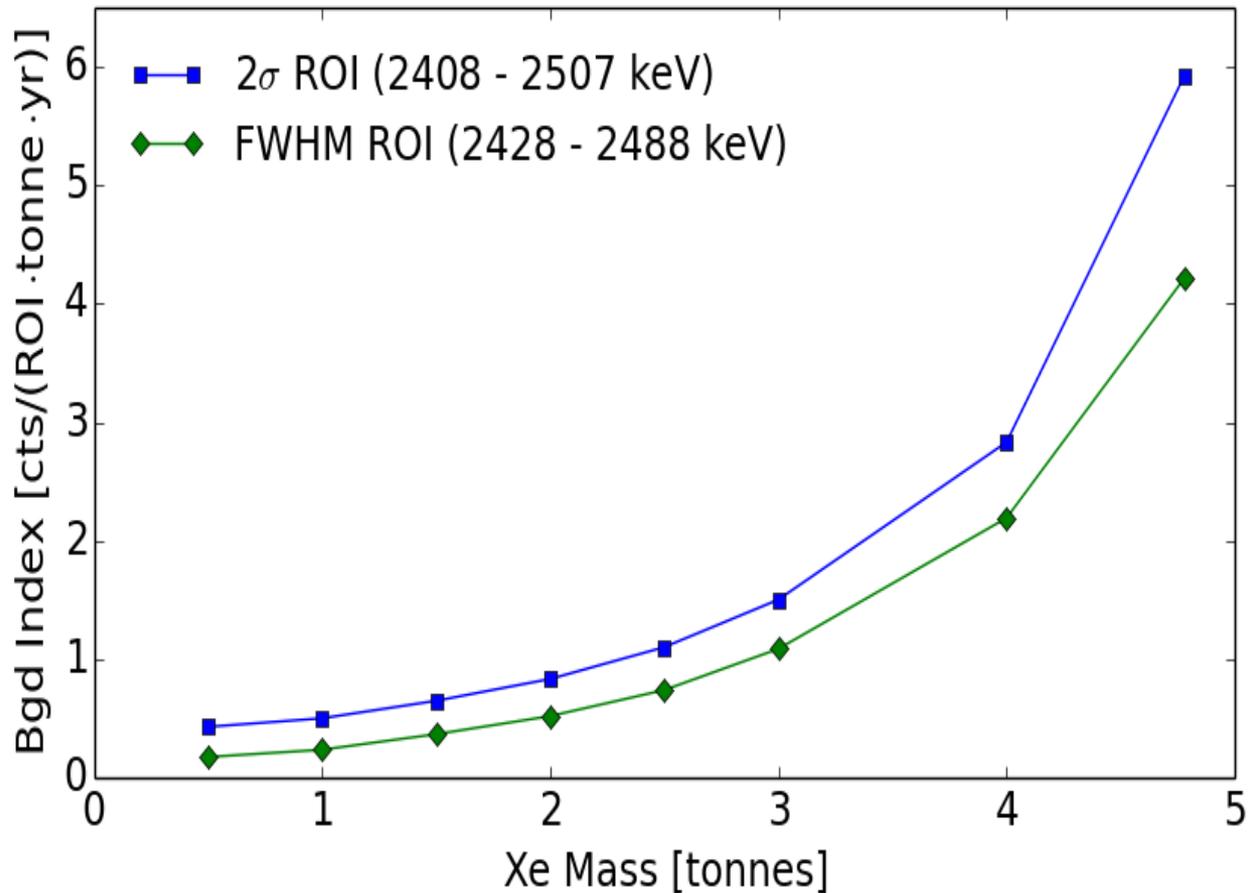
1000kg

500kg



The fit uses standoff distance information in the optimal way

Event location: background index



While nEXO background index by itself is very good...
the fit optimally uses the background vs standoff distance

EXO copper

R&D on (and later qualification of) low background materials is in full swing for nEXO.

A note on the copper that is the dominant background from the TPC vessel:

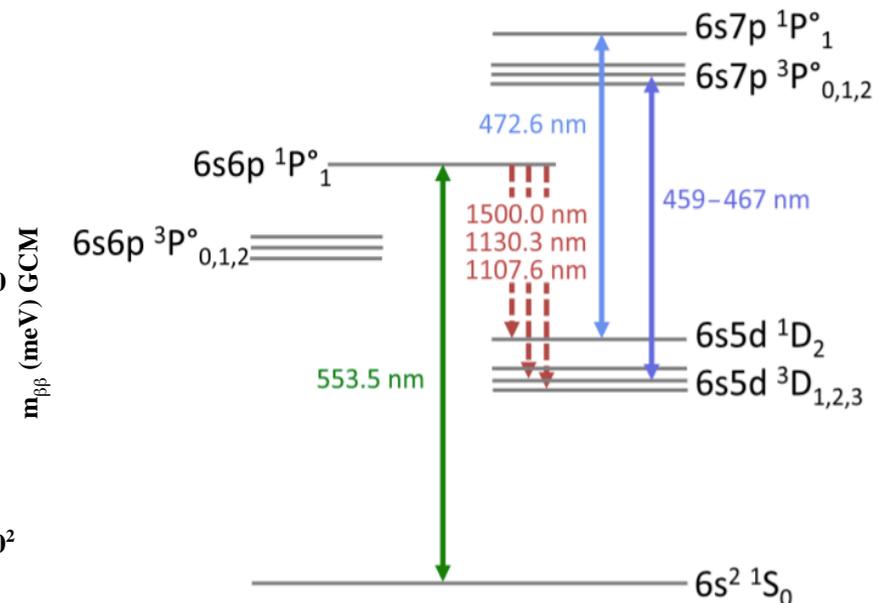
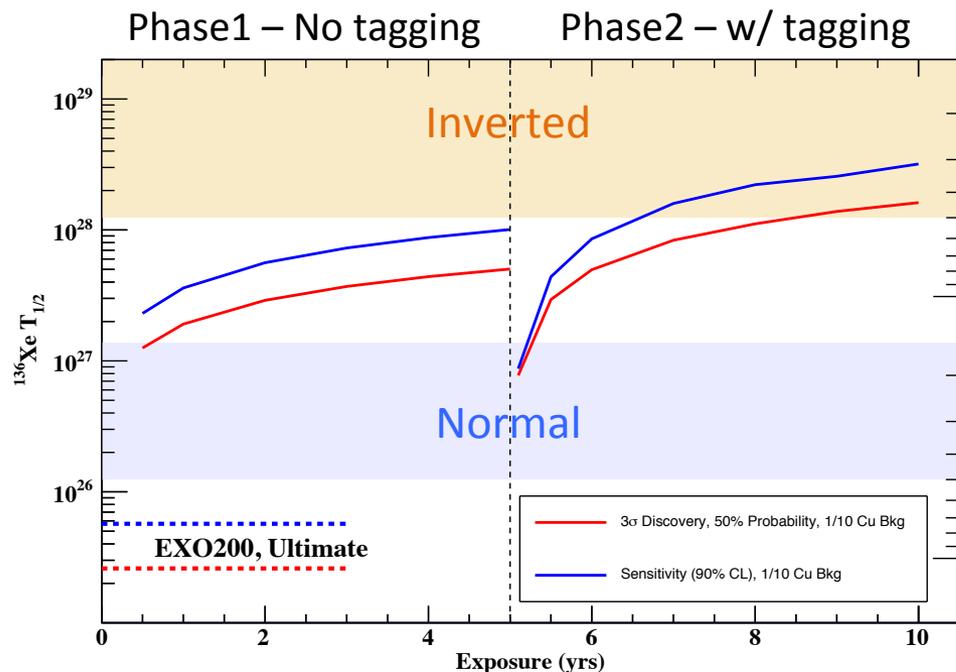
| | ~U, Th (ppt) |
|--|--------------|
| EXO-200 ICPMS measurement (Aurubis copper) | < 6, <14 |
| EXO-200 measurement (Aurubis process) | < 4 |
| nEXO measurement of Aurubis copper | < 1 |
| PNNL measurement of electroformed Cu | ~ 0.01 |

Study in progress of the Aurubis process seems to indicate that 0.1 ppt may very well be already achieved.

Barium tagging



Barium Tagging: identify barium daughter at $0\nu\beta\beta$ decay site for complete background elimination



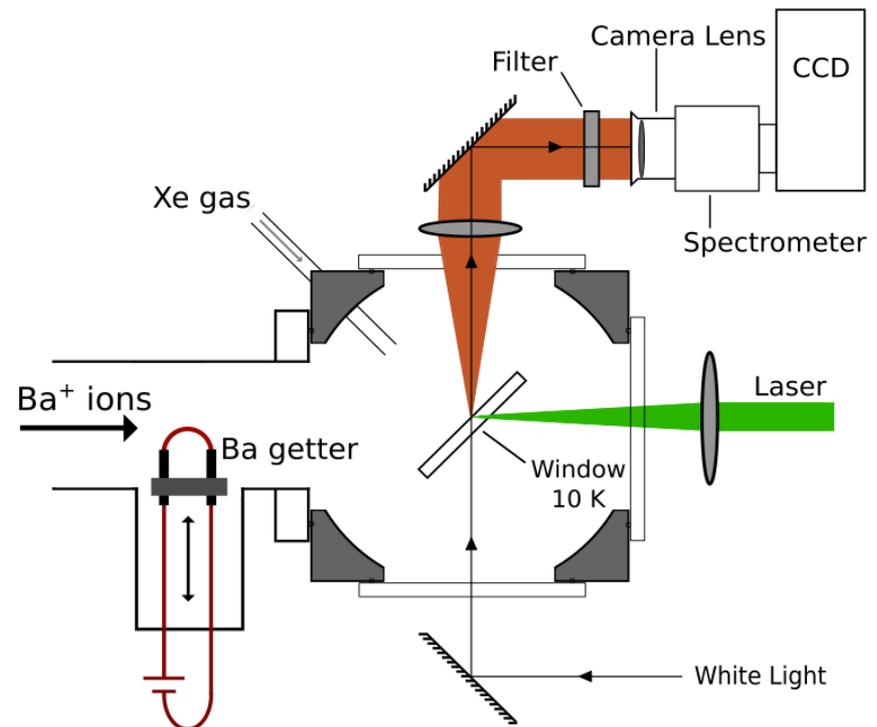
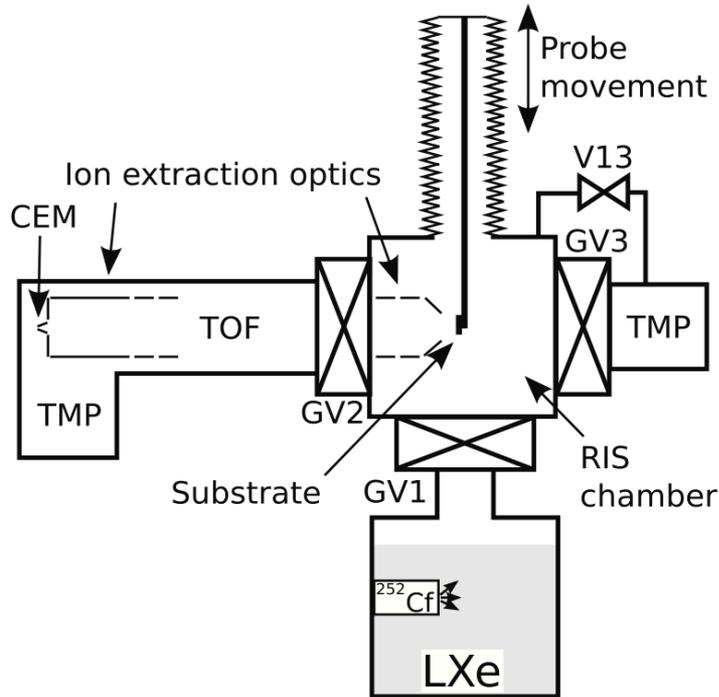
Positive identification of Ba made via atomic transitions

nEXO Ba tagging R&D

- Trigger on ionization/scintillation
- In real time estimate if 0ν candidate
- Insert probe and electrostatically attract Ba-ion (**plate** or **freeze it**)
- Several tagging methods being explored (two of which are):

Remove → desorb → RIS → CEM

Spectroscopy in SXe matrix

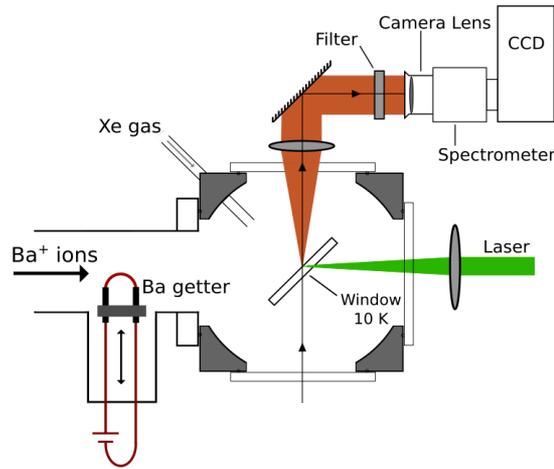
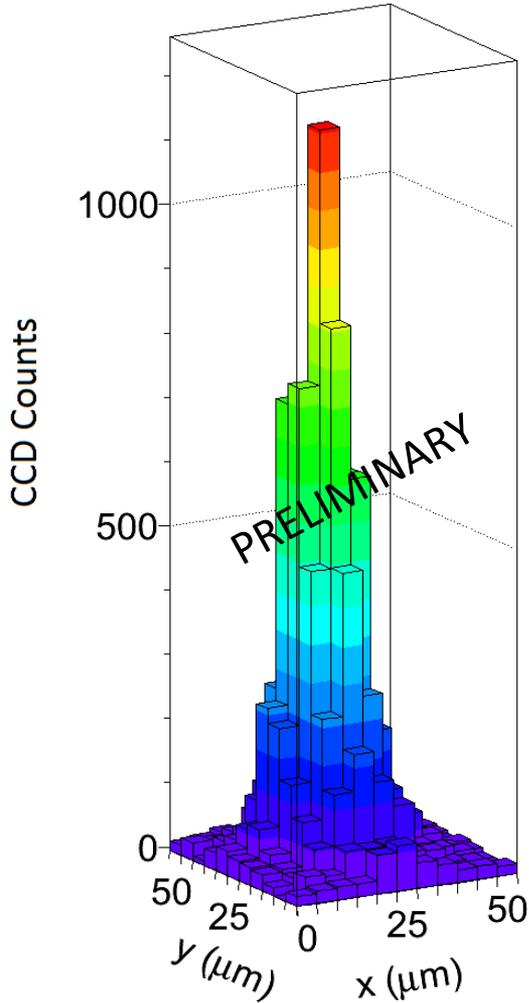


K. Twelker et al. Rev Sci Inst 85 (2014) 095114

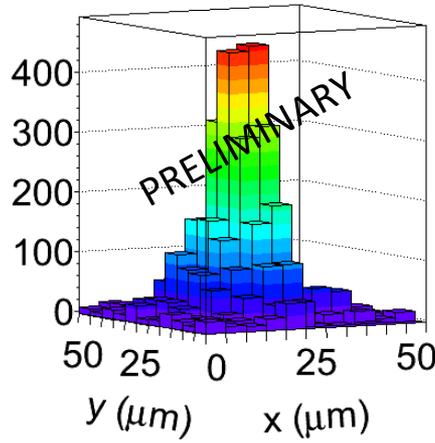
B. Mong et al. Phys Rev A 91 (2015) 022505

Ba Tagging: Identifying atoms in solid xenon

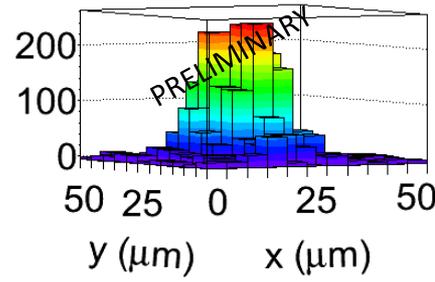
≤ 27 atoms



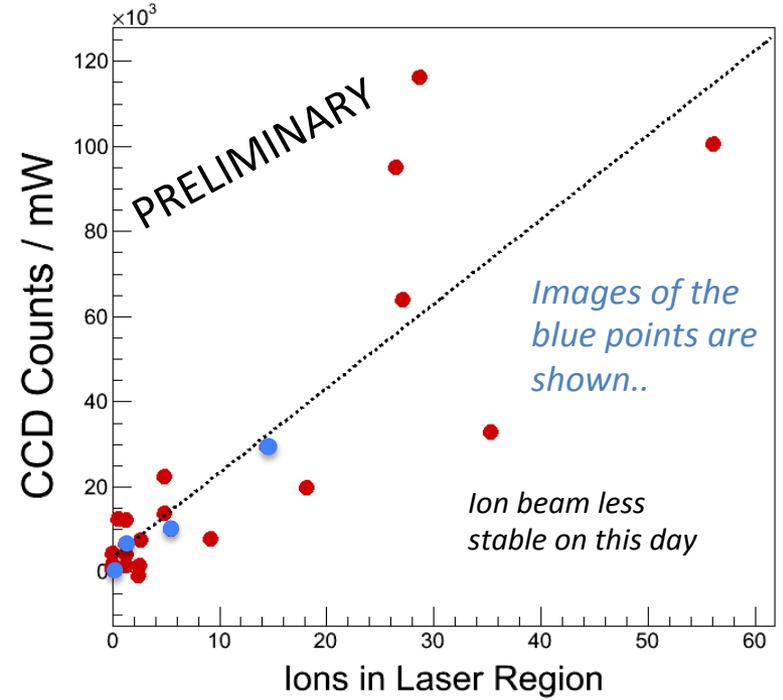
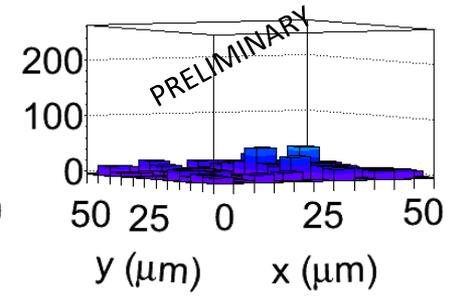
≤ 9 atoms



≤ 3 atoms



0 atoms



Summary

- nEXO is a 5-ton LXe $0\nu\beta\beta$ experiment that will cover the inverted mass hierarchy
- nEXO is a very flexible and cost effective detector with a clear upgrade path to address possible future science scenarios making the best use of the enriched isotope
- The multi-parameter approach tested in EXO-200 will be even more powerful in the larger nEXO detector
ionization density \times energy resolution \times multiplicity \times event location
- A large number of R&D projects are currently underway to verify scale up is feasible

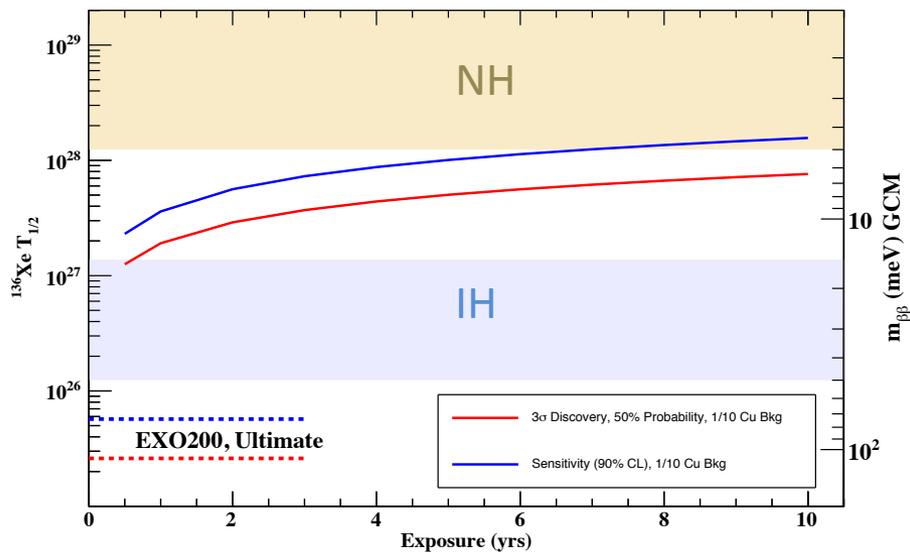


University of Alabama, Tuscaloosa AL, USA — T Didberidze, M Hughes, A Piepke, R Tsang
University of Bern, Switzerland — J-L Vuilleumier
Brookhaven National Laboratory, Upton NY, USA — M Chiu, G De Geronimo, S Li, V Radeka, T Rao, G Smith, T Tsang, B Yu
California Institute of Technology, Pasadena CA, USA — P Vogel
Carleton University, Ottawa ON, Canada — Y Baribeau, M Bowcock, M Dunford, M Facina, R Gornea, K Graham, P Gravelle, R Killick, T Koffas, C Licciardi, K McFarlane, R Schnarr, D Sinclair
Colorado State University, Fort Collins CO, USA — C Chambers, A Craycraft, W Fairbank Jr, T Walton
Drexel University, Philadelphia PA, USA — E Callaghan, MJ Dolinski, YH Lin, E Smith, Y-R Yen
Duke University, Durham NC, USA — PS Barbeau, G Swift
University of Erlangen-Nuremberg, Erlangen, Germany — G Anton, R Bayerlein, J Hoessl, P Hufschmidt, A Jamil, T Michel, T Ziegler
IBS Center for Underground Physics, Daejeon, South Korea — DS Leonard
IHEP Beijing, People's Republic of China — G Cao, W Cen, X Jiang, H Li, Z Ning, X Sun, T Tolba, W Wei, L Wen, W Wu, J Zhao
ITEP Moscow, Russia — V Belov, A Burenkov, A Karelin, A Kobayakin, A Kuchenkov, V Stekhanov, O Zeldovich
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Indiana University, Bloomington IN, USA — JB Albert, S Daugherty, TN Johnson, LJ Kaufman, G Visser, J Zettlemoyer
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TRIUMF, Vancouver BC, Canada — J Dilling, P Gumplinger, R Krücken, F Retière, V Strickland

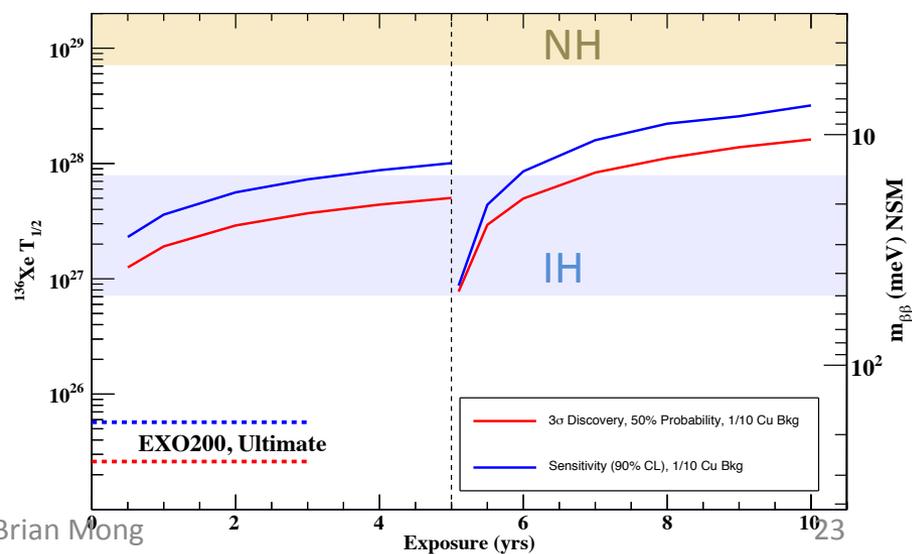
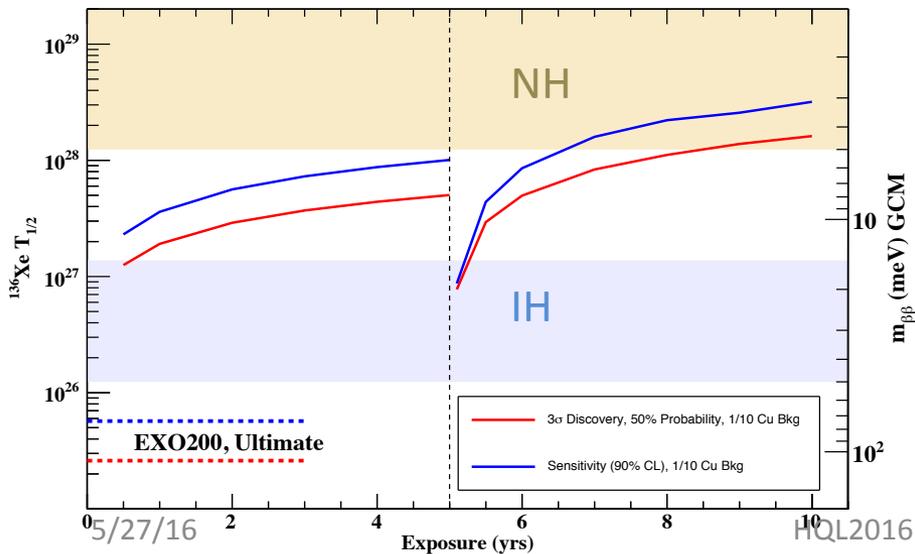
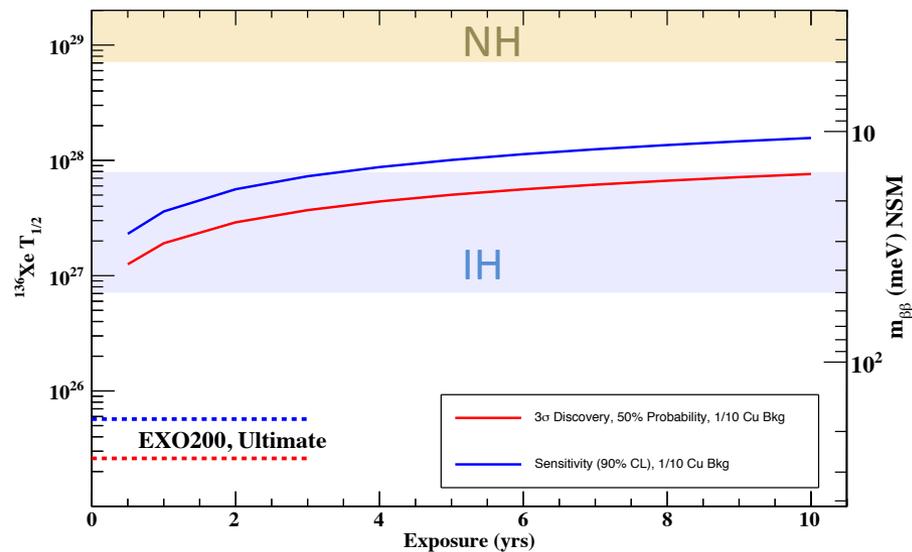


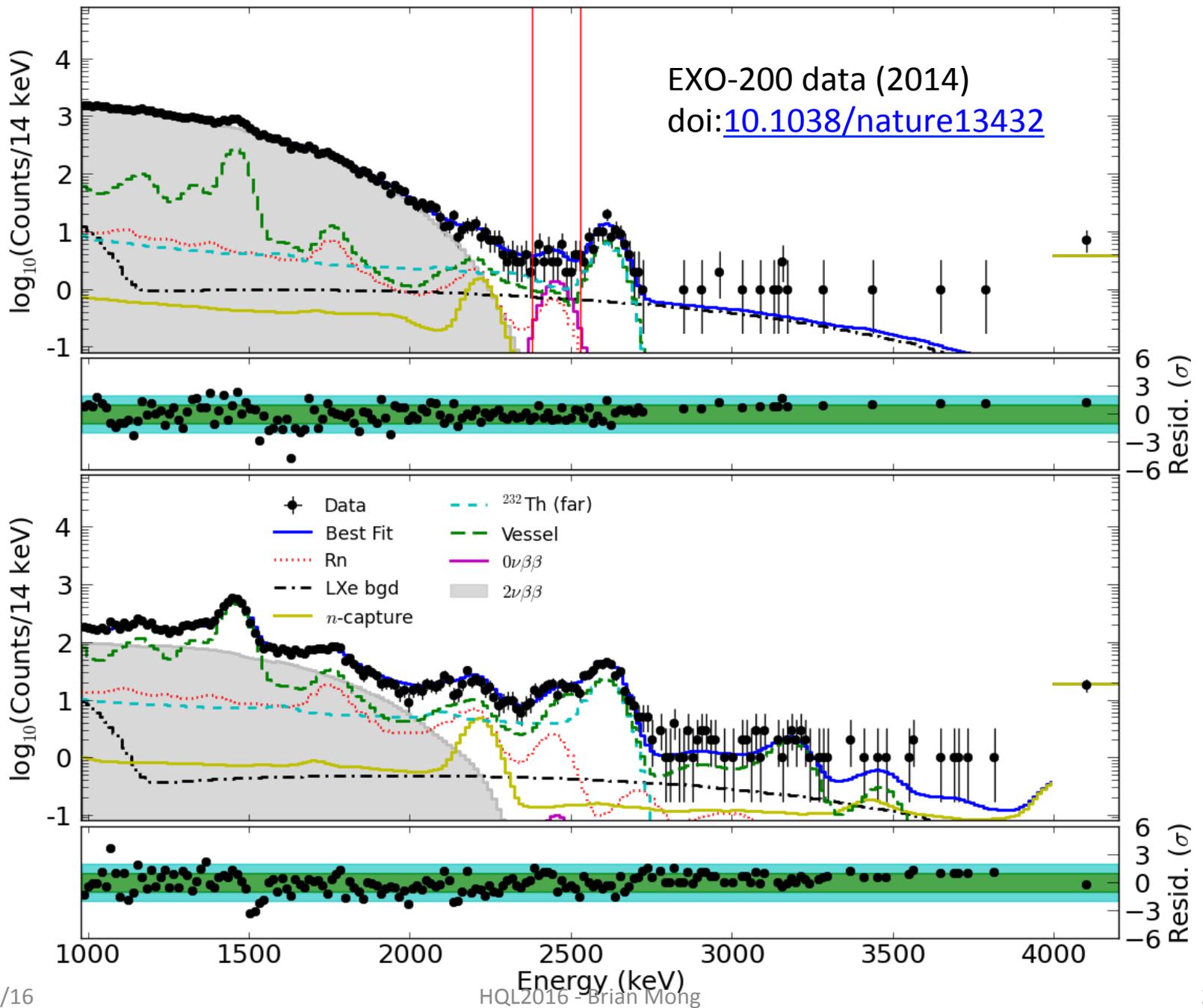
Sensitivities for worst and best case nuclear matrix elements

Best case (GCM)

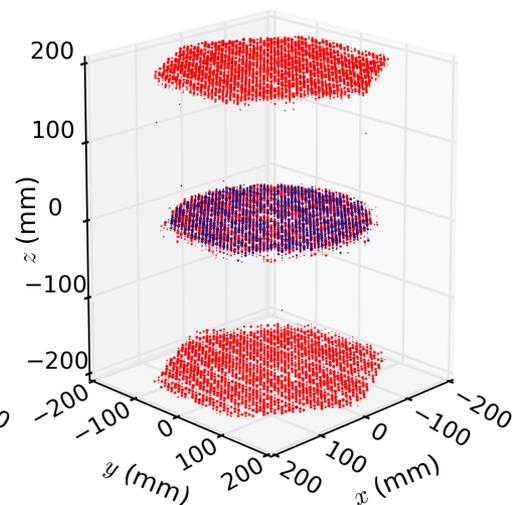
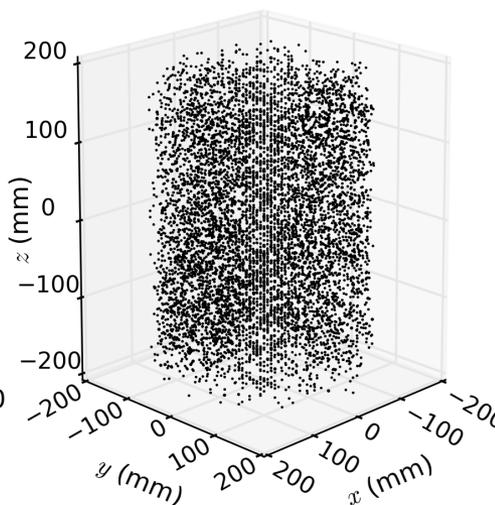
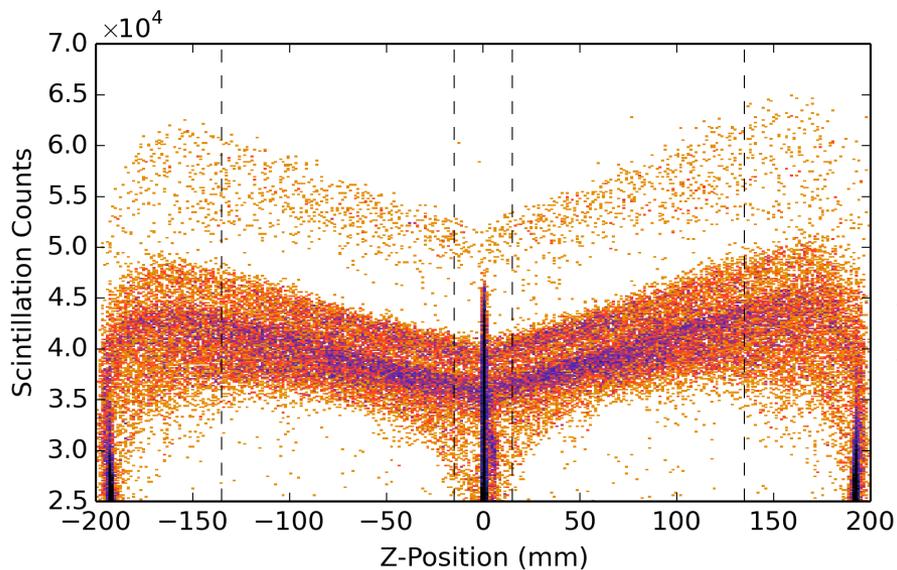
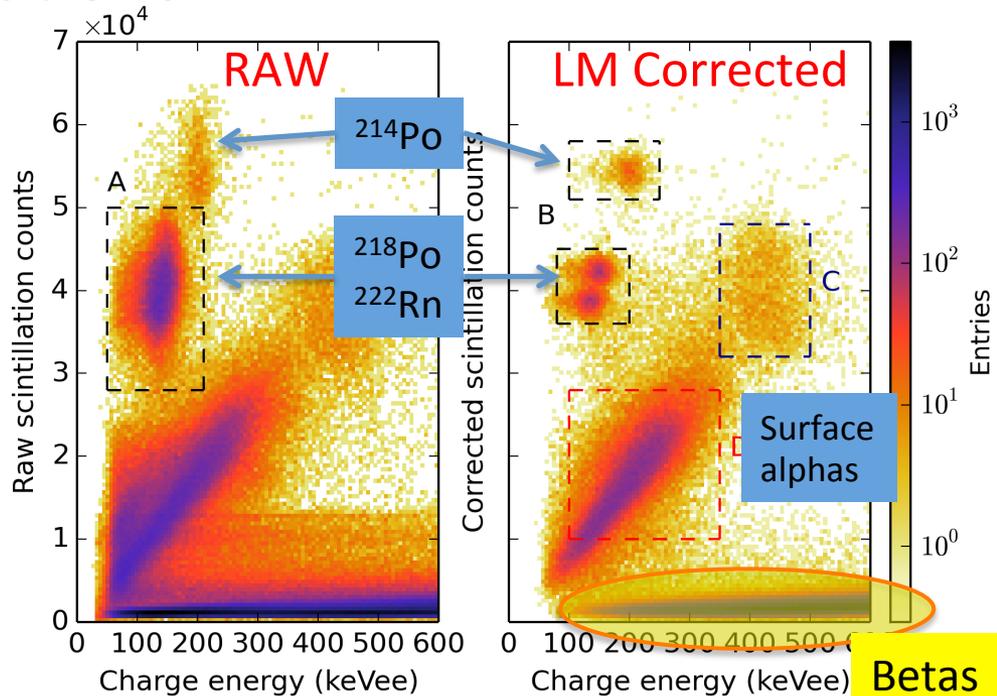
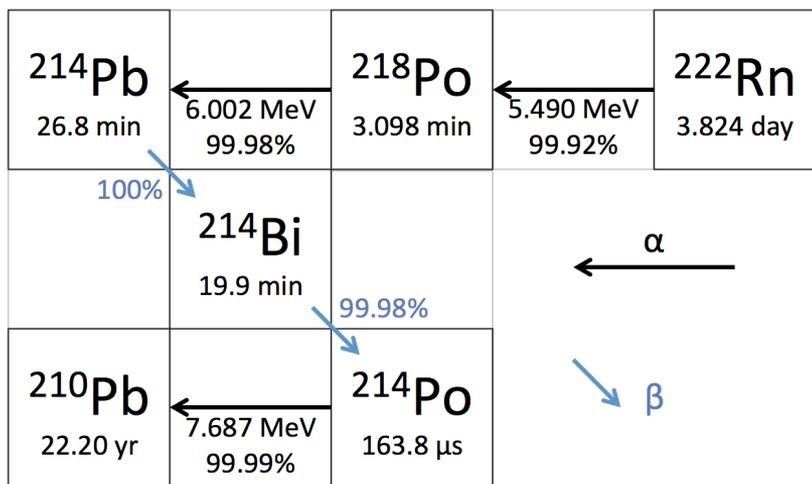


Worst case (NSM)





Alpha energy and event locations



^{214}Bi ion fraction (β decay)

Measure ratio of alpha events in LXe

(Assume 50.3(3.0)% of ^{214}Pb are ions)

$$\frac{A_{\text{Po}}^{214}}{A_{\text{Rn}}^{222}}$$

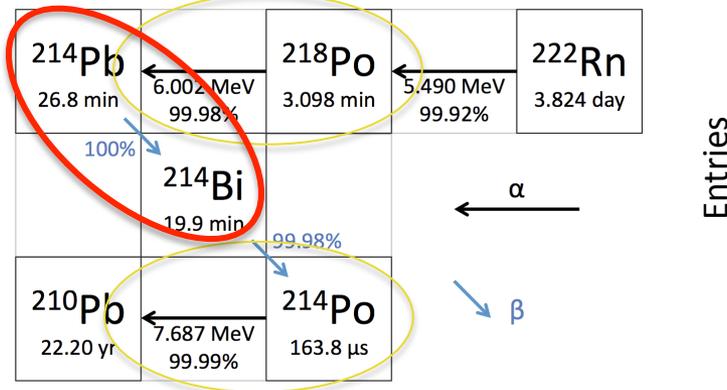
Neutrals Decay in flight

$$\frac{A_{\text{Po}}^{214}}{A_{\text{Po}}^{218}} = (1 - f'_\alpha + f'_\alpha \epsilon_{\text{Pb}})(1 - f_\beta + f_\beta \epsilon_{\text{Bi}})$$

f_β is daughter ion fraction of β decay

Measure 2 alpha decays in LXe fid. vol.
Deduce beta decay ion fraction

$$f_\beta: 76.4 \pm 5.7 \%$$

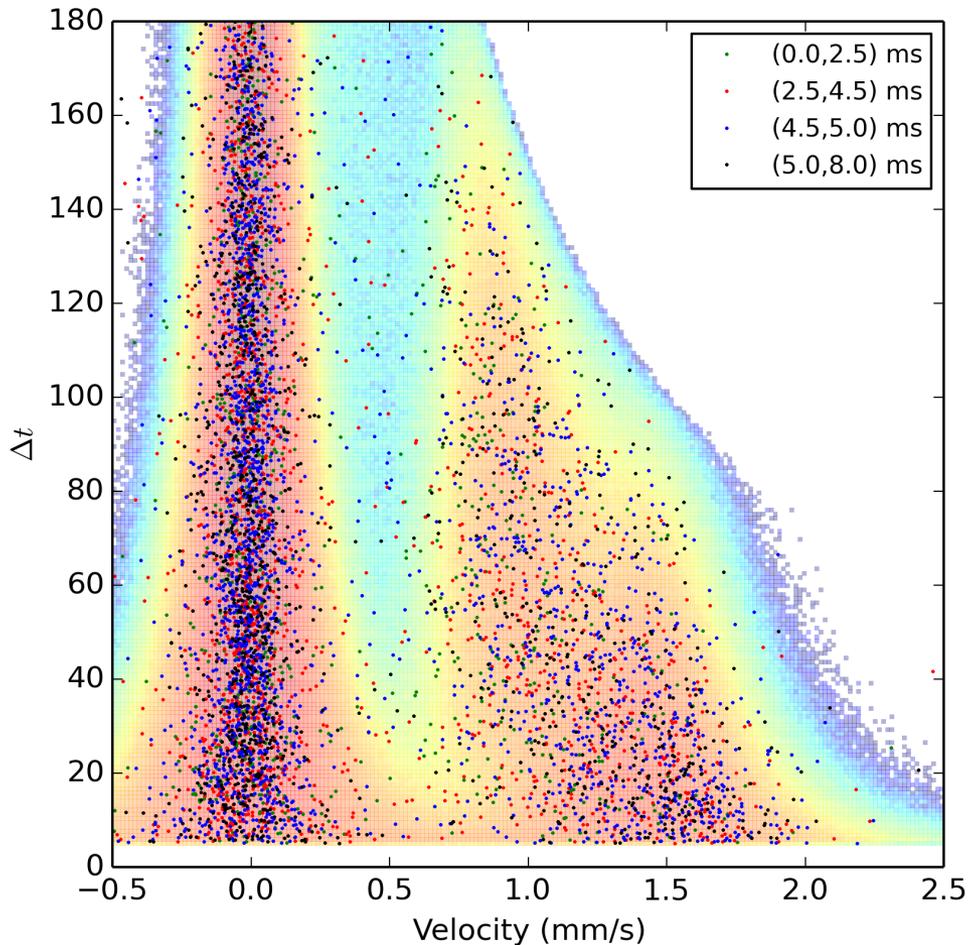


Good news for
barium tagging!

J. B. Albert et al. (EXO-200 Collaboration)

[Phys. Rev. C 92, 045504](https://arxiv.org/abs/1508.01211)

Using ^{222}Rn - ^{218}Po events to determine ion drift parameters



MC Model Parameters:

$$v_1 = 1.48 \pm 0.01 \text{ mm/s}$$

$$v_2 = 0.83 \pm 0.01 \text{ mm/s}$$

$$C = 12600 \pm 660$$

$$N = 6^{+4.9}_{-1.7} * 10^5$$

$$D = 0.61 \pm 0.04 \text{ mm}^2/\text{s}$$

Mobility of ^{218}Po

$$\mu_1 = 0.390 \pm 0.006 \text{ cm}^2/(\text{keV}\cdot\text{s})$$

$$\mu_2 = 0.219 \pm 0.004 \text{ cm}^2/(\text{keV}\cdot\text{s})$$

other ions are 0.13-0.28 $\text{cm}^2/(\text{keV}\cdot\text{s})$

C is ratio of reaction lifetime to electron lifetime:

$$(t_e \sim 3 \text{ ms}, t_{\text{reaction}} \sim 40\text{s})$$

*N is ratio or neutralization time to electron lifetime ($N=50 * C$):*

An important result for Ba tagging: ($t_e \sim 3 \text{ ms}, t_{\text{neut}} > 1000\text{s}$)