

New Results from RENO & Future RENO-50 Project

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Heavy Quarks and Leptons
Virginia Tech
May 22-27, 2016



RENO Collaboration



Reactor Experiment for Neutrino Oscillation

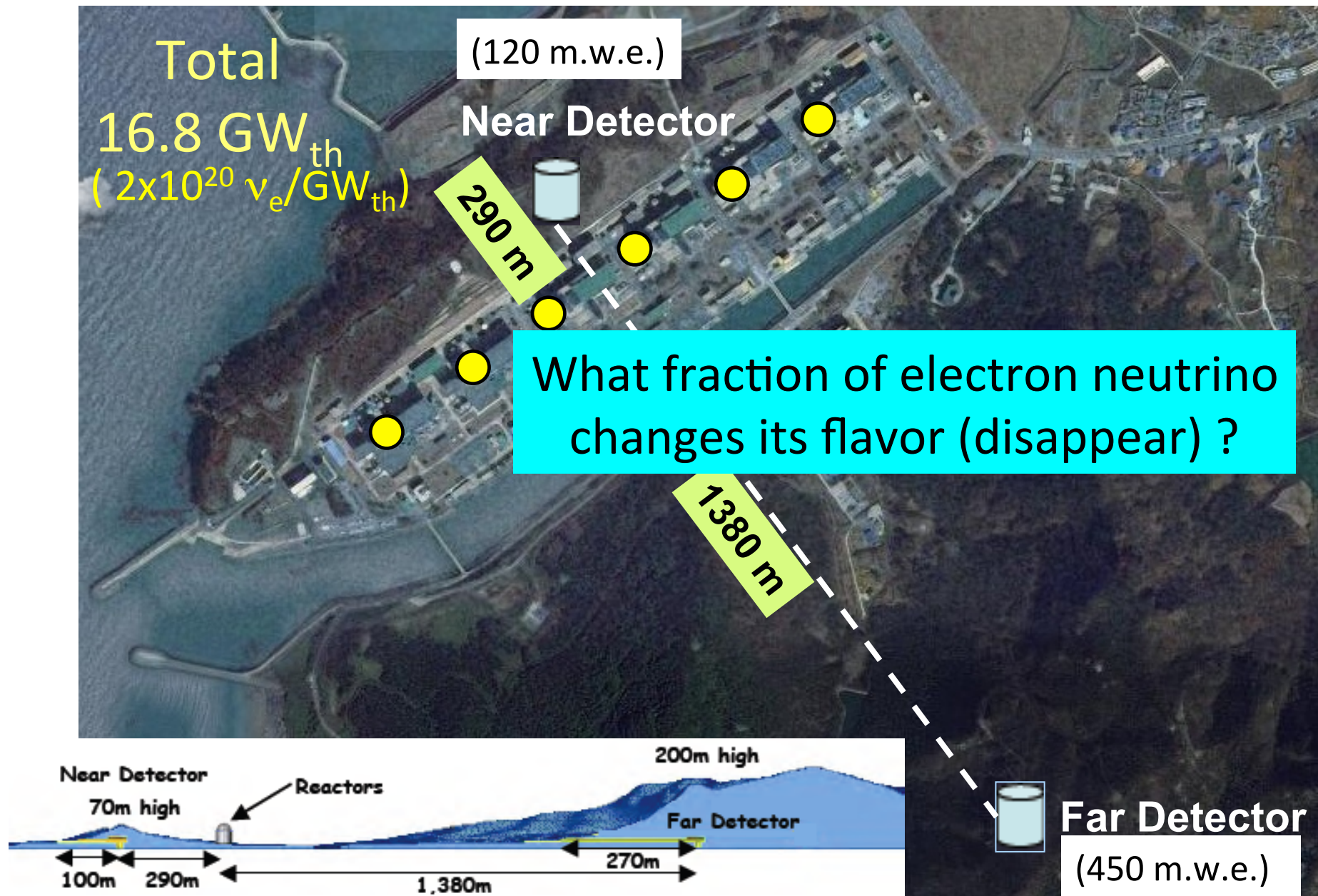
(~30 members in 10 institutions)

- Chonnam National University
- Dongshin University
- GIST
- Gyeongsang National University
- IBS
- Kyungpook National University
- Sejong University
- Seoul National University
- Seoyeong University
- Sungkyunkwan University

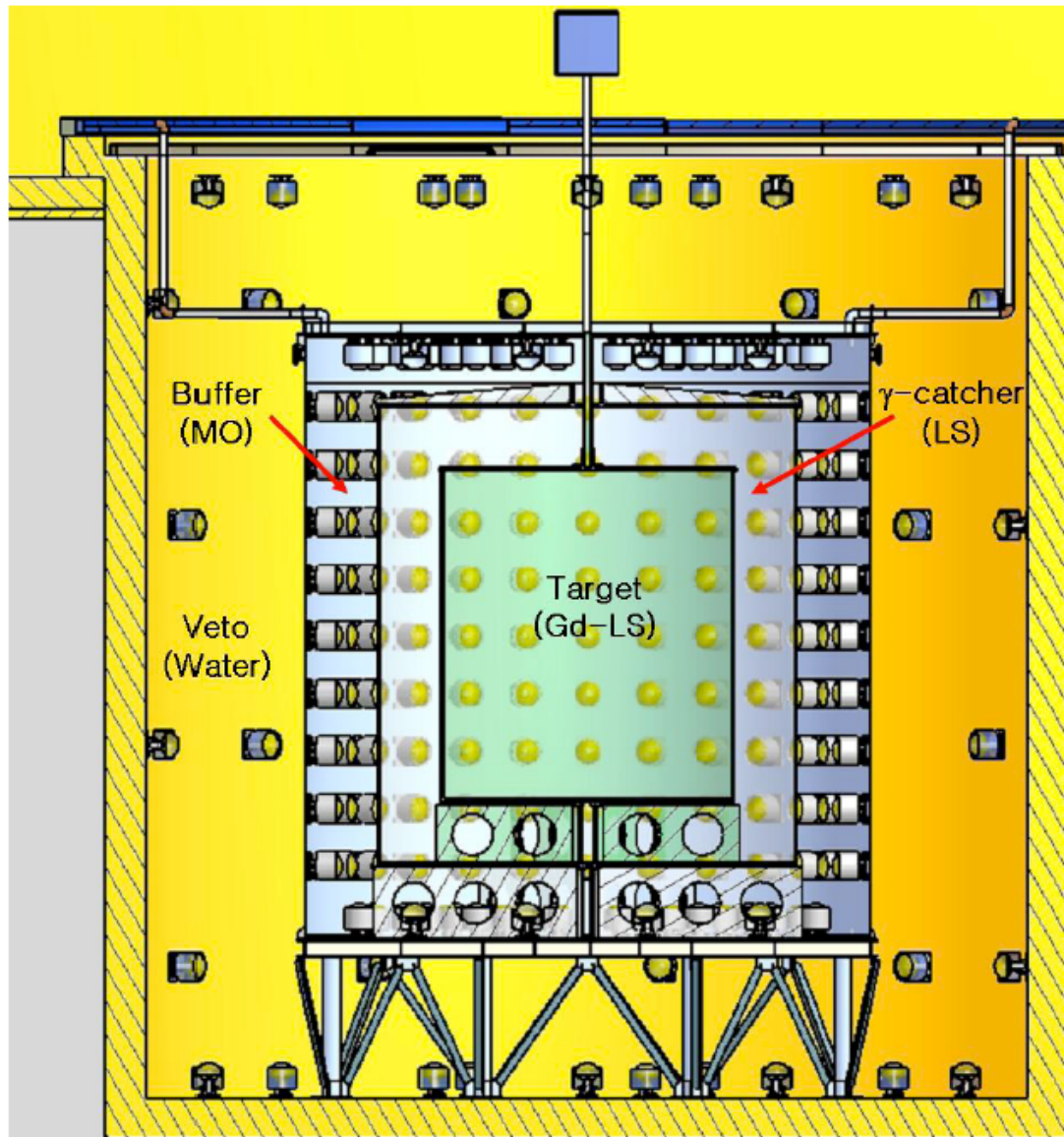
- Total cost : \$10M
- Start of project : 2006
- The first reactor experiment running with both near & far detectors from Aug. 2011



RENO Experimental Setup



The RENO Detector



4 enclosed cylinders

- **Target** : 16 ton Gd-LS

(R=1.4m, H=3.2m)

- **Gamma Catcher** :

30 ton LS

(R=2.0m, H=4.4m)

- **Buffer** : 65 ton mineral oil

(R=2.7m, H=5.8m)

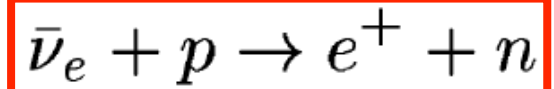
- **Veto** : 350 ton water

(R=4.2m, H=8.8m)

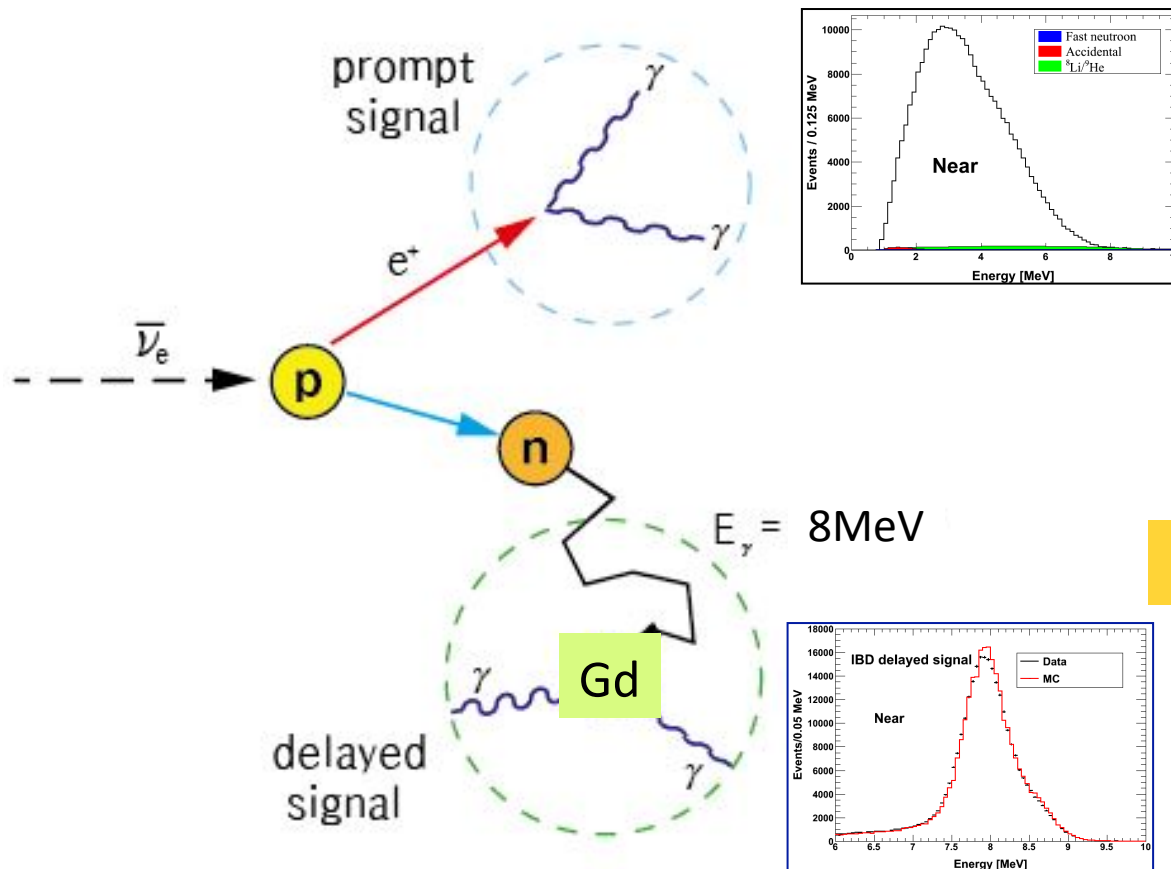
-- 354 ID 10 " PMTs

-- 67 OD 10" PMTs

Detection Principle of Reactor Neutrinos



IBD process



or

Gd capture

Delayed signal

$\sim 30\ \mu\text{s}$

$\sim 8\text{ MeV}$

H capture

Delayed signal

$\sim 200\ \mu\text{s}$

$\sim 2.2\text{ MeV}$

- Prompt signal (e^+) : 1 MeV 2γ 's + e^+ kinetic energy ($E = 1\sim 10\text{ MeV}$)
- Delayed signal (n) : 8 MeV γ 's from neutron's capture by **Gd** in $\sim 30\ \mu\text{s}$
or 2.2 MeV γ 's from neutron's capture by **H** in $\sim 200\ \mu\text{s}$

RENO Data-taking Status

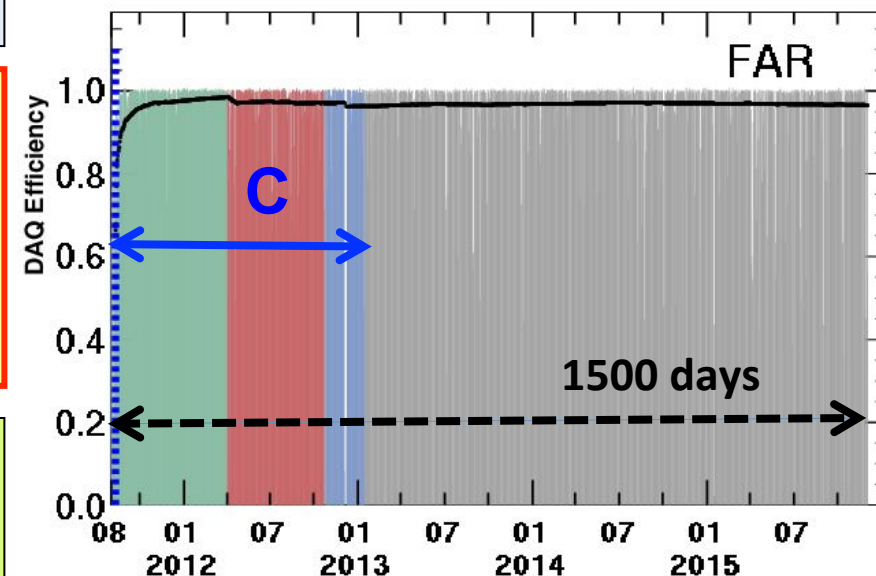
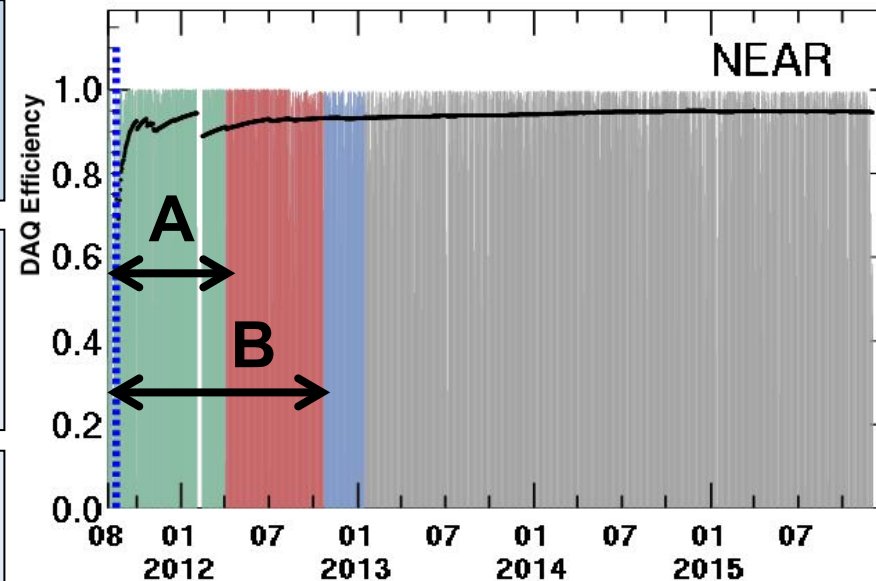
- Data taking began on Aug. 1, 2011 with both near and far detectors.
(DAQ efficiency : ~95%)

- A (220 days) : First θ_{13} result**
[11 Aug, 2011~26 Mar, 2012]
PRL 108, 191802 (2012)

- B (403 days) : Improved θ_{13} result**
[11 Aug, 2011~13 Oct, 2012]
NuTel 2013, TAUP 2013, WIN 2013

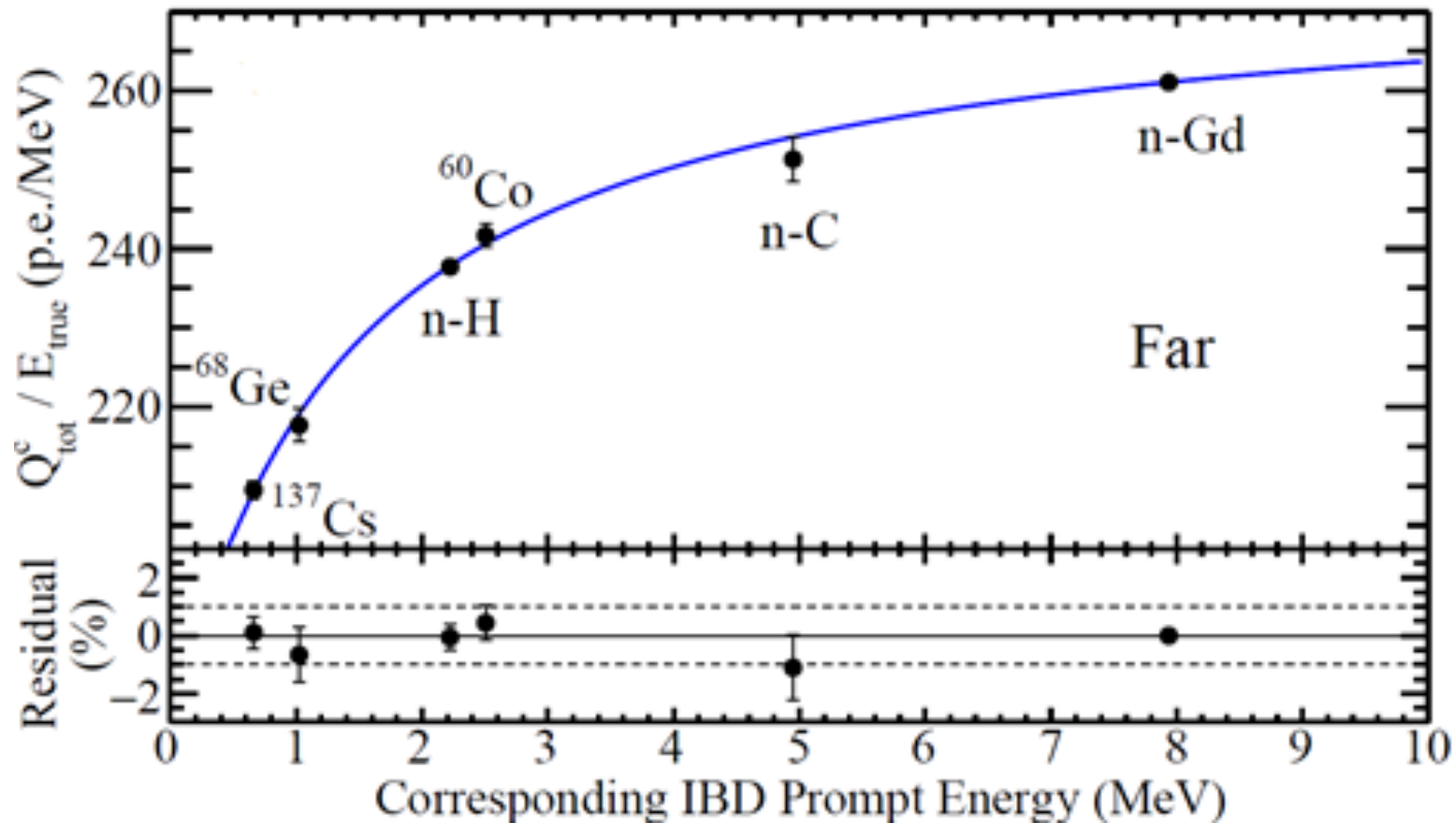
- C (~500 days) : New result**
Shape+rate analysis (θ_{13} and $|\Delta m_{ee}^2|$)
[11 Aug, 2011~21 Jan, 2013]
→ Sterile neutrino search in progress,
n-H analysis in progress

- Total observed reactor neutrino events as of today (**1500 days**):
~ 1.5M (Near), **~ 0.15M** (Far)



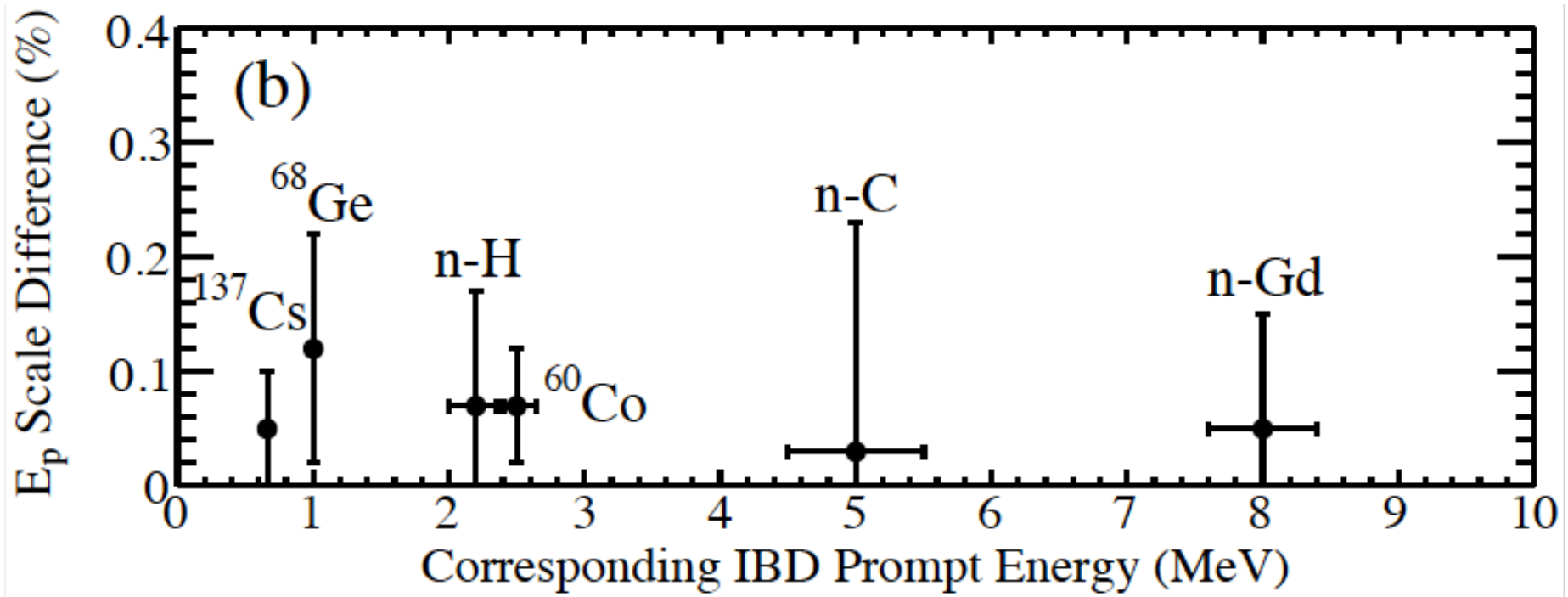
Energy Calibration from γ -ray Sources

- Non-linear response of the scintillation energy is calibrated using γ -ray sources.
- The visible energy from γ -ray is corrected to its corresponding positron energy.



Fit function : $E_{\text{vis}}/E_{\text{true}} = a - b/(1 - \exp(-cE_{\text{true}} - d))$

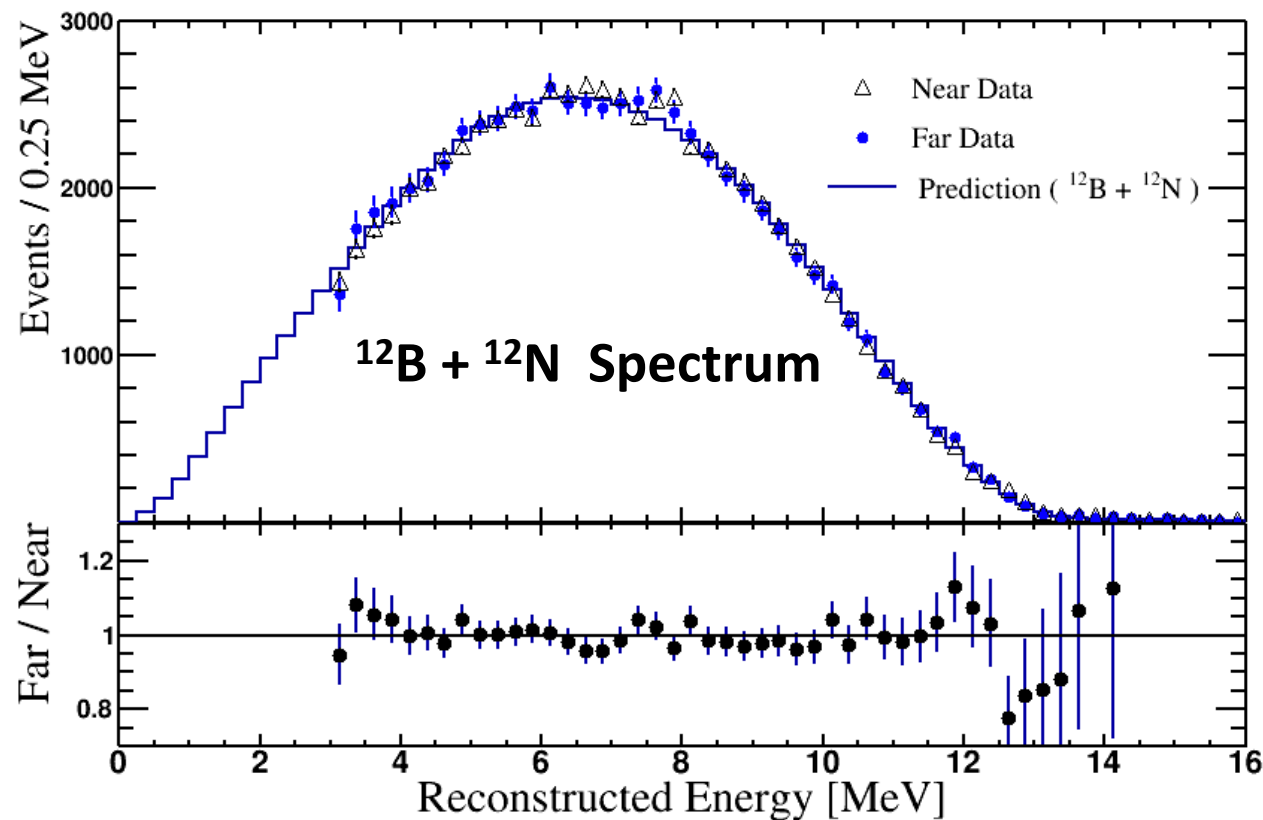
Energy Scale Difference between Near & Far



Energy scale uncertainty: $< 0.15 \%$

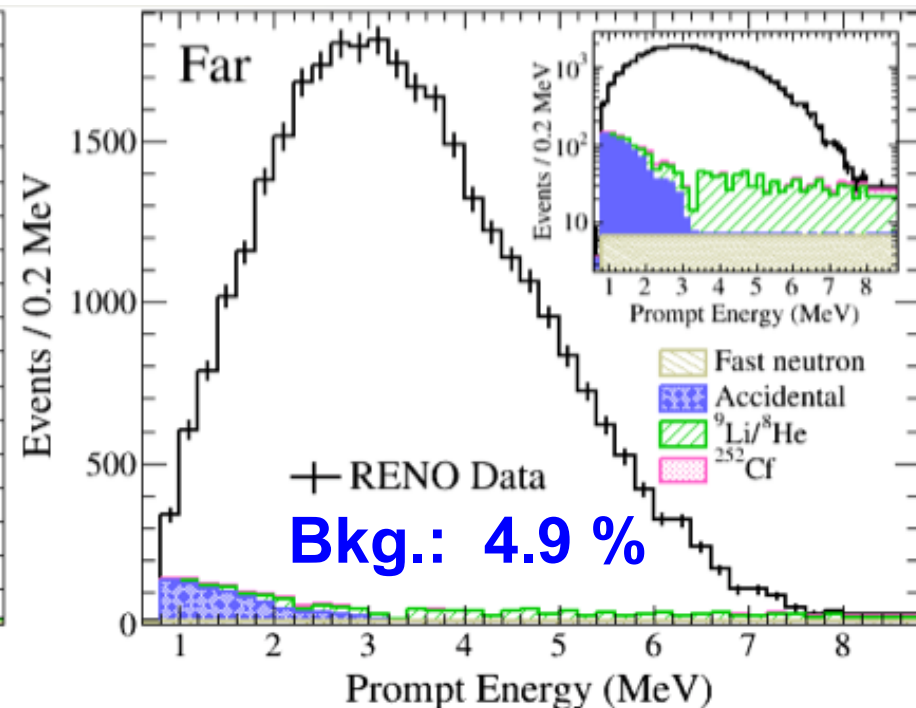
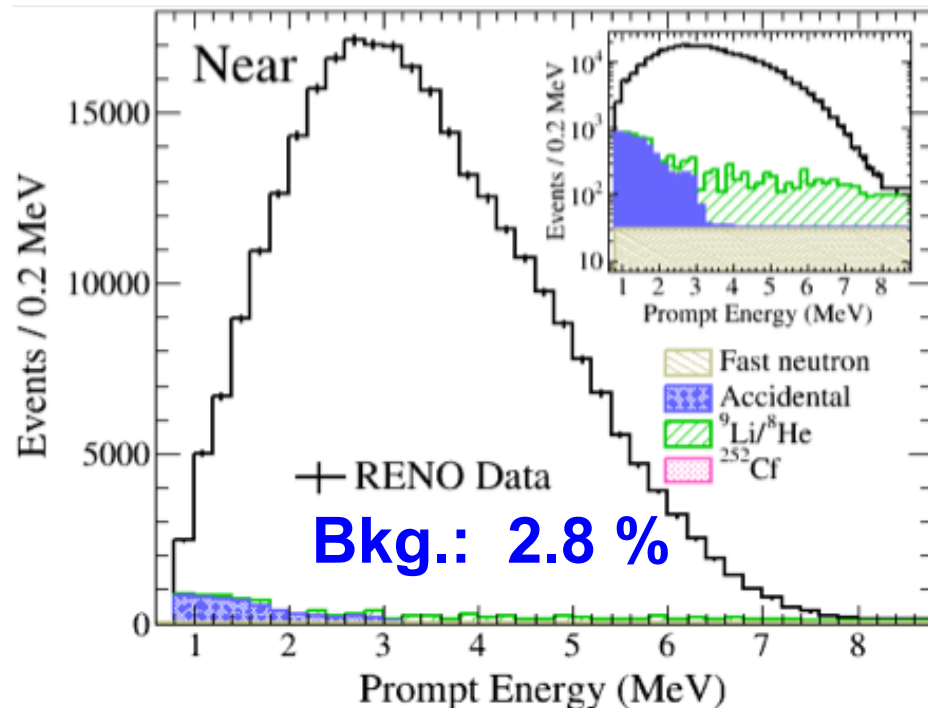
^{12}B Energy Spectrum

- Electron energy spectrum from β -decays from ^{12}B and ^{12}N , which are produced by cosmic-muon interactions.



Good agreement between data and MC spectrum!

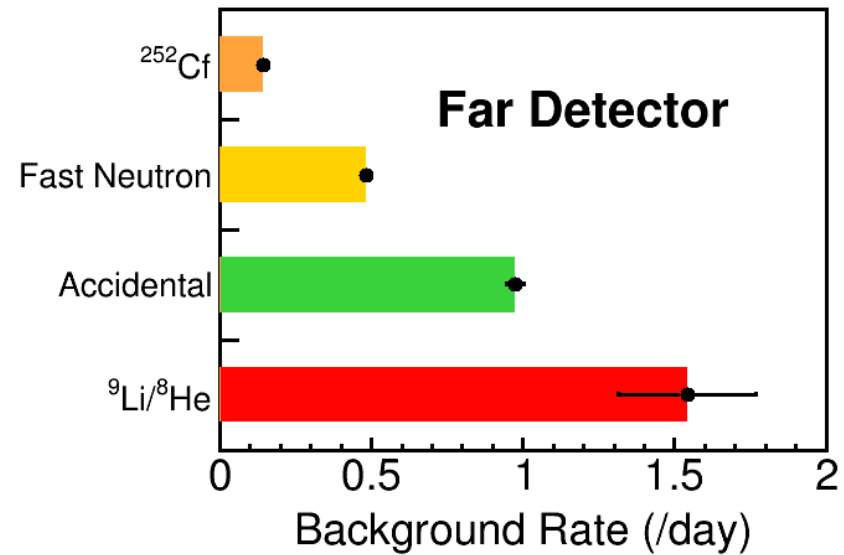
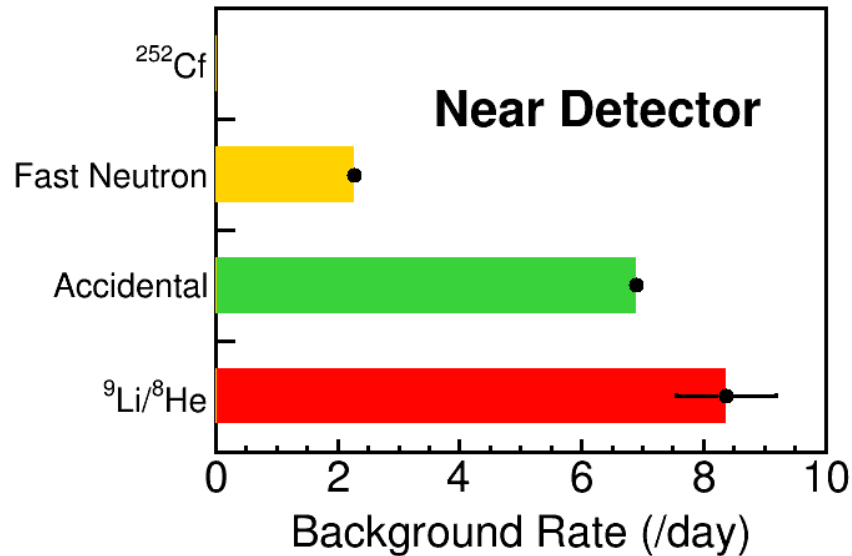
Prompt Energy Spectra of IBD Candidates



Near Live time = 458.49 days
of IBD candidate = 290,775
of background = 8,041 (2.8 %)

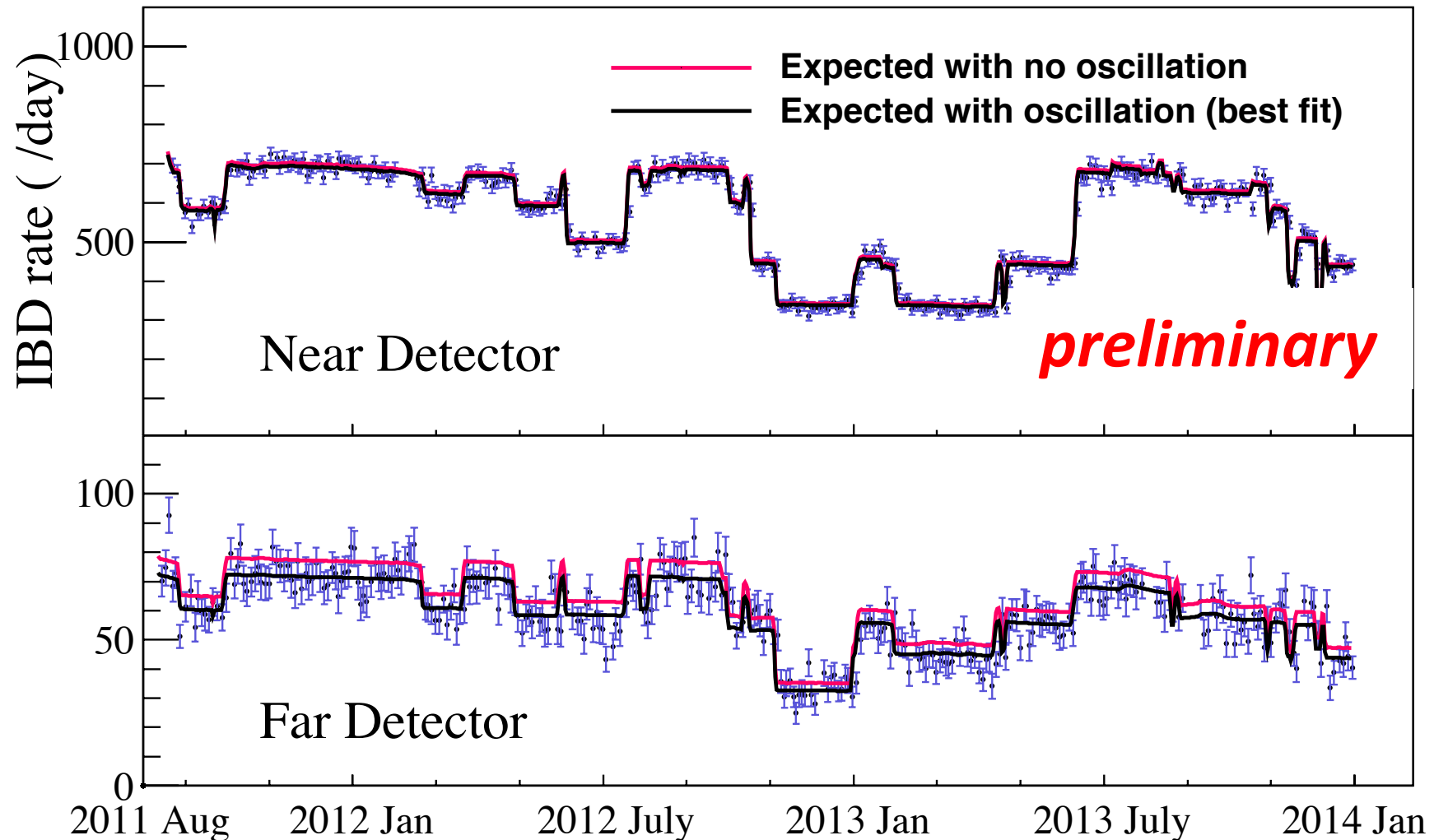
Far Live time = 489.93 days
of IBD candidate = 31,541
of background = 1540 (4.9 %)

IBD Candidates & Background



	Near	Far
DAQ live time [days]	458.49	489.93
IBD candidates	290755	31541
Total BKG rate [/day]	17.54 ± 0.83	3.14 ± 0.21
IBD rate [/day] after BKG subtraction	616.67 ± 1.44	61.24 ± 0.42

Observed Daily Averaged IBD Rate



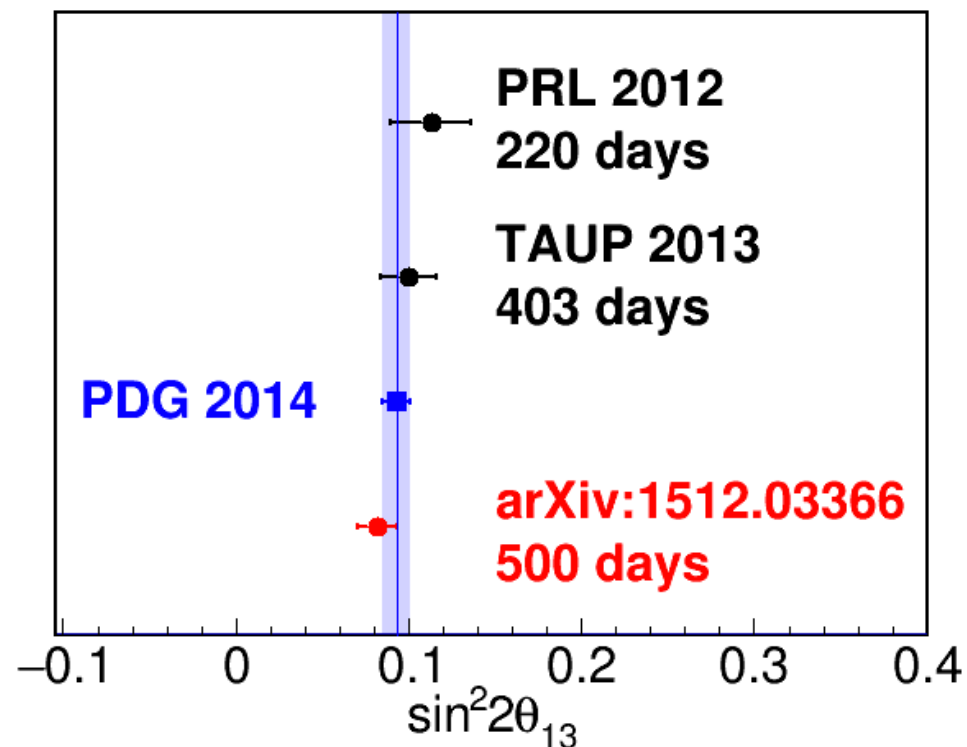
- Good agreement with observed rate and prediction.
- Accurate measurement of thermal power by reactor neutrinos

New θ_{13} Measurement by Rate-only Analysis

Rate-only
new result

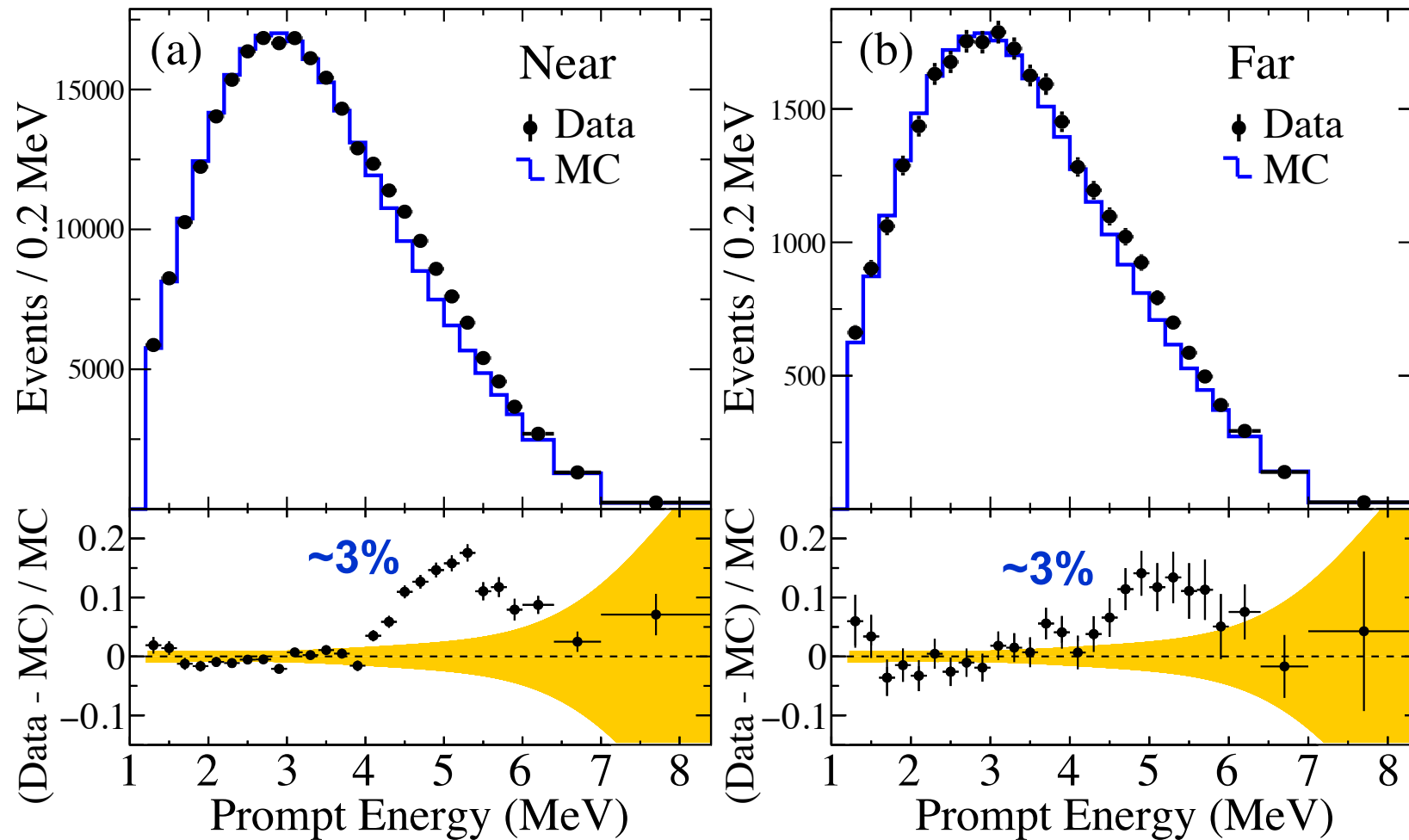
$$\sin^2 2\theta_{13} = 0.087 \pm 0.009(\text{stat.}) \pm 0.007(\text{syst.})$$

By minimizing $\chi^2 = \frac{(O^{F/N} - T^{F/N})^2}{(U)^2} + \text{Pull_Terms}$

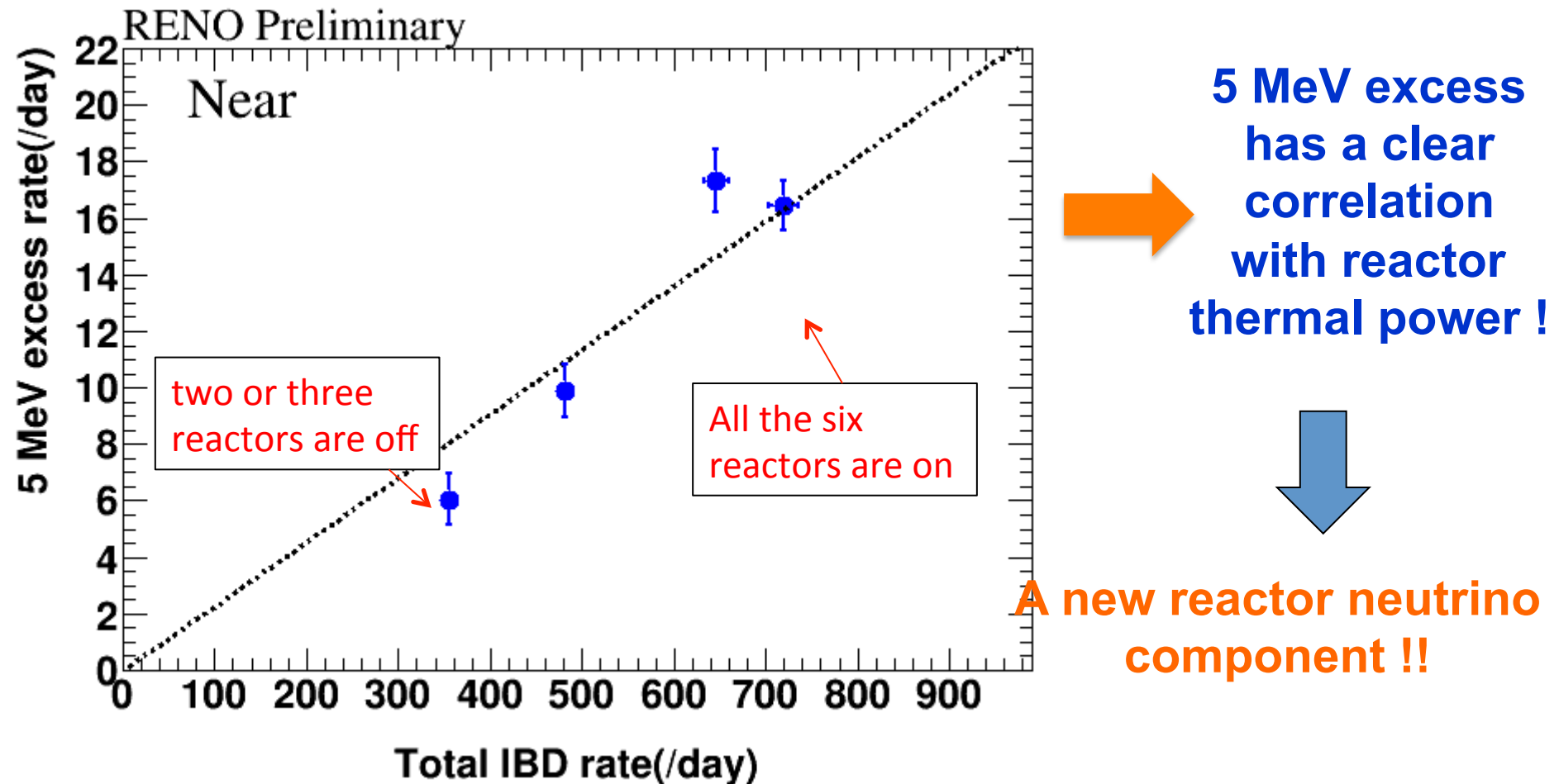


Observation of an excess at 5 MeV

arXiv:1511.05849.v2



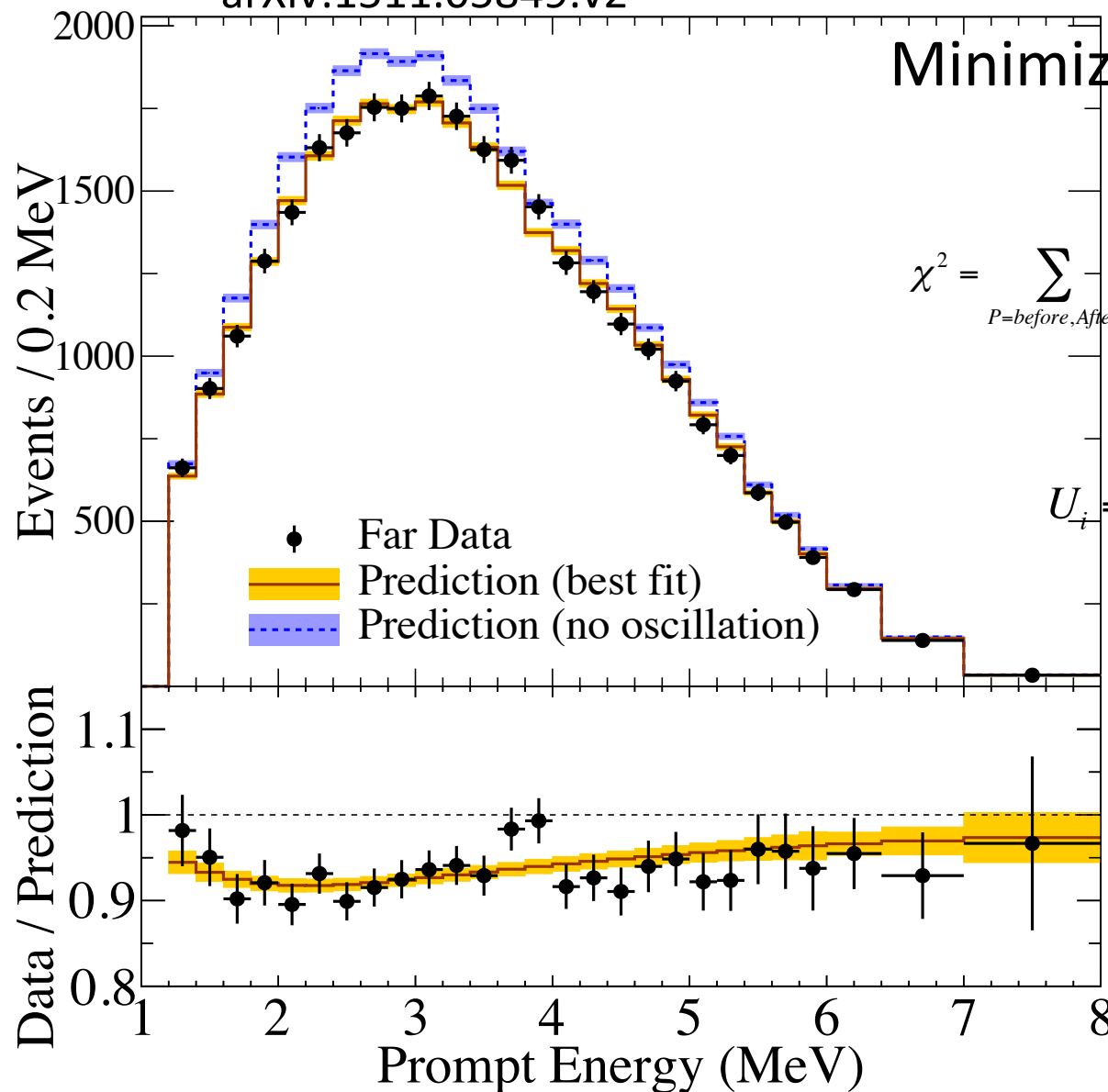
Correlation of 5 MeV Excess with Reactor Power



- ** Recent ab initio calculation [D. Dwyer and T.J. Langford, PRL 114, 012502 (2015)] :
- The excess may be explained by addition of eight isotopes, such as ^{96}Y and ^{92}Rb

Far/Near Shape Analysis for $|\Delta m_{ee}^2|$

arXiv:1511.05849.v2

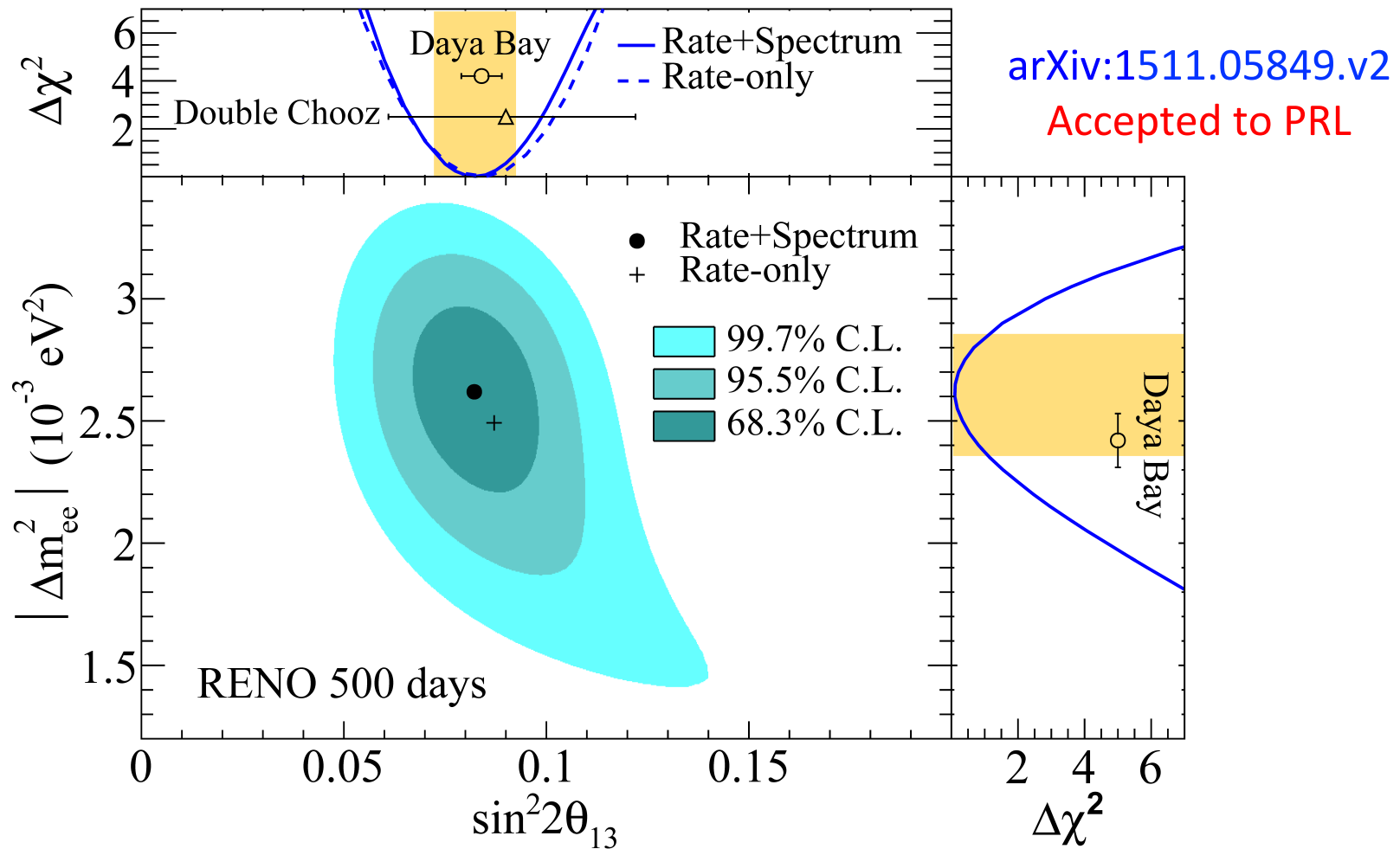


Minimize χ^2 Function

$$\chi^2 = \sum_{P=\text{before, After}} \left\{ \sum_{i=1 \sim N_b} \frac{\left(\frac{N_{obs}^{F,P,i}}{N_{obs}^{N,P,i}} - \frac{N_{Exp}^{F,P,i}}{N_{Exp}^{N,P,i}} \right)^2}{(U_i)^2} \right\} + Pull_Terms$$

$$U_i = \frac{N_{obs}^{F,i}}{N_{obs}^{N,i}} \cdot \sqrt{\frac{N_{obs}^{F,i} + N_{bkg}^{F,i}}{(N_{obs}^{F,i})^2} + \frac{N_{obs}^{N,i} + N_{bkg}^{N,i}}{(N_{obs}^{N,i})^2}}$$

Results from Spectral Fit

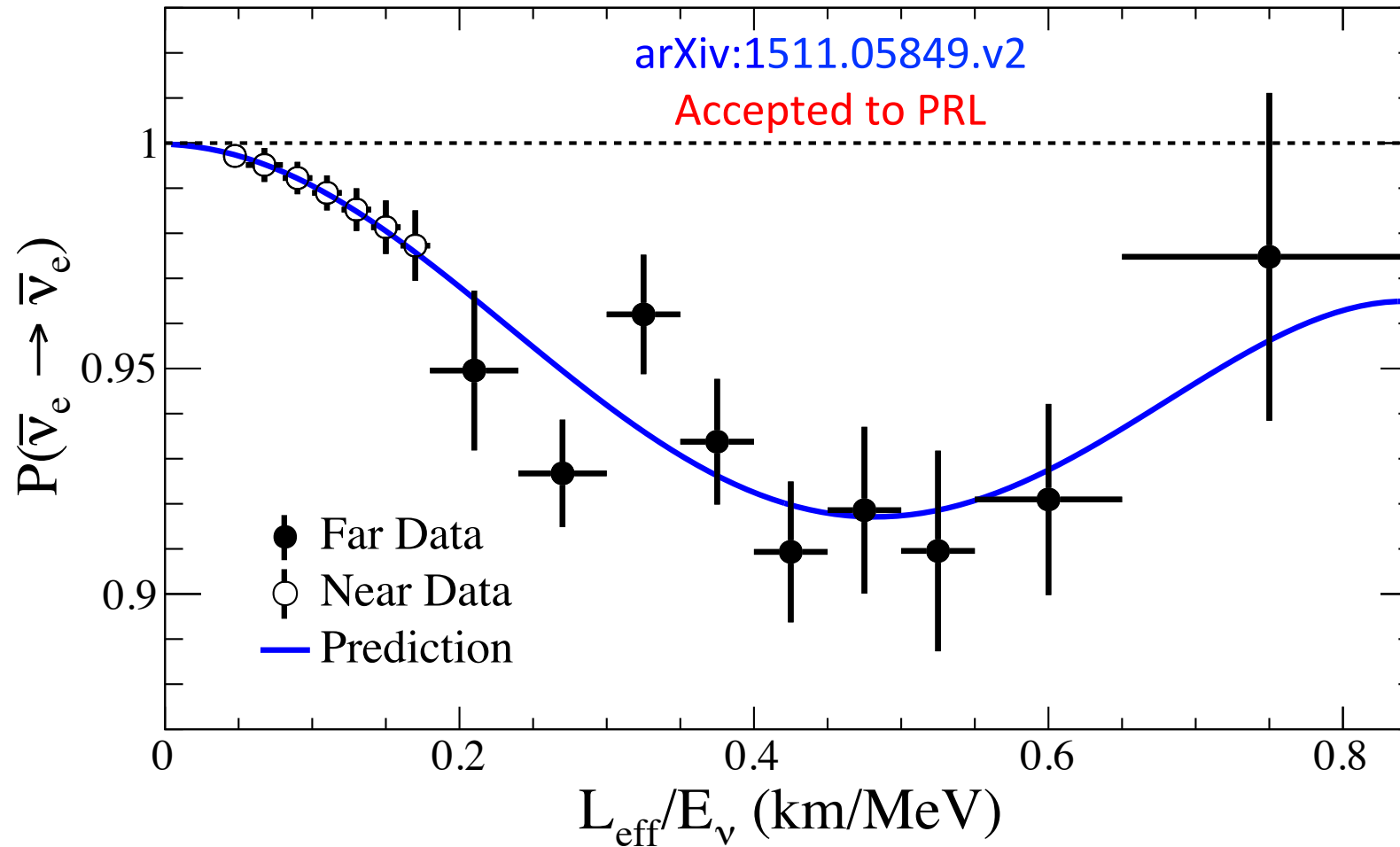


Rate+shape
new results

$$\sin^2 2\theta_{13} = 0.082 \pm 0.009(\text{stat.}) \pm 0.006(\text{syst.})$$

$$|\Delta m_{ee}^2| = 2.62^{+0.21}_{-0.23}(\text{stat.})^{+0.12}_{-0.13}(\text{syst.}) (\times 10^{-3} eV^2)$$

L/E Plot



$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) \approx 1 - \sin^2 2\theta_{13} \sin^2 \left(\Delta m_{ee}^2 \frac{L}{4E_\nu} \right)$$

Projected Sensitivity of θ_{13} & $|\Delta m_{ee}^2|$

$$\sin^2 2\theta_{13} = 0.082 \pm 0.010$$

(13% precision)

(~500 days)



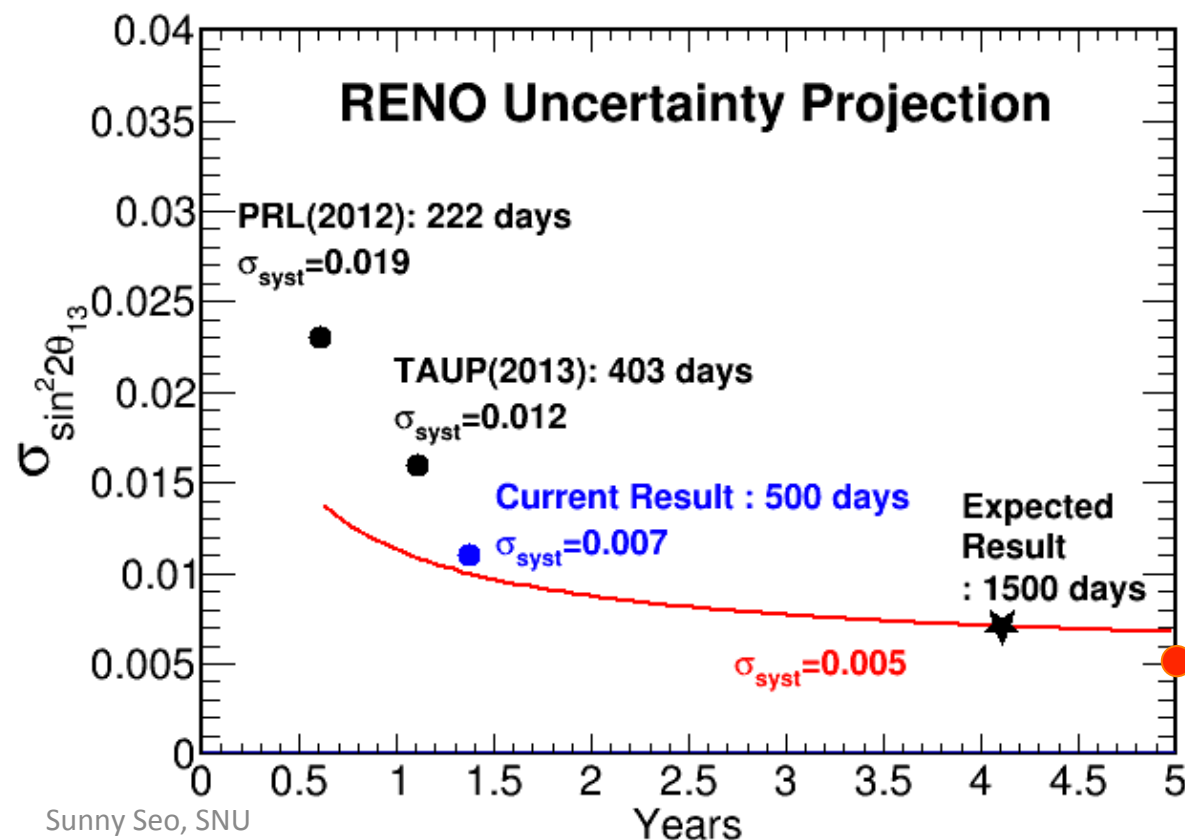
$$\pm 0.005$$

(5 % precision)

(5 years of data)

* Expected precision of $|\Delta m_{ee}^2|$: $\sim 0.1 \times 10^{-3} \text{ eV}^2$

(4% precision)



(5 % precision)

(sensitivity goal of θ_{13})

RENO Summary

- We obtained new results by rate+shape analysis (500 days).

$$\sin^2 2\theta_{13} = 0.082 \pm 0.009(\text{stat}) \pm 0.006(\text{syst})$$

$$|\Delta m_{ee}^2| = 2.62_{-0.23}^{+0.21}(\text{stat.})_{-0.13}^{+0.12}(\text{syst.}) (\times 10^{-3} eV^2)$$

 Accepted to PRL
(issue of May 27, 2016)

- Observed an excess at 5 MeV in reactor neutrino spectrum

- Measurement of θ_{13} from n-H IBD analysis

$$\sin^2 2\theta_{13} = 0.103 \pm 0.014(\text{stat}) \pm 0.014(\text{syst}) \quad (\text{preliminary})$$

- Sterile neutrino search result will be available soon.

- Goals: ~5% accuracy within 2 years for both $\sin(2\theta_{13})$ and $|\Delta m_{ee}^2|$

Overview of RENO-50

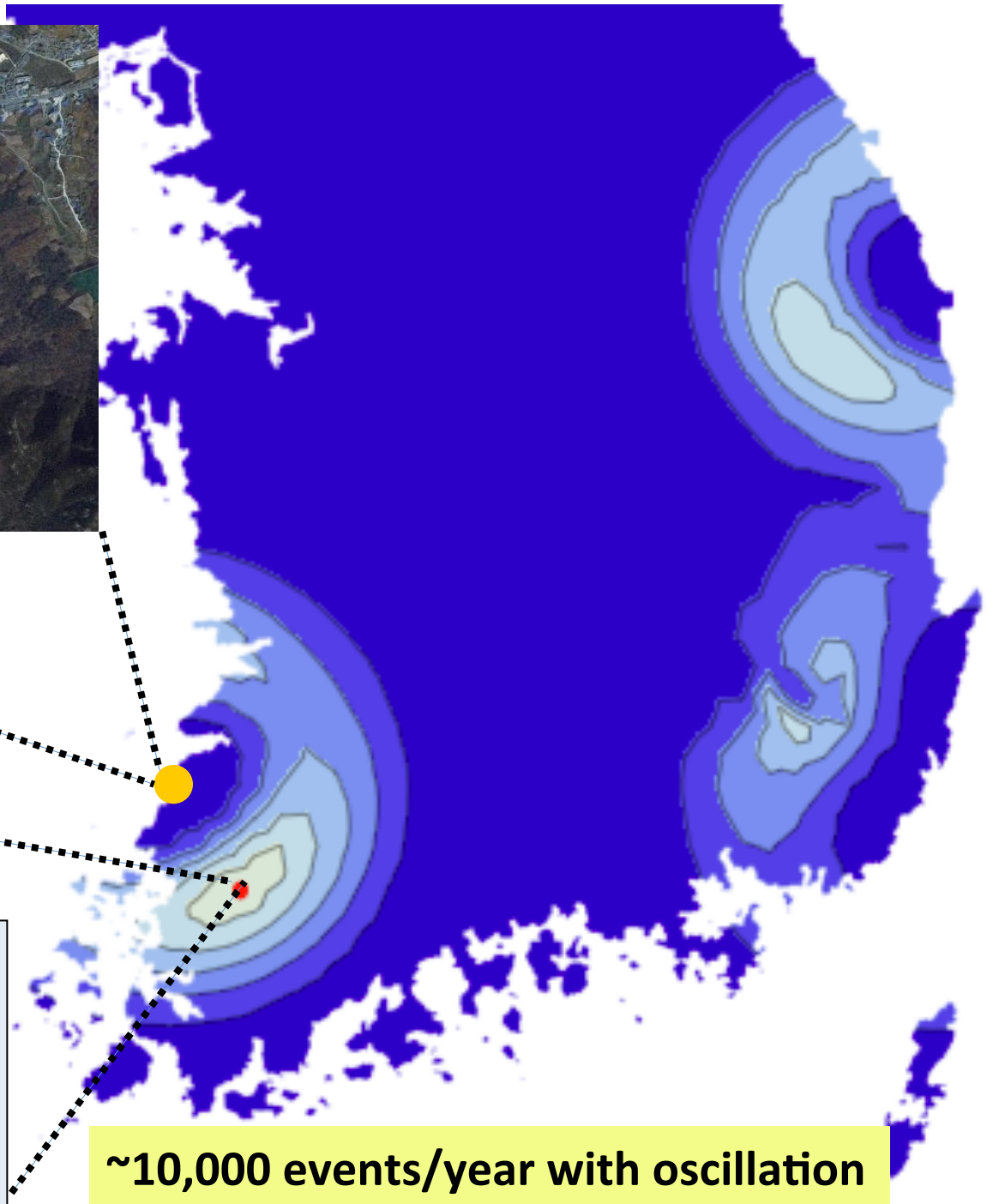
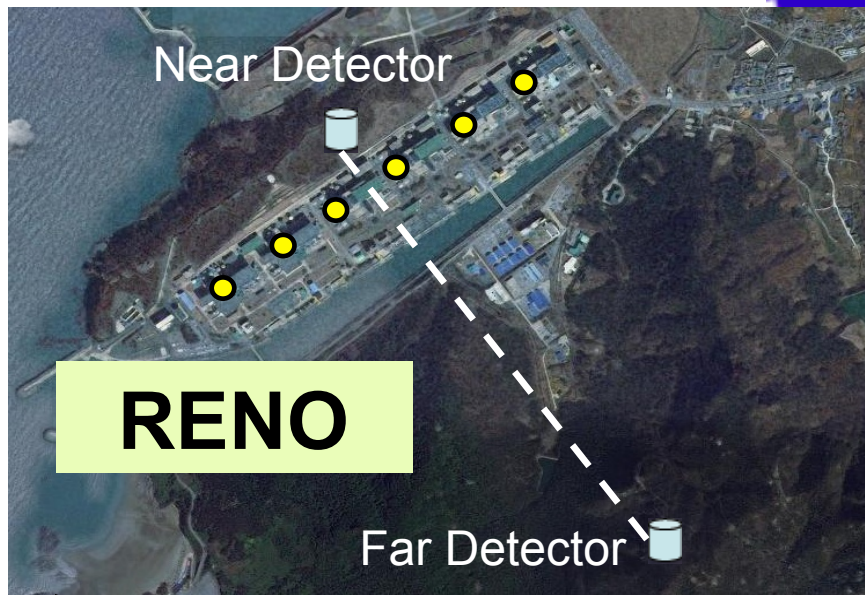
- **RENO-50** : An underground detector consisting of 18 kton ultra-low-radioactivity liquid scintillator & 15,000 20" PMTs, at 50 km away from the Hanbit(Yonggwang) nuclear power plant

RENO can be used as Near detector for RENO-50 (reduce sys. error)

- **Goals** : - Determination of neutrino mass hierarchy
- High-precision measurement of θ_{12} , Δm^2_{21} and Δm^2_{ee}
– Neutrino astronomy, Geo ν , sterile ν search, etc.

- **Budget** : \$ 100M for 6 year construction
(Civil engineering: \$ 15M, Detector: \$ 85M)

- **Schedule** : 2016 ~ 2021 : Facility and detector construction
2022 ~ : Operation and experiment



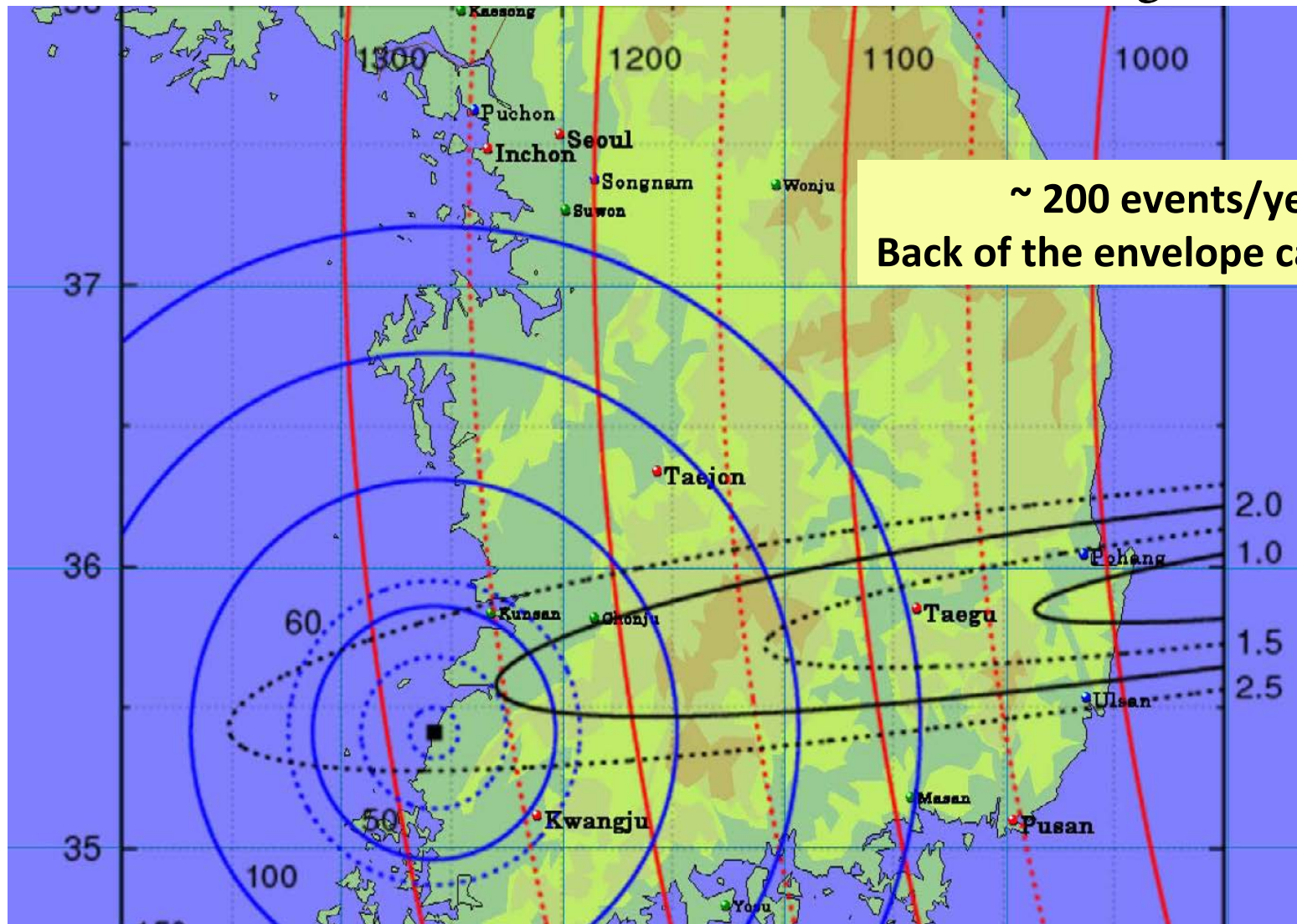
RENO-50

18 kton LS Detector
~47 km from YG reactors
Mt. Guemseong (450 m)
~900 m.w.e. overburden

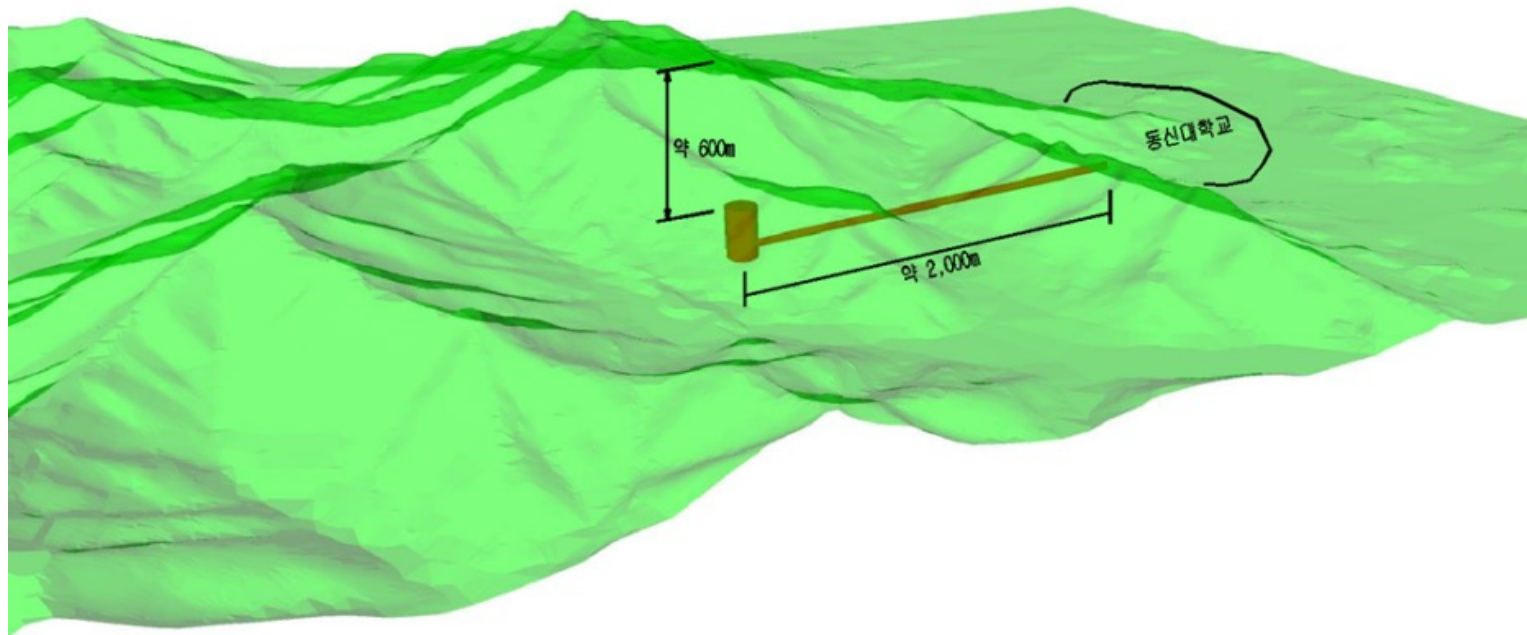
~10,000 events/year with oscillation

J-PARC neutrino beam

Dr. Okamura & Prof. Hagiwara



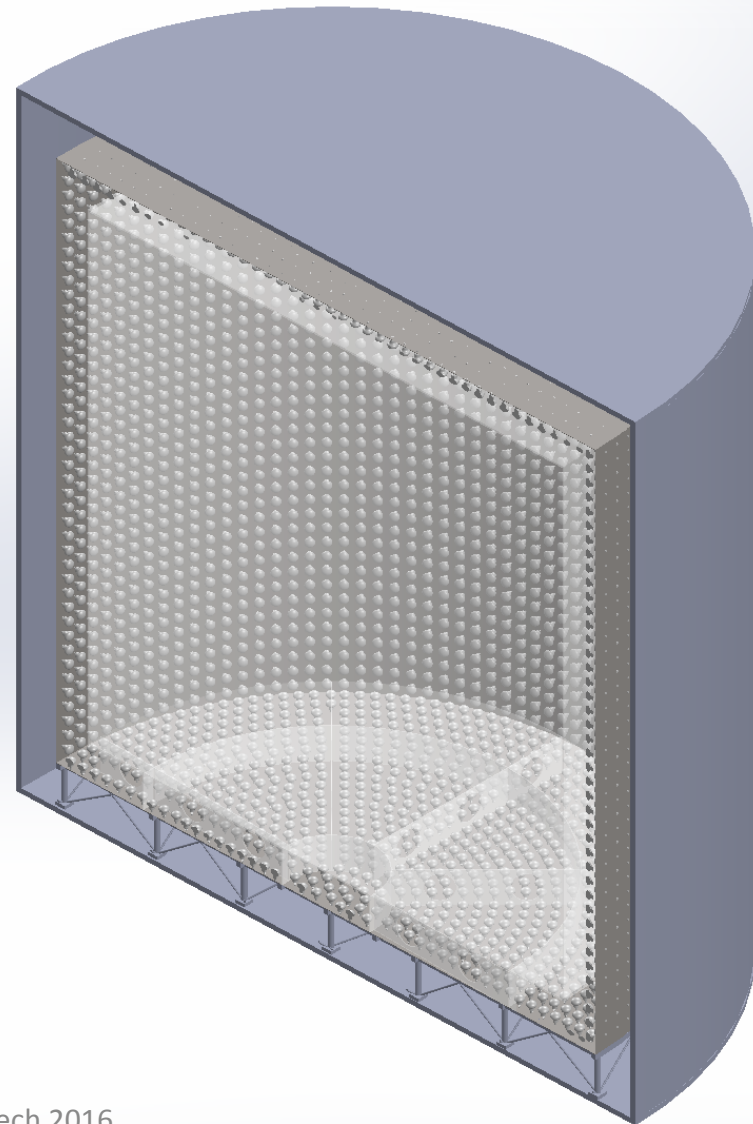
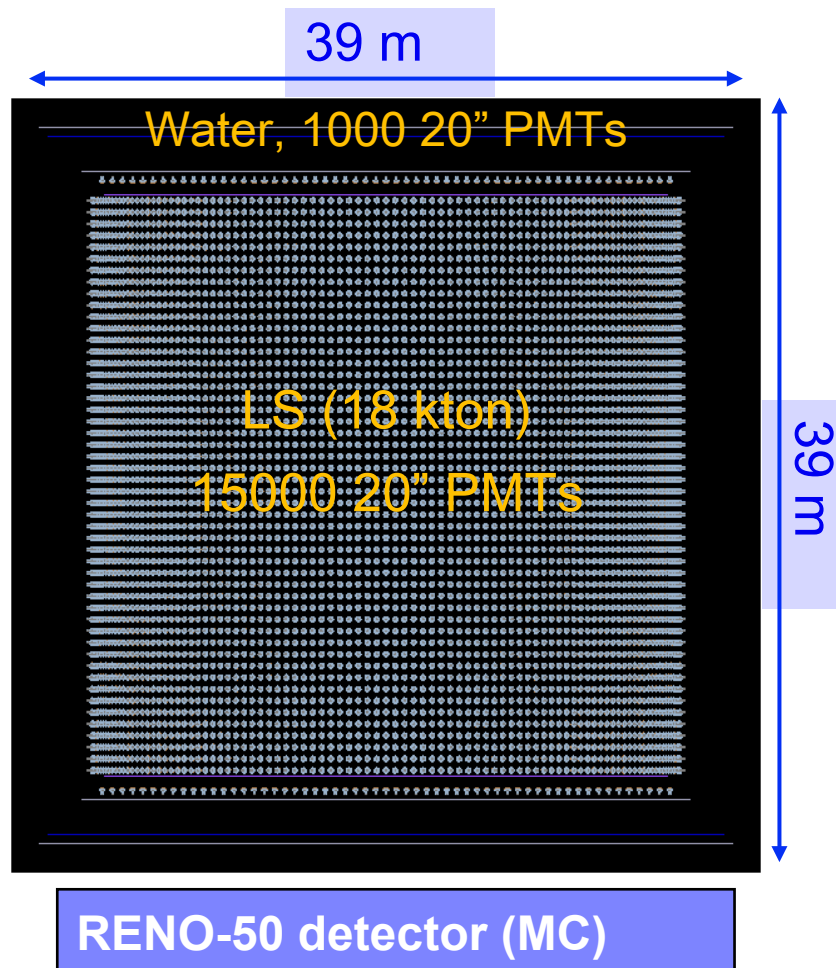
Geological Survey for Underground Facility



Cost estimation for RENO-50 underground facility (in progress)

- Geological survey for design of tunnel and experimental hall
- Cost estimation to be obtained soon

Conceptual Design of RENO-50 Detector



Status of Funding and R&D

- An R&D funding (US \$ 2M in 3 years, 2015~2017) is given by the Samsung Science & Technology Foundation.

- A proposal has been submitted to obtain construction funding.

**→ International Neutrino community's supports
will greatly enhance our opportunities !**

- A domestic symposium and an international workshop were held in 2013 to discuss the feasibility and physics opportunities.

- R&D is in progress for LS, PMT, DAQ, MC and detector design, in order to prepare a Technical Design Report (TDR).

- International collaboration is expected to be formed.
You are welcome to join us for R&D and detector construction!

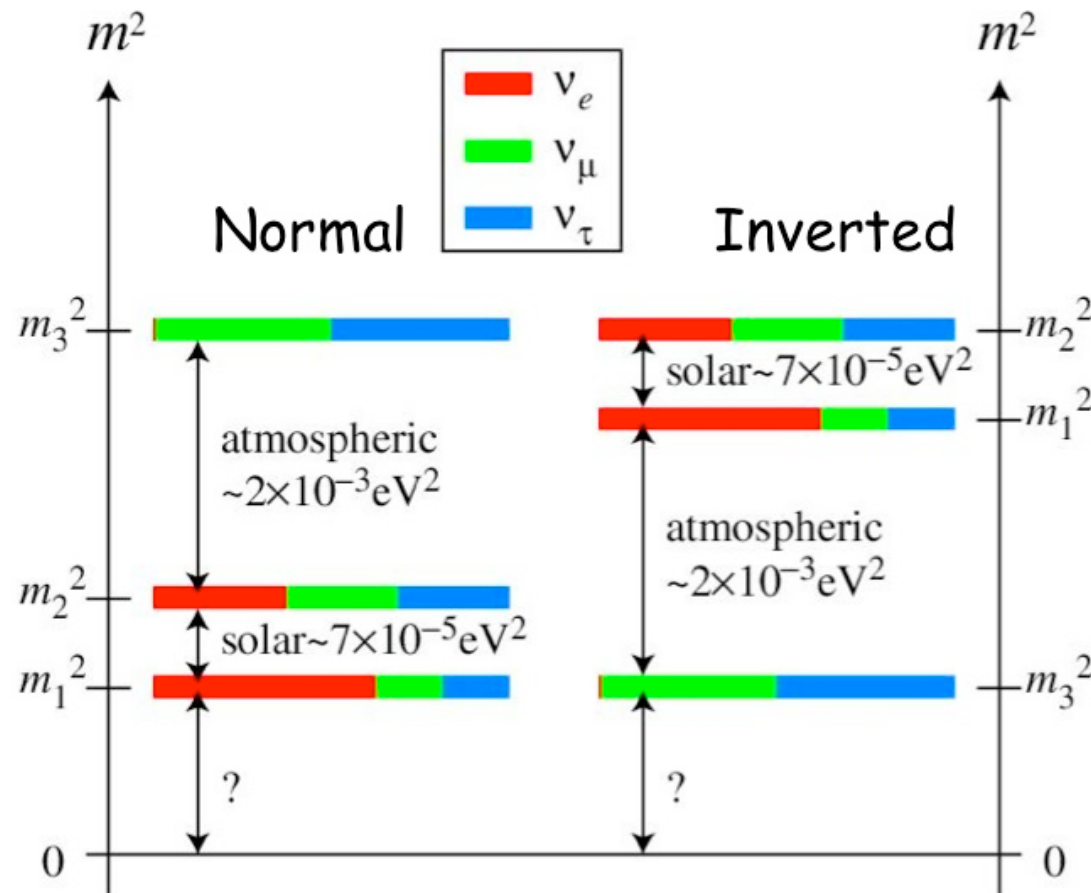
Schedule

- 2016 : Group organization
Detector simulation & design
Geological survey
- 2017 ~ 2018 : Civil engineering for tunnel excavation
Underground facility ready
Structure design
PMT evaluation and order,
Preparation for electronics, HV, DAQ & software tools,
R&D for liquid scintillator and purification
- 2019 ~ 2021 : Detector construction
- 2022 ~ : Data taking & analysis



Thank you !

What are the mass ordering of the three neutrinos ?



Reactor:

RENO-50, JUNO

Accelerator:

DUNE,
Hyper-K etc.

Atmosphere:

PINGU, ORCA

Physics Goals with RENO-50

- **Determination of neutrino mass hierarchy (very challenging)**
~3 σ sensitivity from ~10 years of data

- **Precise measurement of θ_{12} , Δm^2_{21} and $|\Delta m^2_{ee}|$**

$$\frac{\delta \sin^2 \theta_{12}}{\sin^2 \theta_{12}} < 1.0\%(1\sigma)$$

($\leftarrow 5.4\%$)

$$\frac{\delta \Delta m^2_{21}}{\Delta m^2_{21}} < 1.0\%(1\sigma)$$

($\leftarrow 2.4\%$)

$$\frac{\delta \Delta m^2_{ee}}{\Delta m^2_{ee}} < 1.0\%(1\sigma)$$

($\leftarrow 2.8\%$)

- **Neutrino burst from a Supernova in our Galaxy**
~5,600 events @ 8 kpc

- **Geo-neutrinos : ~ 1,500 geo-neutrinos for 5 years**
- Study the heat generation mechanism inside the Earth

- **Solar neutrinos** : with ultra low radioactivity detector
- Test MSW effect on neutrino oscillation and solar models

- **Sterile neutrino searches** : reactor ν , radio-sources, IsoDAR ν

- **Detection of J-PARC beam : ~200 events/year**

R&D in Progress

(1) Development of DAQ electronics

- Specification for *dead time free, high sensitivity and high speed signal processing*
- Prototype boards to be tested

(2) Develop techniques of LS purification

- Reduction of LS radioactivity to 10^{-16} g/g of U and Th
- Removal of LS impurities for attenuation length of ~ 25 m
- Several methods applied for investigation and evaluation
- Efforts on high sensitive measurement of radioactive concentration and optical parameters in LS

(3) Mechanical design of detector

- Detailed drawing of mechanical parts in progress
- MC simulation to estimate the performance

R&D in Progress

(4) Measurement of radioactivity for the detector materials

- Evaluate radioactive contamination of detector parts using a high purity Ge detector
- Estimate event rate contribution of those contaminations

(5) Measurement device for absolute LS attenuation length

- Developed a long pipe device with a laser source and a PMT
- Upgrade of the device in progress



- An R&D funding (US \$2M for 3 years of 2015-2017) is given by the Samsung Science & Technology Foundation.
- Efforts on obtaining a full construction fund